

1 **On the pitfalls of Airy isostasy and the isostatic gravity anomaly in general**

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9 **SUMMARY**

10 Isostatic gravity anomalies provide a measure of the Earth's gravity field free from the
11 gravitational attractions of the topography and its isostatic compensation, most commonly
12 represented by a variation in the depth of a compensating density contrast, for example the
13 Moho. They are used by both geodesists and geophysicists alike, though often for different
14 purposes. Unfortunately though, the effect of subsurface loading on the lithosphere renders
15 transfer function (admittance) methods unusable when surface and subsurface loads coexist.
16 Where they exist, subsurface loads are often expressed in the Bouguer anomaly but not in the
17 topography, and it is shown here that this phase disconnect cannot be faithfully represented
18 by either real- or complex-valued analytic admittance functions. Additionally, many studies
19 that employ the isostatic anomaly ignore the effects of the flexural rigidity of the lithosphere,
20 most often represented as an effective elastic thickness (T_e), and assume only Airy isostasy,
21 i.e. surface loading of a plate with zero elastic thickness. The consequences of such an
22 omission are studied here, finding that failure to account for flexural rigidity and subsurface
23 loading can result in (1) over- or underestimates of both inverted Moho depths and dynamic
24 topography amplitude, and (2) underestimates of the size of topographic load that can be
25 supported by the plate without flexure. An example of the latter is shown over Europe.

26 Finally, it is demonstrated how low values of the isostatic anomaly variance can actually be
27 biased by these anomalies having low power at the long wavelengths while still possessing
28 high power at middle to short wavelengths, compared to the corresponding Bouguer anomaly
29 power spectrum. This will influence the choice of best-fitting isostatic model if the model is
30 chosen by minimization of the isostatic anomaly standard deviation.

31

32 **Key words:** Gravity anomalies and Earth structure; Europe; Dynamics: gravity and tectonics;
33 Dynamics of lithosphere and mantle; Lithospheric flexure.

34

35 **1 INTRODUCTION**

36 According to the principle of isostasy, topographic features on the Earth's surface must be
37 compensated to some extent by subsurface mass-density anomalies, as in Archimedes'
38 principle. The simplest models of isostatic compensation are the Airy-Heiskanen (Airy 1855;
39 Heiskanen 1931) and Pratt-Hayford (Pratt 1855; Hayford 1909) (e.g. Heiskanen & Moritz
40 1967; Watts 2001). The Airy-Heiskanen (or just 'Airy') model compensates variations in
41 topography by variations in the relief of the crust-mantle interface (the Moho) about a mean
42 compensation depth, where the crust and mantle have spatially uniform densities; for
43 example, higher topography has a thicker crust below it. In the Pratt-Hayford (or just 'Pratt')
44 model, topographic variations are compensated by lateral variations in crustal density, with a
45 flat compensation depth (Moho), so that higher topography is modelled as having lower
46 density relative to its surroundings. These two models are commonly called 'local isostasy' to
47 distinguish them from the regional isostatic model of compensation, originally proposed by
48 Vening Meinesz (1931).

49

50 The regional model is similar to the Airy model in that compensation is achieved by
51 deflection of the Moho, where the density of the crust is constant. The difference between
52 them lies in the flexural rigidity (D) of the crust and its ability to support loads mechanically.
53 In the Airy model, the crust has no such mechanical strength ($D = 0$); a point surface
54 topographic load is compensated by deflection of the Moho directly underneath it and
55 nowhere else. In the Vening Meinesz model ($D > 0$), the stresses caused by the point load are
56 regionally distributed about the point, and the load is both compensated by the Moho
57 deflection, and supported by the mechanical strength of the crust; the Moho deflection is thus
58 less than it would be under Airy isostasy. Modern interpretations of the Vening Meinesz
59 model discuss the flexure of the elastic portion of the lithosphere, being the crust and
60 uppermost mantle, or portions thereof, rather than the crust alone, where the lithosphere
61 'floats' on an inviscid asthenosphere (e.g. Watts 2001). In most models, the primary density
62 contrast providing compensation of loads is still assumed to lie at the Moho (e.g. Forsyth
63 1985), since the density contrast at the lithosphere-asthenosphere boundary is an order of
64 magnitude smaller, $\sim 40 \text{ kg m}^{-3}$ as opposed to $\sim 500 \text{ kg m}^{-3}$ at the Moho (Cordell *et al.* 1991).
65

66 These isostatic compensation mechanisms can be used to apply corrections to gravity
67 anomalies by modelling and removing the gravity effect of the subsurface mass-density
68 anomalies compensating the topography. In the Pratt model this will be the gravity effect of
69 the lateral density variations in the crust; in the Airy and Vening Meinesz models the
70 correction will model the density anomalies caused by deflection of the Moho; while the
71 Vening Meinesz model must also account for the fact that mechanical support implies a
72 shallower compensation depth. When the isostatic correction thus derived is subtracted from
73 the Bouguer anomaly the result is the isostatic anomaly, which should be small if the model
74 is a fair description of the actual compensation and its parameters, and the Bouguer

75 correction and/or topographic reduction have properly accounted for the presence of any
76 mass-density anomalies above the geoid.

77

78 Non-zero isostatic anomalies show that the actual compensation mechanism differs from that
79 of the assumed model. Such differences can be attributed to relatively minor concerns such as
80 incorrect choice of densities or compensation depths in the model, or major concerns such as
81 a wrong choice of model where the effects of mechanical or dynamic support are
82 misinterpreted as under- or overcompensation in a local isostatic model (e.g. Simpson *et al.*
83 1986). However, non-zero isostatic anomalies can also reveal compensated intralithospheric
84 density anomalies with no topographic expression; regional variations in effective elastic
85 thickness (T_e), or loading at one value of T_e followed by erosional unloading at another;
86 mantle density variations and the dynamic support of the lithosphere due to convective
87 processes in the upper mantle; the viscoelastic response to time-varying loads, such as glacial
88 isostatic adjustment; and the deeper signals from the mantle and core (e.g. Forsyth 1985;
89 Simpson *et al.* 1986; Ussami *et al.* 1993; Kaban *et al.* 2004).

90

91 Geophysicists have largely used isostatic anomalies for investigations of Earth structure.
92 Some of these studies have assumed Airy isostasy when computing isostatic anomalies (e.g.
93 Karner & Watts 1982; Simpson *et al.* 1986; Ussami *et al.* 1993; Kaban *et al.* 1999; Tiwari *et*
94 *al.* 2003). Other workers have invoked the flexural rigidity (or its geometric analogue, the
95 effective elastic thickness, T_e) in isostatic anomaly computation (e.g. Walcott 1970; Dorman
96 & Lewis 1972; Karner & Watts 1983; Watts *et al.* 1995; Jordan & Watts 2005; Harmon *et al.*
97 2006; Wyer & Watts 2006; Watts & Moore 2017). Several studies have sought to estimate
98 lithospheric parameters such as interface depths and densities by minimizing the isostatic
99 anomaly variance (e.g. Sünel 1985; Martinec 1993, 1994a,b; Kaban *et al.* 2004; Sjöberg

100 2009; Bagherbandi & Sjöberg 2012), a topic that will be investigated in Section 3.

101 Additionally, many studies assume Airy isostasy when computing dynamic topography (e.g.

102 England & Molnar 2015; Molnar *et al.* 2015), a topic that will be addressed in Section 4.3.

103

104 In the geodetic community, however, workers have largely persisted with Airy, and to a

105 lesser extent Pratt models of local isostasy (e.g. Pavlis & Rapp 1990; Martinec 1994b;

106 Balmino *et al.* 2012; Hirt *et al.* 2012; Aitken *et al.* 2015), often because their analyses

107 concern very large, even global study areas where Airy or Pratt models would be more

108 appropriate in describing isostasy than regional models. However, the inability of the Airy

109 model to correctly represent the actual state of compensation is well documented. For

110 instance, Lewis & Dorman (1970) and Dorman & Lewis (1972) assumed Airy isostasy and

111 inverted an isostatic response function to find the density contrasts compensating the North

112 American topography. In addition to the expected shallow positive density contrast their

113 model required an unrealistic negative density contrast (i.e., a decrease in density with depth)

114 within the upper mantle (~400 km depth). Banks *et al.* (1977) adequately remodelled Dorman

115 & Lewis's (1972) data using a regional compensation model with a non-zero elastic

116 thickness, concluding that any local model with negative densities can be replaced by a

117 regional model with compensating positive densities at shallower depths (Banks *et al.* 1977;

118 Simpson *et al.* 1986; Cordell *et al.* 1991).

119

120 Oceanic lithosphere provides cases where local isostasy does not apply. As Sünkel (1985)

121 notes, oceanic crust is typically 7 km thick on average but much deeper compensation levels

122 are required to prevent the oceanic anticroots of the Airy model extending above the seafloor

123 (e.g. Rapp 1982). In contrast, the mechanical support provided by regional isostatic models

124 enables compensation levels to be placed at much more realistic depths, closer to the Moho.

125 For example, Louden & Forsyth (1982) analysed the compensation of the Kane fracture zone,
126 and where the Airy model required a 30 km thick oceanic crust, a flexural model with a more
127 realistic 6 km thick crust gave a better fit to observed admittance data. Nevertheless, Pratt-
128 style compensation is thought to apply at mid-ocean ridges, though with the density increase
129 away from the ridge occurring in the sub-crustal mantle rather than the crust, while Airy
130 isostasy would apply at continent-ocean boundaries (Watts 2007).

131

132 And recently, using spherical harmonics, Watts & Moore (2017) fitted two isostatic models
133 to the mid- to high-degrees ($12 \leq n \leq 400$) of the global EGM2008 free air anomaly (Pavlis *et*
134 *al.* 2012). When using an Airy model the best fit was provided by an unrealistic
135 compensation (Moho) depth of 61 km. In contrast, a better fit was obtained with a flexural
136 model of compensation depth 30 km and $T_e = 34$ km.

137

138 Hence, this study aims to test the limits of the applicability of the Airy isostatic compensation
139 model, but also comments on the utility of isostatic gravity anomalies in general. In Section 2
140 some theory is presented regarding the computation of isostatic anomalies. In Section 3,
141 synthetic Bouguer anomaly and topography models are generated with known plate
142 parameters (such as compensation depth, elastic thickness, etc.); isostatic anomalies are then
143 computed from these synthetic models over a whole range of plate parameters, using the
144 spectral methods described in Section 2. In this fashion many assumptions regarding isostatic
145 models and gravity anomalies can be tested. Section 4 then discusses the implications of the
146 findings in Section 3, but also comments on the isostatic and dynamic support of surface
147 topography in the context of Airy versus flexural isostasy.

148

149 **2 ESTIMATION OF ISOSTATIC GRAVITY ANOMALIES**

150 2.1 Methods favoured by geophysicists

151 The isostatic anomaly (Δg_I) is the difference between the Bouguer anomaly (Δg_B) and the
152 gravitational attraction of the compensation model (δg_C):

$$153 \quad \Delta g_I = \Delta g_B - \delta g_C \quad (1)$$

154 Dorman & Lewis (1970) proposed that the Earth acts as a linear filter, and given an observed
155 (post-flexural) topography, h , it was possible to determine the gravitational effect of the
156 isostatic compensation of that topography by filtering the topography with an ‘isostatic
157 response function’, q . In the 1-D Cartesian space (x) domain they wrote

$$158 \quad \delta g_C(x) = q(x) * h(x) \quad (2)$$

159 where the $*$ indicates space-domain convolution. Hence, by the convolution theorem:

$$160 \quad \delta g_C(x) = \mathbf{F}^{-1} [Q(k)H(k)] \quad (3)$$

161 where \mathbf{F}^{-1} is the inverse Fourier transform operator, k is wavenumber (spatial frequency), and
162 $Q = \mathbf{F}[q]$ is the admittance which is the wavenumber-domain transfer function from
163 topography to gravity, discussed below. Thus, eq. (1) can be written as

$$164 \quad \Delta g_I = \Delta g_B - \mathbf{F}^{-1} [QH] \quad (4)$$

165 While Dorman & Lewis (1970) called Δg_I the “geologic effect”, Lewis & Dorman (1970)
166 identified it with the isostatic anomaly, as did Watts (1978), McNutt (1980), Simpson *et al.*
167 (1986), and Kaban *et al.* (1999), who all used such transfer function techniques. Lewis &
168 Dorman (1970) also tested eq. (2) for non-linear terms in the topography, h , by correlating the
169 resultant isostatic anomaly with terms h^n , for $n = 2 - 5$, finding insignificant correlation in
170 their North American study area.

171

172 Using equations of flexure and Parker’s (1972) formula, one can derive analytic equations for
173 the theoretical admittance for many loading models, for use in eq. (4) (e.g. Kirby 2014).

174 Those considered in this study are described next. The expression for the Bouguer admittance
175 for Airy isostasy is:

$$176 \quad Q_A(k) = -2\pi\mathcal{G}\Delta\rho_0 e^{-kz_m} \quad (5)$$

177 (e.g. Forsyth 1985; Watts 2001) where \mathcal{G} is the gravitational constant (Table 1), z_m is the
178 depth to the compensating interface (commonly assumed to be the Moho), and $\Delta\rho_0 = \rho_c - \rho_f$
179 where ρ_c is the density of an incompressible single-layer crust and ρ_f is the density of the
180 overlying fluid, either air or water.

181

182 More general loading models consider the flexural rigidity of the lithosphere, D , which is
183 most commonly expressed in terms of an effective elastic thickness (T_e) where

$$184 \quad D = \frac{ET_e^3}{12(1-\nu^2)} \quad (6)$$

185 (e.g., Watts 2001) and where E is Young's modulus and ν is Poisson's ratio (see Table 1 for
186 the values of these constants). The elastic lithosphere may include numerous crustal layers
187 and the uppermost mantle, depending on the tectonic regime and rheological properties of
188 these strata (Burov & Diament 1995). Although most flexural models place the depth of
189 compensation at the Moho (being a compositional boundary with a large density contrast),
190 the elastic lithosphere may extend into the uppermost mantle where its boundary with the
191 underlying asthenosphere is rheological in nature (with a much lower density contrast as
192 noted in Section 1). It should be noted that T_e does not, in general, describe a physical
193 thickness or depth – rather it is a geometric analogue of the flexural rigidity – though under
194 certain conditions, T_e can equal the mechanical thickness of the lithosphere (e.g. Burov &
195 Diament 1995).

196

197 Banks *et al.* (1977) derived an expression for the Bouguer admittance corresponding to initial
 198 subaerial loading on the surface of a plate with non-zero flexural rigidity ('surface' or 'top
 199 loading'):

$$200 \quad Q_T(k) = -2\pi\mathcal{G}\Delta\rho_0 e^{-kz_m} \left(1 + \frac{Dk^4}{g\Delta\rho_1}\right)^{-1} = -2\pi\mathcal{G}\Delta\rho_0 e^{-kz_m} \xi(k)^{-1} \quad (7)$$

201 where g is the gravity acceleration, $\Delta\rho_1 = \rho_m - \rho_c$, and ρ_m is the density of an inviscid mantle
 202 underlying the crust. Eq. (7) defines the variable $\xi(k)$. Surface loading constitutes an initial
 203 load emplaced at the Earth's topographic surface by, for example, orogenesis or volcanism. A
 204 second flexural model considers loading that takes place within the lithosphere, for instance,
 205 from magmatic underplating, igneous intrusions, or during sedimentary basin formation.
 206 McNutt (1983) derived a Bouguer admittance equation describing the case when initial
 207 loading occurs within the plate ('subsurface' or 'bottom loading'):

$$208 \quad Q_B(k) = -2\pi\mathcal{G}\Delta\rho_0 e^{-kz_m} \left(1 + \frac{Dk^4}{g\Delta\rho_0}\right) = -2\pi\mathcal{G}\Delta\rho_0 e^{-kz_m} \phi(k) \quad (8)$$

209 which defines the variable $\phi(k)$.

