

School of Earth and Planetary Science

On the Genesis and Significance of the  
Archean-hosted Carlow Castle Cu-Co-Au  
Deposit, Pilbara Craton, Western Australia

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This thesis is presented for the Degree of  
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## Declaration

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## Works published during the course of this thesis

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## Research

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|---|---|-------------|-------------|
| <b>Project Title:</b>                     | On the Genesis and Significance of the Archean-hosted Carlow Castle Cu-Co-Au Deposit, Pilbara Craton, Western Australia |             |             |
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| 2     | <b>Fox, D.</b> , Spinks, S., Barham, M., Kirkland, C.L., Pearce, M.A., Aspandiar, M., Birchall, R., and Mead, E. (2021) Working up an apatite: Enigmatic Mesoarchean hydrothermal Cu-Co-Au mineralization in the Pilbara Craton. <i>Economic Geology</i> 116, 1561-1573. <b>Published.</b> | Journal article                                      |
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17

## 18 Abstract

19 Significant hydrothermal Cu-Co-Au deposits are generally restricted to Proterozoic or younger  
20 terranes, occurring overwhelmingly in sediment-hosted Cu-Co deposits. As such, the 2018  
21 discovery of significant Cu-Co-Au mineralisation at *Carlow Castle* within a Paleo-  
22 Mesoarchean greenstone belt in the Pilbara Craton of northwest Western Australia was notable.  
23 However, little study had been conducted on the deposit previously. This thesis provides the  
24 first mineralogical, geochronological, and isotopic constraints on the formation of the Carlow  
25 Castle Cu-Co-Au deposit.

26 Petrographic analysis demonstrates that Cu-Co-Au mineralisation from Carlow Castle occurs  
27 within quartz-carbonate veins through heavily hydrothermally altered wall-rocks within the  
28 Regal Thrust; a regionally extensive crustal-scale thrust fault that likely acted as a critical  
29 conduit for mineralising fluids. Mineralisation is defined by two distinct ore mineral  
30 assemblages. Ore mineral Assemblage One is defined primarily by pyrite and chalcopyrite,  
31 whilst Assemblage Two is defined by the presence of cobaltite, chalcocite, and electrum.  
32 Additional minor and trace phases include pyrrhotite, uraninite, and several telluride and  
33 selenide minerals. Mineral paragenetic relationships appear to reflect an increasing sulphur  
34 fugacity defined by a transition from selenides and tellurides through to sulphosalts and  
35 subsequently base-metal sulphides. Additionally, electrum appears to be paragenetically  
36 associated with cobaltite, suggesting Au was transported simultaneously with Co and Cu; most  
37 likely via chloride complexation.

38 In-situ U-Pb dating of syn-mineralisation hydrothermal apatite via laser ablation ICP-MS dates  
39 ore formation to  $2957 \pm 67$  Ma. This age makes Carlow Castle the oldest known hydrothermal  
40 Cu-Co deposit on Earth, coincident with the late stages of the formation of the proximal De  
41 Grey Superbasin (3066-2919 Ma) during post-orogenic extension and back-arc rifting related  
42 to subduction at the northwest margin of the Pilbara Craton. The hydrothermal alteration  
43 mineral assemblage of Carlow Castle is dominated by calc-silicate minerals which collectively  
44 define a classic propylitic alteration assemblage. This propylitic alteration assemblage is  
45 characteristic of ore-formation at peak temperatures of  $\sim 300^\circ\text{C}$  and alkaline fluid conditions.  
46 These constraints demonstrate continuity of Cu-Co mineralisation at moderate hydrothermal  
47 conditions well into the Archean and may suggest that Archean greenstone belts globally may  
48 be prospective for analogous deposits.

49 Multiple sulphur isotope analysis of sulphide mineralisation from Carlow Castle reveals a  
50 conspicuous lack of ‘Archean sulphur’ within the mineral system, with  $\Delta^{33}\text{S}$  values within a  
51 very narrow range between -0.17‰ and 0.12‰. Additionally,  $\delta^{34}\text{S}$  values range between -  
52 3.4‰ and 6.1‰; with the majority of these data ranging between 2‰ and 4‰. This is  
53 consistent with an isotopically heavy seawater sulphate contribution to the ore system. The  
54 unique absence of mass-independent sulphur isotope fractionation within the system is  
55 explained through a connection to contemporaneous oxidative conditions during deposition of  
56 glacial diamictites within the Pongola Basin of South Africa around 2950 Ma with similarly  
57 attenuated  $\Delta^{33}\text{S}$  signatures. Thermodynamic modelling of hydrothermal Cu and Co solubility,  
58 indicates significant metal mobilisation necessitates sulphate-stable oxidised conditions. These  
59 findings support emerging models of the evolution of the Archean atmosphere, which suggest  
60 that it was subject to periods of transient oxygenation following the evolution of oxygenic  
61 photosynthesis.

