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3 Diverse Permian magmatism in the Tarim Block, NW China:  
4 genetically linked to the Permian Tarim mantle plume?  
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## Abstract

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Zircon U-Pb ages and geochemical data are reported for the Piqiang oxide-bearing ultramafic-mafic complex, the Bachu mafic dyke swarm, the Yingan and Kaipazileike basalts and the Halajun A-type granites in the Tarim Block, Northwest China. The Piqiang complex and the Halajun A-type granites were emplaced at ca. 276 Ma and ca. 278 Ma, respectively. Together with previously reported geochronological data, the diverse intrusive and extrusive rocks in Tarim show a peak age at ca. 275 Ma. Elemental and Nd isotope geochemistry suggests that the spatially and temporally related Piqiang complex (including some dolerite dykes or stocks) and the Halajun A-type granites were formed via crystal fractionation/accumulation of a common plume-derived parental mafic magma (melting degree > 10%), coupled with variable extents of crustal contamination. Crystal fractionation/accumulation in one or several magma chambers resulted in the diversity of rocks types. The Bachu mafic dyke swarm shares a similar mantle source with the intrusive rocks in the Piqiang-Halajun area but with a relatively lower degree of partial melting (~5%). In contrast, the basalts were derived from a time-integrated, enriched lithospheric mantle source as suggested by their high-Ti, LREE- and LILE-enriched trace element signature and negative  $\epsilon_{\text{Nd}}(\text{T})$  values (-2.0 ~ -2.6). The synchronous yet diverse range of Permian igneous rocks in Tarim can best be accounted for by a Permian mantle plume, which is about 15 Ma earlier than the Emeishan plume in southwestern China.

**Key words:** ultramafic-mafic complex; mafic dyke swarm; A-type granites; basalts; Permian Tarim LIP; mantle plume; NW China

## 47 **1. Introduction**

48

49 Mantle plumes are thermal anomalies that rise from the lower mantle or even the  
50 core–mantle boundary (Campbell and Griffiths, 1990). The ascent of a mantle plume can  
51 cause melting in the mantle and the crust, resulting in the formation of diverse igneous rocks  
52 (Bryan and Ernst, 2008; Xu et al., 2008). The diversity of extrusive and intrusive rocks  
53 genetically related to a single mantle plume is well documented in the Siberian traps  
54 (Czamanske et al., 1995; Fedorenko and Czamanske, 1997; Arndt et al., 1998, 2003), the  
55 Emeishan large igneous province (ELIP) (Xu et al., 2004, 2007, 2008; Zhou et al., 2002,  
56 2008; Wang et al., 2008) and even in some Precambrian large igneous provinces (Ernst et al.,  
57 2008; Li X.H. et al., 2003; Li Z.X. et al., 2008). Such diversity has been attributed to the  
58 variability in mantle sources, variable degrees of plume-lithosphere interaction, variable  
59 degrees of crustal melting and assimilation, or a combination of these processes (Arndt et al.,  
60 1998, 2003; Zhou et al., 2008, 2009; Wang et al., 2009).

61 Recent work suggests that a possible mantle plume was responsible for the early to middle  
62 Permian large igneous province in Tarim and the southern part of the Central Asian Orogenic  
63 Belt (CAOB) (Zhou et al., 2004; Chen et al., 2006; Borisenko et al., 2006; Mao et al., 2008;  
64 Pirajno et al., 2008, 2009; Polyakov et al., 2008; Zhang et al., 2008, 2010; Zhou et al., 2009;  
65 Tian et al., 2010) (Fig. 1), which was termed the Tarim LIP by Borisenko et al. (2006; also  
66 known as Bachu LIP or Tarim-Bachu LIP, *e.g.*, Zhang et al., 2008, 2010; Pirajno et al., 2009).  
67 However, the genetic links between the intrusive and extrusive rocks in the Tarim LIP,  
68 especially those in the Tarim Block, have not yet been established and thus the processes that  
69 formed the diverse igneous rocks are not well understood. In this study, we carried out



70 geological, geochronological and geochemical analyses on the Piqiang oxide-bearing  
71 intrusive complex, the Halajun granite plutons (pluton □ and pluton □), the Bachu mafic  
72 dyke swarm and basalts in the Tarim Block, NW China. The aims of the study were to  
73 constrain the timing of the emplacements and to address the petrogenesis of these diverse  
74 igneous rocks by identifying possible sources involved in magma generation and deciphering  
75 their relationship to the mantle plume.

76

## 77 **2. Regional geology and petrography**

78

79 The Tarim Block in northwestern China covers an area of more than 600,000 km<sup>2</sup>. It is  
80 one of the least known continental blocks in Asia due to its extensive coverage by desert.  
81 Nevertheless, several important phases of igneous activities have been identified in Tarim,  
82 *i.e.*, the Neoarchaeon, early Palaeoproterozoic, Neoproterozoic and early Permian events (Hu  
83 et al., 2000; Lu et al., 2008; Zhang et al., 2007, 2008, 2010). Among these igneous activities,  
84 the early Permian phase is the latest and was considered to be related to mantle plume activity  
85 (Pirajno et al., 2008; Mao et al., 2008; Zhou et al., 2009; Zhang et al., 2008, 2010; Tian et al.,  
86 2010).

87 The study region is close to Bachu in the northwestern part of the Tarim Block (Figs. 1  
88 and 2). Some geological, geochronological and geochemical data have previously been  
89 reported on the ultramafic-mafic-syenite complex, mafic dykes and syenite plutons (or dykes)  
90 in the Bachu region, and on the basalts in the Yingan and Kaipazileike sections (marked as  
91 KF1 and KF2 in Fig. 1) (Rui et al., 2002; Jiang et al., 2004a, 2004b; Yang et al., 1995, 2005,

92 2006, 2007; Li et al., 2007; Zhou et al., 2009). However, most of these data are published in  
93 the Chinese literature and precise geochronology and systematic geochemical data are still  
94 scarce, and no age and geochemical data have yet been reported for the oxide  
95 (magnetite)-bearing Piqiang ultramafic-mafic complex and the nearby Halajun granites.

96

### 97 *2.1 The Piqiang ultramafic-mafic complex*

98 The oxide (magnetite)-bearing Piqiang intrusive complex crops out ca. 120 km northeast  
99 of Atushi City (Fig. 2a). It intrudes Devonian sedimentary rocks with the contact dipping 70 -  
100 80° toward the interior of the complex, and has an outcrop area of ca. 25 km<sup>2</sup> (Fig. 2a).  
101 Thin-section examinations and field observations, together with previous field mapping and  
102 petrographic studies (*e.g.*, Rui et al., 2002), indicate that the complex is composed mainly of  
103 gabbro (accounting for ca. 95% of the outcrop) with minor olivine-bearing gabbro and  
104 dolerite (Fig. 2b). Most of the crystalline rocks are cumulates (except for the dolerite). The  
105 gabbros are medium- to coarse-grained, and consist of clinopyroxene (30–50%), plagioclase  
106 (40–45%) and variable amounts of orthopyroxene (5-15%), magnetite (5-20%, mostly  
107 vanadium titanite - magnetite) and olivine (1-10%). Accessory minerals include apatite and  
108 zircon. Magnetite is commonly disseminated in the gabbros but in places it occurs as veins or  
109 blocks, forming economic orebodies (Fig. 2b). The dolerite is fine-grained or microcrystalline  
110 and consists mainly of plagioclase (45-55%), clinopyroxene (35-45%) and Ti-Fe oxides  
111 (5-10%). No phenocrysts have been observed.

112 One sample (08KT01) collected from coarse-grained gabbro in the Piqiang complex  
113 (08KT01, 40°24'42"N, 77°38'10"E) was selected for U-Pb zircon dating. Eleven gabbro

114 cumulates and two dolerite samples were chosen for geochemical analyses.

115

## 116 *2.2 The Bachu mafic dyke swarm*

117 Mafic dykes around the Wajilitage complex intrude the upper Devonian sedimentary  
118 rocks with variable strikes (Fig.1, Fig. 2c). The local geological map (1:5000) shows that the  
119 dykes also intrude the Wajilitage layered ultramafic-mafic-syenite complex (Fig. 2c) and  
120 some of them intermingle with the complex. At several outcrops bimodal dykes, consisting of  
121 dolerite and syenite, have been observed. The petrography of the mafic dykes around  
122 Wajillitage is similar to that in the Mazaertage area (also known as the Xiaohaizi area) north  
123 of the Wajilitage complex (Jiang et al., 2004a; Yang et al., 2007; Zhou et al., 2009).  
124 According to Jiang et al. (2004a) and our own field observations, the mafic dyke swarm in  
125 the Mazaertage area mainly crops out around the Xiaohaizi gabbro-syenite complex with  
126 variable strike directions (Fig. 1), as with the mafic dyke swarm in the Wajilitage area (this  
127 study). The widths of the dykes from both areas range mostly between ~2 m and ~30 cm,  
128 with a few exceptions of over 5 m wide. Sixteen samples were collected from the Wajilitage  
129 dolerite dykes for geochemical analyses (Fig. 2c).

130

## 131 *2.3 The Halajun granite plutons*

132 Two granite plutons crop out at Halajun (termed Halajun pluton I and II), near the Piqiang  
133 complex (Fig. 2a). Pluton □, with a ca. 40 km<sup>2</sup> outcrop area, intrudes the upper  
134 Carboniferous sedimentary rocks with the contact dipping 50-60° outward (Fig. 2a). A  
135 hornfels belt, 10 – 30 m wide, developed along the margin of the pluton. The granites are

136 pinkish in color and medium- to coarse-grained. The main minerals are alkali feldspar  
137 (60–65%), quartz (35-40%), hornblende (~2 %) and brown biotite (~1%). Accessory minerals  
138 include zircon, apatite and Fe-Ti oxides. Pluton □ has an outcrop area of ca. 50 km<sup>2</sup>  
139 surrounded by the desert. This pluton is characterized by a coarse-grained texture (up to 1 cm)  
140 and a rather uniform mineralogical composition. The rock-forming minerals are quartz  
141 (~40%), alkali feldspar (~60%) and very minor hornblende. The coarse-grained texture of this  
142 pluton and the coverage of its margins by the desert indicate that the pluton is likely much  
143 larger than its outcrop area.

144 One sample each from Halajun pluton □ (08KT02: 46°35'8"N, 90°11'57"E) and pluton □  
145 (08KT03: 46°15'34"N, 90°48'45"E) were selected for zircon U-Pb geochronology. Additional  
146 samples were collected from the two plutons for geochemical analyses.

147

#### 148 *2.4 Permian basalts in Yingan and Kaipazileike*

149 The Permian basaltic sections in Yingan and Kaipazileike have been described by Jiang  
150 et al. (2004b) and Zhou et al. (2009). The rock successions have been subdivided into the  
151 Kupukuziman Formation (KF1) and the overlying Kaipazikelei Formation (KF2) (Fig. 1).  
152 KF1 is dominated by basaltic flows interlayered with valcaniclastic rocks, and KF2 is mainly  
153 composed of basalts with minor fine-grained volcanic pyroclastics. Twelve basalt samples  
154 (08KT01 – 08KT12) from KF1 and three basalt samples (08KT13 – 08KT15) from KF2 were  
155 collected during this study. All the samples are of greenish color and cryptocrystalline or  
156 microcrystalline. They consist of needle-like plagioclase (40-50%), clinopyroxene (30-40%,  
157 mostly replaced by actinolite) and Ti-Fe oxides (10-20%).

158

159 **3. Analytical methods**

160

161 Zircon separation was carried out first using conventional magnetic and density  
162 techniques to concentrate non-magnetic, heavy fractions. Zircon grains were then  
163 hand-picked under a binocular microscope. Zircon ages were analyzed using the SHRIMP  
164 U-Pb method. Zircon grains were cast on an epoxy mount, and polished to expose the crystals  
165 in half for analyses. The inner structures of the zircon grains were documented with  
166 transmitted and reflected light photomicrographs and cathodoluminescence (CL) images. The  
167 mount was then vacuum-coated with high-purity gold. U-Th-Pb analyses were conducted at  
168 the Beijing SHRIMP Center, Chinese Academy of Geological Sciences using standard  
169 operating conditions (Williams, 1998). U-Th-Pb ratios were determined relative to the  
170 TEMORA standard zircon, and the absolute abundances of U and Th were determined using  
171 the SL13 standard zircon. Measured compositions were corrected for common Pb using  
172 non-radiogenic  $^{204}\text{Pb}$ , and an average crustal composition (Stacey and Kramers, 1975)  
173 appropriate to the age of the mineral. Software SQUID 1.0 and ISOPLOT (Ludwig, 1999,  
174 2001) were used for data processing. The weighted mean ages are quoted at 95% confidence  
175 level. U-Pb zircon data are presented in Data Repository Table DR1.

176 Geochemical and isotope measurements were carried out at the Guangzhou Institute of  
177 Geochemistry, Chinese Academy of Sciences (GIG CAS). Major elements were analyzed  
178 using a Rigaku ZSX100e XRF following the analytical procedure described by Li et al.  
179 (2004). Analytical precision is generally better than 2%. Trace elements were analyzed using

180 a Perkin-Elmer Sciex ELAN 6000 ICP-MS following the procedure described by Li et al.  
181 (2002). In-run analytical precisions for most elements were better than 3-5%. Major and trace  
182 elements data are presented in Data Repository Table DR2.

183 Nd isotopes were determined using a Micromass Isoprobe multi-collector ICPMS  
184 (MC-ICPMS) at GIG CAS, following the procedure described by Li et al. (2004). Measured  
185  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios were normalized to  $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$ . The reported  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios  
186 were adjusted to the Shin Etsu JNdi-1 standard  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512115$ . Sm-Nd isotopic  
187 compositions and calculated parameters are listed in Table 1.

188

## 189 **4 Results**

190

### 191 *4.1 U-Pb zircon ages*

192

#### 193 *4.1.1 The Piqiang ultramafic-mafic complex*

194 Zircon grains from sample 08KT01 are mostly euhedral, transparent and colorless. They  
195 are 200-300  $\mu\text{m}$  long, with length to width ratios of 3–5. Euhedral concentric zoning is  
196 common in most crystals under CL images, typical of a magmatic origin. No relict cores have  
197 been observed. Eighteen analyses were conducted on 18 zircon grains. Uranium  
198 concentrations range from 61 to 531 ppm, Th from 39 to 524 ppm, and Th/U ratios from 0.42  
199 to 1.19 (Data Repository Table DR1). Among the eighteen analyses, three analyses (spots 2.1,  
200 7.1 and 13.1) have higher common lead and their  $^{206}\text{Pb}/^{238}\text{U}$  ages deviate significantly from  
201 the other analyses. These three spots were excluded from the age calculation. The remaining  
202 15 analyses yielded a weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  age of  $276 \pm 4$  Ma (MSWD = 2.8) (Fig. 3a).

203

204 *4.1.2 The Halajun granites (plutons □ and □)*

205 Euhedral zircons from sample 08KT02 are up to 100-200  $\mu\text{m}$  in length, and have length to  
206 width ratios up to 4:1. Most zircons are transparent and colourless and a few are brown and  
207 turbid. In CL images no inherited cores have been observed. Euhedral concentric zoning is  
208 common in most crystals. Fourteen analyses were obtained on 14 grains (Data Repository  
209 Table DR1). U and Th concentrations are 80–169 ppm and 34–95 ppm, respectively. Th/U  
210 ratios vary from 0.39 to 0.70. Among the fourteen analyses, spot 1.1 has a large error and was  
211 rejected from the age calculation. The remaining thirteen analyses have consistent  $^{206}\text{Pb}/^{238}\text{U}$   
212 ages within error and yielded a weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  age of  $278 \pm 3$  Ma (MSWD = 0.2)  
213 (Fig. 3b). This age is interpreted as the emplacement age of the Halajun pluton I.

214 Zircons from sample 08KT03 are euhedral, up to 400  $\mu\text{m}$  long, and have length to width  
215 ratios up to 4:1. They are transparent and colourless, without inherited cores. Euhedral  
216 concentric zoning is common, suggesting a magmatic origin. Seventeen analyses were  
217 obtained on 17 grains (Data Repository Table DR1). U concentrations range from 80 to 840  
218 ppm, Th from 30 to 233 ppm and Th/U ratios vary within the range of 0.27–0.70. Spot 1.1  
219 has a  $^{206}\text{Pb}/^{238}\text{U}$  age of  $570 \pm 11$  Ma and it is likely a xenocryst (a much larger grain  
220 compared with other zircons). The other 16 analyses agree within errors and yielded a  
221 weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  age of  $278 \pm 3$  Ma (MSWD = 1.36) (Fig. 3c). This age is  
222 interpreted as the crystallization age of the Halajun pluton □.

223

## 224 4.2 Elemental geochemistry

225

### 226 4.2.1 Hydrothermal alteration effects

227 Most samples from the Piqiang complex, the Halajun granites (plutons □ and □) and the  
228 basalts have relatively low LOI (loss on ignition), thus their elemental compositions likely  
229 remained immobile since their emplacement. On the other hand, the Bachu mafic dykes have  
230 relatively high LOI (>2%; Data Repository Table DR2), suggesting that they may have been  
231 altered to various degrees during post-emplacement processes. Generally, alteration processes  
232 do not significantly affect high field strength elements (HFSE: Th, Nb, Zr, Hf, Ti, Y) and rare  
233 earth elements (REE) (Pearce, 1992), as is evidenced by the positive correlations of Zr versus  
234 La and Nb for the studied samples (Fig. 4). Thus, the variations of HFSE and REE can be  
235 used to characterize the magmatic processes (see the following section). On the contrary,  
236 LILEs (large ion lithophile elements), such as K, Na and Rb, do not show a correlation with  
237 the immobile elements, suggesting that they have been variably affected by alteration  
238 processes (Fig. 4).