210

211 Forsyth (1985) unified the surface and subsurface loading regimes by assuming independence
 212 of the initial loading processes. From his formulation an expression for the 'combined
 213 loading' Bouguer admittance can be derived:

$$214 \quad Q_{TB}(k) = -2\pi\mathcal{G}\Delta\rho_0 e^{-kz_m} \left(\frac{\xi + \phi f^2 r^2}{\xi^2 + f^2 r^2}\right) \quad (9)$$

215 (e.g. Ito & Taira 2000) where $r = \Delta\rho_0/\Delta\rho_1$, and where $f(k)$ is the ratio of the initial
 216 subsurface to surface load amplitudes (Forsyth 1985). Considering uniform (wavenumber-
 217 independent) values of f , when $f=0$ eq. (9) becomes eq. (7) and describes surface-only initial
 218 loading; when $f \rightarrow \infty$, eq. (9) becomes eq. (8) and describes subsurface-only initial loading. In

219 practice, $f(k)$ is computed from the recovered initial loads in Forsyth's (1985) method of T_e
220 estimation, though a uniform value of f can be used in theoretical analytic admittance
221 equations such as eq. (9) (e.g. Kirby 2014). Note that when $T_e = 0$ km, the parameters ξ and
222 ϕ , eqs (7) and (8), both equal 1 and the combined admittance, Q_{TB} in eq. (9), reduces to the
223 Airy admittance formula, eq. (5), showing that the effects of surface and subsurface loading
224 upon a zero-rigidity plate are indistinguishable.

225

226 **2.2 Methods favoured by geodesists**

227 In the geodetic community, the problem is typically framed in terms of a spherical harmonic
228 expansion of the topography, with derivation of expressions for the gravitational potential of
229 the topography and its isostatic compensation derived from Newton's law of gravitation. The
230 potential of the compensation has invariably been derived assuming Airy isostasy (e.g. Kaula
231 1967; Lachapelle 1976; Rapp 1982, 1989; Rummel *et al.* 1988; Kuhn 2003), though some
232 researchers have considered the effects of flexural rigidity in their models, both indirectly
233 (e.g. Sünkel 1985; Rummel *et al.* 1988; Abd-Elmotaal 1993; Kuhn 2003), and directly (e.g.
234 Watts & Moore 2017).

235

236 In their formulations, Sünkel (1985) and Rummel *et al.* (1988) use a Gaussian smoothing
237 operator to generalise the Airy compensation model to the regional (Vening Meinesz) model,
238 rather than employing an expression derived from physical principles, such as ξ in eq. (7),
239 though Rummel *et al.* (1988) do note the correspondence. Both studies estimate the
240 parameters of the Gaussian function using least squares minimization of the resultant isostatic
241 anomaly. Abd-Elmotaal (1993) and Kuhn (2003) both compute Vening Meinesz isostatic
242 anomalies using space domain solutions of the plate bending equations, following Brochie &
243 Silvester (1969). It should be noted, however, that if computation areas are small enough so

244 that Earth curvature effects are minimal, then spherical solutions of the plate bending
245 equations should give identical results to planar solutions.

246

247 A technique that has reappeared in the literature in the past decade is a development of a
248 concept originally proposed by Vening Meinesz (1931) that Moho depths can be obtained
249 from the Bouguer anomaly. This so-called “inverse problem of isostasy” was built upon by
250 Moritz (1990) and later Sjöberg (2009) who dubbed it the VMM (Vening Meinesz-Moritz)
251 inverse problem. As noted by Moritz (1990) the method is the spherical equivalent of
252 Parker’s (1972) planar method to determine the gravity field due to an undulating, subsurface
253 density contrast – or rather its inverse. However, despite the name ‘Vening Meinesz’
254 implying regional isostatic compensation, the VMM (and Parker) methods make no
255 assumptions about, or even invoke, a particular method of isostatic compensation. Thus the
256 method, while useful, cannot be said to describe any state of isostasy because it simply
257 assumes that the entire Bouguer anomaly is due solely to undulations of the Moho density
258 contrast; the surface topography – essential to any isostatic model – does not feature in the
259 formulation. Recently, Eshagh (2016) showed that the VMM method is a generalisation of
260 the Airy-Heiskanen model, while Eshagh (2018) has attempted a reconciliation of the VMM
261 method with more established methods of modelling flexural isostasy.

262

263 **3 TESTS ON SYNTHETIC ISOSTATIC ANOMALIES**

264 **3.1 Generation of the synthetic data**

265 In order to test how well the isostatic anomaly can be retrieved from real data, one can use
266 synthetic plate models with known parameters. Since the work of Macario *et al.* (1995),
267 synthetic testing of T_e -estimation methods is now well-established (see Kirby (2014) for a
268 summary). Initial surface and subsurface loads are represented by random, fractal surfaces

269 (Saupe 1988; Macario *et al.* 1995), and for uniform- T_e plates the post-loading Bouguer
 270 anomaly and surface topography are found by solving the flexural equation with the Fourier
 271 transform (e.g. Kirby 2014). Unless otherwise noted, in all experiments here the synthetic
 272 models were generated using a single-layer crust of density 2800 kg m^{-3} and thickness 35 km,
 273 overlying a mantle of density 3300 kg m^{-3} (Table 1). When invoked, the initial subsurface
 274 load was emplaced at the base of the crust, 35 km depth.

275

276 Synthetic Bouguer anomaly and topography pairs were generated from combined loading on
 277 plates with T_e and f values described in Sections 3.2 and 3.3, following Macario *et al.* (1995).
 278 Since the quantity f is unbounded $[0, \infty)$ results are presented in terms of F which is bounded
 279 $[0, 1]$ and is the fraction of the initial subsurface load to the total initial load amplitude:

$$280 \quad F = \frac{f}{1+f} \quad f = \frac{F}{1-F} \quad (10)$$

281 Surface-only initial loading is given by $f=0$ and $F=0$ at all wavenumbers; subsurface-only
 282 initial loading is given by $f=\infty$ and $F=1$ at all wavenumbers; equal combined loading is
 283 given by $f=1$ and $F=0.5$ at all wavenumbers.

284

285 Once the Bouguer anomaly and topography pairs were generated, the compensation attraction
 286 (δg_C) was computed from eq. (3) by multiplying the final topography Fourier transform (H)
 287 by a theoretical Bouguer admittance function (Q_{th}), and inverse Fourier transforming, thus

$$288 \quad \delta g_C(\mathbf{x}) = \mathbf{F}^{-1} [Q_{th}(k)H(\mathbf{k})] \quad (11)$$

289 where Q_{th} could be any of the analytic functions given by eqs (5), (7), (8) or (9). Note that
 290 isotropic admittances were used, as functions of $k = |\mathbf{k}|$. The parameters of the admittance
 291 functions (crust and mantle densities and crustal thickness) were identical to those of the
 292 synthetic model (unless otherwise specified). An isostatic anomaly (Δg_I) was then recovered

293 by subtracting the compensation attraction from the synthetic model Bouguer anomaly as in
294 eq. (4). Hence, because there are no non-flexural signals in the synthetic Bouguer anomaly
295 and topography, if the correct compensation model (admittance function) is chosen then the
296 compensation attraction should exactly reproduce the Bouguer anomaly, yielding uniformly
297 zero isostatic anomalies.

298

299 **3.2 Tests of varying T_e and F**

300 Here, 10,100 synthetic Bouguer anomaly and topography pairs were generated from
301 combined loading on plates with T_e ranging from 0 to 100 km, in steps of 1 km (101 values),
302 and F ranging from 0 to 0.99, in steps of 0.01 (100 values).

303

304 *3.2.1 Assuming Airy isostasy*

305 In the first test ('test A'), the compensation attraction was computed for each of the 10,100
306 synthetic models from its topography using the Airy isostatic admittance function, i.e.

307 $\delta g_C = \mathbf{F}^{-1}[Q_A H]$ with Q_A given by eq. (5). The densities and compensation depth, z_m , were
308 set equal to the synthetic model values, given above. The isostatic anomaly corresponding to
309 a synthetic model (F, T_e) value was then computed using eq. (4), and its standard deviation
310 (σ_{IA}) determined. Rather than plotting the standard deviations directly, they were normalized
311 (to $\bar{\sigma}_{IA}$) by the Bouguer anomaly standard deviation at the model (F, T_e) value and then
312 plotted, shown in Fig. 1(a). This normalization was performed to place a proper perspective
313 on apparently large isostatic anomalies, which might not be large compared to the
314 corresponding Bouguer anomalies. Thus $\bar{\sigma}_{IA}$ is interpreted as the error in using the chosen
315 compensation model, in this case the Airy model, i.e. assuming $T_e = F = 0$, when the reality is
316 that T_e and F are not necessarily zero. As noted in the caption to Fig. 1, the maximum value

317 of $\bar{\sigma}_{IA}$ in Fig. 1(a) is 2.14, occurring when the model T_e is 100 km and there is no subsurface
 318 load ($F = 0$). An alternative way of phrasing this is that the total amplitude of the isostatic
 319 anomaly at these T_e and F values is 214 per cent that of the corresponding Bouguer anomaly.
 320

321 Fig. 1(a) shows that when the synthetic models were generated with $T_e = 0$ km (for any value
 322 of the initial subsurface load fraction F), the isostatic anomalies were uniformly zero,
 323 confirming that the compensation attraction perfectly reproduces the Bouguer anomaly, and
 324 that the method used to recover isostatic anomalies performs adequately. Recall from Section
 325 2.1 that when $T_e = 0$ km, surface and subsurface loading are indistinguishable in their results,
 326 explaining why the locus of zero standard deviation extends from $F = 0$ to $F = 1$.
 327

328 Fig. 1(a) suggests that, when the increased power of the Bouguer anomalies is accounted for,
 329 the Airy model is appropriate for many combinations of the plate's actual T_e and F values (
 330 $\bar{\sigma}_{IA} < 10$ per cent, for example). But both of these interpretations are misleading, as Fig. 2(a)
 331 shows, which plots the power spectra of the Bouguer and isostatic anomalies. Fig. 2(a)
 332 demonstrates that in regions of the (F, T_e) -space where $\bar{\sigma}_{IA}$ is low but $T_e > 0$ and F is high
 333 (e.g. $T_e = 25$ km, $F = 0.75$), the low values of $\bar{\sigma}_{IA}$ are caused by low relative isostatic
 334 anomaly power at mid-to-long wavelengths only, and at mid-to-short wavelengths, the
 335 Bouguer and isostatic anomalies actually have equal power. That is, the low values of $\bar{\sigma}_{IA}$ in
 336 Fig. 1(a) are biased by the large difference in long wavelength power. And because the
 337 Earth's gravity and topography naturally have red power spectra (high power at long
 338 wavelengths, low power at short), the high isostatic anomaly power at short wavelengths,
 339 relative to the Bouguer anomaly, contributes much less to the overall whole-spectrum power
 340 difference, by several orders of magnitude. Thus, minimization of the isostatic anomaly

341 standard deviation alone (as done by several studies, noted in Section 1) will give misleading
342 conclusions as to the actual state of compensation/support, even when the Bouguer anomaly
343 standard deviation is accounted for. Instead, the spectra of Bouguer and isostatic anomalies
344 over a region should be compared (and one should use Bouguer and not free air anomalies in
345 the comparison since the gravity effect of the topography has been removed when generating
346 both Bouguer and isostatic anomalies – making them compatible – but not removed from free
347 air anomalies). And it is only when the isostatic anomaly power spectrum is much less than
348 that of the Bouguer anomaly at all wavelengths that the conclusion should be made that the
349 assumed compensation mechanism is a faithful representation of the actual compensation
350 mechanism. Estimates of the isostatic parameters (e.g. T_e) needed to choose or refine the
351 compensation model can be readily obtained from spectral methods (e.g. Kirby 2014).

352

353 Fig. 2(a) shows that when subsurface loading dominates ($F > 0.5$) the Airy isostatic
354 anomalies have equal power to the Bouguer anomalies at mid-to-short wavelengths, even for
355 relatively low-rigidity plates ($T_e \leq 25$ km). Furthermore, their spectral content becomes
356 higher-power and longer-wavelength as T_e increases. If the plate is strong and F close to 1,
357 the initial subsurface loads will not cause a surface (topographic) deflection and their
358 presence will not be measurable in the topography Fourier transform, H ; therefore, the
359 compensation attraction, $\mathbf{F}^{-1}[Q_A H]$, will be small, reflecting only any low-amplitude initial
360 surface loads that may be present. Since the Bouguer anomaly will be large, due to the large
361 subsurface loads, the isostatic anomaly will also be large.

362

363 Fig. 2(a) also shows that, when surface loading dominates ($F < 0.5$), the Airy isostatic
364 anomalies can have greater power than the Bouguer anomalies at the shorter wavelengths, a
365 phenomenon that becomes more pronounced when T_e is large and subsurface loading is

366 reduced. In this case, the surface loads are supported by the plate's rigidity, so H has higher
367 power than it would on a weak plate, leading to an excess of power in the compensation
368 attraction by virtue of $\mathbf{F}^{-1}[Q_A H]$, which propagates into the isostatic anomalies.

369

370 *3.2.2 Assuming surface loading*

371 In the second test ('test B'), the compensation attraction was computed for each of the 10,100
372 synthetic models from its topography using the surface-loading admittance function, i.e.

373 $\Delta g_C = \mathbf{F}^{-1}[Q_T H]$ with Q_T given by eq. (7) using the T_e value of the synthetic model. The

374 densities and compensation depth, z_m , were set equal to the synthetic model values, given

375 above. The isostatic anomaly was then computed using eq. (4), and its normalized standard

376 deviation determined and plotted in Fig. 1(b) at the location given by the model (F, T_e) value.

377 The normalized standard deviation is interpreted here as the error in assuming only initial

378 surface loading; i.e., using the correct T_e , but assuming $F = 0$, when the reality is that F is not

379 necessarily zero. As noted in the caption to Fig. 1, the maximum value of $\bar{\sigma}_{IA}$ in Fig. 1(b) is

380 0.36, occurring when the model T_e is 100 km and there is almost no surface load ($F = 0.99$).

381 So whereas the assumption of Airy isostasy gave rise to isostatic anomalies having

382 amplitudes at 214 per cent of the Bouguer anomaly, when surface-only loading is assumed

383 the isostatic anomalies are much smaller, reaching only 36 per cent of the Bouguer anomaly

384 amplitude.

385

386 First, Fig. 1(b) shows that when the synthetic models were generated with $F = 0$ (for any

387 value of T_e), the isostatic anomalies were uniformly zero, showing that the compensation

388 attraction perfectly reproduces the Bouguer anomaly, and that the method used to recover

389 isostatic anomalies performs adequately as long as T_e is known. Zero isostatic anomalies also

390 occur when $T_e = 0$ km, for any value of F because surface and subsurface loading are
391 indistinguishable in their results, as explained before.