62 This thesis presents the first application of Cu isotope analysis to an Archean Cu ore deposit.  
63 Analysis of 19 Archean primary Cu sulphide ore samples and 11 secondary supergene Cu ore  
64 samples reveal that primary ore samples display light isotopic fractionation ( $\delta^{65}\text{Cu} = -0.80$  to  
65  $0.00\text{‰}$ ) and supergene Cu ore samples are comparably isotopically heavy ( $\delta^{65}\text{Cu} = -0.50$  to  
66  $0.62\text{‰}$ ). Across primary Cu ore samples a Rayleigh fractionation process is recorded and  
67 modelled, defined by evolution of the hydrothermal system toward isotopically heavy and Cu-  
68 poor compositions. This is coeval with a decline in alteration temperature. This modelling also  
69 supports a mafic igneous source of ore metals, consistent with other significant hydrothermal  
70 Co-rich deposits globally. Relatively limited heavy isotopic fractionation within supergene Cu  
71 ore samples is explained through in-situ oxidative weathering of vein-hosted Cu sulphides,  
72 with minimal Cu transport. These findings demonstrate that Cu isotope analysis may be a  
73 valuable tool to understand ore genesis in Archean hydrothermal Cu deposits.

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400

# 401 Chapter 1

## 402 Introduction

403 The utilisation of Earth's mineral resources has been vital to the industrial development of  
404 civilisation. However, as technology has evolved, the diversity of mineral resources utilised by  
405 the global economy has grown substantially and the relative importance of individual mineral  
406 commodities has varied significantly Schulz, DeYoung Jr, Seal Ii, and Bradley (2017). As the  
407 global economy evolves and the transition from fossil fuel to renewable energy sources  
408 continues, many mineral commodities, of previously niche importance, are forecast to become  
409 critically important (Arrobas, Hund, McCormick, Ningthoujam, & Drexhage, 2017; Schulz et  
410 al., 2017). Because of the variable nature of many renewable energy sources, a necessity for  
411 battery storage has been identified in order to enable this renewable energy transition (Leonard,  
412 Michaelides, & Michaelides, 2020). As such, several 'battery metals'; such as Co, Li, Ni, and  
413 Cu have emerged as potentially critical to the future economy due to their primary importance  
414 in battery production (Mudd et al., 2018). A significant increase in global demand for these  
415 battery metals is widely projected, necessitating the discovery of new sources to supply these  
416 metals (Arrobas et al., 2017; Mudd et al., 2018). In the case of cobalt, the importance of  
417 discovering new deposits is compounded by widely publicised supply risks (Horn et al., 2021;  
418 Slack, Kimball, & Shedd, 2017a). Exploration efforts for these ore deposits can be optimised  
419 through understanding the full range of geological processes that produce them (McCuaig,  
420 Hronsky, Kelley, & Golden, 2014). The development of more informed exploration strategies  
421 and novel exploration techniques for these ore deposits is especially important given the  
422 declining discovery rates for new deposits of all varieties (Groves & Santosh, 2015). To  
423 understand these ore-forming processes, this thesis presents a multifaceted analysis of *Carlow*  
424 *Castle*; a distinctive Cu-Co-Au deposit in the Pilbara Craton in north-western Western  
425 Australia.