239

### 240 4.2.2 The Piqiang ultramafic-mafic complex

241 Samples from the Piqiang complex show large elemental compositional variations  
242 (39.1-52.2 wt.% SiO<sub>2</sub>, 7.8-24.3 wt.% Al<sub>2</sub>O<sub>3</sub>, 2.7-14.0 wt.% MgO, 6.9-11.8 wt.% CaO, 0.6-5.6  
243 wt.% TiO<sub>2</sub>, 5.5-25.0 wt.% Fe<sub>2</sub>O<sub>3</sub><sup>(t)</sup>, 1.01-3.52 wt.% Na<sub>2</sub>O, 0.11-1.83 wt.% K<sub>2</sub>O, 0.01-0.37  
244 wt.% P<sub>2</sub>O<sub>5</sub>; Data Repository Table DR2, and Fig. 4), which may be related to crystal  
245 fractionation/accumulation. They straddle the line separating the alkaline and sub-alkaline  
246 fields on both the total alkali versus silica diagram and the incompatible elements



247 classification diagram (Fig. 5). In Harker diagrams (Fig. 4), as SiO<sub>2</sub> decreases, MgO and  
248 Fe<sub>2</sub>O<sub>3</sub> increase while Na<sub>2</sub>O and K<sub>2</sub>O decrease. Al<sub>2</sub>O<sub>3</sub>, TiO<sub>2</sub>, and CaO show more scatter (Fig.  
249 4). In the AFM (Na<sub>2</sub>O+K<sub>2</sub>O-FeO-MgO) diagram, the rocks follow a tholeiitic trend (Fig. 5).

250 The cumulate rocks (magnetite-bearing gabbro and olivine-bearing gabbro) from the  
251 Piqiang complex have low and variable compatible element contents, e.g., Cr concentrations  
252 ranging from 0.5 to 260 ppm, Ni from 0.1 to 173 ppm and V from 176 to 841 ppm (Data  
253 Repository Table DR2). Large variations in REE contents ( $\Sigma$ REE = 16-149 ppm) could be  
254 ascribed to different abundances of cumulus and accessory minerals such as apatite and  
255 zircon because of their high partition coefficients for REEs. Nonetheless, all the cumulate  
256 samples are uniformly enriched in LREE relative to HREE with (La/Yb)<sub>N</sub> ranging from 2 to 8.  
257 All but one sample (KL01-6) show variable positive Eu anomalies ( $\delta$ Eu =1.1-3.4) (Fig. 6,  
258 where both the chondrite and primitive mantle values are from Sun and McDonough, 1989).  
259 Incompatible elements also exhibit variable abundances in the primitive mantle-normalized  
260 spider-diagrams, reflecting different abundances of cumulus and accessory minerals (Fig. 6).  
261 The rocks show positive anomalies in Ba, Sr and Ti relative to neighbouring elements and  
262 depletion in P relative to Nd and Sm.

263 Two dolerite samples (KL01-14, KL01-15) have relatively high total REE ( $\Sigma$ REE =  
264 155-187 ppm) and insignificant Eu anomalies ( $\delta$ Eu  $\approx$  0.95). However, they exhibit REE  
265 patterns and trace element spider-diagrams similar to the cumulate samples, except that they  
266 have higher concentrations than the cumulate samples (Fig. 6). Taking into account their  
267 coexistence on outcrops, this suggests a common origin for the dolerites and the cumulate  
268 gabbros.

269

#### 270 4.2.3 The Bachu mafic dykes

271 The Bachu mafic dykes have low SiO<sub>2</sub> (41.9 - 45.1%, mostly <45%), MgO (2.7-6.1%), Cr  
272 (<60 ppm), Ni (<100 ppm) contents, variable total alkali (3.4-9.9%), MnO (0.05-0.20%),  
273 Al<sub>2</sub>O<sub>3</sub> (12.4-16.0%), CaO (6.5-9.8%), Fe<sub>2</sub>O<sub>3</sub><sup>T</sup> (7.6-17.2%) and relatively high TiO<sub>2</sub>  
274 (2.2-5.2%), P<sub>2</sub>O<sub>5</sub> (0.39-0.97%), Sr (390-1607 ppm) and Ba (330-2500 ppm). They plot within  
275 the alkali basalt field on both the TAS and the Nb/Y versus Zr/TiO<sub>2</sub> diagrams (Fig. 5a, b).  
276 Nonetheless, on the AFM diagram of Miyashiro (1974) they exhibit a typical tholeiitic trend  
277 (Fig. 5c). With decreasing SiO<sub>2</sub>, TiO<sub>2</sub>, (Fe<sub>2</sub>O<sub>3</sub>)<sup>T</sup>, MgO and CaO increase while Al<sub>2</sub>O<sub>3</sub>  
278 decreases. K<sub>2</sub>O, Na<sub>2</sub>O and P<sub>2</sub>O<sub>5</sub> show no obvious correlation with SiO<sub>2</sub> (Fig. 4).

279 The Bachu mafic dykes have high total REE abundances (181-477 ppm) with La<sub>N</sub> = 7-26.  
280 On chondrite-normalized REE plots, they display enriched LREE patterns with relatively  
281 constant La<sub>N</sub>/Sm<sub>N</sub> (2.3-4.4) and Sm<sub>N</sub>/Yb<sub>N</sub> (3.1-6.6) ratios and insignificant Eu anomalies  
282 ( $\delta\text{Eu} = 0.97\text{-}1.0$ ) (Fig. 6). Their trace element spider diagram is characterized by significant  
283 negative Sr anomalies and negative to positive Ti anomalies, probably due to Ti-bearing  
284 magnetite fractionation/accumulation (Fig. 6).

285

#### 286 4.2.4 The Halajun granites

287 The Halajun granite □ and □ are highly silicic in composition with SiO<sub>2</sub> ranges of 71.7%  
288 – 75.2 % and 77% – 78.2%, respectively. Their A/CNK values vary between 0.95-1.02,  
289 indicating a weak metaluminous to weak peraluminous nature. There is a significant negative  
290 correlation between SiO<sub>2</sub> and total alkali (Fig. 5a), suggesting extensive crystal fractionation

291 of alkali feldspar. Probably due to this fractionation, these samples show a sub-alkaline  
292 signature (Fig. 5a). However, HFSE may not have been affected by fractionation given their  
293 relatively high contents. On the  $10000 \cdot \text{Ga}/\text{Al}$  versus Zr diagram (Fig. 7a), these samples plot  
294 in the field of typical A-type granite. Specifically, they show A<sub>1</sub>-type characteristics  
295 according to the classification of Eby (1992). In this regard, the Halajun granites resemble the  
296 syenite plutons and dykes in the Bachu area (Fig. 7b).

297 The Halajun granite □ and □ have essentially the same trace element compositions except  
298 that pluton □ shows more evolved characteristics, with lower  $\delta\text{Eu}$  and deeper troughs for Ba,  
299 Sr, P and Ti than those of the pluton I (Fig. 6). The more evolved nature of pluton □ is also  
300 highlighted by its relatively high SiO<sub>2</sub> and low TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, Fe<sub>2</sub>O<sub>3</sub>, CaO, MgO contents  
301 compared to pluton □ (Data Repository Table DR2).

302

#### 303 *4.2.5 The Permian basalts*

304 The Permian basalts have very similar geochemical signatures except that the KF2  
305 samples have higher SiO<sub>2</sub> contents than the KF1 samples. All the samples are characterized  
306 by high total alkali ( $\text{Na}_2\text{O} + \text{K}_2\text{O} = 4.15\text{-}5.26\%$ ) over a relatively narrow SiO<sub>2</sub> range of  
307 44.1-51.2%. They plot within the alkaline basalt field in the TAS diagram (Fig. 5a),  
308 consistent with their high Nb/Y ratios (Fig. 5b). Their high Fe<sub>2</sub>O<sub>3</sub><sup>T</sup> contents show a typical  
309 tholeiitic trend in the FeO (t) – Na<sub>2</sub>O+K<sub>2</sub>O – MgO triangle diagram (Fig. 5c). On the Harker  
310 diagrams, Al<sub>2</sub>O<sub>3</sub>, TiO<sub>2</sub>, CaO and MgO increase with decreasing SiO<sub>2</sub>, while Fe<sub>2</sub>O<sub>3</sub><sup>T</sup>, K<sub>2</sub>O,  
311 Na<sub>2</sub>O and P<sub>2</sub>O<sub>5</sub> show more scattered variations (Fig. 4). As for REE and incompatible trace  
312 elements, the basalts have high total REE contents ( $\Sigma\text{REE} = 180\text{-}233$  ppm) and REE

313 distribution patterns characterized by enriched LREEs and insignificant Eu anomalies  
314 ( $\text{La}_N/\text{Yb}_N = 6-8$ ,  $\text{La}_N/\text{Sm}_N = 2.3-2.7$ ,  $\text{Sm}_N/\text{Yb}_N = 2.8-3.3$ ,  $\delta\text{Eu} = 0.95-1.0$ ) (Fig. 6). Normalized  
315 to primitive mantle, they exhibit coherent incompatible element distribution patterns  
316 characterized by positive Ba and negative Sr and Nb (Ta) anomalies ( $\text{Nb}/\text{La} = 0.6-0.8$ ,  $\delta\text{Sr} =$   
317  $0.3-0.5$ ) (Fig. 6).

318

#### 319 *4.3 Whole-rock Sm-Nd isotopic compositions*

320 The Piqiang samples show variable Sm and Nd contents and  $^{147}\text{Sm}/^{144}\text{Nd}$  ratios (0.1144 -  
321 0.1920). Their  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios range from 0.512436 to 0.512666 and  $\epsilon_{\text{Nd}}(t)$  values from  
322 -1.05 to 2.08 (Table 1). The Bachu mafic dykes have similar  $^{147}\text{Sm}/^{144}\text{Nd}$  (0.1144 - 0.1262)  
323 and high  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios of 0.512524 to 0.512762 ( $\epsilon_{\text{Nd}}(t) = 0.56 - 5.02$ ). Basaltic samples  
324 have consistent  $^{147}\text{Sm}/^{144}\text{Nd}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios and similar  $\epsilon_{\text{Nd}}(t)$  values of -2.05 to -2.57.  
325 The Halajun plutons □ and □ have slightly different  $\epsilon_{\text{Nd}}(t)$  values, with  $\epsilon_{\text{Nd}}(t)$  values for  
326 pluton □ averaging ca. -2.0 and pluton □ ca. -0.3. In spite of their different  $\epsilon_{\text{Nd}}(t)$  values, the  
327 two plutons have similar  $T_{2\text{DM}}$  (~1.1 Ga) (Table 1).

328

### 329 **5. Petrogenesis**

330

#### 331 *5.1 The Piqiang complex*

332 Though the rocks from the Piqiang complex appear to be crystal cumulates, they show  
333 enrichments in LREE and LILE (Fig. 6). This, and the  $\epsilon_{\text{Nd}}(t)$  range of -1.0 to 2.1, are  
334 comparable to some coeval ultramafic-mafic complexes in Tianshan and southern Altay (Han

335 et al., 1999, 2004, 2006; Li et al., 1998; Zhou et al., 2004). Their high Nb/La ratios  
336 (mostly >1.0), high-Ti and Fe-rich (*e.g.*, high Fe/Mn ratios of 70-116 and the presence of  
337 large magnetite deposits in the Piqiang complex) characteristics argue against a derivation  
338 from a recently metasomatized sub-continental lithospheric mantle (SCLM). This is because  
339 simple partial melting of metasomatized SCLM could not produce a Fe-rich high-Ti basaltic  
340 magma (Baker et al., 1997; Davies and Blanckenburg, 1995; Xu et al., 2001, 2003; Zhang et  
341 al., 2009). On the other hand, a sub-lithospheric mantle-derived magma, affected by crustal  
342 contamination and crystal fractionation/accumulation, could satisfy the abovementioned  
343 elemental and Nd isotopic characteristics. First, the dolerites mingled with the cumulates  
344 have trace element characteristics similar to that of OIB despite their having slightly higher  
345 Rb, Ba and Th contents and slightly negative Nb ( $Nb/La = 0.98-1.0$ ) and positive Zr-Hf  
346 anomalies (Fig. 6). Second, the weak negative Nb anomaly and higher Th/La and  $(Th/Yb)_N$   
347 ratios can be ascribed to crustal contamination. The large range of  $\epsilon_{Nd}(t)$  for the cumulates,  
348 and the deviation from the OIB field of some samples in  $(Nb/La)_{pm}$  vs.  $(Zr/Nd)_{pm}$ , Th/Yb vs.  
349 Nb/Ta,  $(Th/Yb)_N$  vs.  $(Nb/Th)_N$  and Th/La vs. Nb/La diagrams (Fig. 8), suggest variable  
350 crustal assimilation during magma ascent. Third, the Piqiang complex shares most  
351 petrographic, mineralization (large magnetite deposits) and geochemical features with the  
352 coeval Wajilitage complex in the Bachu area (Figs. 1 and 2c), which was interpreted as being  
353 formed via crystal fractionation/accumulation of an OIB-like primitive magma (Zhang et al.,  
354 2008).

355 The wide ranges of  $Mg^\#$ , MgO, SiO<sub>2</sub> and low values of some compatible trace elements  
356 (*e.g.*, Cr < 260 ppm, Ni < 180 ppm) indicate that the Piqiang complex underwent extensive

357 crystal fractionation/accumulation. Negative correlations of SiO<sub>2</sub> with Fe<sub>2</sub>O<sub>3</sub> and MgO, and  
358 positive correlation between Fe<sub>2</sub>O<sub>3</sub> and MgO, point to fractionation of olivine and  
359 clinopyroxene (Figs. 5, 8). CaO and Al<sub>2</sub>O<sub>3</sub> exhibit more scattered variation while CaO/Al<sub>2</sub>O<sub>3</sub>  
360 decreases with increasing SiO<sub>2</sub> (Figs. 5, 8). Along with the variable positive Eu and Sr  
361 anomalies (Fig. 6), these features are consistent with fractionation/accumulation of  
362 plagioclase. Negative P anomalies and positive correlation between P<sub>2</sub>O<sub>5</sub> and La indicate  
363 fractionation of apatite. In contrast, a significant positive Ti anomaly and deviation from the  
364 WPB area to the Ti end in the Ti/100-Zr-Y/3 triangle diagram (Fig. 9) argues for the  
365 cumulation of Ti-bearing magnetite. This process may result in the formation of magnetite  
366 ore bodies in the complex (Fig. 9).

367 Crystal fractionation could significantly affect REE and HFSE contents. However, some  
368 trace element ratios are insensitive to this process if the paired elements have similar partition  
369 coefficients; such ratios can be used to deduce the melting degree and the depth of the  
370 primary magma (*e.g.*, Pearce, 2008). In Zr/Nb vs. Ce/Y diagram of Hardarson and Fitton  
371 (1991) and the Sm/Yb vs. La/Sm diagram of McKenzie and O’Nions (1991) (Fig. 10), the  
372 Piqiang dolerite samples with Nb/La >1.0 plot on the partial melting line of spinel-garnet  
373 lherzolite. Taking into account that crustal contamination would have elevated the REE ratios  
374 of the dolerites (Sm/Yb and La/Sm ratios of the average continental crust are 2.1 and 11, after  
375 Rudnick and Gao, 2003), we infer that the degree of partial melting for the Piqiang complex  
376 was greater than 10% (Fig. 10).

377

378 *5.2 The Bachu mafic dyke swarm*

379 The coherent major and trace element variations for these dykes are consistent with their  
380 derivation from a common parental magma. As shown in Figure 8, the less contaminated  
381 samples have  $(\text{Nb/La})_{\text{pm}}$ ,  $(\text{Zr/Nd})_{\text{pm}}$ , Nb/Ta and Th/Yb ratios similar to those of OIB. In fact,  
382 the trace element distribution patterns for the Bachu mafic dykes are strikingly similar to that  
383 of OIB. All these features, along with previously reported Nd-Sr isotopic data (Jiang et al.,  
384 2004a), indicate a sub-lithospheric mantle source for these dykes (Zhang et al., 2008).  
385 Several samples deviate from OIB toward the crustal end members (possibly the lower crust,  
386 see the Th/La vs. Nb/La diagram in Fig. 9), indicating variable assimilation of crustal  
387 materials.

388 Their low contents of compatible elements (e.g., Cr <60 ppm, Ni <70 ppm) and MgO ( $\text{Mg}^{\#}$   
389 = 33-38) suggest that the Bachu mafic dykes have undergone extensive crystal fractionation  
390 either in the magma chamber and/or during magma ascent. The decreases in MgO and  $\text{Fe}_2\text{O}_3$   
391 with increasing  $\text{SiO}_2$  (Fig. 5) indicate crystal fractionation of olivine; the negative correlation  
392 between  $\text{SiO}_2$  and  $\text{CaO/Al}_2\text{O}_3$  (Fig. 8) suggests the involvement of clinopyroxene  
393 fractionation. The variable negative P anomaly and the positive correlation between La and  
394  $\text{P}_2\text{O}_5$  are consistent with fractionation of apatite (Fig. 8). Decreases in  $\text{TiO}_2$  with increasing  
395  $\text{SiO}_2$ , and the positive or negative Ti anomalies, argue for fractionation/accumulation of Ti-Fe  
396 oxides (titanomagnetite), which is confirmed by the presence of Ti-Fe oxides observed in thin  
397 sections. Weak Eu anomalies and slight depletion of Sr in trace elements spider diagrams  
398 argue for minor crystal fractionation of plagioclase.

399 Although the Bachu mafic dykes and the Piqiang complex were likely derived from a  
400 similar OIB-like mantle source, they have distinguishable geochemical characteristics.