392

393 As discussed above, the power spectra reveal more information than the standard deviation
394 plots. The power spectra for test B are shown in Fig. 2(b). When initial loading is more
395 surface than subsurface ($F < 0.5$), a comparison of Figs 2(a) and 2(b) shows that the
396 assumption of surface loading with a knowledge of the actual T_e generally results in lower-
397 power isostatic anomalies than the assumption of Airy isostasy, especially in the mid-to-short
398 wavelengths. Not shown in Fig. 2(b) are the isostatic anomaly spectra for $F = 0$, which are
399 uniformly zero at all wavelengths, as expected.

400

401 In contrast, when initial loading is more subsurface than surface ($F > 0.5$), the assumption of
402 surface loading results in higher-power long wavelength isostatic anomalies compared to
403 Airy isostasy, by as much as two orders of magnitude. In conclusion, knowledge of T_e will
404 not necessarily provide an improved compensation model.

405

406 3.2.3 Assuming combined loading

407 In the third test ('test C'), the compensation attraction was computed for each of the 10,100
408 synthetic models from its topography using the combined-loading admittance function, i.e.

409 $\delta g_c = \mathbf{F}^{-1} [Q_{TB} H]$ with Q_{TB} given by eq. (9) using the T_e and F values of the synthetic model.

410 The densities and compensation depth, z_m , were set equal to the synthetic model values, given
411 above. The isostatic anomaly was then computed using eq. (4), and its normalized standard
412 deviation determined and plotted in Fig. 1(c) at the location given by the model (F, T_e) value.

413 The normalized standard deviation is interpreted here as the error in assuming combined
414 loading; i.e., using the correct T_e and F . In other words, there should be no error.

415

416 In tests A and B the non-zero isostatic anomalies were explained as arising from a lack of
417 knowledge about the actual compensation mechanisms; incorrectly assuming Airy isostasy in
418 test A, and surface-loading in test B, when the reality was combined loading of a rigid plate.
419 However, in test C, since the model T_e , F , z_m and density values are known and used to
420 compute the compensation attraction, it would be reasonable to expect that the correct
421 compensation model (the correct theoretical admittance) had been used, giving uniformly
422 zero $\bar{\sigma}_{LA}$ over Fig. 1(c). While this is not seen, the $\bar{\sigma}_{LA}$ values are somewhat smaller for test
423 C than test B [with a maximum value of $\bar{\sigma}_{LA}$ in Fig. 1(c) of 30 per cent at $T_e = 100$ km, $F =$
424 0.3], suggesting that some knowledge of subsurface loading is an improvement over the
425 assumption of surface loading only, especially for almost complete surface-only ($F < 0.05$)
426 and subsurface-only ($F > 0.95$) initial loading. Nevertheless, when the power spectra are
427 analysed (Fig. 2c), the short-wavelength isostatic anomaly power is considerable for $T_e \geq 10$
428 km and $F \geq 0.25$. Compared to the results from Airy isostasy (test A, Fig. 2a), the long-
429 wavelength isostatic anomaly power from combined loading is several orders of magnitude
430 lower when F is low or high, but not with intermediate values ($0.25 \leq F \leq 0.75$) when the
431 power is similar at all wavelengths. Importantly, the power is not uniformly zero, when it
432 should be.

433

434 Evidently, further analysis is required, and this analysis should focus upon the role of
435 subsurface loading, since in the tests conducted so far the isostatic anomaly power increases
436 with increasing subsurface load amplitude, for a given T_e . The investigation should also focus
437 on the admittance, since, as discussed in Section 3.2.1, the lack of subsurface load signal in
438 the surface topography transform, H , for high T_e leads to inaccurate compensation attractions

439 by virtue of $\mathbf{F}^{-1}[Q_{th}H]$. Indeed, a full explanation can only be made by studying the complex
 440 nature of the admittance.

441

442 3.2.4 Complex admittance

443 When used in flexural studies, the admittance is assumed to be real-valued. However, in
 444 general – when many processes are operating – it will be complex (e.g. Forsyth 1985; Kirby
 445 & Swain 2009; Kirby 2014). This can be seen in the formula to estimate the admittance. The
 446 observed Bouguer admittance, Q_{obs} , is estimated from observed data (real or synthetic) via a
 447 formula of the type

$$448 \quad Q_{obs}(\mathbf{k}) = \frac{\langle G(\mathbf{k})H(\mathbf{k})^* \rangle}{\langle H(\mathbf{k})H(\mathbf{k})^* \rangle} \quad (12)$$

449 (e.g. Forsyth 1985) where G is the Fourier transform of the Bouguer anomaly, H is the
 450 Fourier transform of the topography, the asterisk denotes complex conjugation, and the angle
 451 brackets indicate some averaging process which can give an isotropic (1-D) $Q(k)$, or an
 452 anisotropic (2-D) $Q(\mathbf{k})$ (see Kirby (2014) for a summary). In general, the numerator of eq.
 453 (12) will be complex, as shown ahead.

454

455 Consider two noise-free, independent processes, for instance surface (T) and subsurface (B)
 456 loading. Let $G_T = Q_T H_T$, and $G_B = Q_B H_B$, in the Fourier domain, where the Q s are real-
 457 valued, given by eqs (7) and (8), and all variables are functions of wavevector \mathbf{k} . Then the
 458 final gravity will be $G = G_T + G_B$, and the final topography will be $H = H_T + H_B$ (Forsyth
 459 1985). Substituting these expressions in eq. (12), the combined admittance is

$$460 \quad Q_{TB}' = \frac{\langle (Q_T H_T + Q_B H_B)(H_T + H_B)^* \rangle}{\langle (H_T + H_B)(H_T + H_B)^* \rangle} \quad (13)$$

461 Letting $H_T = |H_T|e^{i\alpha_T}$ and $H_B = |H_B|e^{i\alpha_B}$, where the phases α are functions of wavenumber,
 462 it can be shown that the combined admittance is indeed complex:

$$463 \quad Q_{TB}' = \frac{\langle Q_T |H_T|^2 + Q_B |H_B|^2 + (Q_T + Q_B) |H_T| |H_B| \cos \theta \rangle + i \langle (Q_T - Q_B) |H_T| |H_B| \sin \theta \rangle}{\langle |H_T|^2 + |H_B|^2 + 2 |H_T| |H_B| \cos \theta \rangle} \quad (14)$$

464 where the phase difference $\theta = \alpha_T - \alpha_B$ is also wavenumber-dependent.

465

466 Eq. (14) can be turned into a theoretical analytic equation for admittance in the following
 467 manner. Write eq. (14) as

$$468 \quad Q_{TB}' = \frac{Q_1' + Q_2'}{Q_3'} \quad (15)$$

469 and treat the real part of the numerator (Q_1') initially:

$$470 \quad Q_1' = \langle Q_T |H_T|^2 + Q_B |H_B|^2 + (Q_T + Q_B) |H_T| |H_B| \cos \theta \rangle \quad (16)$$

471 In his paper, Forsyth (1985) introduced the initial loading ratio, f , as the ratio of the weight of
 472 the applied load at the Moho to that of the applied load on the surface, and expressed it as

$$473 \quad f(\mathbf{k}) r |H_T(\mathbf{k})| = \xi |H_B(\mathbf{k})| \quad (17)$$

474 where ξ is defined in eq. (7), and r in Section 2.1. Thus, using eq. (17), eq. (16) becomes

$$475 \quad Q_1' = \langle Q_T |H_T|^2 + Q_B f^2 r^2 \xi^{-2} |H_T|^2 + (Q_T + Q_B) f r \xi^{-1} |H_T|^2 \cos \theta \rangle \quad (18)$$

476 The averaging is typically performed over the 360° of an annulus in the wavevector, \mathbf{k} , space,

477 with many annuli spanning the space, yielding an isotropic quantity as a function of

478 wavenumber modulus, $k = |\mathbf{k}|$. The wavenumber-dependent functions ξ , Q_T and Q_B are

479 smooth and slowly varying analytic functions, so if the annuli are very narrow, spanning a

480 very small range of k , they can be treated as constants and taken out of the averaging. Thus

481 eq. (18) can be written as

482 $Q_1' = Q_T \langle |H_T|^2 \rangle + Q_B r^2 \xi^{-2} \langle f^2 |H_T|^2 \rangle + (Q_T + Q_B) r \xi^{-1} \langle f |H_T|^2 \cos \theta \rangle$ (19)

483 Taking the loading ratio, $f(\mathbf{k})$, and phase difference, $\theta(\mathbf{k})$, out of the averaging procedure is
 484 perhaps harder to justify since they are dependent on the data and most likely will not be
 485 smooth and slowly varying but instead potentially highly variable functions of wavevector \mathbf{k} .
 486 Additionally, the phase difference contains the phase information of H_T and H_B . Therefore in
 487 order to obtain an all-purpose analytic expression for Q_{TB}' , one that is data-independent akin
 488 to eqs (7) – (9), one must make three assumptions. First, Kirby & Swain (2009) reasoned that
 489 if the amplitudes and phases of the data are independent then one can write

490 $\langle f(\mathbf{k}) |H_T(\mathbf{k})|^2 \cos \theta(\mathbf{k}) \rangle \approx \langle f(\mathbf{k}) |H_T(\mathbf{k})|^2 \rangle \langle \cos \theta(\mathbf{k}) \rangle$. Their second assumption was that if

491 the amplitudes of the surface and subsurface processes are independent then $\langle f(\mathbf{k}) |H_T(\mathbf{k})|^2 \rangle$

492 $\approx f(k) \langle |H_T(\mathbf{k})|^2 \rangle$. Third, they assumed that if the phase difference is independent of

493 azimuth, then $\langle \cos \theta(\mathbf{k}) \rangle \approx \cos \theta(k)$. Hence, with these assumptions, eq. (19) becomes

494 $Q_1' = [Q_T + Q_B f^2 r^2 \xi^{-2} + (Q_T + Q_B) f r \xi^{-1} \cos \theta] \langle |H_T|^2 \rangle$ (20)

495

496 Applying the same treatment to the imaginary part of the numerator of eq. (14) and its
 497 denominator gives

498 $Q_2' = i(Q_T - Q_B) f r \xi^{-1} \sin \theta \langle |H_T|^2 \rangle$ (21)

499 and

500 $Q_3' = [1 + f^2 r^2 \xi^{-2} + 2 f r \xi^{-1} \cos \theta] \langle |H_T|^2 \rangle$ (22)

501

502 A final observation made by Kirby & Swain (2009), made on synthetic data, was that
 503 correlated initial loads had the property $\langle \cos \theta \rangle \approx 1$, while randomly correlated loads had
 504 $\langle \cos \theta \rangle \approx 0$, implying that the correlated loading regime is characterised by $\theta = 0^\circ$ and the
 505 randomly-correlated regime by $\theta = 90^\circ$ (see also Wieczorek (2007)). Since in this study the
 506 synthetic loads are all randomly correlated we can use $\cos \theta = 0$ and $\sin \theta = 1$ in eqs (20) –
 507 (22), which makes eq. (15) become

$$508 \quad Q_{TB}' = \frac{(Q_T + Q_B f^2 r^2 \xi^{-2}) \langle |H_T|^2 \rangle + i(Q_T - Q_B) f r \xi^{-1} \langle |H_T|^2 \rangle}{(1 + f^2 r^2 \xi^{-2}) \langle |H_T|^2 \rangle} \quad (23)$$

509 Cancelling the $\langle |H_T|^2 \rangle$ terms, and using eqs (7) and (8), the analytic expression can be
 510 written as

$$511 \quad Q_{TB}'(k) = -2\pi \mathcal{G} \Delta \rho_0 e^{-kz} \left(\frac{\xi + \phi f^2 r^2 + i f r (1 - \phi \xi)}{\xi^2 + f^2 r^2} \right) \quad (24)$$

512 Note how the real part of eq. (24) is identical to eq. (9), the (real) combined-loading
 513 admittance.

514

515 The (complex) 2-D observed admittance of the synthetic models was computed in
 516 wavenumber space using eq. (12) and using Slepian multitapers (see Fig. 3 caption). Rather
 517 than plotting the 2-D observed admittance estimates as a function of $\mathbf{k} = (k_x, k_y)$, they were
 518 plotted (without averaging around annuli) on 1-D graphs as functions of their radial
 519 wavenumber $k = |\mathbf{k}|$ as green dots in Fig. 3. The blue curves in Fig. 3 show the real and
 520 imaginary theoretical admittance functions from eq. (24).

521

522 Figs 3(a) and (b) show that when the synthetic Bouguer anomaly and topography were
 523 generated from a plate with $T_e = 0$ km (Airy isostasy, implying $F = 0$) the real and imaginary

524 observed admittance estimates agreed very well with the theoretical curves. When they were
525 generated from surface-loading only ($F = 0$) on a plate of $T_e = 40$ km, there was also good
526 agreement (Figs 3c and d). Note how the observed imaginary parts in Figs 3(b) and (d) are
527 almost uniformly zero at all wavelengths, agreeing with the theoretical predictions of eq. (24)
528 when $f = F = 0$.

529

530 When the synthetic Bouguer anomaly and topography were generated from combined loading
531 ($F = 0.5$) of a plate with $T_e = 40$ km, the real observed admittance estimates show a certain
532 degree of scatter about the theoretical curve (Fig. 3e), particularly at medium to short
533 wavelengths, which is expected and arises from random correlations between the two
534 synthetic loads (Kirby & Swain 2008). In contrast to the Airy and surface loading models, the
535 observed admittance in the combined loading case gains a non-zero imaginary part (Fig. 3f),
536 in agreement with eqs (14) and (24) because there are now two independent processes in
537 action. But it can be seen that the imaginary observed admittance estimates in Fig. 3(f) do not
538 follow the imaginary theoretical admittance curve; they cluster about the zero-admittance
539 axis, rather than the theoretical imaginary curve. [Interestingly, there are no estimates that lie
540 outside the curve (and its negative reflection about the axis), so it could be said that the
541 theoretical curve provides an envelope within which all observed imaginary admittance
542 estimates fall.]

543

544 This mismatch between the observed and theoretical imaginary admittance provides the
545 answer to the problem raised above: why is the isostatic anomaly for combined loading non-
546 zero (Fig. 1c), when the loading model is known exactly? First note that with any type of
547 noise-free synthetic flexure model, in order to retrieve zero isostatic anomalies the
548 compensation attraction must be perfectly recovered from a formula of the type

549 $\delta g_C = \mathbf{F}^{-1}[Q_{th}H]$, and for this the admittance Q needs to be an accurate representation of the
550 actual synthetic model plate filter, where, given initial loads the plate filter produces a gravity
551 anomaly and topography. Whether this filter is manifested as flexural equations or as
552 admittances is unimportant. Indeed, if the observed (complex) admittance, Q_{obs} , was
553 computed from the synthetic Bouguer anomaly and topography through eq. (12) and then a
554 compensation attraction retrieved via $\delta g_C = \mathbf{F}^{-1}[Q_{obs}H]$, the synthetic Bouguer anomaly and
555 retrieved compensation attraction would be almost identical (but not exactly due to multitaper
556 averaging when computing the admittance). This would apply to real-Earth data too, except
557 here Q_{obs} would contain much more than just isostatic information.