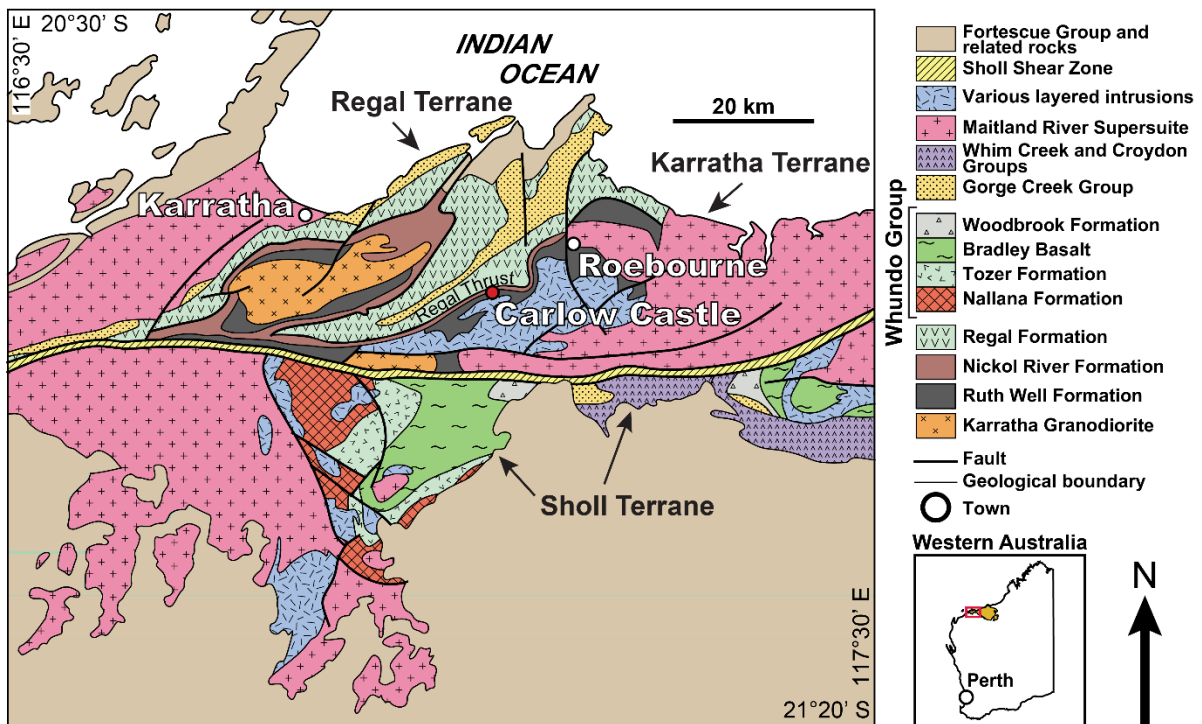
426 Carlow Castle Cu-Co-Au was chosen as the subject of this study as it represents a unique style  
427 of hydrothermal vein-hosted Cu-Co-Au mineralisation within a Paleo-Mesoarchean greenstone  
428 belt, which had not previously been subject to any detailed study. As the vast majority of  
429 significant Cu-Co deposits occur in terranes of Proterozoic age or younger (Petavratzi et al.,  
430 2019; Slack et al., 2017a), Carlow Castle presents an opportunity to understand the ore-forming

431 processes required to produce significant hydrothermal Cu-Co mineralisation in Archean  
432 greenstone belts. Understanding these ore-forming processes could be important in more fully  
433 developing the understanding of Cu-Co deposit genesis, particularly with regard to the  
434 geological setting and temporal distribution of hydrothermal Cu-Co deposits. This is  
435 particularly relevant to broad-scale prospectivity analysis for Cu-Co mineralisation, where  
436 future exploration for Cu-Co mineralisation may need to look toward unconventional settings  
437 to discover the next generation of deposits. This thesis provides the first analysis of Carlow  
438 Castle Cu-Co-Au deposit's mineralisation and the first insights into the key ore-forming  
439 processes responsible for its unique style of mineralisation.

## 440 1.1 Background

### 441 1.1.1 Carlow Castle Cu-Co-Au deposit

442 Carlow Castle copper-cobalt-gold (Cu-Co-Au) deposit is situated in the Pilbara Craton of  
443 Western Australia and has an inferred resource estimate of 14.3 million tonnes (Mt) at 0.4%  
444 Cu, 0.05% Co, and 0.7 g/t Au as of May 2021, updated from 8 Mt at 0.6% Cu, 0.08% Co, and  
445 1.6 g/t Au as of 2019 (Artemis Resources Limited, 2021). Carlow Castle was first discovered  
446 in 1872 and sporadically mined on a small scale throughout the late 19<sup>th</sup> and early 20<sup>th</sup> centuries  
447 (Hickman, 1983). However, significant underlying Cu-Co-Au mineralization was discovered  
448 during drilling conducted in 2017 (Artemis Resources Limited, 2017). Given the recency of  
449 this discovery, no substantive research had been conducted on the nature and origin of Carlow  
450 Castle. The only existing references to Carlow Castle in the literature prior to this project were  
451 passing mentions in reviews and other literature focussed on the Geology of the Pilbara Craton  
452 more broadly (Hickman, 2016; Hickman, Huston, Van Kranendonk, & Smithies, 2006;  
453 Ruddock, 1999b). Carlow Castle occurs in a series of steeply dipping veins through the Ruth  
454 Well and Nickol River Formations (Figure 1.1), a sequence of Paleoproterozoic mafic volcanic  
455 and marine sedimentary rocks, proximal to the Regal Thrust; a crustal-scale orogenic thrust  
456 fault (Hickman, 1983; Ruddock, 1999b). As such, it had been suggested that Carlow Castle  
457 may have been an orogenic Cu-Co-enriched Au deposit (Hickman, 2016). However, whilst  
458 Carlow Castle does occur through an orogenic structure (Regal Thrust); providing a maximum  
459 mineralisation age, its temporal relationship to orogenesis was largely unknown due to the lack  
460 of absolute age constraints. As such, the ore-forming tectonic event that produced Carlow  
461 Castle was largely unconstrained.