401 Compared to the Piqiang dolerites, the Bachu mafic dykes are more enriched in LREE than  
402 HREE with higher  $(La/Yb)_N$  (Fig. 6), have a more alkaline affinity given their higher Nb/Y  
403 and Zr/TiO<sub>2</sub> ratios (Fig. 3), and have higher  $\epsilon_{Nd}(T)$  (Table 1). These differences could be  
404 attributed to different degrees of partial melting and different amounts of crustal assimilation.  
405 Among the Piqiang dolerites, Bachu dolerites and the basalts, the Bachu dolerites have the  
406 highest Ti/Y (~500 for Piqiang dolerites, 600-1000 for Bachu dolerites and 400-600 for  
407 basalts),  $(La/Yb)_N$ ,  $(Sm/Yb)_N$  and Nb/Y ratios. Figure 10 shows that the primitive magma for  
408 the Bachu dolerites likely originated from 3~5% partial melting of a garnet-bearing  
409 asthenospheric mantle source. In the  $(Tb/Yb)_N$  vs.  $(Yb/Sm)_N$  diagram (after Zhang et al.,  
410 2009, figure not presented), the less contaminated samples plot on the ~5% degree partial  
411 melting line of garnet-spinel lherzolite. We thus suggest that the Bachu dolerites formed by  
412 ~5% partial melting of a sub-lithospheric mantle source.

413

### 414 *5.3 The Halajun A-type granites*

415 Two main models exist for the origin of A-type granites. (1) Fractional crystallization of  
416 mafic rocks, with or without substantial assimilation of crustal rocks. This model requires  
417 extreme fractionation, which may be possible by magmatic differentiation in one or several,  
418 progressively shallower magma chambers, prior to the final emplacement of the magma  
419 (Bonin et al., 2004, 2007). (2) Partial melting of pre-existing rocks induced by magmatic  
420 underplating. Possible sources of crustal melts include metaigneous or metasedimentary  
421 lower or mid-crustal rocks and juvenile mafic underplates (Whalen et al., 1987; Frost et al.,  
422 2001; Martin, 2006), variably fractionated and hybridized by repeated injections of mantle



423 magmas.

424 Partial melting of pre-existing hornblende-bearing calc-alkaline granitoids and  
425 metasedimentary rocks (e.g., the Archaean TTG rocks or granulites in Tarim, Hu et al., 2000)  
426 at temperatures >900 °C and pressures <4 kbar (e.g., Patiño Douce, 1997) is excluded for the  
427 formation of the Halajun granites due to their high Nb/La, low Y/Yb and Yb/Ta ratios and  
428 relatively high  $\epsilon_{\text{Nd}}(\text{T})$  and juvenile  $T_{\text{DM}}$  ( $T_{\text{DM}} = 0.9\text{-}1.6$  Ga and  $T_{2\text{DM}} = 1.1\text{-}1.2$  Ga , Table 1  
429 and Fig. 11) (Eby, 1990). On the other hand, no mafic microgranitoid enclaves (MME) have  
430 been found in the Halajun granites, making the mixing of mantle-derived mafic magma and  
431 crustal-derived silicic melts unlikely.

432 Some A-type intrusions in the Emeishan LIP have been ascribed to the melting products  
433 of gabbroic cumulates recently underplated in the lower crust (Luo et al., 2007; Xu et al.,  
434 2008), and this model has been supported by experimental work (e.g., Hay and Wendlandt,  
435 1995). The  $\epsilon_{\text{Nd}}(\text{T})$  values and most geochemical features of the Halajun granites also favor a  
436 derivation by melting of a relatively young, juvenile basaltic crustal component (e.g., high  
437 Nb/La, low Y/Yb and Yb/Ta ratios, Fig. 11) rather than an early Precambrian mafic lower  
438 crust (e.g., Hu et al., 2000). However, several lines of evidence argue against this model. (1)  
439 Geochemical modeling shows that Zr vs. Sr and V variations are useful for discriminating  
440 between fractional crystallization and melting processes (Peccerillo et al., 2003). Silicic  
441 magmas derived from batch melting of mafic rocks generally have Sr >100 ppm and V >50  
442 ppm (Zhong et al., 2007). However, extensive crystal fractionation of a mafic primitive  
443 magma could quickly decrease Sr and V contents in the early stage and quickly increase Zr  
444 contents in the late stage in the residue melts because of their very different partition

445 coefficients (*i.e.*,  $D_{Zr} = 0.05$ ;  $D_V = 4.0$ ;  $D_{Sr} = 2.5$ ) (Peccerillo et al., 2003). Both plutons □  
446 and □ have very low Sr (<100 ppm for pluton □ and <10 ppm for pluton □) and V (<10 ppm)  
447 contents while their Zr ranges from 164 ppm to 367 ppm (Table DR2). Therefore, the silicic  
448 magma could not have been generated by a single-stage partial melting of mafic rocks  
449 concurrent with crystal fractionation. (2) No inherited zircon cores have been found in CL  
450 images and by SHRIMP analyses.

451 The Halajun A-type granites are spatially, temporally and geochemically closely related  
452 to the Piqiang ultramafic-mafic complex. Along the northern margin of Tarim, there are  
453 several ca. 275 Ma ultramafic-mafic-(syenites) intrusive complexes and voluminous mafic  
454 dykes (Rui et al., 2002; Yang et al., 2007; Zhang et al., 2008; this study). Syenites in the  
455 Wajilitage complex were interpreted to be the products of intensive fractionation of OIB-like  
456 basaltic magma (Li et al., 2001; Yang et al., 2007; Zhang et al., 2008). As the Halajun  
457 granites share very similar geochemical compositions with the syenites in the Wajilitage  
458 complex, except that they are more evolved, it is reasonable to suggest that they were the  
459 products of intensive fractionation from the same primitive mafic magma as the Piqiang  
460 complex and the Bachu mafic dykes, although there exists a clear chemical “Daly gap”  
461 between the mafic rocks and the granites. Peccerillo et al. (2003) demonstrated that with a  
462 steady fall in temperature, fractionating magmas generally pass rapidly through the  
463 intermediate stages, *i.e.*, producing relatively small amounts of intermediate melts. On the  
464 other hand, in a fractionating, continuously fed magma chamber, the silicic melts will occupy  
465 the top of the reservoir and the mafic magmas will pond at the bottom. The silicic materials  
466 sitting at the top of the chamber would be preferentially emplaced into the crust to form

467 A-type granites, leaving part of the mafic magma at depth and this magma crystallizes to  
468 form the mafic to ultramafic cumulates. The presence of strong aero-magnetic anomalies at  
469 Atushi, Bachu and Akesu, in combination with the occurrences of ultramafic-mafic complexes  
470 and mafic dyke swarms in the Piqiang and Bachu areas, strongly argue for an  
471 ultramafic-mafic intrusive belt along the north margin of the Tarim Block (Rui et al., 2002).  
472 In contrast to the voluminous mafic rocks along this belt, the volume of granites is relatively  
473 small, suggesting that crystal fractionation is a reasonable model because only small amounts  
474 of granitic magma could be generated from mafic magmas by intensive fractionation.

475 A-type granites derived from crystal fractionation of basaltic magma generally have high  
476 Zr saturation temperatures ( $T_{Zr}$ , *e.g.*, Zhong et al., 2007). The calculated  $T_{Zr}$  values for the  
477 Halajun plutons □ and □ are 800-830°C and 780-860°C, respectively (after Watson and  
478 Harrison, 1983), which are significantly lower than that of basaltic magma-fractionated  
479 A-type granites within the Emeishan LIP (Zhong et al., 2007). We note that the Halajun  
480 plutons are the products of extreme fractionation. The significant Zr troughs in the trace  
481 elements distribution patterns (Fig. 6) indicate zircon crystallization from the magma.  
482 However, the coeval syenites at Bachu may represent the least fractionated primary magma  
483 for the A-type granites because the Bachu syenites and the Halajun A-type granites share  
484 common geochemical signatures (Zhang et al., 2008).  $T_{Zr}$  values for the Bachu syenites range  
485 from 890°C to 1010°C, consistent with those of basaltic magma-fractionated A-type granites  
486 in the Emeishan LIP (Zhong et al., 2007).

487 The small but distinguishable difference in Nd isotopic compositions between the  
488 granites and the Piqiang complex probably resulted from crustal (especially the Archaean

489 crust) contamination during differentiation as negative Nb anomalies have been observed in  
490 several samples from this pluton (08KT02-2, 3, 4).

491 Significant fractional crystallization may have been associated with the formation of the  
492 Halajun granites. This is evidenced by the strong depletions in Ba, Sr, P, Ti and Eu as shown  
493 in the spider-diagrams and REE patterns (Fig. 6). The course of crystal fractionation from  
494 mafic magma to intermediate and then to silicic magma is consistent with that of the syenites  
495 in the Wajilitage complex (Zhang et al., 2008). Nevertheless, fractionation of potassic  
496 feldspar may have been more prominent than fractionation of plagioclase in the late stage as  
497 suggested by the significant Ba depletion, especially for pluton II (Fig. 6).

498

#### 499 *5.4 The Permian basalts*

500 Despite their relatively large SiO<sub>2</sub> range (44.1% to 51.2%), the Permian basalts show a  
501 narrow range of REE and trace element compositions (Fig. 6). Their high ratios of Ti/Y  
502 (550-660) and La<sub>N</sub>/Yb<sub>N</sub> (7-8), and negative  $\epsilon_{Nd}(t)$  (-2.1 to -2.6), preclude the basalts being  
503 derived from a depleted mantle source. Due to their spatial and temporal relationships with  
504 the Bachu mafic dyke swarm (Fig. 1), the basalts possibly shared the same asthenospheric  
505 mantle source as that of the mafic dykes (Zhou et al., 2009). In this scenario, the negative  
506  $\epsilon_{Nd}(t)$  values of the basalts require significant crustal assimilation. However, the low SiO<sub>2</sub> and  
507 high MgO (up to 8.0-8.7%), Fe<sub>2</sub>O<sub>3</sub> and TiO<sub>2</sub> contents of some analyzed samples (*e.g.*,  
508 samples with SiO<sub>2</sub> below 45%) argue against significant addition of crustal material into the  
509 primitive magma. Therefore, the negative  $\epsilon_{Nd}(t)$  (-2.1 ~ -2.6) values may largely reflect the  
510 Nd isotopic compositions of the magma source. A feasible source satisfying all the above

511 elemental and Nd isotopic characteristics is a long-term enriched lithospheric mantle (Farmer  
512 et al., 2003; Comin-Chiaramonti et al., 1997; Jiang et al, 2004b). We note that the analyzed  
513 samples in this study have moderate negative Nb anomalies ( $Nb/La = 0.6-0.8$ ), which could  
514 have resulted from subduction-related metasomatism during the late Mesoproterozoic to early  
515 Neoproterozoic rather than during the Palaeozoic because (1) an active continental margin  
516 may have existed along the northern margin of the Tarim Block during the Mesoproterozoic  
517 to earliest Neoproterozoic (Liu et al., 1996; Chen et al., 2004; Zhang et al., 2009); (2) during  
518 the Palaeozoic, a passive continental margin appears to have persisted along the northern  
519 margin of the Tarim Block (see detailed discussions in Rui et al., 2002); and (3)  
520 Mesoproterozoic  $T_{DM}$  of the basalts (1.3 Ga - 1.5 Ga) may indicate the timing of the  
521 metasomatism.

522 As shown in Figures 5 and 8,  $Al_2O_3$ , CaO, MgO and  $TiO_2$  decrease with increasing  $SiO_2$   
523 and MgO and  $CaO/Al_2O_3$  are negatively correlated with  $Fe_2O_3$  and  $SiO_2$ , respectively. All  
524 these features indicate that olivine and plagioclase, rather than clinopyroxene, were the  
525 dominant phases of the crystal fractionation.

526 The basalts have very coherent Zr/Nb (10-11) and Ce/Y ( $\sim 2.0$ ) ratios. In the Zr/Nb vs.  
527 Ce/Y diagram (Fig. 10a), the basalts plot on the primitive garnet-spinel lherzolite partial  
528 melting line with ca. 5-10% melting degrees, which is consistent with the La/Sm vs. Sm/Yb  
529 diagram (Fig. 10b). Thus, we conclude that the basalts were derived from 5-10% partial  
530 melting of the lower part of a time-integrated enriched lithospheric mantle source and  
531 underwent intensive crystal fractionation of olivine and plagioclase.

532

## 533 **6. Geodynamic implications**

534

### 535 *6.1 Age data of the Permian igneous rocks in the Tarim Block*

536

537 Table 2 lists all the reliable age data for the Permian igneous rocks in the Tarim Block  
538 (mostly in northwestern part of Tarim; see Fig. 1). These ages indicate that despite their  
539 compositional diversity, all the intrusive rocks (i.e., mafic dykes, ultramafic-mafic complexes,  
540 syenite dykes and granites) were formed coevally at *ca.* 275 Ma. Nonetheless, the extrusive  
541 rocks display two age peaks, *i.e.*, *ca.* 290 Ma and *ca.* 275 Ma. This scenario is similar to that  
542 in the Sangtanghu and Tuha basins north of Tianshan (Zhou et al., 2006; Zhang et al., 2010).  
543 The age data, in combination with their spatial distribution shown in Figure 1, indicate that  
544 the intrusive and extrusive rocks in northern Tarim are temporally and spatially related to  
545 each other, with the exception of an earlier phase of basaltic eruptions at *ca.* 290 Ma, which  
546 could be the early phase of the same tectono-magmatic event (see discussions below).

547

### 548 *6.2 Lithospheric and sub-lithospheric mantle melting above a common mantle plume*

549

550 The Early Permian magmatic rocks in the study region display a great variety in lithology  
551 and petrology, ranging from ultramafic, basalt, dolerite, syenite to granite and all of them  
552 exhibit typical within-plate affinities (Figs. 7c and 9). Despite the different petrogenetic  
553 processes associated with them, these diverse rocks share some common features. (1) Except  
554 for minor *ca.* 290 Ma basalts, they were largely emplaced simultaneously at *ca.* 275 Ma  
555 (Table 2). As mentioned above, the two pulses of basalt eruption in Tarim are similar to that

556 in the Tuha and Santanghu basins just north of Tianshan (Fig. 1). According to the regional  
557 geology (Xijiang BGMR, 1993; Zhou et al., 2006), the volume of the ca. 290 Ma basalts is  
558 much smaller than that of the ca. 275 Ma basalts. Moreover, voluminous mafic dykes,  
559 ultramafic-mafic complex, syenite and A-type granites are temporally and spatially related  
560 with the ca. 275 Ma basaltic eruptions (Fig. 1). Therefore, we suggest that the Permian  
561 igneous activity peaked at ca. 275 Ma after a minor ca. 290 Ma basaltic event that possibly  
562 represented the earlier phase of the same Permian igneous activity. (2) the Permian igneous  
563 events, with granites, syenites, and ultramafic rocks forming igneous complexes, were closely  
564 related in space (Fig. 1). (3) These magmas may have been derived from a common  
565 sub-lithospheric source, with the exception of some basalts which may have originated from  
566 the lithospheric mantle. It is therefore reasonable to conclude that the generation of these  
567 diverse rocks was governed by a common geodynamic process. In particular, such features  
568 are typical of large igneous provinces (Coffin and Eldholm, 1994; Bryan and Ernst, 2008;  
569 Ernst et al., 2008).

570 As pointed out by Zhang et al. (2010), geochronological and stratigraphic studies of the  
571 Permian basalts in Tarim and surrounding areas show the likely existence of a Permian large  
572 igneous province in Tarim and the western part of the CAO (i.e., the Tarim LIP). This  
573 igneous activity would have occurred ca. 15 My before the Emeishan LIP in southwestern  
574 China (Zhou et al., 2002, 2008; Xu et al., 2004) and 25 My before the ca. 250 Ma Siberian  
575 traps (Reichow et al., 2009). Such a sudden fare up of plume activity in the Permian may  
576 represent the early stage of the dipolar Pangean and SW Pacific superplumes due to  
577 circum-Pangea subduction and mantle avalanches (Li and Zhong, 2009). The occurrence of

578 diverse igneous rocks during the Permian represents the products of this plume event.

579 We illustrate in Figure 12 a possible model for the formation of the spatially and  
580 temporally related oxide-bearing ultramafic-mafic-(syenite) complexes (e.g., the Piquang,  
581 Wajilitage and Mazaertage complexes), mafic dyke swarms, A-type granites (syenites) and  
582 basalts along the northern margin of the Tarim Block, based in part on the ideas developed by  
583 Zhou et al. (2008) and Xu et al. (2008) for the deposits and intrusive rocks in the Emeishan  
584 LIP. In this model, the Fe-rich magmas that gave rise to the Fe-Ti-rich rocks and their ore  
585 deposits were probably derived from a sub-lithospheric mantle source, which was enriched  
586 not only in Fe but also in incompatible elements, and most possibly, in volatiles (H<sub>2</sub>O and  
587 CO<sub>2</sub>) (Zhou et al., 2008). Because of the dissolved volatiles, this magma had a relatively low  
588 density and passed through the density discontinuities and reached the lower to middle crust.  
589 Crystal fractionation in one or several progressively shallowing magma chambers, prior to  
590 the final emplacement of the magma, took place during this passage (Bonin et al., 2007).  
591 Crystal fractionation produced the magma that was parental to both the Fe-Ti-rich ultramafic  
592 complexes and the A-type granites. The formation of the ore minerals was linked to the  
593 formation of an unusually Fe-rich magma but the trigger for this process needs further  
594 investigation.

595 Unlike the ultramafic-mafic complexes, the basalts in Tarim were likely derived from  
596 partial melting of a time-integrated enriched lithospheric mantle source. The emplacement of  
597 enriched CLM-derived magmas before the sub-lithospheric mantle-derived magmatism is  
598 consistent with lower melting temperatures of enriched components compared to the “dry”  
599 mantle (Turner et al., 1996; Xu, 2001). The trigger for the melting of the enriched CLM



600 components may be related either to convective heating associated with underplating of mafic  
601 magmas from the sub-lithospheric mantle, or to conductive heating associated with an  
602 upwelling plume and lithospheric thinning. Thus, both the extrusive and intrusive rocks in  
603 Tarim could have genetically been related to a Permian upwelling mantle plume (Fig. 12).