558

559 Consequently, in the surface-only loading model the real theoretical admittance curves are a
560 very good fit to the real observed admittance estimates (Fig. 3c), so that when the theoretical
561 admittance function, eq. (7), is used with the topography Fourier transform in eq. (11), or

562 $\delta g_C = \mathbf{F}^{-1}[Q_T H]$, the retrieved compensation attraction exactly reproduces the actual

563 Bouguer anomaly. That is, performing $\mathbf{F}^{-1}[Q_T H]$ with a theoretical admittance gives almost

564 the same results as performing $\mathbf{F}^{-1}[Q_{obs} H]$ with an observed admittance; both yield a

565 compensation attraction that exactly matches the Bouguer anomaly (for these noise-free,

566 surface-loading synthetic models). The same is true for the Airy isostatic case.

567

568 In the combined-loading models the compensation attraction is also obtained using a real-

569 valued theoretical analytic admittance, $\delta g_C = \mathbf{F}^{-1}[Q_{TB} H]$ with Q_{TB} given by eq. (9). But even

570 though Q_{TB} is a fair fit to the real, observed admittance (Fig. 3e), use of a real-valued analytic

571 admittance will always yield antiphase (180°) compensation attraction and topography when

572 in fact the combined model synthetic Bouguer anomaly and topography actually have random

573 phase due to the independence of the initial surface and subsurface loads. This means that the
574 difference $\Delta g_B - \delta g_C$ will not account for the out-of-phase harmonics between Δg_B and h , and
575 will be non-zero. Since $\Delta g_B - \delta g_C$ is interpreted as the isostatic anomaly (eq. (1)), it appears
576 that there exists a non-zero isostatic anomaly even in this noise-free example. The conclusion
577 is that the compensation model has failed, even if T_e , F , densities and depths have actually
578 been estimated accurately.

579

580 And unfortunately the situation cannot be remedied by use of a complex analytic admittance,
581 which would at least generate out-of-phase gravity and topography; for example, using Q_{TB}'
582 from eq. (24) to generate the compensation attraction from $\delta g_C = \mathbf{F}^{-1} \left[Q_{TB}' H \right]$. While the real
583 part of Q_{TB}' is a fair fit to the real, observed admittance (Fig. 3e), its imaginary part is
584 definitely not (Fig. 3f), so $\mathbf{F}^{-1} \left[Q_{TB}' H \right]$ will not yield anything remotely resembling the
585 compensation attraction. Therefore, while useful in providing a theoretical understanding of
586 phase relationships between gravity and topography, complex analytic admittance formulae
587 such as eq. (24) have little practical value. This is most likely due to the assumptions made
588 regarding the phase difference θ when deriving eq. (24).

589

590 In summary, if both surface and subsurface loading are present, then the correct
591 compensation and isostatic anomalies cannot be estimated using spectral methods. Complex
592 analytic admittance functions do not faithfully represent the imaginary admittance, while real
593 analytic admittance functions cannot correctly predict the phase difference between gravity
594 and topography.

595

596 3.3 Tests of varying compensation depth

597 This experiment sought to ascertain the error in estimating compensation depth, when an Airy
598 isostatic model is used to interpret gravity and topography data generated from a plate with T_e
599 > 0 km. As described in Section 3.2, 10,100 synthetic Bouguer anomaly and topography pairs
600 were generated from combined loading on plates with T_e ranging from 0 to 100 km, in steps
601 of 1 km (101 values), and F ranging from 0 to 0.99, in steps of 0.01 (100 values).
602 Lithospheric parameters were set as: crust density 2800 kg m^{-3} , crust thickness (and
603 compensation depth) 35 km, mantle density 3300 kg m^{-3} , as usual. The compensation
604 attraction was then recovered from the topography using the Airy admittance function, eq. (5)
605 , 196 times, using depth to compensation values (z_m) ranging from 5 to 200 km, in steps of 1
606 km. From each of the 1,979,600 ($= 101 \times 100 \times 196$) compensation attractions the
607 corresponding isostatic anomaly was computed using eq. (4), and its normalized standard
608 deviation ($\bar{\sigma}_{IA}$) calculated (an example for $F = 0$ is shown in Fig. 4a). For each model with a
609 certain (F, T_e) value, the value of z_m that gave the smallest $\bar{\sigma}_{IA}$ was determined, and plotted in
610 Fig. 4(b).

611

612 Both plots in Fig. 4 show how compensation depth can be overestimated if an Airy model is
613 assumed, when loading is surface-only. For example, consider a region of the Earth where the
614 actual, but unknown, T_e is 40 km, and one wishes to find the compensation depth when
615 erroneously assuming that $T_e = 0$ km. If one used a method to find the compensation depth
616 such as that presented here, i.e. by finding that compensation depth over a range of values
617 that minimized the isostatic anomaly standard deviation, one would retrieve $z_m \approx 110$ km
618 from Figs 4(a) and (b), a value greater than its true value by 75 km. For combined-loading
619 scenarios ($F > 0$), Fig. 4(a) no longer applies and one must turn to Fig. 4(b). Now, if the

620 actual T_e were 40 km but the internal load fraction $F = 0.5$, then the estimated compensation
621 depth would be approximately 50 km.

622

623 Fig. 4(b) shows that the largest overestimates of the actual compensation depth occur when
624 the initial loading is predominantly at the surface ($F < 0.5$), and then when T_e is high.

625 Furthermore, large underestimates of compensation depth occur when there exist significant
626 subsurface loads ($F > 0.8$) in a plate of high T_e . Indeed, compensation depth is only recovered
627 accurately in a very narrow range of (F, T_e) values, $T_e < 5$ km, or $0.65 < F < 0.7$,
628 approximately, as shown by the red contours in Fig. 4(b).

629

630 **4 DISCUSSION**

631 **4.1 Subsurface loading**

632 Studies have shown that that subsurface loading exists in many regions worldwide, and that it
633 plays a key role in flexure and in the compensation of surface loads (e.g. Karner & Watts
634 1983). However, it has been shown in Section 3.2.4 that combined surface and subsurface
635 loading cannot be realistically accounted for in the determination of isostatic anomalies. This
636 is because the admittance, when multiplied by the observed surface topography, does not
637 accurately model the compensation at the surface of subsurface loads generated during
638 combined loading and flexure of a plate of $T_e > 0$ km. Since subsurface loading, and its effect
639 upon the surface topography and gravity field, is part of the flexural-isostatic process, the
640 question of whether an isostatic anomaly is meaningful must be raised.

641

642 The detrimental effect of subsurface loads upon isostatic anomaly estimation is not unknown.
643 Such loads can be emplaced or generated during many tectonic events such as the obduction
644 of oceanic crust, the development of accretionary wedges and the mobilization of thrust

645 sheets and nappes during continental convergence, emplacement of subsurface plutons during
646 volcanic activity, intra-crustal thrusts, and dense downgoing slabs (e.g. Karner & Watts 1983;
647 Forsyth 1985; Jordan & Watts 2005). They may be shallow, such as crustal blocks of
648 different compositions, sedimentary basins, or igneous intrusions, or deep, such as density
649 anomalies due to crustal underplating, lithospheric thermal anomalies or deeper
650 compositional variations (Zuber *et al.* 1989). Most importantly, they are often expressed in
651 the Bouguer anomaly but not in the topography, leading to incorrect Airy isostatic anomalies
652 which are derived under the assumption that the topography is the only load acting on the
653 lithosphere (e.g. Watts & Talwani 1974; McNutt 1980; Karner & Watts 1983; Forsyth 1985).
654 For example, in their study of isostasy at orogenic belts, Karner & Watts (1983) showed that
655 the maximum crustal thickness in the Alps does not occur under the highest elevations as
656 Airy isostasy would predict, but rather under the large subsurface loads that characterise the
657 Ivrea zone. The rigidity of the lithosphere here prevents hydrostatic adjustment of the
658 subsurface density anomalies, but isostatic balance is still maintained.

659

660 The difficulty, if not impossibility, of accounting for subsurface loads when estimating
661 isostatic anomalies using transfer function (admittance) methods has been acknowledged
662 qualitatively by several authors (e.g. McNutt 1980; Ussami *et al.* 1993; Watts *et al.* 1995;
663 Harmon *et al.* 2006). This study has quantified and confirmed that. However, subsurface
664 loads can be included in an isostatic anomaly if the computation of their gravity effect is
665 performed in the space domain, as Karner & Watts (1983) did. Space domain modelling is,
666 however, comparatively tedious compared with admittance methods in that it involves
667 forward modelling the subsurface density distributions, perhaps constrained by independent
668 (e.g. seismic) data, and adjusting T_e to match the observed gravity field.

669

670 One possible approach would be to acknowledge but ignore subsurface loads, and use
671 surface-loading models only (e.g. Ussami *et al.* 1993; Watts & Moore 2017). And this might
672 be considered appropriate, given that the tests conducted in this paper have shown that the
673 real-valued combined loading admittance function, eq. (9), does not give correct isostatic
674 anomalies even when T_e , F and the compensation depth are known (test C, Fig. 1c).
675 However, the tests have also shown that the assumption of only surface loading when both
676 surface and subsurface loads are present can, in certain scenarios, yield larger isostatic
677 anomalies than when Airy isostasy is assumed (compare Figs 1a and b). It is worth noting
678 though that accurate knowledge of T_e , F and compensation depth (test C, Fig. 1c) does yield
679 smaller average isostatic anomalies than when these parameters are unknown, even though
680 the spatial distribution of computed isostatic anomalies may not reflect their actual
681 distribution.

682

683 **4.2 Coherence transition wavelength**

684 The assumption of local isostatic compensation underestimates the ability of the lithosphere
685 to mechanically support topographic loads, resulting in an overestimation of the depth to
686 compensation, as noted in Section 1, and demonstrated in Section 3.3. In many studies it is
687 often assumed that most surface topographic loads are isostatically compensated, or at least
688 that mechanically supported loads are of such a small wavelength as to be irrelevant to that
689 study's conclusions (e.g. Kaban *et al.* 1999, 2004). Often, this 'transition' wavelength (from
690 compensated to supported topography) is assumed to be uniform across the study area, with
691 an arbitrarily chosen value (e.g. Martinec 1994b; Kaban *et al.* 2004; Bagherbandi *et al.*
692 2015). Such assumptions are misplaced, however, as the transition wavelength depends upon
693 T_e and initial loading ratio (f).

694

695 One possible method to separate compensated from supported topography is to analyse the
 696 wavenumber-domain coherence between Bouguer anomalies (G) and topography (H):

$$697 \quad \gamma^2(k) = \frac{\left| \langle G(\mathbf{k})H(\mathbf{k})^* \rangle \right|^2}{\langle G(\mathbf{k})G(\mathbf{k})^* \rangle \langle H(\mathbf{k})H(\mathbf{k})^* \rangle} \quad (25)$$

698 (e.g. Forsyth 1985). The Bouguer coherence will have values close to zero at short
 699 wavelengths because such small loads can be adequately supported by the mechanical
 700 strength of the plate (Fig. 5): small surface loads retain a topographic signature but do not
 701 generate a Bouguer anomaly because the compensating interface (e.g. the Moho) does not
 702 flex; small subsurface loads generate a Bouguer anomaly but do not produce a surface
 703 topography by flexure. At the other end of the spectrum, long wavelength loads cannot be
 704 mechanically supported, and the loads (both surface and subsurface) are hydrostatically
 705 compensated; the Bouguer coherence is hence unity. There is, then, a transition wavelength
 706 (λ_T), at which the Bouguer anomaly and topography transition from being incoherent (with
 707 wavelength $\lambda < \lambda_T$) to coherent ($\lambda > \lambda_T$), which can also be interpreted as the wavelength at
 708 which the loads transition from being supported to compensated.

709
 710 By way of example, Fig. 6(b) shows a map of the value of the predicted Bouguer coherence
 711 transition wavelength over Europe and surrounding seas, where λ_T is taken to be the
 712 wavelength at which the coherence has a value of 0.5. The predicted coherence was estimated
 713 by application of the fan-wavelet adaptation of Forsyth's (1985) method (Kirby & Swain
 714 2008) to EGM2008 gravity data (Pavlis *et al.* 2012), rock-equivalent topography from the
 715 Earth2014 model (Hirt & Rexer 2015), and the depths and densities of the CRUST1.0 model
 716 (Laske *et al.* 2013). The resulting T_e map (Fig. 6a) broadly agrees with the estimate obtained
 717 by Pérez-Gussinyé & Watts (2005) using the multitaper method, and is not discussed here.

718

719 Note that the map in Fig. 6(b) shows the transition wavelength of both (post-flexure) surface
720 and subsurface loads, because the predicted coherence (and hence λ_T) is computed by
721 Forsyth's (1985) method that assumes that both types of load are present. Fig. 6(b) shows that
722 the transition wavelength from compensated to supported topography is highly variable, with
723 the weaker lithosphere of western and southern Europe much less able to mechanically
724 support large loads than the stronger eastern European and Asian lithosphere. The lithosphere
725 in western and southern Europe is able to support loads with wavelengths <400 km
726 approximately, while the stronger lithosphere to the east can adequately support loads with
727 wavelengths up to ~ 1200 km.

728

729 **4.3 Implications for dynamic topography studies**

730 Under the reasoning that the Earth's actual topography is the sum of isostatic and dynamic
731 components, models of isostatic compensation are frequently used to isolate dynamic
732 topography (e.g. Forte *et al.* 1993; Perry *et al.* 2002; Boschi *et al.* 2010; Komut *et al.* 2012;
733 Bagherbandi *et al.* 2015). To date Airy isostasy has been the model of choice. Assuming Airy
734 isostasy, the undulations of a seismically-determined Moho can be used to predict the
735 isostatic topography that the Moho undulations compensate, as shown below.

736

737 Consider an incompressible crust of uniform density ρ_c overlying an inviscid mantle of
738 greater uniform density ρ_m , and itself overlain by a fluid of lesser density ρ_f , either air or
739 water. An applied surface load also of density ρ_c then deflects the crust resulting in surface
740 topography $h(x)$ and Moho relief of $w(x)$. Under Airy isostatic compensation, the pressure
741 generated from displacement of the fluid by the surface topography must balance the pressure
742 generated by displacement of the mantle by the Moho, thus

$$743 \quad -\Delta\rho_0 g h(x) = \Delta\rho_1 g w(x) \tag{26}$$

744 where g is the gravity acceleration, $\Delta\rho_0 = \rho_c - \rho_f$ and $\Delta\rho_1 = \rho_m - \rho_c$, giving

$$745 \quad w(x) = -\frac{\Delta\rho_0}{\Delta\rho_1} h(x) \quad (27)$$

746 This is the Airy isostatic case, and it is instructive to work in the wavenumber domain by
747 taking the Fourier transform of eq. (27):

$$748 \quad H(k) = -\frac{\Delta\rho_1}{\Delta\rho_0} W(k) \quad (28)$$

749 where capital letters indicate the function's Fourier transform. For example, 1 km of Moho
750 relief (W) compensates 0.179 km of subaerial surface topography (H), at any wavelength of
751 relief anomaly (using density values from Table 1).

752

753 $W(k)$ can be obtained from a seismic Moho model, then eq. (28) used to find $H(k)$ the derived
754 isostatic topography. Then, the difference between the actual, observed topography (from a
755 DEM for example) and the isostatic topography is interpreted as the dynamic topography.
756 Different authors have attributed different phenomena to explain the dynamic processes that
757 support the topography (e.g. Molnar *et al.* 2015), and it is not the purpose of this paper to
758 comment on those. However, as Molnar *et al.* (2015) point out, the error on most seismic
759 models of the Moho is at least 2 km and often more than 5 km, which by eq. (28) imparts an
760 error of at least 0.4 – 0.9 km on the isostatic, and thus dynamic, topography when derived in
761 this manner. This error, they say, is often greater than the estimate.