462  
 463 Figure 1.1 - Geologic map of the northwest Pilbara Craton. The Pilbara Craton is  
 464 highlighted in orange within the inset map. Modified after Van Kranendonk, Hickman,  
 465 Smithies, Nelson, and Pike (2002).

466 1.1.2 Formation of the West Pilbara Superterrane

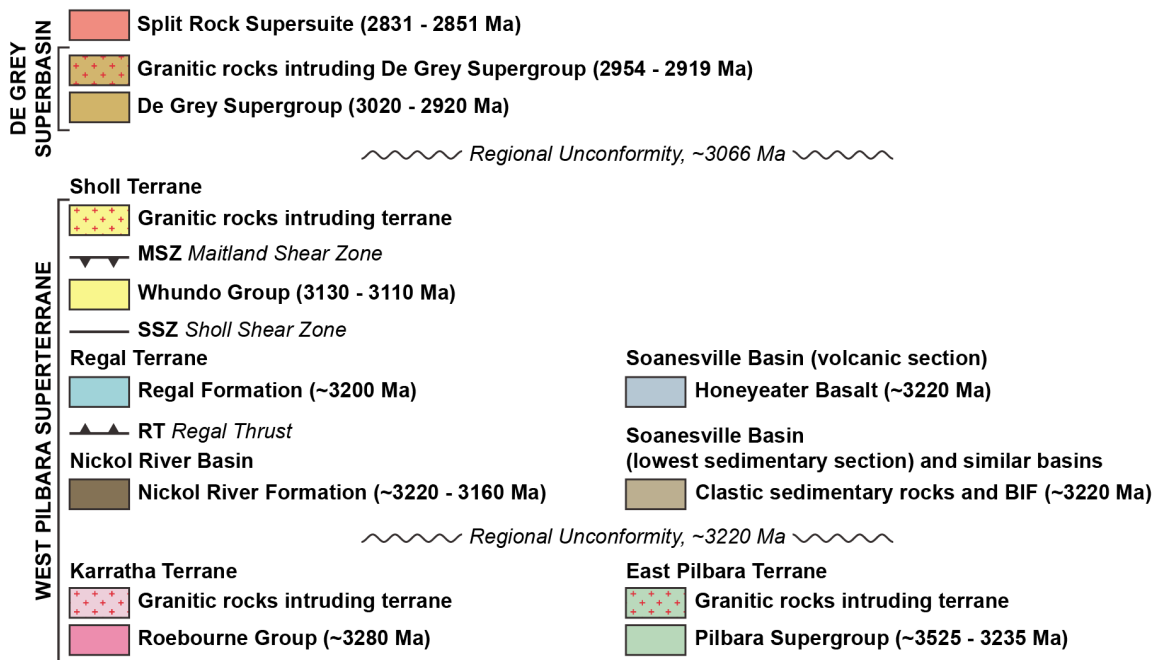
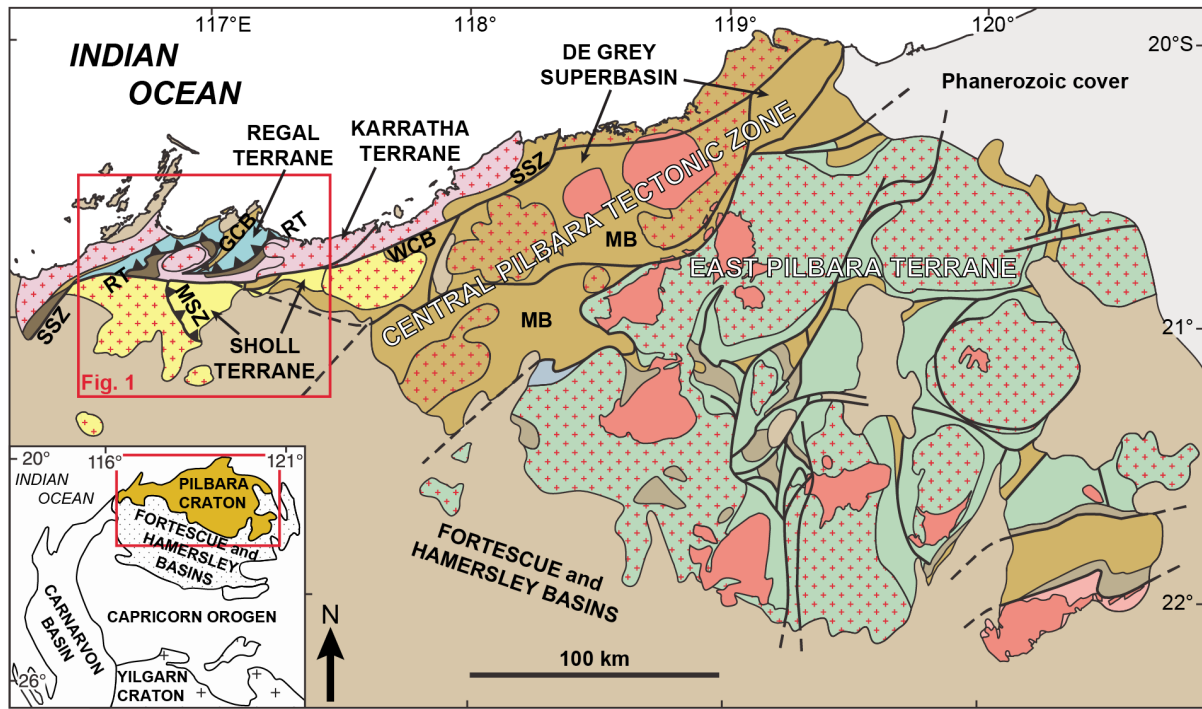
467 Carlow Castle occurs within the West Pilbara Superterrane, an accreted tectonic unit composed  
 468 of several Paleo-Mesoarchean terranes. The West Pilbara Superterrane is sub-divided into the  
 469 Karratha Terrane (3280-3236 Ma), Nickol River Basin (3220-3160 Ma), Regal Terrane (3200-  
 470 3160 Ma), Sholl Terrane (3130-3093 Ma), and portions of the Elizabeth Hill Supersuite (3068-  
 471 3066 Ma) (Hickman, 2016). Within this framework, Carlow Castle Cu-Co-Au deposit is  
 472 located primarily through the Karratha Terrane and Nickol River Basin. However, its  
 473 structurally-hosted nature within the Regal Thrust, which defines the boundary between the  
 474 Regal and Karratha Terranes, complicates this slightly. The oldest of these terranes; the  
 475 Karratha Terrane most likely represents a rifted portion of the early Paleoproterozoic Pilbara  
 476 Craton, which separated from the East Pilbara Terrane around 3280 Ma during a transition  
 477 from vertical tectonics to more modern-style horizontal tectonics (Hickman, 2021; Hickman et  
 478 al., 2006; Hickman & Van Kranendonk, 2012). During this period of rifting, the Nickol River  
 479 Basin formed on a passive margin at the edge of the rifted Karratha Terrane, hosting a  
 480 sedimentary succession composed predominantly of shale, marine carbonates, and banded iron  
 481 formations (Hickman, 2012; Kiyokawa, Taira, Byrne, Bowring, & Sano, 2002). As a result of