604

## 605 **7 Conclusions**

606

607 Based on new data as well as existing information on Permian magmatic activities in  
608 Tarim, we draw the following conclusions.

609 (1) The oxide-bearing ultramafic-mafic-(syenite) complexes (Piqiang, Wajilitage and  
610 Mazhaertage), a mafic dyke swarm, and A-type granites were all emplaced at ca. 275 Ma  
611 within the interior of the Tarim Block. The voluminous basalts in the Tarim, Jungar,  
612 Santanghu and Tuha basins are spatially and temporally related to the intrusive rocks. All  
613 these igneous rocks were the likely products of a Permian mantle plume (the Tarim plume).  
614 The ca. 290 Ma basaltic eruptions in both the Tarim basin and the region north of Tianshan  
615 may represent an earlier phase of the same large igneous event.

616 (2) Although the intrusive and extrusive rocks are spatially and temporally related to each  
617 other, they were derived from different mantle sources. The intrusive rocks and mafic dykes  
618 in Tarim were formed via crystal fractionation in one or several, progressively shallowing  
619 magma chambers, from a sub-lithospheric mantle-derived primitive basaltic magma. In  
620 contrast, the basalts in Tarim were derived from a long-term enriched lithospheric mantle  
621 source.

622

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629 #xxx.

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907

908 **Fig Captions**

909 **Fig. 1** A geotectonic sketch map of the Tarim Block and part of the Central Asian Orogenic  
910 Belt (CAOB) in Xinjiang, showing the distribution of the Permian basalts,  
911 ultramafic-mafic intrusions, mafic dykes and A-type granites, with their ages shown in  
912 Ma (grey, blue and blank stars represent the locations of dated mafic-ultramafic  
913 intrusions, basalts and mafic dykes, and granitic intrusions, respectively). Insert in the  
914 lower-left corner shows the locations of the CAOB and the main crustal blocks to the  
915 north and south of the CAOB. The locations of the studied Piqiang complex, Halajun  
916 granite plutons, Bachu mafic dyke swarm, and basalts are shown. Discussions about the  
917 dividing line between the two mantle provinces (or domains) can be found in Zhang et  
918 al. (2010). KF1 – the Kupukuziman Formation at the Yingan section; KF2 – the  
919 Kapaizileike Formation at the Kaipaizileike section; M – the Mazaertage complex; W –  
920 the Wajilitage complex; P – the Piqiang complex.

921 **Fig. 2** (a) Simplified geological map of the Piqiang complex and the Halajun granite plutons  
922 (plutons I and II); (b) detailed geological map of part of the Piqiang complex showing  
923 the rock types of the complex; (c) detailed geological map of the mafic dyke swarm  
924 around the Wajilitage complex showing the broadly radiating distribution (based on the  
925 1:5000 geological mapping by local geologists).

926 **Fig. 3** Concordia plots of U-Pb zircon results of the studied intrusive rocks: (a) the Piqiang  
927 complex, (b) the Halajun granite pluton □, (c) - the Halajun granite pluton □.

928 **Fig. 4** Binary Harker diagrams ( $\text{SiO}_2$  versus  $\text{Al}_2\text{O}_3$ ,  $\text{TiO}_2$ ,  $\text{Fe}_2\text{O}_3$ ,  $\text{CaO}$ ,  $\text{MgO}$ ,  $\text{Na}_2\text{O}$ ,  $\text{K}_2\text{O}$ ,  
929  $\text{P}_2\text{O}_5$  and Zr versus Rb, V, La and Nb) for the Piqiang complex, Bachu mafic dyke



930 swarm and basalts (see detailed discussion in the text).

931 **Fig. 5**  $\text{SiO}_2$  vs.  $\text{K}_2\text{O} + \text{Na}_2\text{O}$  (a) and  $\text{Nb}/\text{Y}$  vs.  $\text{Zr}/\text{TiO}_2$  (b) classification diagrams (after Floyd  
932 et al., 1977) and AFM ( $\text{Na}_2\text{O} + \text{K}_2\text{O} - \text{FeO}^{\text{T}} - \text{MgO}$ ) diagram (c) showing the tholeiitic  
933 trends of the Piqiang complex, the Bachu mafic dyke swarm and the Permian basalts  
934 ( $\text{FeO}^{\text{T}} = \text{Fe}_2\text{O}_3^{\text{T}}/1.111$ ).

935 **Fig. 6** Chondrite-normalized REE patterns and primitive mantle-normalized incompatible  
936 element spider-diagrams for the studied rocks. The normalization values are from Sun  
937 and McDonough (1989).

938 **Fig. 7**  $\text{Zr}$  vs  $10000\text{Ga}/\text{Al}$  discrimination diagram showing that the intermediate and felsic  
939 sub-groups from the Halajun granite plutons are A-type granites (after Whalen et al.,  
940 1987), where I = I-type, S = S-type and M = M-type granitoids (a); Nb-Y-Ce  
941 discrimination diagram indicating  $A_1$  characteristics of the Halajun granite plutons  
942 (after Eby, 1992) (b) and Y vs. Nb diagram showing their typical within-plate chemical  
943 characteristics (c) (see detailed discussions in the text).

944 **Fig. 8** Binary Harker diagrams for rocks of the Piqiang intrusive complex, Bachu mafic dykes  
945 and basalts. PM – primitive mantle, UC – upper crust, MC – middle crust, LC – lower  
946 crust, OIB – oceanic island basalt, N-MORB – normal middle ocean ridge basalt (see  
947 details in the text).

948 **Fig. 9**  $\text{Ti}/100 - \text{Zr} - \text{Y} \times 3$  triangle diagram showing the typical intraplate geochemical  
949 signatures of the Piqiang complex, the Bachu mafic dykes, and the basalts (after Pearce  
950 and Cann, 1973).

951 **Fig. 10**  $\text{Ce}/\text{Y}$  vs.  $\text{Zr}/\text{Nb}$  (a) and  $\text{La}/\text{Sm}$  vs.  $\text{Sm}/\text{Yb}$  (b) plots of the Piqiang mafic dykes, less

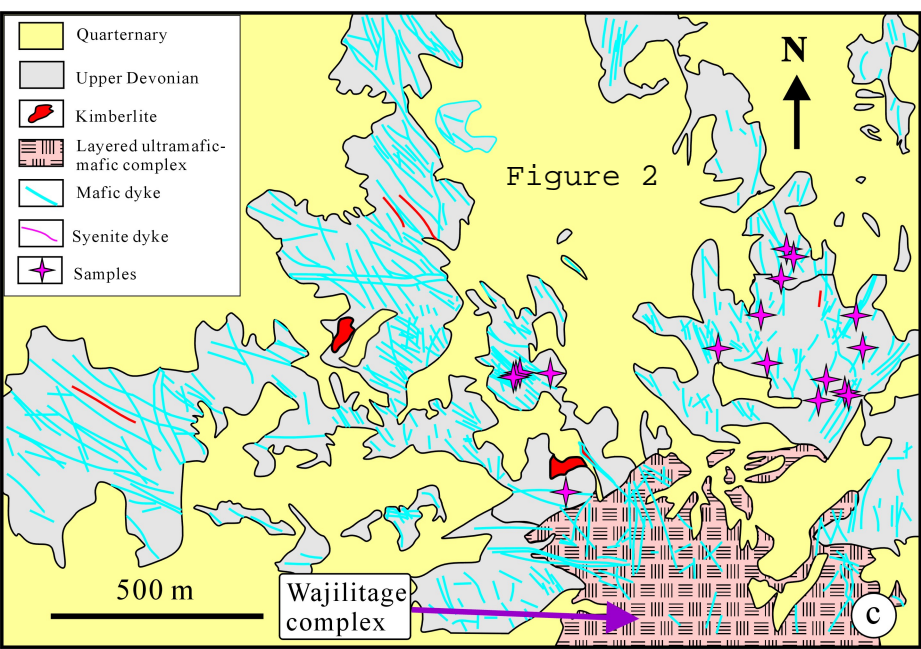
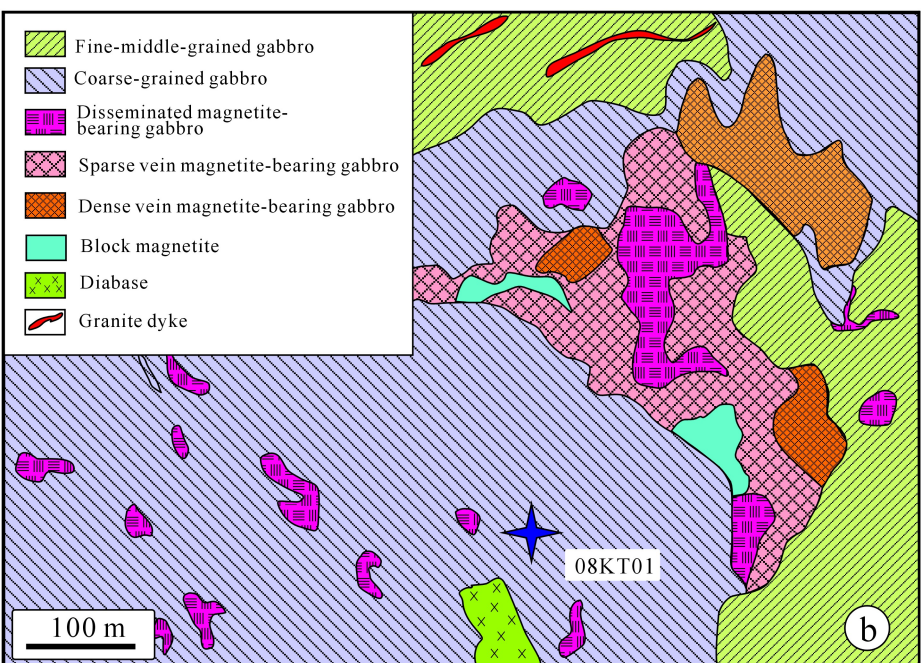
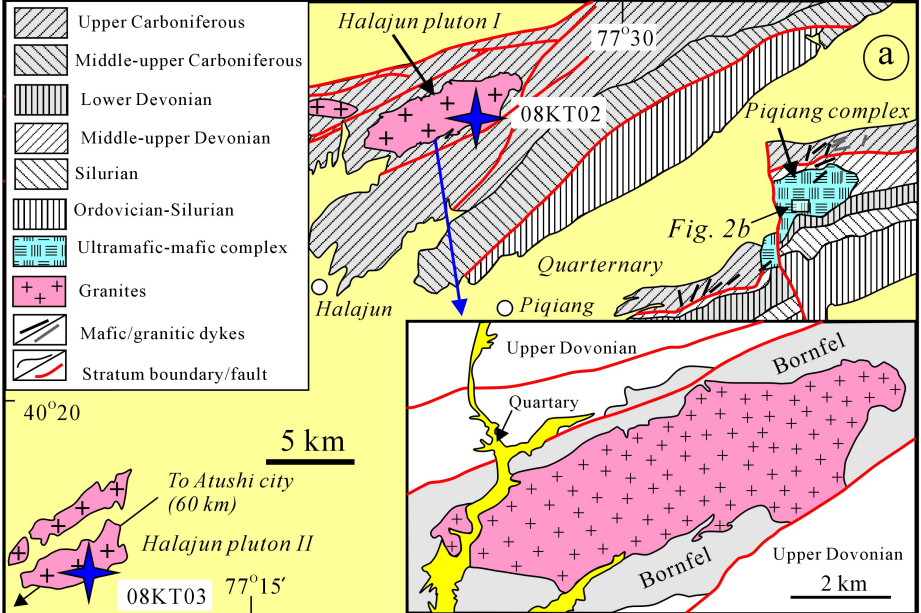
952 contaminated Bachu mafic dykes and basalts along the northern margin of Tarim.  
953 Mantle compositions: GD—depleted garnet lherzolite; GP—primitive garnet lherzolite;  
954 SD—depleted spinel lherzolite; SP—primitive spinel lherzolite. Numbers on lines  
955 refer to percentages of melt (after Hardarson and Fitton, 1991; McKenzie and O’Nions,  
956 1991).

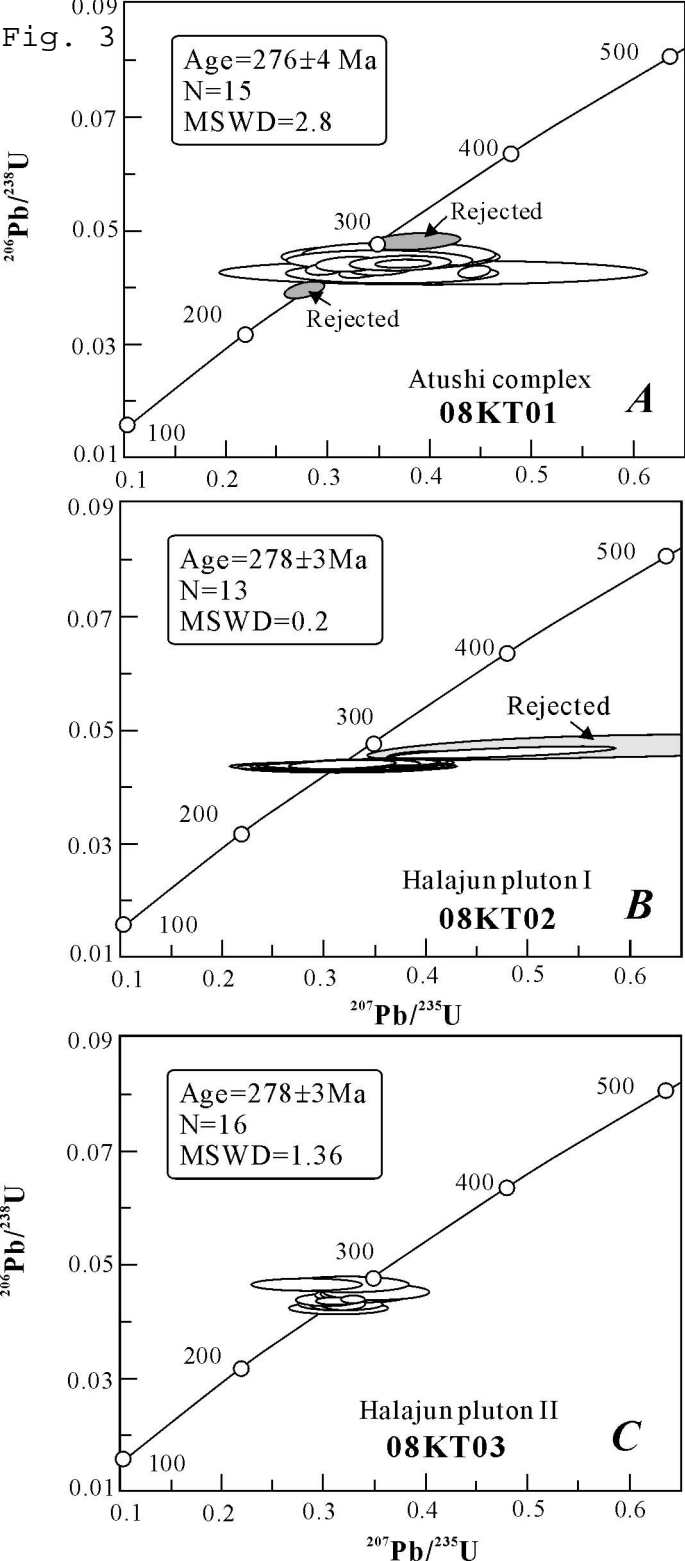
957 **Fig. 11** A Y/Yb vs. Yb/Ta diagram showing that the Halajun A-type granites plot in the OIB –  
958 like mafic magma source (after Eby, 1990).

959 **Fig. 12** A generalized model for the two magma series and associated ore deposit types  
960 within the Tarim LIP, NW China.