762

763 Another source of error arises from the omission of plate rigidity in solutions. The
764 contribution of flexural rigidity to such studies of dynamic topography can be estimated by
765 considering the amount of surface topography that is compensated or supported by a
766 specified Moho relief anomaly, in an extension of the Airy isostasy case discussed above.

767 When the plate possesses non-zero rigidity, two processes must be considered. The first
 768 process, surface loading, asks ‘when an initial surface load applied to a plate generates a
 769 Moho relief anomaly of amplitude 1 km after flexure, what is the amplitude of the post-
 770 flexural surface topography?’ The solution is obtained in a similar manner to the Airy case,
 771 above, but now the forces produced by the load are laterally distributed, and a fourth-order
 772 derivative term representing the bending stress (e.g. Watts, 2001) must be introduced into eq.
 773 (26):

$$774 \quad -\Delta\rho_0 g h_T(x) = \Delta\rho_1 g w_T(x) + D \frac{\partial^4 w_T(x)}{\partial x^4} \quad (29)$$

775 Now using the subscript ‘T’ to denote surface loading, if D is spatially uniform, eq. (29) can
 776 be solved by taking its Fourier transform:

$$777 \quad -\Delta\rho_0 g H_T(k) = \Delta\rho_1 g W_T(k) + D k^4 W_T(k) \quad (30)$$

778 or:

$$779 \quad H_T(k) = \frac{-(D k^4 + \Delta\rho_1 g)}{\Delta\rho_0 g} W_T(k) \quad (31)$$

780

781 Eq. (31) is plotted in Fig. 7(a) for several values of T_e , as a function of load wavelength ($\lambda' =$
 782 $2\pi / k'$) where the load in this case is a delta function in the Fourier domain with value -1 km
 783 at wavenumber k' and zero elsewhere (i.e. a sinusoid in the space domain), or $W_T(k) = -\delta(k -$
 784 $k')$. Under Airy isostasy ($D = 0$), $W_T = -1$ km of Moho relief compensates $H_T = 0.179$ km of
 785 surface topography, for any wavelength of Moho relief and surface topographic expression,
 786 as noted above. However, as load wavelength decreases and/or T_e increases, it becomes
 787 harder and harder to flex the plate sufficiently to produce 1 km of Moho relief. For example,
 788 under initial surface loading, in order to compensate a 500 km-wavelength Moho relief
 789 anomaly of amplitude -1 km on a plate with $T_e = 80$ km, a surface load of amplitude 4.319

790 km is required, or 4.14 km more than expected for Airy isostasy. Thus, if an Airy model is
 791 used to separate the isostatic from dynamic topography, then the isostatic topography may be
 792 underestimated and the dynamic topography therefore overestimated.

793

794 The second process, subsurface loading, asks ‘when a subsurface load applied to a plate
 795 generates a Moho relief anomaly of amplitude 1 km after flexure, what is the amplitude of the
 796 post-flexural surface topography?’ The subsurface loading analogue of eq. (30) is

$$797 \quad -\Delta\rho_1 g W_B(k) = \Delta\rho_0 g H_B(k) + D k^4 H_B(k) \quad (32)$$

798 (e.g. Forsyth 1985) which becomes

$$799 \quad H_B(k) = \frac{-\Delta\rho_1 g}{D k^4 + \Delta\rho_0 g} W_B(k) \quad (33)$$

800 Eq. (33) is plotted in Fig. 7(b) for several values of T_e , where $W_B(k) = -\delta(k - k')$. Under Airy
 801 isostasy ($D = 0$), $W_B = -1$ km of Moho relief compensates $H_B = 0.179$ km of surface
 802 topography, for any wavelength of Moho relief and surface topographic expression, as for the
 803 surface loading case, above. However, as load wavelength decreases and/or T_e increases, the
 804 surface topographic amplitude decreases to zero. For example, under initial subsurface
 805 loading, a 500 km-wavelength Moho relief anomaly of amplitude 1 km on a plate with $T_e =$
 806 80 km compensates a surface load of amplitude 0.035 km, or 0.144 km less than expected for
 807 Airy isostasy. Thus, if an Airy model is used to separate the isostatic from dynamic
 808 topography, then the isostatic topography may be overestimated and the dynamic topography
 809 therefore underestimated. Note that it is difficult to combine the expressions for surface and
 810 subsurface loading due to the likely phase difference between the two processes.

811

812 Hence it can be seen that determination of the isostatic topography, let alone dynamic
 813 topography, is difficult. In addition to dealing with the error in Moho depth, one has to decide

814 whether surface or subsurface loading dominated the region, and be confident in that decision
815 because, as seen, the two can have very different results. While one could use the loading
816 ratio, f , provided by Forsyth's (1985) coherence method to estimate the relative amounts of
817 each loading type at different wavelengths, one would also have to recreate the initial loads in
818 order to ascertain the phase difference between them. Fortunately, recreating initial loads is
819 achievable via Forsyth's (1985) method (e.g. Bechtel *et al.* 1987), and Lowry *et al.* (2000)
820 have used such an approach when computing dynamic topography in the U.S. Cordillera.
821

822 Many studies write that they avoid the effect of flexural rigidity by low-pass frequency
823 filtering the data in order to operate in harmonics where loads are hydrostatically
824 compensated rather than mechanically supported (e.g. Forte *et al.* 1993; Perry *et al.* 2002;
825 Boschi *et al.* 2010; Komut *et al.* 2012; Bagherbandi *et al.* 2015). This supposition is correct
826 and can be seen in Fig. 7. For a given T_e value, the wavelength at which the $H_T(k)$ or $H_B(k)$
827 curve intersects the Airy ($T_e = 0$) line provides the minimum wavelength at which a dynamic
828 topography study will be free from the effects of flexural rigidity. For example, if $T_e = 80$ km
829 then low-pass filtering the observed topography and Moho relief with a cut-off of 2000 km
830 (spherical harmonic degree <20) will ensure that only Airy compensation signals are present
831 in the data. If T_e is lower, 10 km say, then the cut-off only need be 500 km (spherical
832 harmonic degree <80).

833

834 But as discussed in Section 4.2 the cut-off wavelength in the above-cited studies often seems
835 to be arbitrarily chosen, and sometimes is not chosen *per se* at all, but rather is implied by the
836 resolution of the Moho depth model (e.g. Perry *et al.* 2002; Boschi *et al.* 2010; Komut *et al.*
837 2012). Many studies overcompensate and select cut-off wavelengths far in excess of the
838 minimum needed (e.g. Forte *et al.* 1993; Bagherbandi *et al.* 2015). Three exceptions who

839 have taken elastic thickness into account when choosing cut-off wavelengths are Kaban *et al.*
840 (2004) and Braun *et al.* (2014), albeit differently to here, and Watts & Moore (2017) though
841 their study is of global average spectra.

842

843 **5 CONCLUSIONS**

844 It has been shown here that attempts to determine mechanisms of isostatic compensation of
845 surface topographic features, and thence isostatic gravity anomalies, are prone to failure for
846 several reasons. First, many attempts omit, by accident or design, flexural rigidity from
847 computations. As with any inversion method, omission of an important inversion parameter
848 will affect the values of those remaining parameters chosen for estimation, for example, the
849 depth to compensating interface. Experiments with synthetic models have shown that the
850 flexural rigidity, or its geometric analogue the effective elastic thickness, is indeed an
851 important parameter in isostatic compensation and its omission can affect isostatic anomaly
852 standard deviations by up to 214 per cent of the corresponding Bouguer anomaly standard
853 deviation, for the models considered here. To address these shortcomings, it is recommended
854 that (1) elastic thickness be included, if possible, in isostatic anomaly modelling, and (2) that
855 isostatic anomaly power spectra, rather than standard deviations, relative to the Bouguer
856 anomaly be analysed. The second recommendation is important because isostatic anomalies
857 may have a low standard deviation but still have high power at short wavelengths relative to
858 the Bouguer anomaly. This happens because of the redness of the Earth's gravity and
859 topography spectra: for example, a power difference of 10^3 between Bouguer and isostatic
860 anomalies at long wavelengths contributes much more to the whole-spectrum average power
861 than does a 10^3 power difference at short wavelengths.

862

863 Equally as important as elastic thickness is the role of subsurface loads. Such loads are
864 isostatically compensated and manifested in the present-day topography. If they are present
865 yet ignored and the topography is assumed to comprise only surface loads, then inversion of
866 the topography for isostatic model parameters will, again, yield incorrect parameter estimates.
867 But most importantly, transfer function (admittance) methods cannot properly account for
868 subsurface loads when they coexist with surface loads. This is true even if complex-valued
869 analytic transfer functions are employed that should, in theory, account for phase differences
870 between the two loads, but in practice do not. This suggests that more ingenious methods
871 must be devised in order to obtain more realistic isostatic anomalies.

872

873 Isostatic anomalies can also be inverted for the depth to the compensating density interface,
874 often assumed to be the Moho. Again, if T_e is ignored and an Airy isostatic model assumed,
875 then the Moho depth estimate will always be larger than the reality. This further affects
876 derived values of the dynamic topography when Moho depths are used for its estimation.
877 Coupled with an uncertainty of the role of subsurface loading, omission of T_e from analyses
878 can lead to very large overestimates, or moderate underestimates of the dynamic topography.
879 In addition to dynamic support, T_e plays a strong role in the mechanical support of
880 topography. The size of topographic load that can be supported by a plate, as opposed to
881 hydrostatically compensated, can be found by estimation of the Bouguer coherence transition
882 wavelength.

883

884 Given the noted shortcomings of local isostatic models, one must question their continued
885 use. If T_e and the degree of subsurface loading can be reliably estimated in a region then
886 derived isostatic anomalies will have slightly lower power than those obtained under the
887 assumption of Airy isostasy, especially for high values of T_e and predominantly surface

888 loading (compare Figs 1a and b, 2a and b). However, as no transfer function-based method
889 can ever model the gravitational attraction of the compensation with 100% accuracy (for T_e
890 and $F > 0$), non-zero isostatic anomalies will always be observed, and potentially
891 misinterpreted.

892

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896

897 **REFERENCES**

898 Abd-Elmotaal, H., 1993. Vening Meinesz Moho depths: traditional, exact and approximated,
899 *Manuscr. Geod.*, **18**, 171-181.

900 Airy, G.B., 1855. On the computations of the effect of the attraction of the mountain-masses,
901 as disturbing the apparent astronomical latitude of stations in geodetic surveys, *Phil. Trans.*
902 *R. Soc. Lond.*, **145**, 101-104.

903 Aitken, A.R.A., Altinay, C. & Gross, L., 2015. Australia's lithospheric density field, and its
904 isostatic equilibration, *Geophys. J. Int.*, **203**, 1961-1976.

905 Bagherbandi, M. & Sjöberg, L.E., 2012. Non-isostatic effects on crustal thickness: a study
906 using CRUST2.0 in Fennoscandia, *Phys. Earth planet. Inter.*, **200-201**, 37-44.

907 Bagherbandi, M., Tenzer, R., Sjöberg, L.E. & Abrehdary, M., 2015. On the residual isostatic
908 topography effect in the gravimetric Moho determination, *J. Geodyn.*, **83**, 28-36.

909 Balmino, G., Vales, N., Bonvalot, S. & Briais, A., 2012. Spherical harmonic modelling to
910 ultra-high degree of Bouguer and isostatic anomalies, *J. Geod.*, **86**, 499-520.

911 Banks, R.J., Parker, R.L. & Huestis, S.P., 1977. Isostatic compensation on a continental
912 scale: local versus regional mechanisms, *Geophys. J. R. astr. Soc.*, **51**, 431-452.

913 Bechtel, T.D., Forsyth, D.W. & Swain, C.J., 1987. Mechanisms of isostatic compensation in
914 the vicinity of the East African Rift, Kenya, *Geophys. J. R. astr. Soc.*, **90**, 445-465.

915 Boschi, L., Faccenna, C. & Becker, T.W., 2010. Mantle structure and dynamic topography in
916 the Mediterranean Basin, *Geophys. Res. Lett.*, **37**, L20303, doi:10.1029/2010GL045001.

917 Braun, J., Simon-Labric, T., Murray, K.E. & Reiners, P.W., 2014. Topographic relief driven
918 by variations in surface rock density, *Nat. Geosci.*, **7**, 534-540.

919 Brotchie, J.F. & Silvester, R., 1969. On crustal flexure, *J. geophys. Res.*, **74**, 5240-5252.

920 Burov, E.B. & Diament, M., 1995. The effective elastic thickness (T_e) of continental
921 lithosphere: what does it really mean? *J. geophys. Res.*, **100**(B3), 3905-3927.

922 Cordell, L., Zorin, Y.A. & Keller, G.R., 1991. A decompensative gravity anomaly and deep
923 structure of the region of the Rio Grande rift, *J. geophys. Res.*, **96**(B4), 6557-6568.

924 Dorman, L.M. & Lewis, B.T.R., 1970. Experimental isostasy: 1. Theory of the determination
925 of the Earth's isostatic response to a concentrated load, *J. geophys. Res.*, **75**, 3357-3365.

926 Dorman, L.M. & Lewis, B.T.R., 1972. Experimental isostasy: 3. Inversion of the isostatic
927 Green function and lateral density changes, *J. geophys. Res.*, **77**, 3068-3077.

928 England, P. & Molnar, P., 2015. Rheology of the lithosphere beneath the central and western
929 Tien Shan, *J. geophys. Res.*, **120**, 3803-3823.

930 Eshagh, M., 2016. A theoretical discussion on Vening Meinesz-Moritz inverse problem of
931 isostasy, *Geophys. J. Int.*, **207**, 1420-1431.

932 Eshagh, M., 2018. Elastic thickness determination based on Vening Meinesz-Moritz and
933 flexural theories of isostasy, *Geophys. J. Int.*, **213**, 1682-1692.

934 Forsyth, D.W., 1985. Subsurface loading and estimates of the flexural rigidity of continental
935 lithosphere, *J. geophys. Res.*, **90**(B14), 12,623-12,632.

936 Forte, A.M., Peltier, W.R., Dziewonski, A.M. & Woodward, R.L., 1993. Dynamic surface
937 topography: a new interpretation based upon mantle flow models derived from seismic
938 tomography, *Geophys. Res. Lett.*, **20**, 225-228.

939 Harmon, N., Forsyth, D.W. & Scheirer, D.S., 2006. Analysis of gravity and topography in the
940 GLIMPSE study region: isostatic compensation and uplift of the Sojourn and Hotu Matua
941 Ridge systems, *J. geophys. Res.*, **111**, B11406, doi:10.1029/2005JB004071.

942 Hayford, J.F., 1909. *The Figure of the Earth and Isostasy from Measurements in the United*
943 *States*, pp 178, Government Printing Office.

944 Heiskanen, W.A., 1931. Isostatic tables for the reduction of gravimetric observations
945 calculated on the basis of Airy's hypothesis, *Bull. Géod.*, **30**, 110-153.