482 this period of extension, the Central Pilbara Basin developed as a rift basin between the  
483 Karratha Terrane and the East Pilbara Terrane; composed predominantly of basaltic oceanic  
484 crust now represented as the Regal Terrane within the West Pilbara Superterrane (Hickman,  
485 2004b; Hickman, 2016). A renewed period of convergence between the Karratha and East  
486 Pilbara Terranes, potentially caused by collision of the Karratha Terrane with an unidentified  
487 tectonic plate to the northwest of the Pilbara Craton, initiated from 3160 Ma causing the closure  
488 of the Central Pilbara Basin (Hickman, 2016). This renewed convergence also produced the  
489 Sholl Terrane as a volcanic oceanic island arc, during ocean-ocean convergence and  
490 development of a subduction zone within the Central Pilbara Basin from 3130-3093 Ma  
491 (Hickman & Van Kranendonk, 2012). Convergence continued until ~3070 Ma, when it  
492 culminated in the Prinsep Orogeny with the accretion of the Nickol River Basin and Karratha,  
493 Regal, and Sholl terranes into the West Pilbara Superterrane due to collision with the East  
494 Pilbara Terrane (Hickman, 2016; Hickman et al., 2006). During the Prinsep Orogeny, the Regal  
495 Terrane was thrust over the Karratha Terrane and Nickol River Basin along the Regal Thrust;  
496 the host structure of Carlow Castle (Van Kranendonk, Smithies, Hickman, Wingate, &  
497 Bodorkos, 2010). This orogenic event was also associated with the intrusion of the Elizabeth  
498 Hill granitic Supersuite. Hickman (2016) interpreted there to be a probable genetic relationship  
499 to the Prinsep Orogeny between Carlow Castle and other deposits hosted within the Regal  
500 Thrust. However, Hickman et al. (2006) previously noted that an absolute temporal relationship  
501 between this mineralisation and the Prinsep Orogeny had not been established.