961







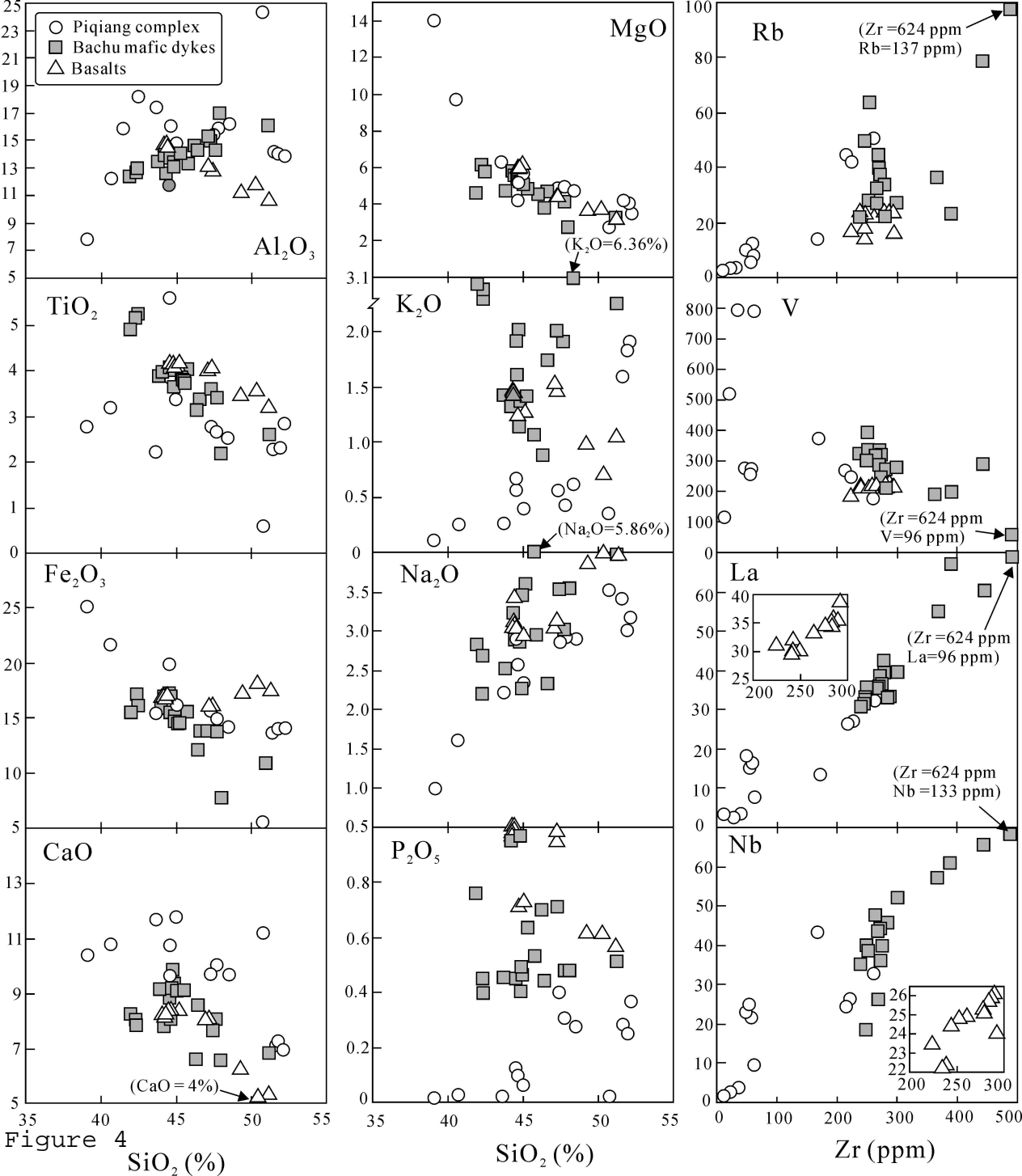


Figure 5

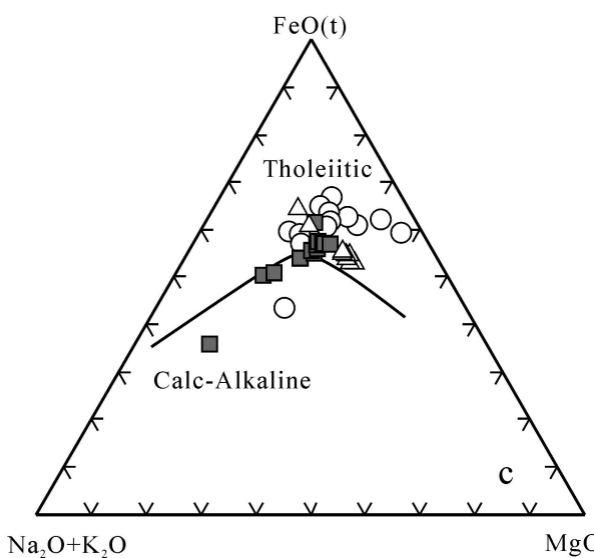
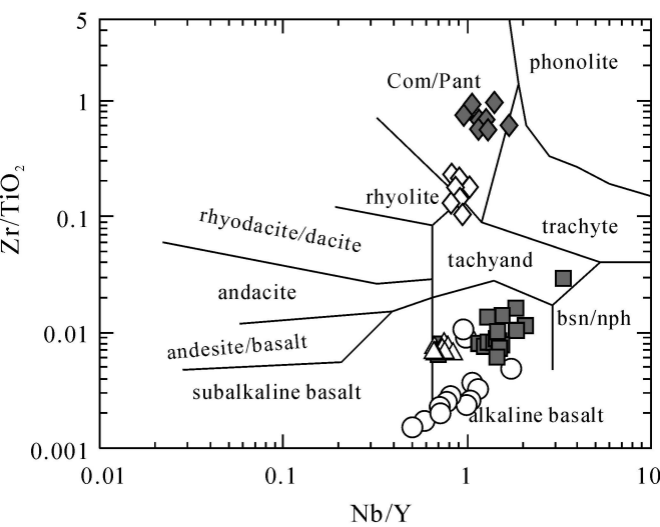
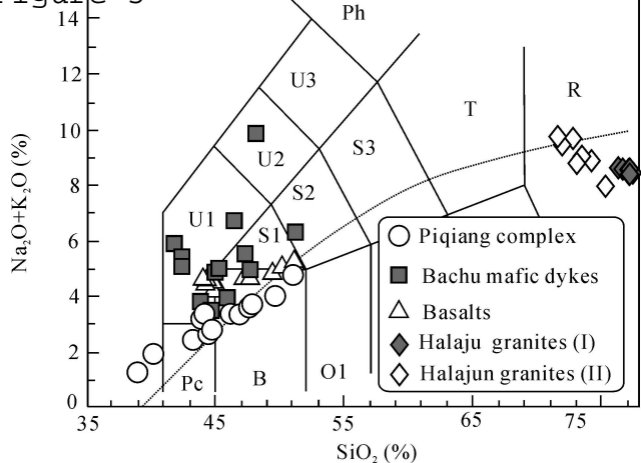
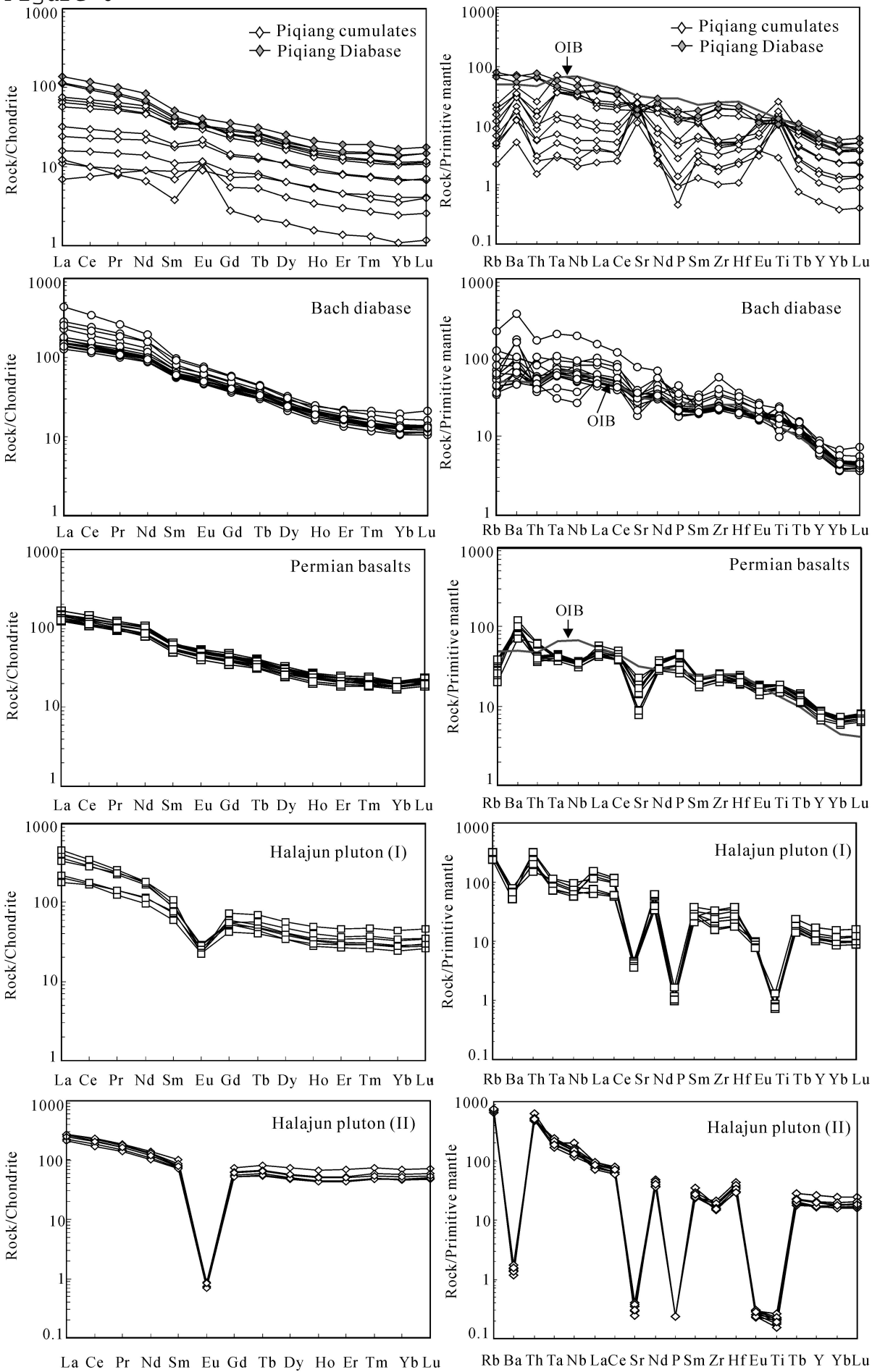


Figure 6





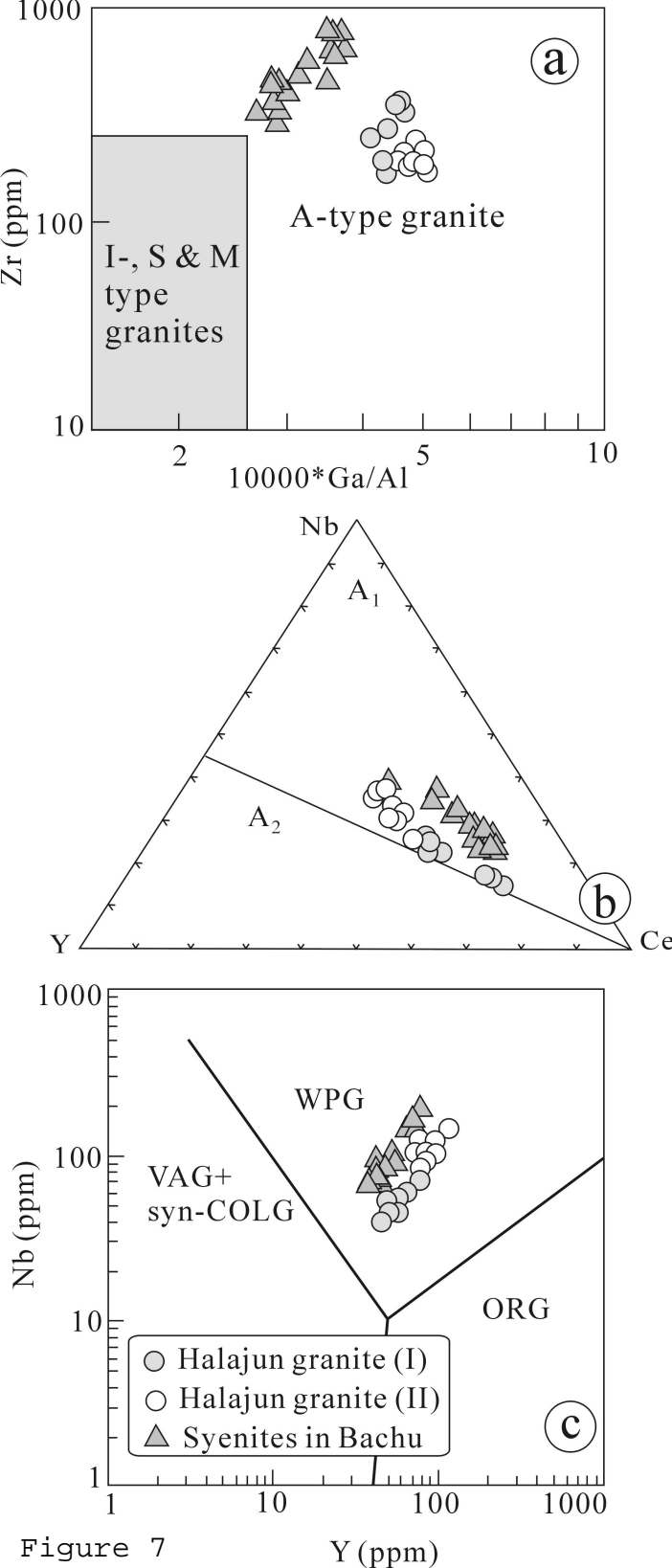


Figure 7

Y (ppm)

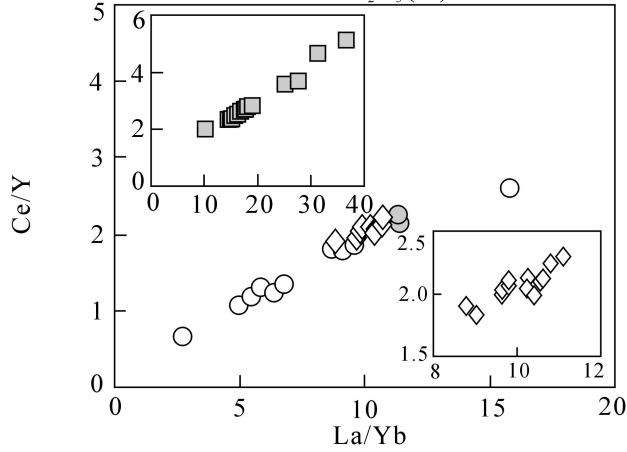
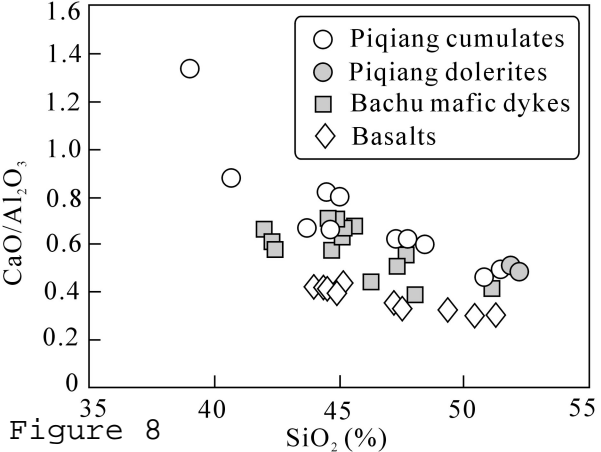
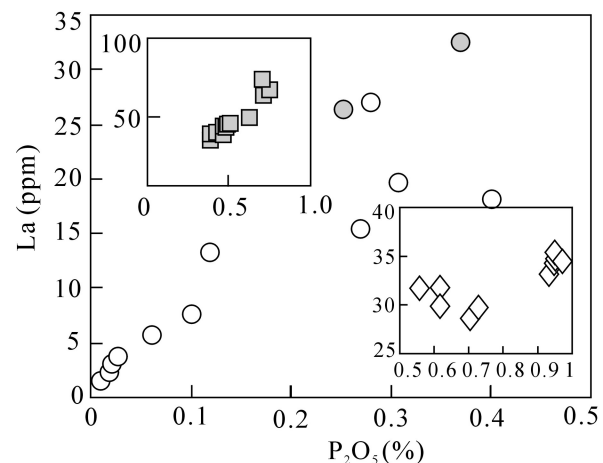
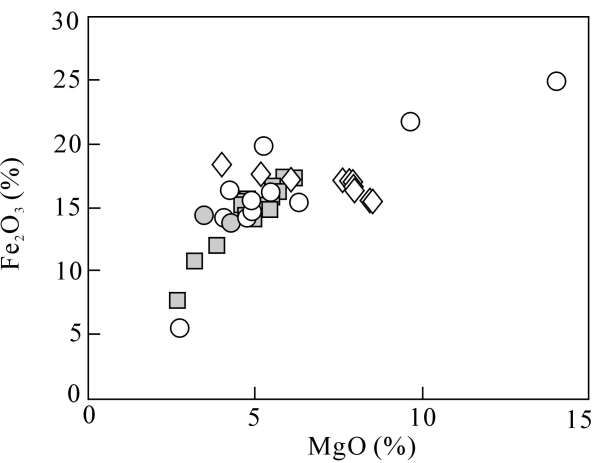
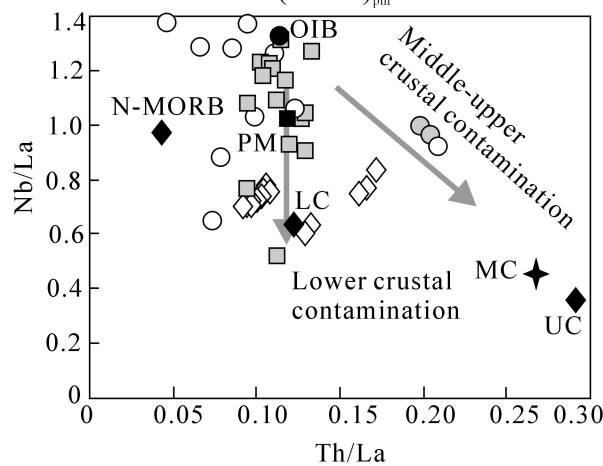
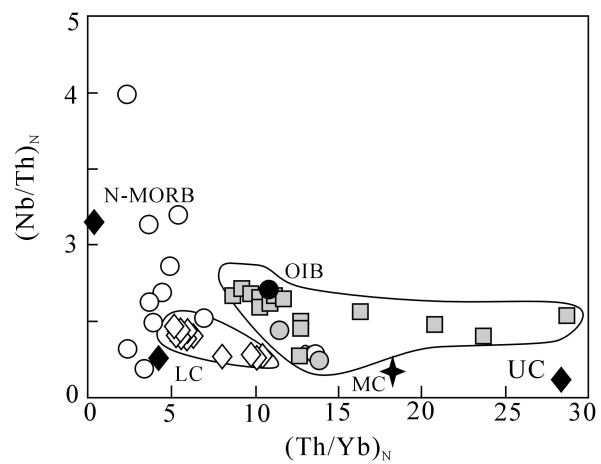
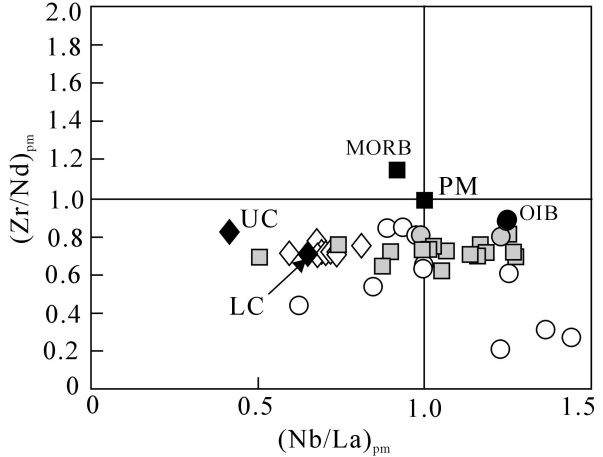
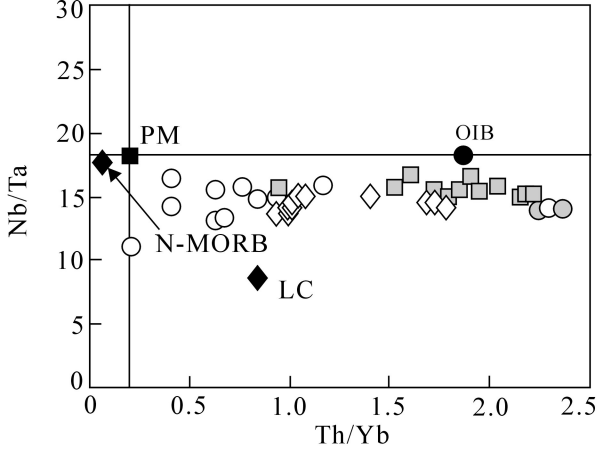
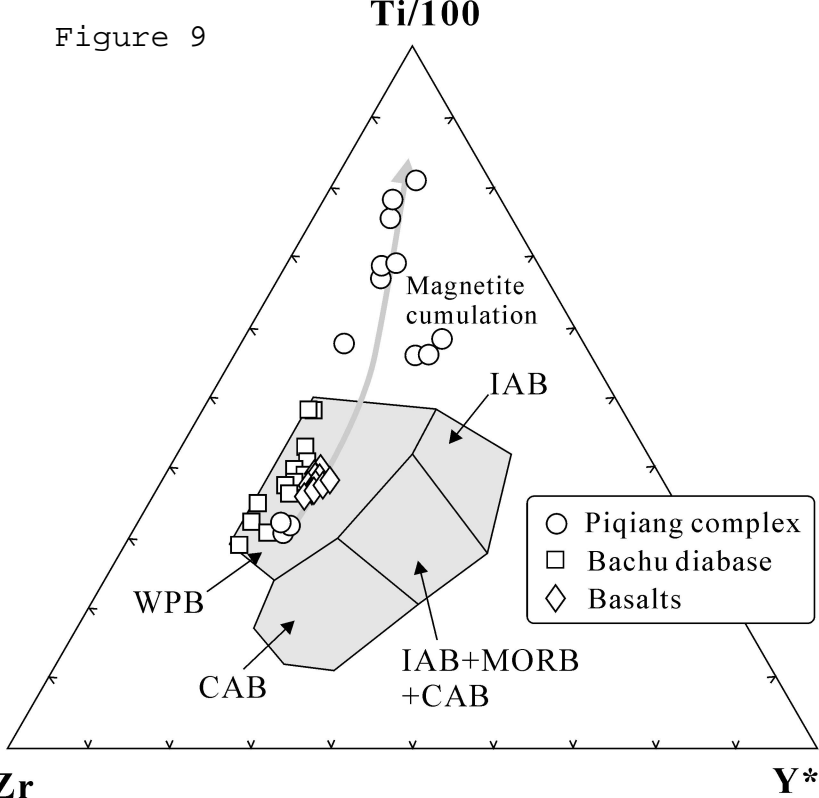