946 Heiskanen, W.A. & H. Moritz, 1967. *Physical Geodesy*, W.H. Freeman.

947 Hirt, C. & Rexer, M., 2015. Earth2014: 1 arc-min shape, topography, bedrock and ice-sheet
948 models – available as gridded data and degree-10,800 spherical harmonics, *Int. J. Appl.*
949 *Earth Observation Geoinformation*, **39**, 103-112.

950 Hirt, C., Kuhn, M., Featherstone, W.E. & Göttl, F., 2012. Topographic/isostatic evaluation of
951 new-generation GOCE gravity field models, *J. geophys. Res.*, **117**, B05407, doi:
952 10.1029/2011JB008878.

953 Ito, G. & Taira, A., 2000. Compensation of the Ontong Java Plateau by surface and
954 subsurface loading, *J. geophys. Res.*, **105**(B5), 11,171-11,183.

955 Jordan, T.A. & Watts, A.B., 2005. Gravity anomalies, flexure and the elastic thickness
956 structure of the India-Eurasia collisional system, *Earth planet. Sci. Lett.*, **236**, 732-750.

957 Kaban, M.K., Schwintzer, P. & Tikhotsky, S.A., 1999. A global isostatic gravity model of the
958 Earth, *Geophys. J. Int.*, **136**, 519-536.

959 Kaban, M.K., Schwintzer, P. & Reigber, C., 2004. A new isostatic model of the lithosphere
960 and gravity field, *J. Geod.*, **78**, 368-385.

961 Karner, G.D. & Watts, A.B., 1982. On isostasy at Atlantic-type continental margins, *J.*
962 *geophys. Res.*, **87**(B4), 2923-2948.

963 Karner, G.D. & Watts, A.B., 1983. Gravity anomalies and flexure of the lithosphere at
964 mountain ranges, *J. geophys. Res.*, **88**(B12), 10,449-10,477.

965 Kaula, W.M., 1967. Geophysical implications of satellite determinations of the Earth's
966 gravitational field, *Space Sci. Rev.*, **7**, 769-794.

967 Kirby, J.F., 2014. Estimation of the effective elastic thickness of the lithosphere using inverse
968 spectral methods: the state of the art, *Tectonophysics*, **631**, 87-116.

969 Kirby, J.F. & Swain, C.J., 2008. An accuracy assessment of the fan wavelet coherence
970 method for elastic thickness estimation, *Geochem. Geophys. Geosyst.*, **9**, Q03022,
971 doi:10.1029/2007GC001773.

972 Kirby, J.F. & Swain, C.J., 2009. A reassessment of spectral T_e estimation in continental
973 interiors: the case of North America, *J. geophys. Res.*, **114**(B8), B08401,
974 doi:10.1029/2009JB006356.

975 Kirby, J.F. & Swain, C.J., 2013. Power spectral estimates using two-dimensional Morlet-fan
976 wavelets with emphasis on the long wavelengths: jackknife errors, bandwidth resolution and
977 orthogonality properties, *Geophys. J. Int.*, **194**, 78-99.

978 Komut, T., Gray, R., Pysklywec, R. & Göğüş, O.H., 2012. Mantle flow uplift of western
979 Anatolia and the Aegean: interpretations from geophysical analyses and geodynamic
980 modelling, *J. geophys. Res.*, **117**, B11412, doi:10.1029/2012JB009306.

981 Kuhn, M., 2003. Geoid determination with density hypotheses from isostatic models and
982 geological information, *J. Geod.*, **77**, 50-65.

983 Lachapelle, G., 1976. A spherical harmonic expansion of the isostatic reduction potential,
984 *Bollettino di Geodesia e Scienze Affini*, **3**, 281-299.

985 Laske, G., Masters, G., Ma, Z. & Pasyanos, M., 2013. Update on CRUST1.0 – a 1-degree
986 global model of Earth's crust, in *Geophysical Research Abstracts. Vol. 15*, Eur. Geosci.
987 Un., Munich, p. EGU2013-2658.

988 Lewis, B.T.R. & Dorman, L.M., 1970. Experimental isostasy, 2: An isostatic model for the
989 USA derived from gravity and topography data, *J. geophys. Res.*, **75**, 3367-3386.

990 Louden, K.E. & Forsyth, D.W., 1982. Crustal structure and isostatic compensation near the
991 Kane fracture zone from topography and gravity measurements: 1. Spectral analysis
992 approach, *Geophys. J. R. astr. Soc.*, **68**, 725-750.

993 Lowry, A.R., Ribe, N.M. & Smith, R.B., 2000. Dynamic elevation of the Cordillera, western
994 United States, *J. geophys. Res.*, **105**(B10), 23,371-23,390.

995 Macario, A., Malinverno, A. & Haxby, W.F., 1995. On the robustness of elastic thickness
996 estimates obtained using the coherence method, *J. geophys. Res.*, **100**(B8), 15,163-15,172.

997 Martinec, Z., 1993. A model of compensation of topographic masses, *Surv. Geophys.*, **14**,
998 525-535.

999 Martinec, Z., 1994a. The density contrast at the Mohorovičić discontinuity, *Geophys. J. Int.*,
1000 **117**, 539-544.

1001 Martinec, Z., 1994b. The minimum depth of compensation of topographic masses, *Geophys.*
1002 *J. Int.*, **117**, 545-554.

1003 McNutt, M., 1980. Implications of regional gravity for state of stress in the Earth's crust and
1004 upper mantle, *J. geophys. Res.*, **85**(B11), 6377-6396.

1005 McNutt, M.K., 1983. Influence of plate subduction on isostatic compensation in northern
1006 California, *Tectonics*, **2**, 399-415.

1007 Molnar, P., England, P.C. & Jones, C.H., 2015. Mantle dynamics, isostasy, and the support of
1008 high terrain, *J. geophys. Res.*, **120**, 1932-1957.

1009 Moritz, H., 1990. The inverse Vening Meinesz problem in geodesy, *Geophys. J. Int.*, **102**,
1010 733-738.

1011 Parker, R.L., 1972. The rapid calculation of potential anomalies, *Geophys. J. R. astr. Soc.*, **31**,
1012 447-455.

1013 Pavlis, N.K. & Rapp, R.H., 1990. The development of an isostatic gravitational model to
1014 degree 360 and its use in global gravity modelling, *Geophys. J. Int.*, **100**, 369-378.

1015 Pavlis, N.K., Holmes, S.A., Kenyon, S.C. & Factor, J.K., 2012. The development and
1016 evaluation of the Earth Gravitational Model 2008 (EGM2008), *J. geophys. Res.*, **117**,
1017 B04406, doi:10.1029/2011JB008916.

1018 Pérez-Gussinyé, M. & Watts, A.B., 2005. The long-term strength of Europe and its
1019 implications for plate-forming processes, *Nature*, **436**, 381-384.

1020 Perry, H.K.C., Eaton, D.W.S. & Forte, A.M., 2002. LITH5.0: a revised crustal model for
1021 Canada based on Lithoprobe results, *Geophys. J. Int.*, **150**, 285-294.

1022 Pratt, J.H., 1855. On the attraction of the Himalaya Mountains, and of the elevated regions
1023 beyond them, upon the plumb-line in India, *Phil. Trans. R. Soc. Lond.*, **145**, 53-100.

1024 Rapp, R.H., 1982. Degree variances of the Earth's potential, topography and its isostatic
1025 compensation, *Bull Géod*, **56**, 84-94.

1026 Rapp, R.H., 1989. The decay of the spectrum of the gravitational potential and the
1027 topography for the Earth, *Geophys. J. Int.*, **99**, 449-455.

1028 Rummel, R., Rapp, R.H., Sünnkel, H. & Tscherning, C.C., 1988. *Comparisons of Global*
1029 *Topographic/Isostatic Models to the Earth's Observed Gravity Field*, Ohio State University
1030 *Report 388, Department of Geodetic Science and Surveying*, Ohio State University.

1031 Saupe, D., 1988. Algorithms for random fractals, in *The Science of Fractal Images*, pp. 71-
1032 136, eds Peitgen, H.-O. & Saupe, D., Springer.

1033 Simons, F.J., Zuber, M.T. & Korenaga, J., 2000. Isostatic response of the Australian
1034 lithosphere: estimation of effective elastic thickness and anisotropy using multitaper
1035 spectral analysis, *J. geophys. Res.*, **105**(B8), 19,163-19,184.

1036 Simpson, R.W., Jachens, R.C., Blakely, R.J. & Saltus, R.W., 1986. A new isostatic residual
1037 gravity map of the conterminous United States with a discussion on the significance of
1038 isostatic residual anomalies, *J. geophys. Res.*, **91**(B8), 8348-8372.

1039 Sjöberg, L.E., 2009. Solving Vening Meinesz-Moritz inverse problem in isostasy, *Geophys.*
1040 *J. Int.*, **179**, 1527-1536.

1041 Sünkel, H., 1985. *An Isostatic Earth Model, Ohio State University Report 367, Department of*
1042 *Geodetic Science and Surveying*, Ohio State University.

1043 Tiwari, V.M., Diament, M. & Singh, S.C., 2003. Analysis of satellite gravity and bathymetry
1044 data over Ninety-East Ridge: variation in the compensation mechanism and implication for
1045 emplacement process, *J. geophys. Res.*, **108**(B2), 2109, doi:10.1029/2000JB000047.

1046 Ussami, N., de Sá, N.C. & Molina, E.C., 1993. Gravity map of Brazil: 2. Regional and
1047 residual isostatic anomalies and their correlation with major tectonic provinces, *J. geophys.*
1048 *Res.*, **98**(B2), 2199-2208.

1049 Vening Meinesz, F.A., 1931. Une nouvelle methode pour la reduction isostatique regionale
1050 de l'intensite de la pesanteur, *Bull Géod.*, **29**, 33-51.

1051 Walcott, R.I., 1970. Isostatic response to loading of the crust in Canada, *Can. J. Earth Sci.*, **7**,
1052 716-727.

1053 Watts, A.B., 1978. An analysis of isostasy in the world's oceans: 1. Hawaiian-Emperor
1054 seamount chain, *J. geophys. Res.*, **83**(B12), 5989-6004.

1055 Watts, A.B., 2001. *Isostasy and Flexure of the Lithosphere*, pp. 458, Cambridge University
1056 Press.

1057 Watts, A.B., 2007. An overview, in *Treatise on Geophysics*, vol. **6**, pp. 1-48, ed Schubert, G.,
1058 Elsevier.

1059 Watts, A.B. & Moore, J.D.P., 2017. Flexural isostasy: constraints from gravity and
1060 topography power spectra, *J. geophys. Res.*, **122**, 8417-8430.

1061 Watts, A.B. & Talwani, M., 1974. Gravity anomalies seaward of deep-sea trenches and their
1062 tectonic implications, *Geophys. J. R. astr. Soc.*, **36**, 57-90.

1063 Watts, A.B., Lamb, S.H., Fairhead, J.D. & Dewey, J.F., 1995. Lithospheric flexure and
1064 bending of the central Andes, *Earth planet. Sci. Lett.*, **134**, 9-21.

1065 Wiczorek, M.A., 2007. Gravity and topography of the terrestrial planets, in *Treatise on*
1066 *Geophysics*, vol. **10**, pp. 165-206, ed Schubert, G., Elsevier.

1067 Wyer, P. & Watts, A.B., 2006. Gravity anomalies and segmentation at the East Coast, USA
1068 continental margin, *Geophys. J. Int.*, **166**, 1015-1038.

1069 Zuber, M.T., Bechtel, T.D. & Forsyth, D.W., 1989. Effective elastic thickness of the
1070 lithosphere and the mechanisms of isostatic compensation in Australia, *J. geophys. Res.*,
1071 **94**(B7), 9353-9367.

1072

1073 Table 1. Values of constants and parameters used in this study. Throughout the article the

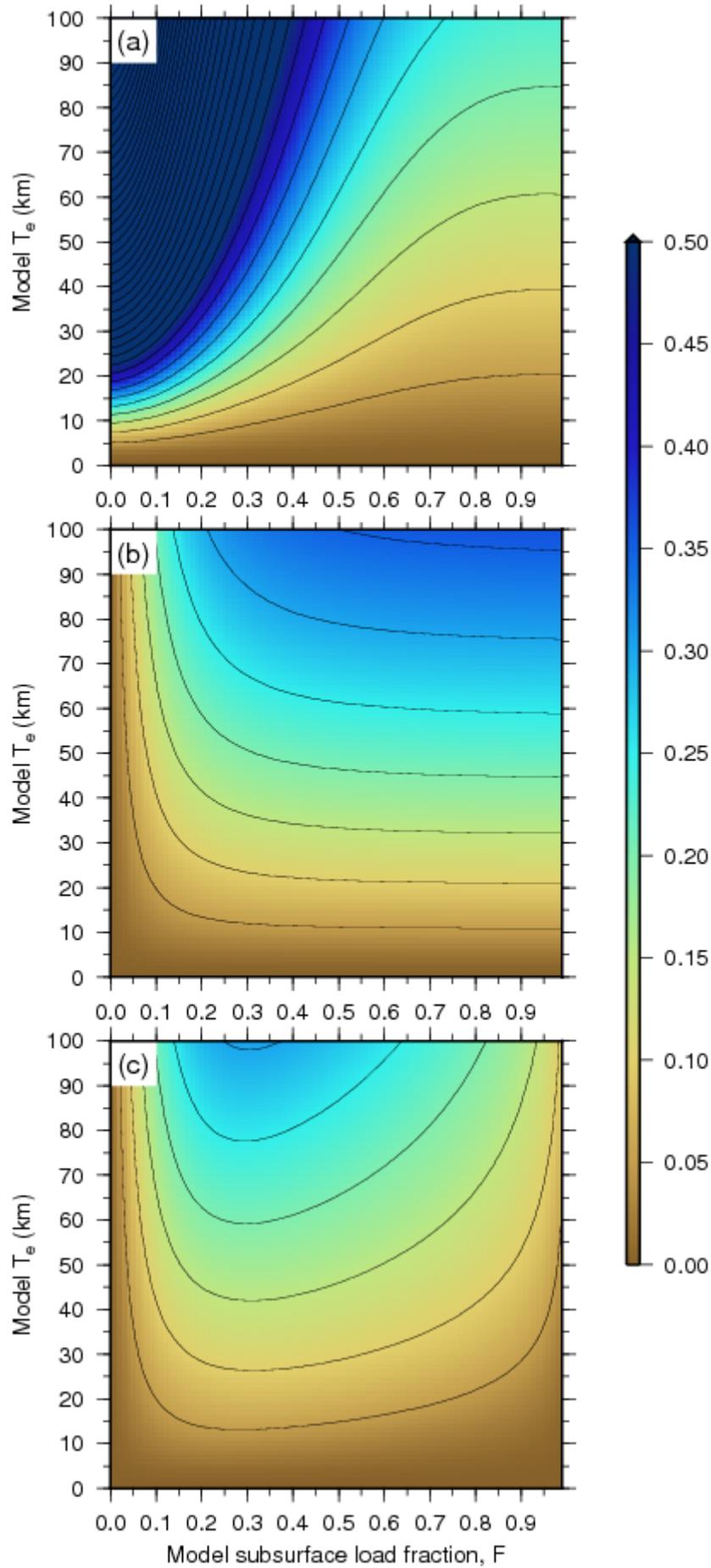
1074 following density contrasts are used: $\Delta\rho_0 = \rho_c - \rho_f$ and $\Delta\rho_1 = \rho_m - \rho_c$.