502 The Prinsep Orogeny and accretion of the West Pilbara Superterrane was followed by a  
503 significant period of crustal extension and the formation of the De Grey Superbasin (3066-  
504 2919 Ma) (Hickman, 2016). The formation of the De Grey Superbasin initiated with a  
505 protracted period of post-orogenic relaxation and subsidence, allowing for the formation of the  
506 Gorge Creek Basin (3066-3015 Ma) within the De Grey Superbasin as a shallow water  
507 epicontinental basin over most of the Pilbara Craton (Hickman, 2016, 2021). Sedimentation  
508 within the Gorge Creek Basin was dominated by deposition of sandstones, conglomeratic units,  
509 and carbonaceous shales in addition to banded iron formation and cherts (Van Kranendonk et  
510 al., 2006). This was followed by the formation of the Whim Creek (3009-2991 Ma) and Mallina  
511 (3015-2931 Ma) basins as later constituents of the De Grey Superbasin. Interpretations about  
512 the tectonic setting responsible for the development of the Whim Creek and Mallina basins  
513 have varied, with the key difference relating to whether basin formation was contemporaneous

514 with proximal subduction at the northwest margin of the Pilbara Craton. The first model  
515 proposed that basin formation was contemporaneous with subduction and that the Whim Creek  
516 Basin represents a continental volcanic arc, composed of volcanic, volcanoclastic, and intrusive  
517 rocks, whilst the Mallina Basin to the southeast represents a back-arc basin dominated by a  
518 thick clastic sedimentary succession (Hickman, 2012, 2016, 2021; Hickman, Smithies, &  
519 Tyler, 2010). The progressive south-eastward younging of granite ages throughout the Pilbara  
520 Craton has been invoked as evidence of subduction at the northwest margin of the Pilbara  
521 Craton (Hickman, 2016; Hickman et al., 2010). Conversely, it has been proposed that these  
522 basins formed due to continental rifting (Van Kranendonk et al., 2006; Van Kranendonk,  
523 Smithies, Hickman, & Champion, 2007; Van Kranendonk et al., 2010). It is suggested that  
524 granitic magmatism was a product of an influx of mantle material beneath the craton due to the  
525 break-off of the slab that was previously subducted from 3130 Ma (Van Kranendonk et al.,  
526 2007). Nonetheless, it is generally agreed that this long-lived period of extension ended with  
527 the beginning of the North Pilbara Orogeny from 2955-2919 Ma (Hickman, 2016, 2021;  
528 Hickman & Van Kranendonk, 2012; Van Kranendonk et al., 2006; Van Kranendonk et al.,  
529 2007; Van Kranendonk et al., 2010). This was most likely caused by continued northwest-  
530 southeast convergence and ongoing subduction (Hickman, 2016, 2021). Convergence was  
531 associated with significant compressional deformation across the West Pilbara Superterrane,  
532 Mallina Basin, and portions of the East Pilbara Terrane, along with intrusive magmatism of the  
533 Sisters Supersuite through the Whim Creek and Mallina Basins and the East Pilbara Terrane  
534 (2954-2919 Ma) (Hickman, 2016, 2021; Van Kranendonk et al., 2007). The North Pilbara  
535 Orogeny represents the last period in the development of the northwest Pilbara Craton of  
536 probable relevance to Carlow Castle Cu-Co-Au deposit's formation.





537

538 Figure 1.2 - Geological map of the Pilbara Craton, modified after Hickman (2016). MB  
 539 = Mallina Basin, WCB = Whim Creek Basin, GCB = Gorge Creek Basin. Note that the  
 540 extent of the area of the northwest Pilbara Craton within Figure 1.1 is indicated.