Figure 8

Figure 9



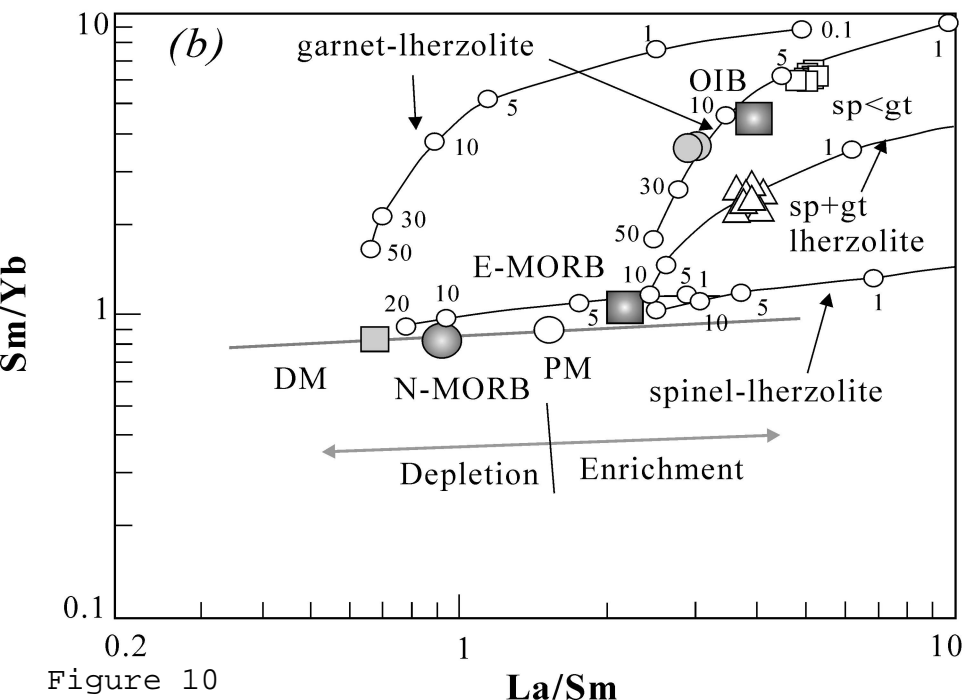
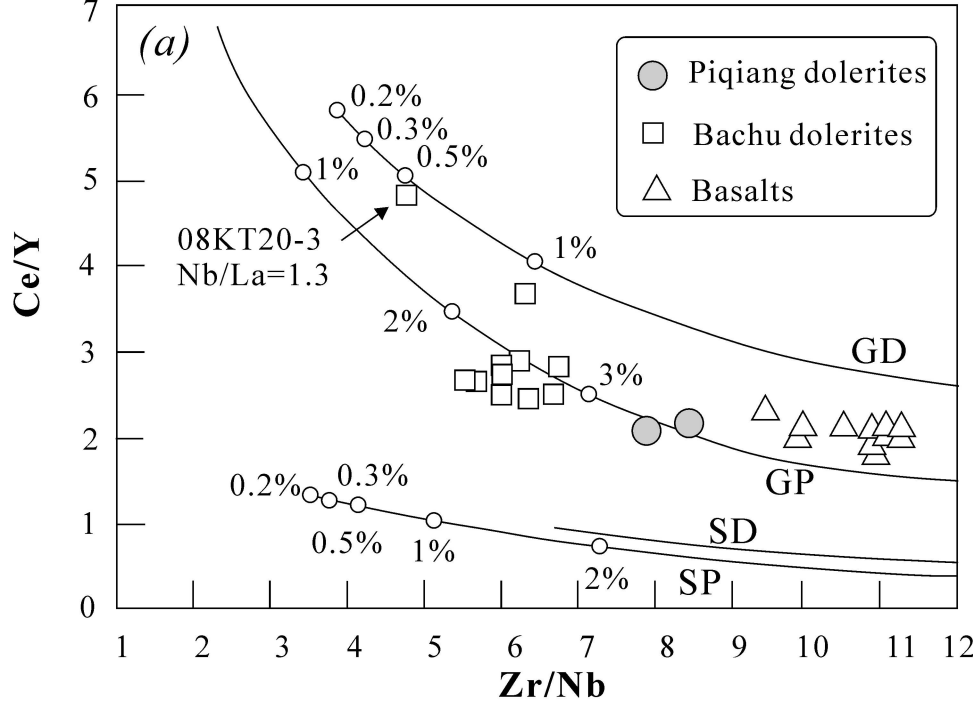
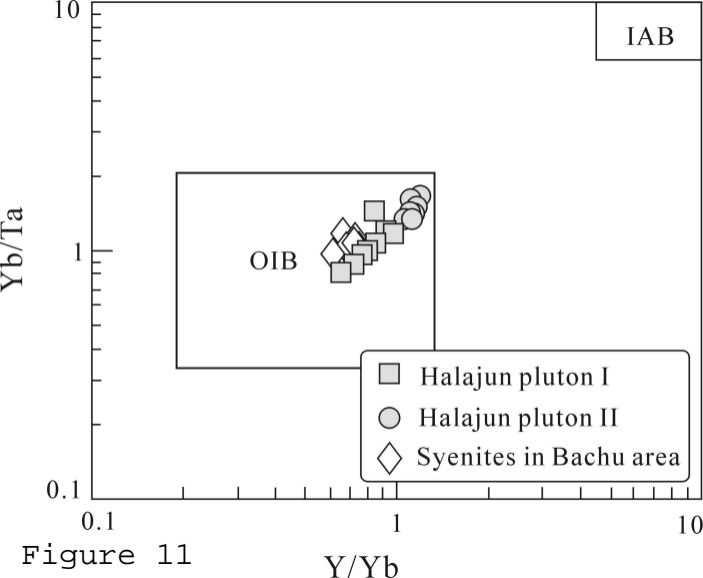


Figure 10

La/Sm



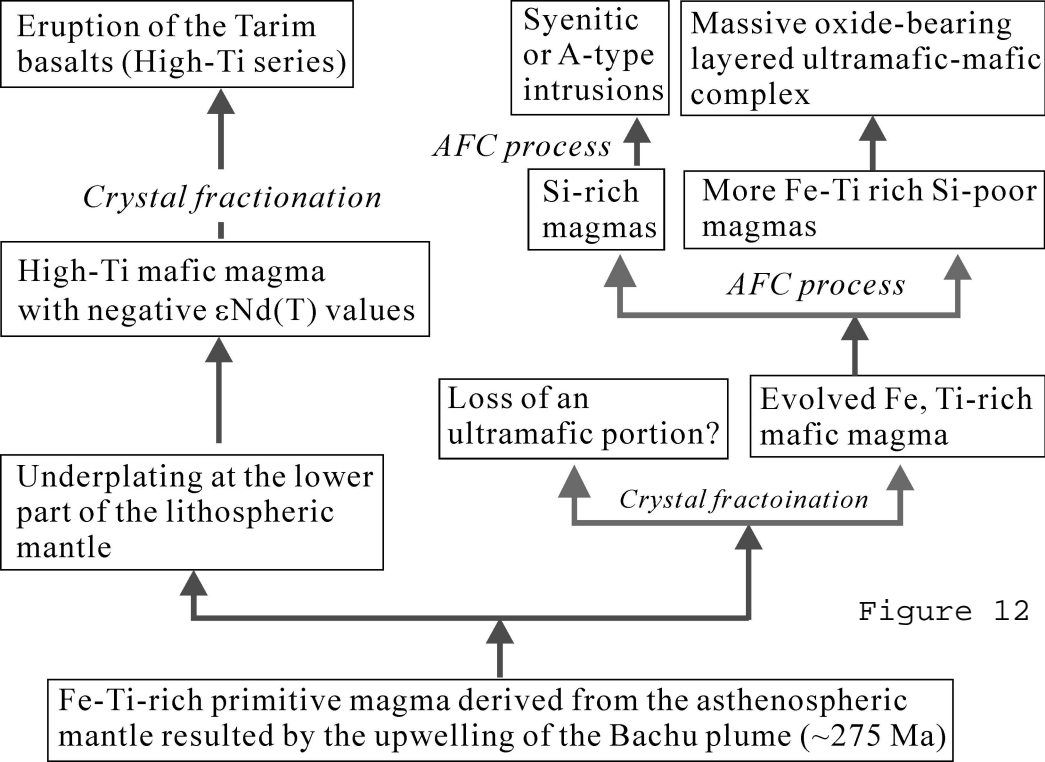


Figure 12

Table.1 SHRIMP U-Pb data for zircons from the Piqiang complex and Halajun granites

Spot	U (ppm)	Th (ppm)	Th/U	$f_{206}^{\#}$ (%)	$^{206}\text{Pb}^*/^{238}\text{U}$ ( $\pm 1\sigma\%$ )	$^{207}\text{Pb}^*/^{235}\text{U}$ ( $\pm 1\sigma\%$ )	$^{207}\text{Pb}^*/^{206}\text{Pb}^*$ ( $\pm 1\sigma\%$ )	$^{206}\text{Pb}/^{238}\text{U}$ Age (Ma) ( $\pm 1\sigma$ ) $\square$	$^{207}\text{Pb}/^{206}\text{Pb}$ Age (Ma) ( $\pm 1\sigma$ ) $\square$	Age				
<i>Piqing ultramafic-mafic complex (08KT01)</i>														
1.1	148	74	0.51	2.24	.0440	2.1	0.48	11.4	0.0785	11.2	307.6	4.9	530	210
2.1	180	121	0.70	4.42	.0437	1.3	0.34	2.7	0.0558	2.4	296.9	4.4	536	178
3.1	172	70	0.42	1.72	.0438	1.5	0.49	8.0	0.0808	7.9	272.3	4.1	316	287
4.1	61	36	0.60	4.08	.0467	1.9	0.38	20.7	0.0587	20.6	277.3	5.6	1160	223
5.1	254	246	1.00	0.77	.0434	1.3	0.30	4.1	0.0505	3.9	276.0	3.6	446	53
6.1	160	136	0.88	0.85	.0442	1.3	0.30	4.2	0.0494	4.1	276.6	4.2	1216	155
7.1	94	73	0.80	2.63	.0456	1.7	0.41	14.0	0.0648	13.9	294.1	5.6	554	449
8.1	455	524	1.19	1.65	.0438	1.6	0.37	8.1	0.0617	7.9	274.1	3.4	220	90
9.1	531	211	0.41	1.58	.0455	1.4	0.30	9.2	0.0478	9.1	279.0	3.4	165	95
10.1	115	59	0.53	0.76	.0403	1.3	0.28	5.7	0.0509	5.6	287.2	4.9	768	293
11.1	120	59	0.51	1.71	.0451	1.6	0.42	15.1	0.0682	15.0	276.6	4.3	662	169
12.1	234	225	0.99	2.28	.0435	1.6	0.41	12.0	0.0681	11.9	287.0	4.0	89	215
13.1	452	329	0.75	6.97	.0436	3.7	0.41	42.3	0.0686	42.1	255.0	3.2	238	129
14.1	150	78	0.54	2.24	.0440	2.1	0.48	11.4	0.0785	11.2	284.1	4.5	875	311
15.1	139	112	0.83	0.42	.0437	1.3	0.34	2.7	0.0558	2.4	274.5	4.4	870	246
16.1	62	39	0.66	1.72	.0438	1.5	0.49	8.0	0.0808	7.9	275.2	10.1	887	870
17.1	128	93	0.75	2.64	.0439	1.8	0.30	14.7	0.0494	14.6	276.9	4.9	169	342
18.1	168	88	0.54	0.75	.0440	1.6	0.37	5.2	0.0614	5.0	277.8	4.3	654	107
<i>Halajun granite ( pluton I, 08KT02)</i>														
1.1	80	44	0.56	2.6	.0460	3.8	0.58	29.6	0.0917	29.3	289.8	10.9	1461	557
2.1	112	56	0.51	1.0	.0434	1.6	0.30	20.4	0.0500	20.3	273.8	4.2	197	472
3.1	145	80	0.57	0.9	.0446	1.3	0.33	14.3	0.0532	14.3	281.3	3.4	338	323
4.1	90	34	0.39	1.1	.0458	1.7	0.48	19.7	0.0757	19.6	288.4	4.8	1088	393
5.1	109	61	0.58	1.0	.0435	1.8	0.32	26.7	0.0539	26.7	274.8	4.8	367	601
6.1	169	64	0.39	0.8	.0438	1.0	0.31	11.2	0.0510	11.2	276.4	2.8	239	258
7.1	155	86	0.57	1.3	.0444	1.5	0.34	11.1	0.0558	11.0	279.8	4.1	443	245
8.1	122	84	0.71	0.9	.0443	1.3	0.32	20.2	0.0530	20.1	279.4	3.6	329	457
9.1	109	73	0.69	1.0	.0436	1.5	0.30	25.8	0.0503	25.8	275.0	4.0	207	598
10.1	146	95	0.67	0.8	.0436	1.5	0.35	18.2	0.0587	18.2	275.2	3.9	556	396
11.1	115	78	0.70	1.4	.0442	1.7	0.33	15.7	0.0534	15.6	278.8	4.6	344	353
12.1	129	79	0.63	0.9	.0427	1.0	0.32	11.6	0.0537	11.6	269.6	2.8	359	261
13.1	118	80	0.70	2.46	.0442	2.3	0.33	14.7	0.0547	14.5	278.6	6.2	400	324
14.1	146	94	0.67	1.20	.0450	2.2	0.37	9.9	0.0589	9.6	284.0	6.0	565	210
<i>Halajun granite ( pluton II, 08KT03)</i>														
1.1	840	217	0.27	0.12	.0922	1.9	0.74	2.3	0.0584	1.2	569.1	10.7	546	25