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Constant	Symbol	Value
Gravitational constant	\mathcal{G}	$6.67259 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$
Young's modulus	E	100 GPa
Poisson's ratio	ν	0.25
Gravity acceleration	g	9.79 ms^{-2}
Fluid density (air or sea water)	ρ_f	0 or 1030 kg m^{-3}
Crust density	ρ_c	2800 kg m^{-3}
Mantle density	ρ_m	3300 kg m^{-3}

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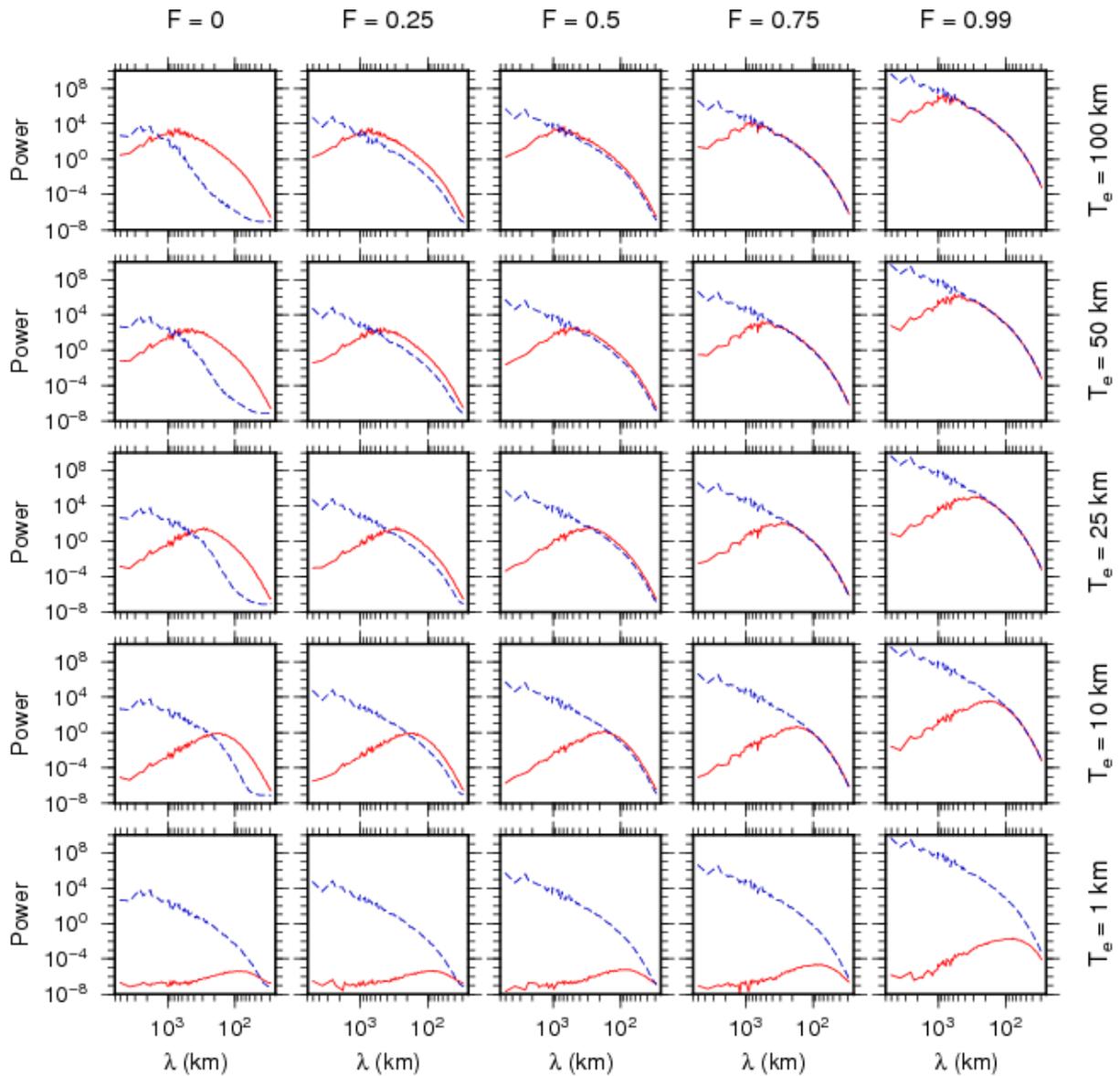
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1080 Figure 1. The ratio of the isostatic anomaly standard deviation to that of the corresponding
1081 Bouguer anomaly ($\bar{\sigma}_{IA}$), from tests A, B and C. (a) Test A: synthetic models generated with
1082 the T_e and F values shown on the axes, and a compensation depth of $z_m = 35$ km;
1083 compensation attraction computed by incorrectly assuming Airy isostasy ($T_e = F = 0$), but the
1084 correct compensation depth. (b) Test B: synthetic models generated with the T_e and F values
1085 shown on the axes, and a compensation depth of $z_m = 35$ km; compensation attraction
1086 computed by incorrectly assuming surface loading ($F = 0$), but with the correct T_e and
1087 compensation depth. (c) Test C: synthetic models generated with the T_e and F values shown
1088 on the axes, and a compensation depth of $z_m = 35$ km; compensation attraction computed by
1089 correctly estimating T_e , F and compensation depth. Contour interval in all images is 0.05, and
1090 the maximum values of $\bar{\sigma}_{IA}$ in each of the panels are: (a) 2.14, (b) 0.36, (c) 0.30.

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1094 Figure 2(a). Power spectra (versus wavelength λ) from some of the models in test A:

1095 synthetic models generated with the T_e and F values shown on the axes, and a compensation

1096 depth of $z_m = 35$ km; compensation attractions computed by incorrectly assuming Airy

1097 isostasy ($T_e = F = 0$), but the correct compensation depth. Bouguer anomaly (blue dashed);

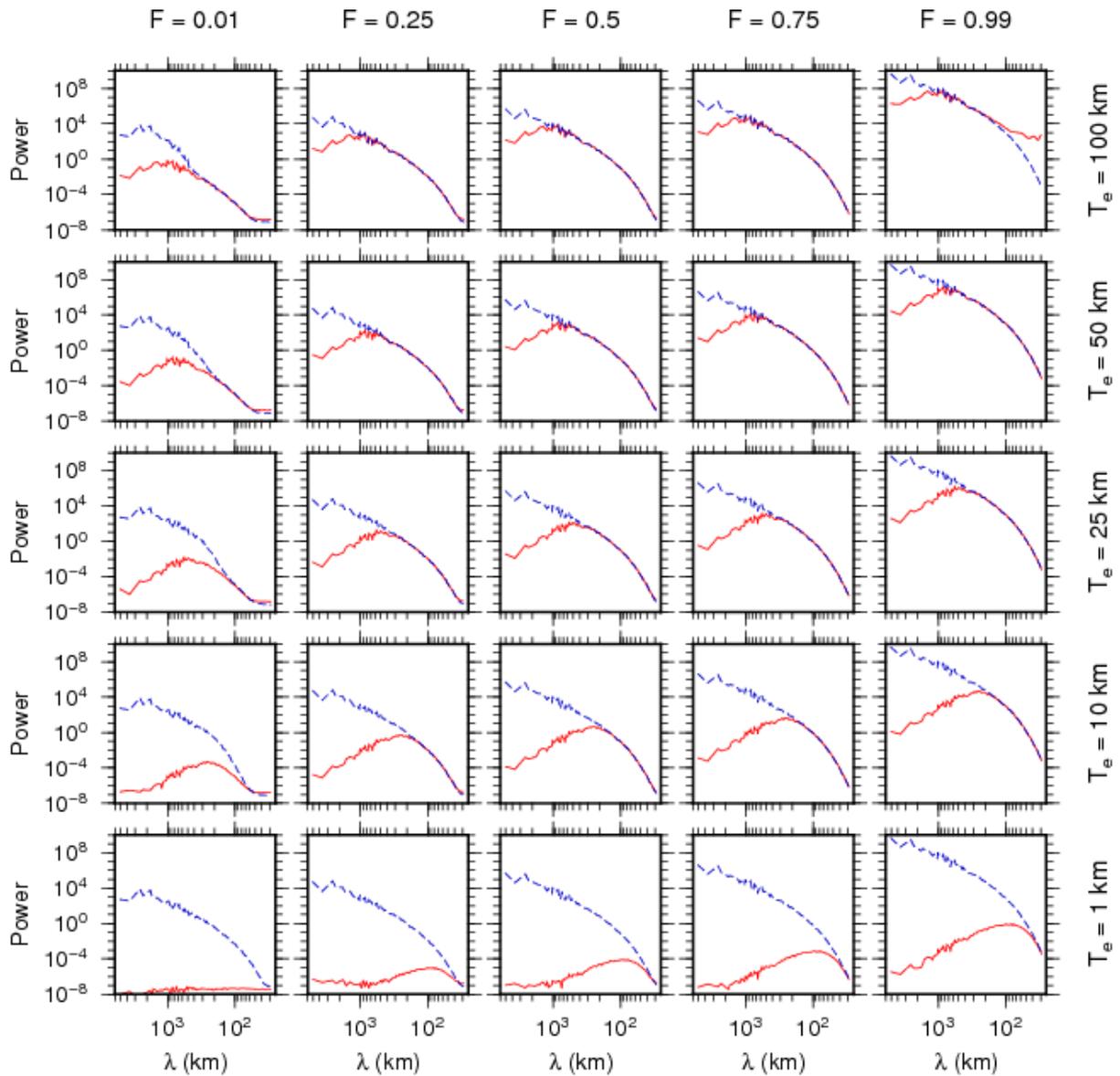
1098 isostatic anomaly (red). The rows and columns correspond to the synthetic model T_e and F

1099 values in Fig. 1(a). The power spectra in this study were estimated by first computing 2-D

1100 power spectra using Slepian multitapers ($K = 1$ taper) of bandwidth parameter $NW = 1$ (e.g.

1101 Simons *et al.* 2000). Then, the 2-D power spectra were azimuthally averaged over annuli in

1102 the wavenumber domain for display as 1-D profiles. The multitaper parameters $K = 1$ and
1103 $NW = 1$ were chosen to maximise the wavenumber-domain resolution of the spectra (e.g.
1104 Kirby and Swain 2013). For these noise-free, synthetic data such low values of K and NW are
1105 acceptable, though for real data more tapers (higher values of K , and therefore also NW) may
1106 be preferable in order to improve the estimation variance (e.g. Simons *et al.* 2000).
1107

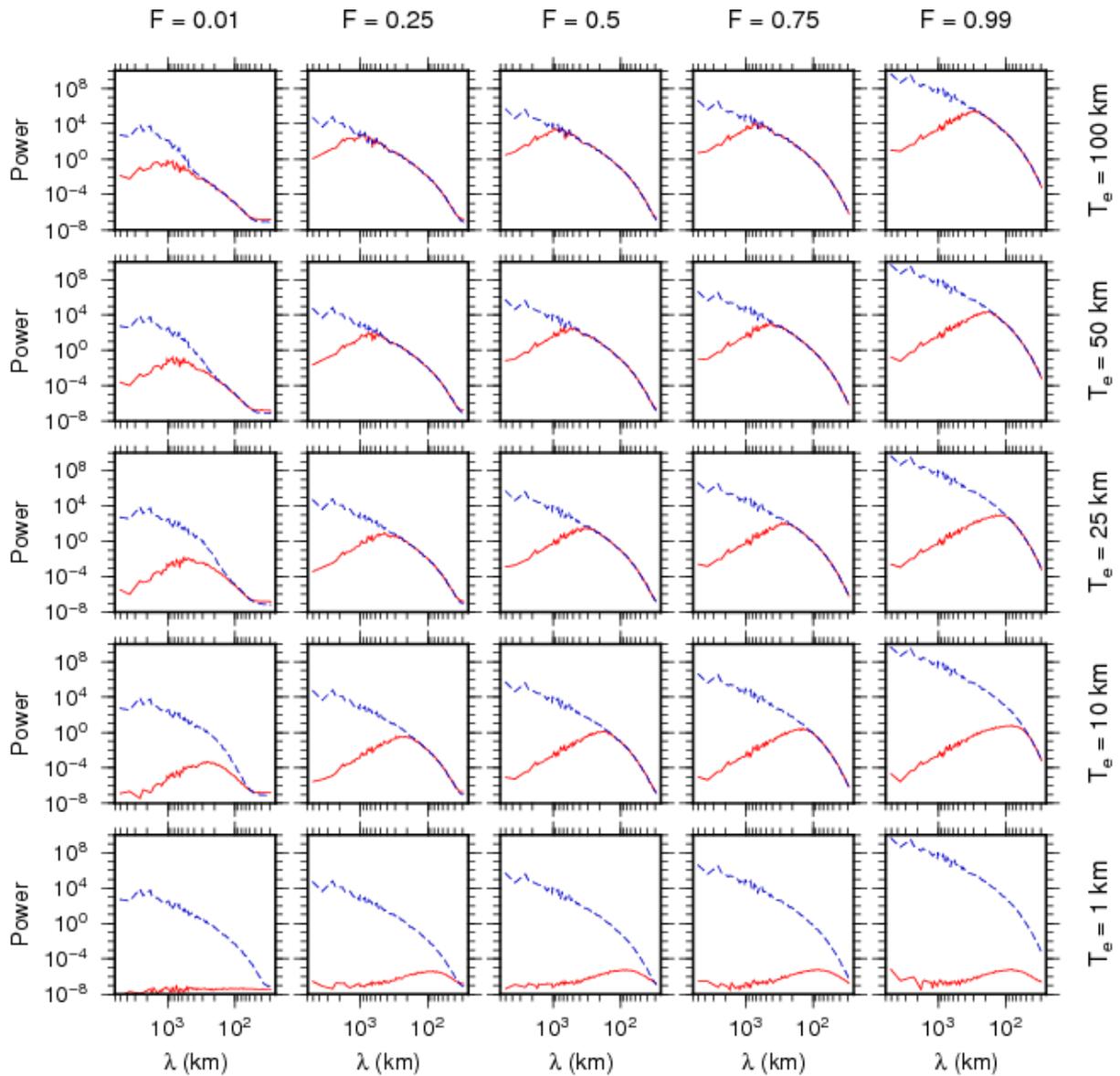


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1110 Figure 2(b). Power spectra (versus wavelength λ) from some of the models in test B:
 1111 synthetic models generated with the T_e and F values shown on the axes, and a compensation
 1112 depth of $z_m = 35$ km; compensation attractions computed by incorrectly assuming surface
 1113 loading ($F = 0$), but with the correct T_e and compensation depth. Bouguer anomaly (blue
 1114 dashed); isostatic anomaly (red). The rows and columns correspond to the synthetic model T_e
 1115 and F values in Fig. 1(b).