541 1.1.3 Hydrothermal copper-cobalt deposits

542 Generally, the occurrence of Co mineralisation with grades sufficient to be considered  
 543 economically viable for primary Co extraction are relatively rare (Hitzman, Bookstrom, Slack,  
 544 & Zientek, 2017). The only ore deposit on Earth where Co is currently mined as a primary

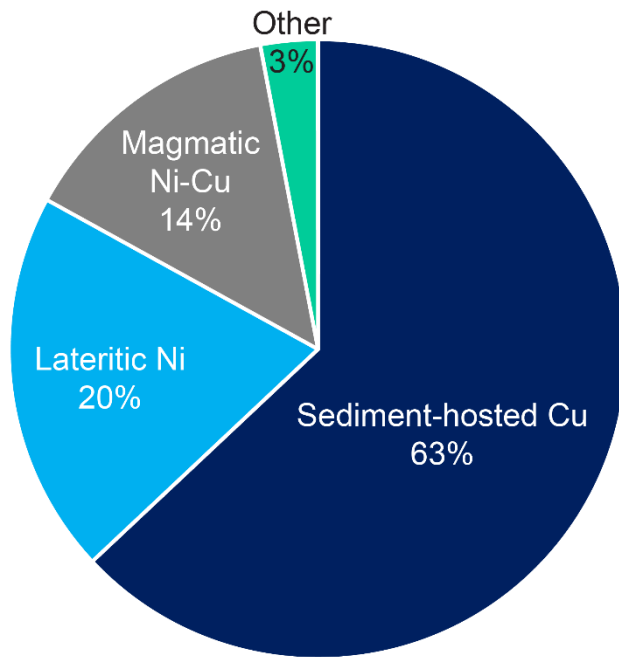


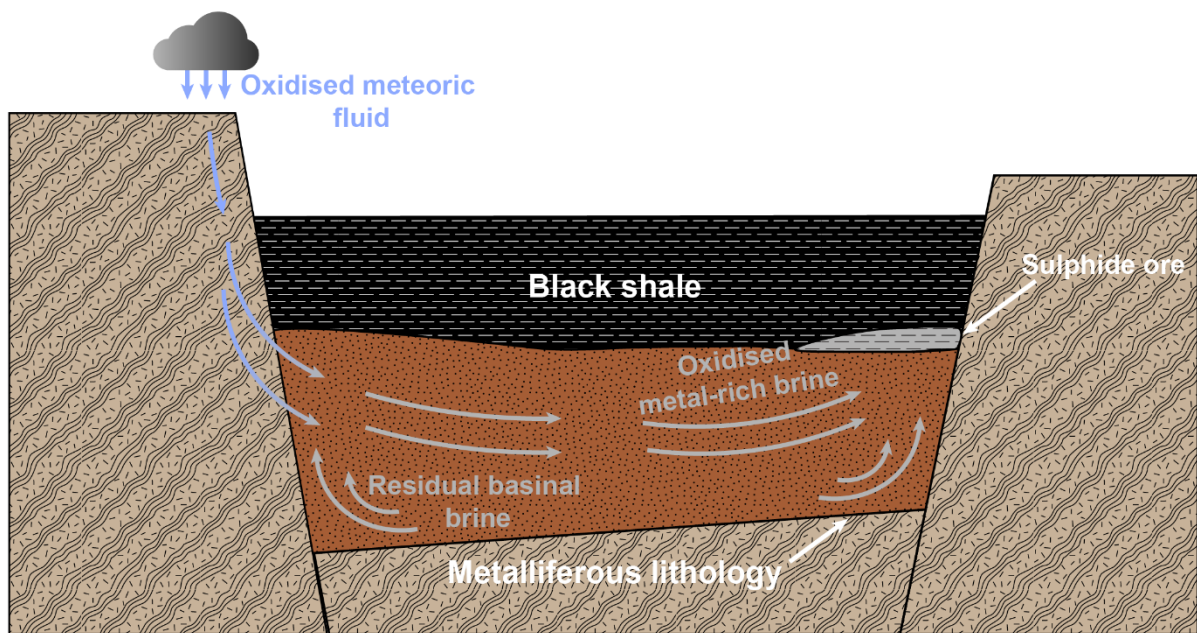
Figure 1.3 - Cobalt production by ore deposit mineralisation style (Petavratzi, Gunn, & Kresse, 2019).

product is Bou Azzer in Morocco; a Neoproterozoic polymetallic vein-hosted deposit with a Co grade of ~1% in 100,000 tonnes of ore (Bouabdellah, Maacha, Levresse, & Saddiqi, 2016; Horn et al., 2021). Instead, Co is typically extracted as a by-product of Cu or Ni mining (Petavratzi et al., 2019; Slack et al., 2017a). Stratiform sediment-hosted deposits are by far the most significant global source of Co; accounting for 63% of global Co production, followed by Ni-Co laterite deposits (20%), magmatic polymetallic sulphide deposits (14%), and other minor deposits (3%) (Figure 1.3;

560 Petavratzi et al., 2019). However, in terms of total assessed recoverable resources, laterites are  
 561 the most significant deposits, accounting for 41.7% of global Co resources, followed by  
 562 sediment-hosted (39%), iron oxide copper gold (IOCG) (10.9%), magmatic sulphide (7%),  
 563 volcanogenic massive sulphide (VMS) (0.8%), and various other deposits (0.8%) (Mudd,  
 564 Weng, Jowitt, Turnbull, & Graedel, 2013). The occurrence of Carlow Castle as structurally  
 565 controlled vein-hosted mineralisation suggests a probable hydrothermal origin (Ruddock,  
 566 1999b). As such, Carlow Castle bears little similarity to a lateritic or magmatic sulphide deposit  
 567 and is more analogous to the sediment-hosted, IOCG, VMS, or other minor Co-bearing styles  
 568 of hydrothermal mineralisation.