2.1	188	80	0.44	0.48	.0439	2.1	0.36	4.5	0.0600	4.0	274.1	5.6	604	87
3.1	115	78	0.70	0.99	.0453	2.2	0.40	9.9	0.0648	9.6	281.0	6.3	767	203
4.1	270	91	0.35	0.29	.0449	2.0	0.37	4.1	0.0599	3.6	280.7	5.6	601	77
5.1	506	233	0.48	0.42	.0438	2.0	0.45	4.3	0.0738	3.9	269.0	5.3	1037	78
6.1	239	104	0.45	0.26	.0443	2.0	0.39	4.1	0.0645	3.5	275.3	5.5	758	74
7.1	80	30	0.39	1.11	.0431	2.4	0.48	10.3	0.0808	10.0	262.7	6.1	1216	197
8.1	197	88	0.46	0.65	.0452	2.0	0.44	7.3	0.0704	7.0	278.5	5.8	939	143
9.1	132	53	0.42	0.51	.0457	2.1	0.47	7.4	0.0745	7.1	280.5	6.0	1055	144
10.1	213	99	0.48	0.12	.0427	2.0	0.41	5.1	0.0688	4.7	263.8	5.4	894	98
11.1	240	94	0.41	0.14	.0432	2.0	0.41	2.9	0.0694	2.1	266.9	5.3	912	43
12.1	235	116	0.51	0.17	.0424	2.0	0.35	3.0	0.0594	2.2	265.2	5.3	580	49
13.1	115	46	0.41	0.64	.0432	2.1	0.45	9.2	0.0753	8.9	265.1	6.0	1076	179
14.1	260	112	0.45	0.10	.0448	2.0	0.41	6.6	0.0665	6.3	277.6	5.7	823	131
15.1	211	95	0.46	0.17	.0433	2.0	0.43	7.3	0.0713	7.1	266.8	5.6	965	144
16.1	160	86	0.55	0.00	.0457	2.1	0.49	7.7	0.0772	7.4	279.2	6.1	1126	147
17.1	340	140	0.43	0.00	.0437	2.0	0.39	6.8	0.0644	6.5	271.7	5.5	756	136

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#  $f_{206}$ : percentage of common  $^{206}\text{Pb}$  in total  $^{206}\text{Pb}$



Table 2 Chemical compositions of the Permian ultramafic-mafic complex, mafic dyke swarm, basalts and granites along the northern margin of the Tarim Block

sample	08KT 01-1	08KT 01-2	08KT 01-3	08KT 01-5	08KT 01-6	08KT 01-8	08KT 01-9	08KT 01-10	08KT 01-11	08KT 01-12	08KT 01-13	08KT 01-14	08KT 01-15
Rock type	MGB	PGB	MQGB	MGB	QGB	OMGB	MGB	MQGB	MGB	MGB	MQGB	DB	DB
<i>Major elements (%)</i>													
SiO <sub>2</sub>	43.63	50.75	47.72	40.63	51.89	39.08	44.97	48.43	44.55	44.58	47.33	51.56	52.18
TiO <sub>2</sub>	2.22	0.62	2.65	3.17	2.33	2.77	3.37	2.53	5.57	3.82	2.75	2.28	2.85
Al <sub>2</sub> O <sub>3</sub>	17.38	24.33	15.91	12.27	14.06	7.77	14.71	16.23	11.74	16.10	15.43	14.13	13.91
Fe <sub>2</sub> O <sub>3</sub>	15.39	5.52	14.89	21.62	14.06	25.04	16.18	14.10	19.81	16.28	15.68	13.87	14.29
MnO	0.12	0.05	0.17	0.18	0.16	0.24	0.16	0.16	0.22	0.14	0.18	0.14	0.19
CaO	11.65	11.19	10.02	10.80	7.21	10.41	11.83	9.75	9.66	10.74	9.69	7.06	6.91
MgO	6.28	2.73	4.90	9.63	4.13	14.02	5.60	4.71	5.27	4.23	4.89	4.24	3.50
K <sub>2</sub> O	0.26	0.35	0.42	0.25	1.83	0.11	0.40	0.61	0.67	0.57	0.55	1.60	1.91
Na <sub>2</sub> O	2.23	3.52	2.91	1.62	3.01	1.01	2.36	2.90	2.57	2.94	2.87	3.40	3.17
P <sub>2</sub> O <sub>5</sub>	0.02	0.02	0.31	0.03	0.25	0.01	0.06	0.28	0.12	0.10	0.40	0.28	0.37
LOI	0.33	0.33	0.31	0.55	0.71	0.73	0.04	0.10	0.49	0.08	0.15	1.09	0.39
Σ	99.5	99.4	100.2	100.8	99.7	101.2	99.7	99.8	100.7	99.57	99.92	99.64	99.67
Mg <sup>#</sup>	41	45	36	43	33	49	37	36	31	30	34	34	29
<i>Trace elements (ppm)</i>													
La	2.62	2.91	16.54	3.81	26.46	1.65	5.71	15.5	13.39	7.67	18.1	27.1	32.4
Ce	6.08	6.07	40.0	9.37	57.8	4.55	14.0	36.5	33.1	18.3	43.2	60.8	72.4
Pr	0.89	0.74	5.61	1.40	7.48	0.79	2.13	5.13	4.83	2.60	6.14	7.91	9.47
Nd	4.19	3.06	25.1	6.47	30.6	4.22	10.01	22.3	21.8	12.02	26.9	32.13	38.9
Sm	1.06	0.58	5.40	1.66	6.20	1.34	2.67	4.85	5.18	2.91	5.79	6.49	7.88
Eu	0.68	0.64	1.95	0.68	1.87	0.52	1.10	1.73	2.08	1.26	2.02	1.90	2.36
Gd	1.13	0.56	5.15	1.72	5.84	1.48	2.82	4.68	5.11	2.92	5.40	6.08	7.27
Tb	0.20	0.08	0.82	0.30	0.96	0.28	0.47	0.76	0.91	0.50	0.86	0.99	1.18
Dy	1.04	0.48	4.43	1.63	5.08	1.60	2.79	4.11	4.83	2.73	4.66	5.40	6.51
Ho	0.19	0.09	0.80	0.31	0.97	0.30	0.53	0.75	0.91	0.50	0.85	0.98	1.20
Er	0.49	0.23	2.06	0.74	2.56	0.76	1.34	1.98	2.42	1.30	2.20	2.57	3.14
Tm	0.07	0.03	0.30	0.11	0.38	0.10	0.19	0.28	0.37	0.18	0.31	0.38	0.48
Yb	0.41	0.18	1.81	0.69	2.32	0.60	1.17	1.78	2.30	1.12	1.91	2.39	2.87
Lu	0.07	0.03	0.29	0.10	0.37	0.10	0.17	0.27	0.37	0.18	0.29	0.37	0.45
Rb	2.79	3.07	5.62	3.22	44.63	1.43	4.29	12.08	14.16	7.59	9.72	41.98	50.62
Ga	20.3	21.0	20.8	18.4	21.7	14.2	20.2	19.8	20.7	24.5	21.1	21.7	23.4
V	522	113	270	796	265	841	961	260	377	786	269	247	176
Cr	103	17.6	16.5	260	3.02	195	171	15.8	0.41	72.5	17.5	3.67	0.53
Ni	93.1	39.7	33.4	173	23.0	163	43.8	32.1	18.3	42.5	34.4	27.6	0.07
Sc	27.0	5.54	22.8	35.2	21.8	44.9	39.9	21.0	38.1	30.6	23.1	21.4	22.0
Sr	491	653	527	310	369	235	415	502	384	491	506	322	391
Ba	88.4	147	225	85.3	506	36.5	122	239	314	206	243	375	467
Th	0.26	0.22	0.75	0.47	5.47	0.13	0.49	1.50	2.15	0.86	1.21	5.51	6.45
U	0.06	0.06	0.33	0.09	1.12	0.07	0.10	0.33	0.42	0.19	0.30	1.17	1.49
Ta	0.21	0.12	1.52	0.30	1.75	0.13	0.52	1.48	2.90	0.62	1.50	1.87	2.37
Nb	2.70	1.88	24.8	3.92	24.5	1.45	7.35	21.9	43.1	9.71	23.3	26.5	32.8

Zr	22.0	11.2	56.3	35.2	215	18.5	51.2	57.9	168	62.4	46.9	221	259
Hf	0.76	0.33	1.57	1.22	5.70	0.69	1.53	1.64	4.40	1.96	1.45	5.74	6.58
Y	4.90	2.35	22.5	7.92	25.7	7.02	13.1	20.3	25.2	13.4	23.6	27.2	33.2

sample	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT	08KT
	13-1	16-1	16-2	16-3	17-1	18-1	18-9	19-1	19-2	19-3	19-4	19-5	20-1	20-2	20-3	21-1
Rock type	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB	DB

*Major elements (%)*

SiO <sub>2</sub>	44.60	47.2	43.79	47.68	44.79	51.14	42.29	45.72	44.49	45.21	44.87	42.35	46.27	47.99	41.89	44.78
TiO <sub>2</sub>	4.05	3.59	3.87	3.41	4.02	2.59	5.15	4.01	3.83	3.81	3.98	5.22	3.12	2.15	4.89	3.63
Al <sub>2</sub> O <sub>3</sub>	14.04	14.95	13.65	14.26	13.21	16.05	13.04	13.34	12.43	14.22	13.52	13.15	14.68	16.94	12.40	13.67
Fe <sub>2</sub> O <sub>3</sub>	17.21	13.90	15.70	13.9	16.7	10.89	17.22	15.76	15.66	14.38	15.12	16.10	12.10	7.64	15.45	14.69
MnO	0.20	0.18	0.19	0.17	0.17	0.14	0.20	0.19	0.15	0.19	0.17	0.17	0.18	0.05	0.15	0.16
CaO	8.06	7.68	9.11	8.06	9.22	6.89	8.01	9.11	8.81	9.09	9.30	7.69	6.50	6.49	8.20	9.83
MgO	5.74	4.92	4.64	4.15	5.54	3.16	6.07	4.59	5.59	4.88	5.14	5.69	3.79	2.70	4.61	5.41
K <sub>2</sub> O	1.61	2.01	1.42	1.91	1.13	2.25	2.96	1.07	1.92	1.42	1.37	2.73	0.88	6.36	3.05	2.02
Na <sub>2</sub> O	2.88	3.52	2.51	3.00	2.29	3.98	2.20	2.96	2.90	3.61	3.45	2.69	5.86	3.54	2.81	2.94
P <sub>2</sub> O <sub>5</sub>	0.97	0.71	0.45	0.48	0.41	0.51	0.39	0.53	0.45	0.63	0.50	0.45	0.70	0.48	0.76	0.47
LOI	0.15	0.82	4.31	2.56	2.08	1.94	2.01	2.32	3.33	2.05	2.16	3.30	5.73	5.20	5.38	1.92
Σ	99.60	99.5	99.64	99.59	99.59	99.55	99.54	99.58	99.55	99.49	99.58	99.55	99.81	99.55	99.58	99.51
Mg <sup>#</sup>	36	37	33	33	36	33	37	33	38	36	36	37	35	37	33	38

*Trace elements (ppm)*

La	33.5	54.8	36.0	39.1	30.0	39.1	33.0	39.6	36.1	42.5	35.7	35.8	67.2	103	60.1	32.4
Ce	76.8	116.8	81.4	86.7	69.3	86.3	75.0	89.8	83.6	96.4	81.0	80.7	147	207	133	74.6
Pr	10.6	15.0	10.8	11.3	9.45	11.2	10.0	12.1	11.2	13.0	10.9	10.7	19.1	25.0	17.8	10.2
Nd	45.7	60.4	45.4	47.5	40.5	46.4	42.1	51.0	47.4	54.2	46.8	44.4	73.9	91.7	73.8	44.0
Sm	9.28	11.44	9.25	9.37	8.46	9.02	8.61	10.4	9.64	10.6	9.71	8.88	12.5	15.0	14.0	9.10
Eu	2.84	3.65	2.94	2.91	2.70	2.75	2.68	3.23	2.96	3.29	3.15	2.74	3.55	4.45	4.18	2.86
Gd	8.68	10.1	8.47	8.21	7.55	7.95	7.36	9.23	8.53	9.23	8.84	7.88	9.93	12.0	11.7	8.17
Tb	1.36	1.45	1.27	1.24	1.16	1.19	1.11	1.37	1.26	1.39	1.32	1.18	1.33	1.68	1.64	1.24
Dy	7.39	6.98	6.43	6.17	5.83	5.97	5.43	6.82	6.34	6.75	6.51	5.81	6.28	8.23	7.89	6.17
Ho	1.36	1.13	1.10	1.08	1.01	1.02	0.91	1.19	1.08	1.18	1.13	0.96	1.07	1.40	1.28	1.07
Er	3.59	2.62	2.64	2.69	2.54	2.61	2.25	2.99	2.67	2.87	2.81	2.38	2.66	3.47	3.13	2.72
Tm	0.53	0.35	0.36	0.38	0.35	0.35	0.30	0.41	0.37	0.41	0.38	0.33	0.37	0.49	0.41	0.37
Yb	3.32	1.94	2.11	2.25	1.95	2.13	1.80	2.39	2.10	2.33	2.27	1.85	2.14	2.83	2.37	2.20
Lu	0.54	0.29	0.31	0.35	0.29	0.33	0.27	0.35	0.31	0.34	0.34	0.29	0.32	0.41	0.35	0.33
Rb	34.0	36.9	33.1	37.3	21.7	44.7	63.6	27.1	39.8	22.7	27.3	50.5	22.9	137	78.1	28.2
Ga	21.6	25.4	25.1	23.6	23.4	24.0	22.6	24.4	23.1	24.0	24.8	22.9	23.7	24.6	25.6	23.2
V	212	199	307	243	326	287	392	278	318	272	328	338	211	96	290	298
Cr	59.4	0.42	0.16	0.33	12.4	12.7	4.90	1.80	21.4	34.7	14.4	0.03	5.10	21.8	1.10	12.8
Ni	67.4	12.4	23.5	17.9	58.2	38.7	38.2	18.8	49.6	48.5	53.5	15.1	14.4	13.8	31.5	56.9
Sc	23.4	16.8	23.4	20.3	24.9	25.0	27.5	19.9	24.9	24.5	24.0	23.3	15.3	3.9	17.6	24.0
Sr	388	700	593	646	534	618	633	572	642	672	729	455	637	1607	806	750
Ba	1187	696	555	660	505	110	664	327	422	457	332	542	318	2496	719	449
Th	3.15	6.97	4.12	4.94	3.50	4.70	3.64	4.54	3.89	4.01	3.64	4.02	8.70	13.9	6.71	3.35
U	0.72	1.72	1.03	1.09	0.82	1.00	0.96	1.03	1.00	1.20	0.94	1.06	2.60	2.83	1.60	0.82
Ta	1.66	2.47	3.06	2.63	2.35	2.36	2.54	3.15	2.85	2.94	2.62	1.24	3.79	8.22	4.30	2.44
Nb	25.8	57.1	47.3	40.2	35.0	36.0	39.8	52.2	44.4	45.6	43.7	18.7	61.1	133	65.6	38.5

Zr	281	366	264	274	239	270	251	299	270	278	266	249	388	624	443	251
Hf	6.53	8.32	6.37	6.65	5.83	6.54	6.19	7.00	6.57	6.62	6.33	6.33	8.88	11.0	9.86	6.04
Y	38.5	31.8	30.8	31.3	28.7	29.5	26.2	34.3	31.1	34.5	33.1	27.7	31.0	40.5	36.8	31.0

sample	08KT 01	08KT 02	08KT 04	08KT 05	08KT 06	08KT 07	08KT 08	08KT 09	08KT 10	08KT 11	08KT 12	08KT 13	08KT 14	08KT 15
Rock type	B	B	B	B	B	B	B	B	<b>B</b>	B	B	B	B	B

*Major elements (%)*

SiO <sub>2</sub>	44.48	44.45	44.10	44.35	44.41	44.31	47.31	47.17	45.12	44.69	44.38	49.32	51.24	50.37
TiO <sub>2</sub>	4.02	4.02	4.00	3.98	3.96	3.95	4.07	4.02	3.74	3.72	3.99	3.41	3.25	3.55
Al <sub>2</sub> O <sub>3</sub>	14.0	14.03	14.23	14.16	14.14	14.28	13.02	13.08	14.24	14.47	14.17	11.22	10.59	11.91
Fe <sub>2</sub> O <sub>3</sub>	17.0	17.01	17.04	16.96	16.90	16.94	16.23	16.24	15.53	15.56	17.05	17.14	17.44	18.15
MnO	0.20	0.20	0.20	0.20	0.19	0.19	0.20	0.21	0.18	0.18	0.19	0.18	0.16	0.13
CaO	5.73	5.73	5.85	5.84	5.69	5.90	4.45	4.68	6.19	5.64	5.81	3.67	3.35	3.72
MgO	7.83	7.98	7.93	7.92	8.12	7.82	7.99	8.05	8.43	8.45	7.91	6.12	5.26	4.00
K <sub>2</sub> O	1.30	1.42	1.34	1.39	1.46	1.34	1.45	1.54	1.26	1.23	1.36	0.97	1.04	0.71
Na <sub>2</sub> O	3.45	2.9	3.19	3.12	3.03	3.16	3.12	3.05	2.91	2.92	3.16	3.87	4.22	4.39
P <sub>2</sub> O <sub>5</sub>	0.97	1.01	0.95	0.95	0.96	0.94	0.99	0.95	0.73	0.71	0.95	0.62	0.56	0.62
LOI	0.43	0.75	0.67	0.70	0.71	0.74	0.79	0.61	1.22	1.99	0.62	3.15	2.60	2.17
Σ	99.5	99.58	99.58	99.58	99.58	99.58	99.61	99.62	99.56	99.55	99.58	99.67	99.72	99.71
Mg <sup>#</sup>	36	36	37	37	36	37	32	33	40	38	36	26	24	26