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1119 Figure 2(c). Power spectra (versus wavelength λ) from some of the models in test C:

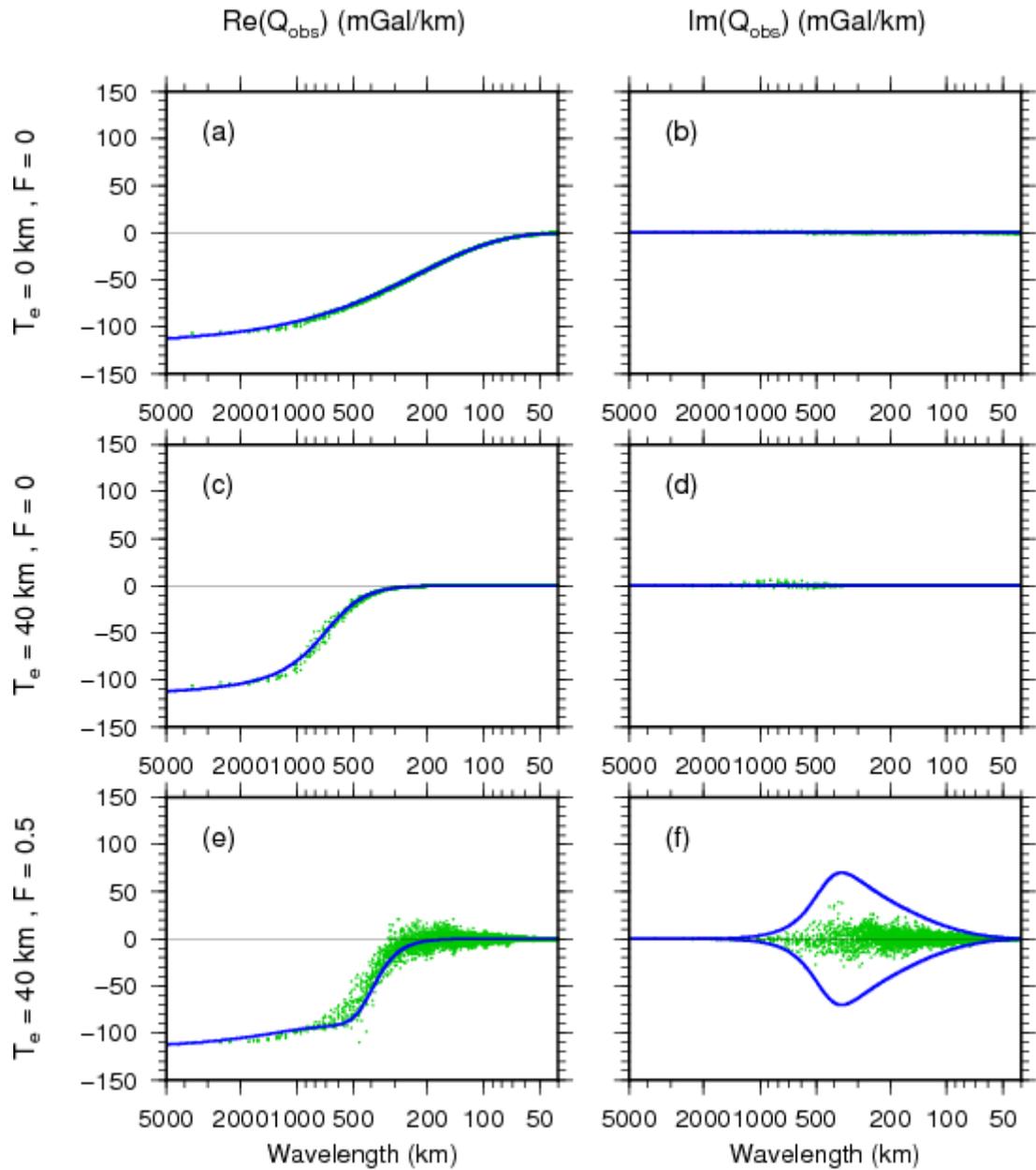
1120 synthetic models generated with the T_e and F values shown on the axes, and a compensation

1121 depth of $z_m = 35$ km; compensation attractions computed by correctly estimating T_e , F and

1122 compensation depth. Bouguer anomaly (blue dashed); isostatic anomaly (red). The rows and

1123 columns correspond to the synthetic model T_e and F values in Fig. 1(c).

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1127 Figure 3. Real and imaginary parts of the observed admittance of the synthetic model

1128 Bouguer anomaly and topography under Airy isostasy (a and b), surface-only loading (c and

1129 d), and combined loading (e and f). The T_e and F values at left are those from which the

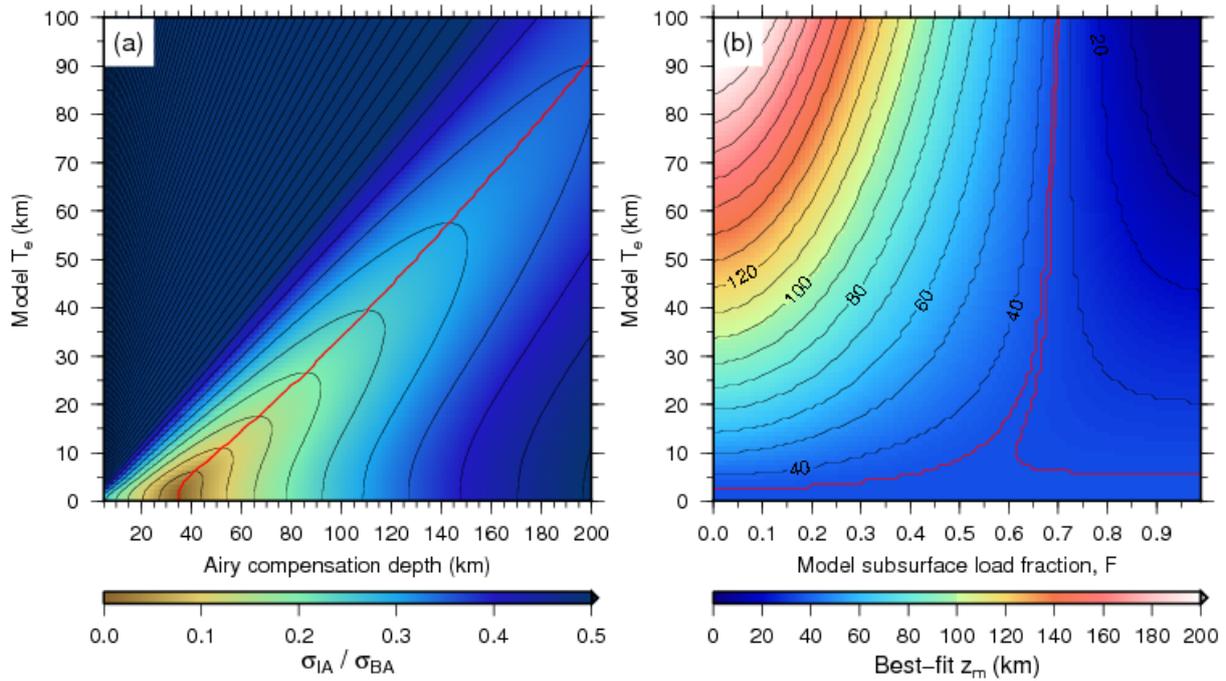
1130 models were generated. The green dots show the estimates of the 2-D real and imaginary

1131 observed admittance plotted as functions of their radial wavenumber $k = |\mathbf{k}|$ (but displayed as

1132 wavelength). The blue curves show the theoretical 1-D real and imaginary admittance from

1133 eq. (24) for each (T_e, F) value indicated at left. The theoretical imaginary admittance curve in

1134 (f) has been reflected about the zero-admittance axis. The auto- and cross-spectra in the
1135 observed admittance were estimated using Slepian multitapers ($K = 3$, $NW = 3$). It was found
1136 that using $K = 1$ taper [as used to compute Fig. 2(a), for example; see Fig. 2 caption]
1137 produced admittance spectra that did not match the theoretical predictions; using 2 or 3, or an
1138 even higher number of tapers for any value of $NW > 2$ resulted in more faithful admittances.
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1142 Figure 4. (a) The ratio of the isostatic anomaly standard deviation to that of the corresponding

1143 Bouguer anomaly ($\bar{\sigma}_{IA}$). Synthetic models were generated using the T_e values shown on the

1144 ordinate, surface-only loading ($F = 0$), and a compensation depth of $z_m = 35$ km;

1145 compensation attraction computed by incorrectly assuming Airy isostasy ($T_e = F = 0$), and

1146 with the compensation depth indicated on the abscissa. The red line in (a) shows the locus of

1147 the minimum standard deviation. Contour interval is 0.05, and the maximum value of $\bar{\sigma}_{IA}$ is

1148 2.26. (b) The value of the best-fitting depth to Moho (z_m , in km) that minimizes the standard

1149 deviation of the isostatic anomalies. The synthetic models were generated with the T_e and F

1150 values shown on the axes (and a compensation depth of 35 km); the compensation (and

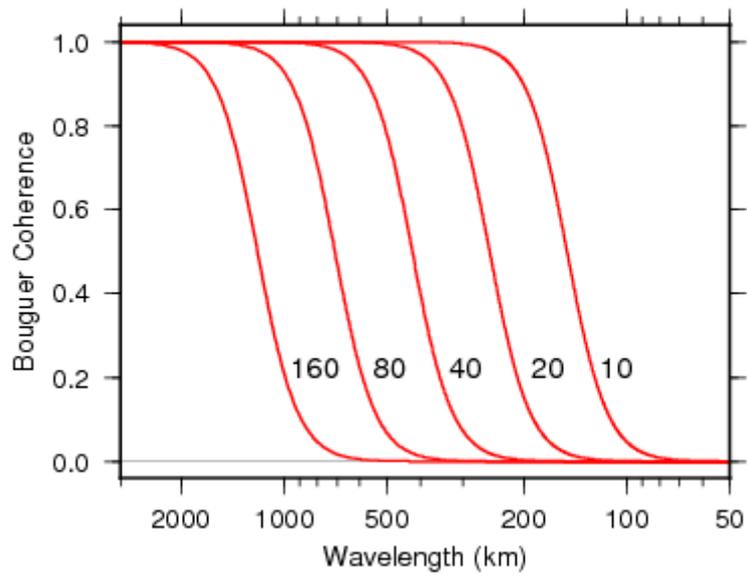
1151 therefore isostatic) anomalies were computed by incorrectly assuming Airy isostasy ($T_e = F =$

1152 0) over a broad range of assumed compensation depths (indicated on the abscissa in Fig. 4a).

1153 The two red contours are 34.5 and 35.5 km, and so mark the regions where the model

1154 compensation (Moho) depth (35 km) is recovered almost exactly. Fig. 4(a) hence represents

1155 the results from a subset of Fig. 4(b), when $F = 0$.



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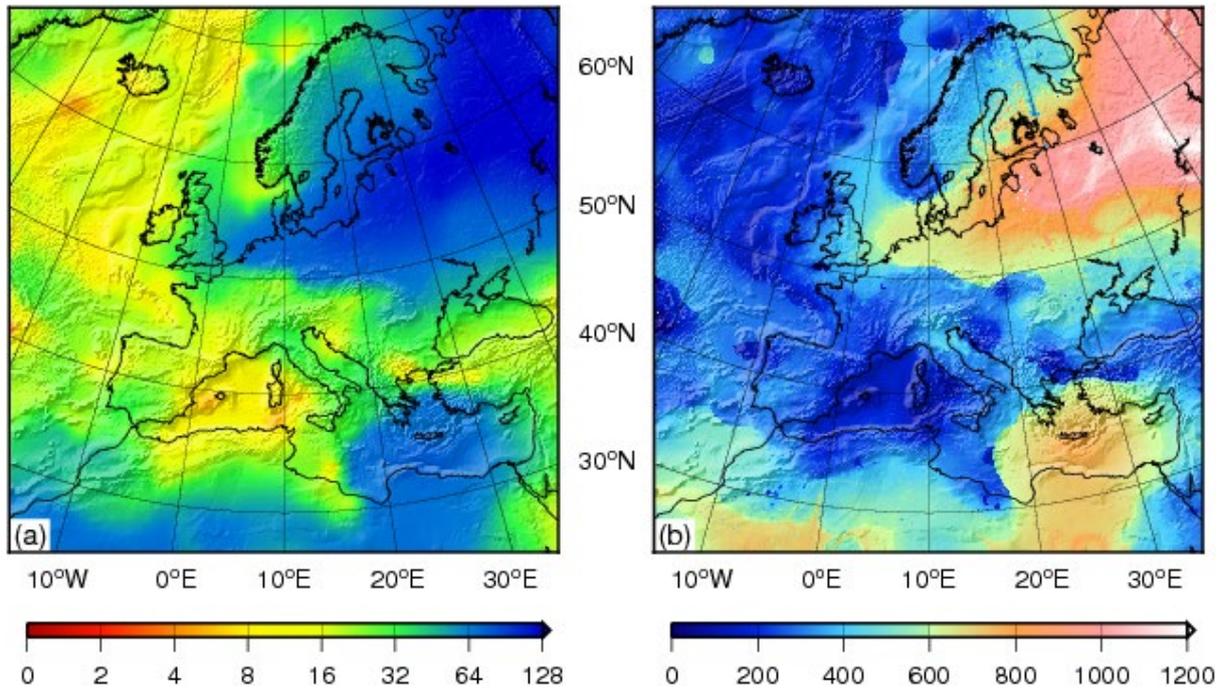
1158 Figure 5. Theoretical curves of the Bouguer coherence for five different T_e values (indicated,

1159 in km). In all cases the initial loading ratio, $f=1$. Note how lithosphere with a high T_e has a

1160 long-wavelength Bouguer coherence rollover (from 1 to 0), and as T_e decreases the rollover

1161 migrates to shorter wavelengths.

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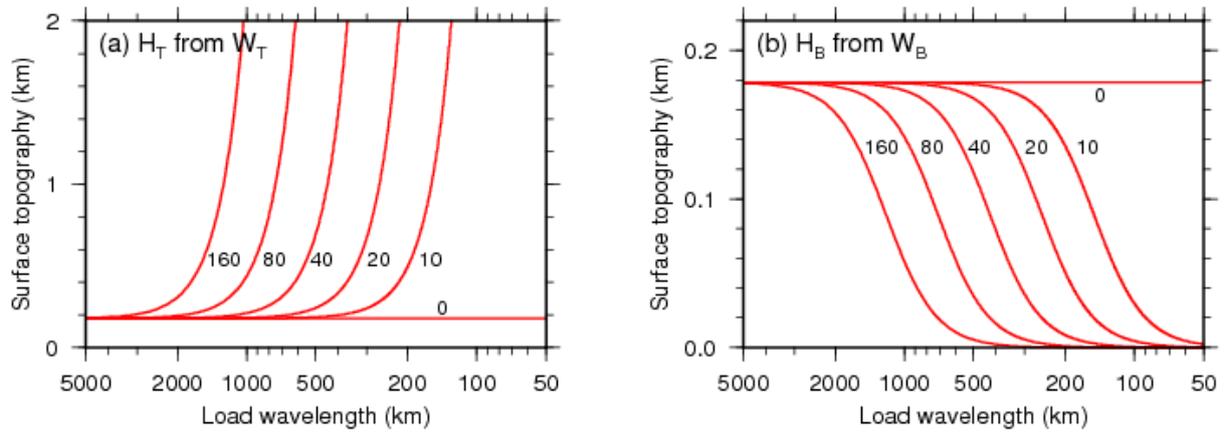
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1165 Figure 6. (a) T_e (km), and (b) predicted Bouguer coherence transition wavelength (km) over

1166 Europe, with topography shaded relief superimposed. Lambert conic conformal projection.

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1170 Figure 7. Surface topographic amplitude as a function of load wavelength, corresponding to a

1171 Moho topographic amplitude of 1 km, for six different T_e values (indicated, in km). (a)

1172 Surface loading, from eq. (31). (b) Subsurface loading, from eq. (33).

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