*Trace elements (ppm)*

La	35.0	35.1	35.0	35.8		33.5	39.4	35.68	29.7	28.9	34.7	32.0	31.9	29.9
Ce	81.0	80.0	80.9	82.4		76.2	88.9	77.31	68.4	66.4	79.5	73.6	69.2	69.2
Pr	11.1	11.07	11.15	11.24		10.37	11.88	10.36	9.42	8.993	11	9.76	9.11	9.34
Nd	48.0	47.3	47.4	48.7		45.2	50.6	44.50	40.6	39.2	47.4	40.8	37.6	39.8
Sm	9.82	9.73	9.76	9.85		9.37	10.11	9.18	8.47	8.07	9.58	8.37	7.66	8.31
Eu	3.12	3.01	3.08	3.03		2.84	2.97	2.84	2.64	2.50	2.91	2.56	2.33	2.55
Gd	10.0	9.25	8.98	9.12		8.70	9.50	8.94	8.11	7.72	8.93	7.98	7.19	7.80
Tb	1.55	1.41	1.39	1.48		1.33	1.45	1.39	1.29	1.26	1.40	1.29	1.17	1.20
Dy	8.39	7.49	7.50	7.84		7.13	7.85	7.85	7.13	6.80	7.72	7.11	6.12	6.51
Ho	1.55	1.39	1.43	1.43		1.31	1.47	1.51	1.36	1.30	1.44	1.33	1.14	1.19
Er	4.12	3.64	3.80	3.86		3.47	3.95	3.92	3.61	3.44	3.95	3.47	3.03	3.26
Tm	0.62	0.54	0.56	0.57		0.50	0.59	0.54	0.54	0.52	0.58	0.50	0.47	0.49
Yb	3.63	3.31	3.57	3.50		3.18	3.64	3.48	3.38	3.18	3.61	3.08	2.87	3.05
Lu	0.60	0.56	0.60	0.55		0.51	0.57	0.53	0.54	0.52	0.58	0.51	0.47	0.49
Rb	23.6	23.4	23.0	24.4		22.2	15.8	23.6	23.8	23.4	24.3	17.8	16.5	12.9
Ga	22.7	20.6	21.6	23.0		20.7	22.3	21.7	22.8	22.9	22.4	16.3	12.8	19.8
V	209	217	221	228		222	222	237	220	208	235	212	181	208
Cr	60.9	63.3	61.1	66.4		66.2	27.8	66.58	78.4	77.7	64.9	10.3	13.9	10.2
Ni	70.4	65.6	74.6	84.3		71.6	32.9	68.54	74.2	79.0	72.2	16.9	18.8	16.9
Sc	23.5	24.5	24.4	25.3		25.3	24.0	24.0	24.1	24.8	19.7	19.1	19.9	23.5

Sr	352	328	420	352				334	310	484	298	303	355	191	164	183
Ba	621	670	689	724				658	768	687	554	548	706	831	707	492
Th	3.61	3.58	3.38	3.55				3.34	5.08	3.02	3.31	3.18	3.36	5.27	5.12	5.15
U	0.82	0.78	0.81	0.80				0.74	1.03	0.73	0.78	0.77	0.77	1.37	1.56	1.45
Ta	1.86	1.70	1.80	1.81				1.64	1.59	1.66	1.51	1.52	1.75	1.70	1.67	1.72
Nb	26.1	25.3	25.0	25.9				24.9	24.0	24.86	22.0	22.1	25.7	24.5	23.5	25.0
Zr	289	276	282	286				262	293	280	239	241	283	243	224	249
Hf	7.61	6.48	7.04	6.84				6.33	7.27	6.13	6.13	5.81	7.00	6.18	5.85	6.14
Y	39.9	37.9	38.8	38.5				36.3	39.4	38.12	35.6	36.2	38.7	36.9	30.4	32.3

sample	08KT0	08KT0	08KT	08KT0	08K	08K	08K	08K	08K	08KT	08KT0	08KT03	08KT	08KT	08KT
	2-1	2-2	02-3	2-4	T02-5	T02-6	T02-7	T03-1	T03-2	03-3	3-4	-5	03-6	03-7	03-8
Rock type	APG	APG	APG	APG	APG	APG	APG	APG	APG	APG	APG	APG	APG	APG	APG

*Major elements (%)*

SiO <sub>2</sub>	71.86	71.70	73.22	74.24	72.89	72.70	75.23	77.67	77.88	77.81	76.76	77.99	77.24	78.22	77.02
TiO <sub>2</sub>	0.19	0.17	0.16	0.19	0.17	0.17	0.28	0.04	0.04	0.05	0.06	0.05	0.05	0.03	0.04
Al <sub>2</sub> O <sub>3</sub>	14.60	14.70	14.44	13.36	13.1	13.9	12.0	12.0	12.0	12.0	12.17	12.18	12.01	12.00	12.42
Fe <sub>2</sub> O <sub>3</sub>	1.91	1.74	1.47	1.76	1.62	1.75	2.73	0.42	0.26	0.40	1.15	0.49	0.77	0.21	0.70
MnO	0.03	0.03	0.03	0.03	0.03	0.03	0.05	0.01	0.01	0.01	0.01	0.01	0.02	0.01	0.01
CaO	1.20	1.14	1.00	0.96	0.89	1.01	1.15	0.53	0.59	0.52	0.57	0.47	0.52	0.49	0.53
MgO	0.15	0.14	0.15	0.15	0.15	0.15	0.18	0.05	0.05	0.08	0.07	0.08	0.05	0.07	0.07
K <sub>2</sub> O	5.30	5.72	5.35	5.19	5.25	5.83	4.56	4.85	4.72	4.88	4.83	4.59	4.86	4.70	5.00
Na <sub>2</sub> O	4.33	4.15	3.73	3.64	3.44	3.86	3.40	3.73	3.79	3.60	3.76	3.50	3.70	3.61	3.58
P <sub>2</sub> O <sub>5</sub>	0.02	0.02	0.02	0.03	0.02	0.02	0.04	0.01	0.00	0.00	0.00	0.00	0.00	0.01	0.00
LOI	0.15	0.21	0.21	0.24	2.20	0.29	0.20	0.42	0.38	0.43	0.43	0.46	0.58	0.48	0.45
Σ	99.74	99.73	99.76	99.79	99.8	99.7	99.8	99.8	99.8	99.8	99.81	99.84	99.80	99.82	99.81

*Trace elements (ppm)*

La	48.11	86.2	96.2	108	42.6	52.4	78.4	56.0	64.7	61.7	65.5	53.1	64.9	49.6	59.0
Ce	103.7	176	194	211	102	108	174	115	137	130	141	113	140	104	127
Pr	13.57	21.2	22.8	24.5	11.8	13.5	22.2	14.1	16.7	15.7	17.6	14.1	16.9	13.3	15.5
Nd	54.6	77.1	80.8	85.7	45.4	52.5	85.4	48.9	58.5	54.9	64.3	50.4	60.6	48.2	55.3
Sm	11.3	13.3	13.0	14.3	9.17	11.8	16.7	10.3	12.5	11.7	15.4	11.2	13.2	10.9	12.2
Eu	1.52	1.68	1.63	1.49	1.32	1.34	1.27	0.04	0.04	0.05	0.05	0.04	0.05	0.04	0.05
Gd	10.6	11.3	10.9	12.3	8.63	11.5	15.0	9.78	12.7	11.4	15.4	10.6	12.8	10.8	12.5
Tb	1.81	1.78	1.64	1.94	1.52	2.11	2.59	1.89	2.42	2.19	3.09	2.04	2.48	2.09	2.41
Dy	10.14	9.76	8.64	10.4	8.72	12.1	14.4	11.5	14.2	12.8	18.8	12.0	14.5	12.3	14.4
Ho	1.88	1.84	1.59	1.99	1.68	2.27	2.78	2.31	2.86	2.49	3.81	2.41	2.94	2.44	2.82
Er	5.19	4.91	4.35	5.62	4.81	6.08	7.72	7.18	8.38	7.40	11.5	7.15	8.75	7.22	8.34
Tm	0.78	0.75	0.65	0.89	0.75	0.95	1.19	1.27	1.38	1.23	1.89	1.24	1.49	1.20	1.38
Yb	4.86	4.59	4.13	5.54	4.78	5.81	7.50	8.82	8.93	7.78	11.9	8.04	9.88	7.94	8.85
Lu	0.74	0.71	0.65	0.87	0.75	0.99	1.17	1.40	1.38	1.19	1.83	1.28	1.50	1.26	1.38
Rb	170	190	183	177	197	207	151	447	441	412	456	416	456	443	470
Ga	31.1	30.1	28	29.3	27.2	28.7	25.9	29.5	30.1	27.9	30.0	28.4	29.71	29.2	29.5
V	2.98	2.43	3.16	4.27	3.01	2.33	4.39	2.45	1.97	3.11	1.35	1.72	1.74	1.60	1.86
Cr	0.05	0.03	0.64	0.45	0.12	0.09	0.05	0.15	0.74	0.62	1.49	0.63	0.11	4.17	0.29
Ni	1.21	0.73	1.64	0.82	0.89	0.18	0.73	0.54	0.22	0.01	1.21	1.55	1.66	4.75	1.37

Sr	88.7	97.8	93.5	87.6	77.9	80.4	75.9	6.23	8.59	7.31	7.34	7.49	5.15	7.84	6.29
Ba	460	551	546	484	465	466	364	9.91	10.5	10.8	9.53	12.2	8.22	9.63	10.8
Th	17.6	20.6	20.3	26.6	17.1	12.8	27.4	41.1	41.7	40.1	53.9	41.8	44.0	45.0	42.9
U	2.83	2.35	2.44	3.98	3.46	5.64	3.80	5.90	9.32	7.32	7.62	9.94	7.79	9.34	5.97
Ta	3.04	3.17	2.93	4.07	3.81	4.51	4.71	6.99	9.17	7.81	8.34	9.74	8.76	8.85	7.67
Nb	44.7	44.2	40.7	52.6	50.2	59.1	69.3	82.4	115	97.5	143	116	105	108	92.1
Zr	367	183	243	314	266	172	367	234	164	176	202	206	210	178	172
Hf	10.6	5.61	7.22	9.81	8.10	5.50	11.9	13.1	9.51	9.00	11.1	11.3	12.0	10.3	9.16
Y	52.8	50.5	45.2	56.4	48.6	61.9	76.7	77.9	91.2	77.4	121	73.9	91.5	78.4	90.7

MGB-magnetite –(bearing) gabbro; MQGB- magnetite and quartz – bearing gabbro; PGB – plagioclase-rich gabbro; QGB – quartz – bearing gabbro; OMGB – olivine and magnetite –bearing gabbro; DB – diabase or dolerite; B- basalt; APG – alkali-feldspar granite

2.1	188	80	0.44	0.48	.0439	2.1	0.36	4.5	0.0600	4.0	274.1	5.6	604	87
3.1	115	78	0.70	0.99	.0453	2.2	0.40	9.9	0.0648	9.6	281.0	6.3	767	203
4.1	270	91	0.35	0.29	.0449	2.0	0.37	4.1	0.0599	3.6	280.7	5.6	601	77
5.1	506	233	0.48	0.42	.0438	2.0	0.45	4.3	0.0738	3.9	269.0	5.3	1037	78
6.1	239	104	0.45	0.26	.0443	2.0	0.39	4.1	0.0645	3.5	275.3	5.5	758	74
7.1	80	30	0.39	1.11	.0431	2.4	0.48	10.3	0.0808	10.0	262.7	6.1	1216	197
8.1	197	88	0.46	0.65	.0452	2.0	0.44	7.3	0.0704	7.0	278.5	5.8	939	143
9.1	132	53	0.42	0.51	.0457	2.1	0.47	7.4	0.0745	7.1	280.5	6.0	1055	144
10.1	213	99	0.48	0.12	.0427	2.0	0.41	5.1	0.0688	4.7	263.8	5.4	894	98
11.1	240	94	0.41	0.14	.0432	2.0	0.41	2.9	0.0694	2.1	266.9	5.3	912	43
12.1	235	116	0.51	0.17	.0424	2.0	0.35	3.0	0.0594	2.2	265.2	5.3	580	49
13.1	115	46	0.41	0.64	.0432	2.1	0.45	9.2	0.0753	8.9	265.1	6.0	1076	179
14.1	260	112	0.45	0.10	.0448	2.0	0.41	6.6	0.0665	6.3	277.6	5.7	823	131
15.1	211	95	0.46	0.17	.0433	2.0	0.43	7.3	0.0713	7.1	266.8	5.6	965	144
16.1	160	86	0.55	0.00	.0457	2.1	0.49	7.7	0.0772	7.4	279.2	6.1	1126	147
17.1	340	140	0.43	0.00	.0437	2.0	0.39	6.8	0.0644	6.5	271.7	5.5	756	136

#  $f_{206}$ : percentage of common  $^{206}\text{Pb}$  in total  $^{206}\text{Pb}$

Table.3 Nd isotope compositions of the coeval diverse igneous rocks along the north margin of Tarim

Sample	Sm (ppm)	Nd (ppm)	$(^{147}\text{Sm}/^{144}\text{Nd})_c$	$(^{143}\text{Nd}/^{144}\text{Nd})_m$	$2\sigma_m$	$\epsilon\text{Nd}(t)$ ( $t=275\text{Ma}$ )
<i>Permian basalts</i>						
08KT01	9.82	48	0.1237	0.512384	0.000008	-2.39
08KT04	9.76	47.4	0.1245	0.512391	0.000007	-2.28
08KT07	9.37	45.2	0.1253	0.512378	0.000007	-2.57
08KT10	8.47	40.6	0.1261	0.512395	0.000007	-2.26
08KT11	8.07	39.2	0.1245	0.512392	0.000007	-2.26
08KT12	9.58	47.4	0.1222	0.512392	0.000007	-2.18
08KT13	8.37	40.8	0.1240	0.512402	0.000007	-2.05
08KT14	7.66	37.6	0.1232	0.512395	0.000008	-2.16
<i>Piqiang ultramafic-mafic complex</i>						
08KT01-1	1.06	4.187	0.1531	0.512666	0.000008	2.08
08KT01-2	0.579	3.06	0.1144	0.512436	0.000008	-1.05
08KT01-5	1.663	6.471	0.1554	0.512588	0.000007	0.48
08KT01-8	1.34	4.22	0.1920	0.512592	0.000010	-0.73
08KT01-9	2.67	10.01	0.1613	0.512536	0.000008	-0.75
08KT01-15	7.884	38.88	0.1226	0.512484	0.000008	-0.40
08KT01-11	5.184	21.82	0.1436	0.512571	0.000006	0.56
<i>Bachu mafic dykes</i>						
08KT16-1	11.44	60.48	0.1144	0.512747	0.000008	5.02
08KT17-1	8.464	40.56	0.1262	0.512674	0.000008	3.18
08KT18-1	9.02	46.4	0.1175	0.512524	0.000008	0.56
08KT18-9	8.61	42.1	0.1236	0.512762	0.000007	4.99

Sample	Sm (ppm)	Nd (ppm)	$^{147}\text{Sm}/^{144}\text{Nd}$ d	$^{143}\text{Nd}/^{144}\text{Nd}_m$	$2\sigma_m$	$T_{\text{DM}}$ Ga	$T_{2\text{DM}}$ Ga	$\epsilon\text{Nd}(t)$ ( $t=275\text{Ma}$ )
08KT02-4	15.4	64.3	0.1448	0.512438	0.000008	1.6	1.2	-2.08
08KT02-5	11.1	50.4	0.1331	0.512431	0.000006	1.4	1.2	-1.81
08KT02-7	10.8	48.1	0.1357	0.512422	0.000007	1.4	1.2	-2.07
08KT03-1	11.2	54.6	0.1240	0.512496	0.000007	1.1	1.1	-0.22
08KT03-3	13.01	80.85	0.0973	0.512439	0.000007	0.9	1.1	-0.39
08KT03-7	16.6	85.4	0.1175	0.512479	0.000007	1.1	1.1	-0.32

$2\sigma_m$  = 2-sigma-mean error.

Table. 4 Age data of the Permian igneous rocks along the northern margin of Tarim

Location	Rock	Mineral	Method	age (Ma)	Reference
Mazhaertage	Syenite	Whole rock	$^{39}\text{Ar}/^{40}\text{Ar}$	$278 \pm 1$	Yang et al., 1996
Mazhaertage	Syenite	Zircon	SHRIMP	$277 \pm 4$	Yang et al., 2006
Keping	Basalt	Whole rock	$^{39}\text{Ar}/^{40}\text{Ar}$	$279 \pm 1$	Chen et al., 1997
Central Tarim	Tuff	Whole rock	$^{39}\text{Ar}/^{40}\text{Ar}$	276-288	XOAC, 2003
Keping	Basalt	Zircon	LA-ICP-MS	$275 \pm 13$	Li et al., 2007
Keping	Tuff	Zircon	LA-ICP-MS	$291 \pm 10$	Li et al., 2007
Kepjng	Gabbro	Zircon	LA-ICP-MS	$274 \pm 15$	Li et al., 2007
Mazhaertage	Diabase	Zircon	LA-ICP-MS	$272 \pm 6$	Li et al., 2007
Mazaertage	Diabase	Zircon	LA-ICP-MS	$282 \pm 3$	Li et al., 2007
Mazaertage	Syenite	Zircon	LA-ICP-MS	$281 \pm 4$	Li et al., 2007
Mazhaertage	Gabbro	Zircon	LA-ICP-MS	$274 \pm 2$	Zhang et al., 2008
South of Akesu	Rhyolite	Zircon	LA-ICP-MS	$277.3 \pm 2.5$	Tian et al., 2010
South of Akesu	Rhyolite	Zircon	LA-ICP-MS	$290.9 \pm 4.1$	Tian et al., 2010
South of Akesu	Rhyolite	Zircon	LA-ICP-MS	$286.6 \pm 3.3$	Tian et al., 2010
South of Akesu	Rhyolite	Zircon	LA-ICP-MS	$282.9 \pm 2.5$	Tian et al., 2010
South of Akesu	Rhyolite	Zircon	SHRIMP	$271.7 \pm 2.2$	Tian et al., 2010
Piqiang complex	Ganbbro	Zircon	SHRIMP	$276 \pm 4$	This study
Halajun pluton □	Granite	Zircon	SHRIMP	$278 \pm 3$	This study
Halajun pluton □	Granite	Zircon	SHRIMP	$278 \pm 3$	This study