569 Sediment-hosted Cu deposits commonly occur at the boundary between underlying oxidised  
 570 terrestrial siliciclastic rocks and overlying reduced organic-rich marine sedimentary rocks  
 571 within a basin sedimentary sequence (Hitzman, Selley, & Bull, 2010). Among sediment-hosted  
 572 Cu deposits, the Kupferschiefer of Central Europe and Central African Copperbelt are by far  
 573 the two largest mining districts, with each containing >200 Mt of total Cu (Hitzman, Kirkham,  
 574 Broughton, Thorson, & Selley, 2005). These mining districts represent classic type-localities  
 575 of sediment-hosted Cu mineralisation (Taylor et al., 2013). Several schemes to sub-divide types  
 576 of sediment-hosted Cu deposits have been proposed based on various deposit characteristics.

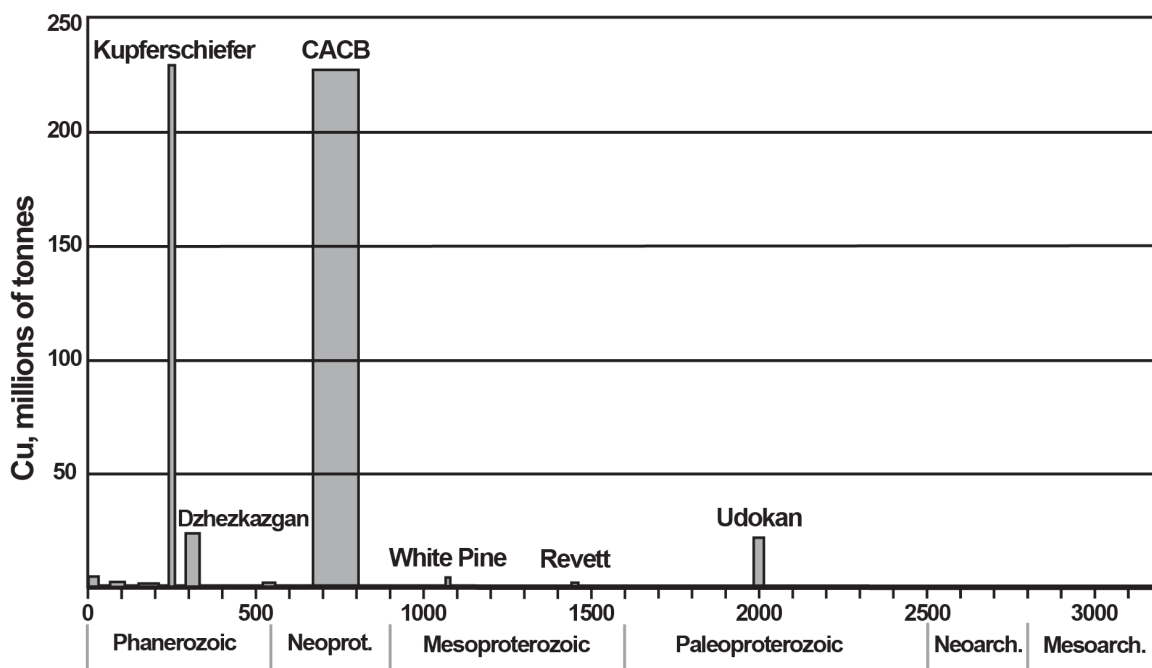
577 However, these ore systems are generally defined by several common characteristics; including  
 578 (i) Cu-rich mafic volcanic or immature mafic-rich siliciclastic sediments to act as a metal  
 579 source, (ii) oxidised, near neutral pH basinal brines to transport dissolved ore metals, (iii)  
 580 preferential migration of basinal brines due to sediment dewatering, (iv) physical traps for ore-  
 581 bearing fluids, e.g., impermeable beds or structures, and (v) reduced organic-rich or pyritic  
 582 sedimentary rocks to induce in-situ reductive ore precipitation from an oxidised ore-bearing  
 583 brine (Figure 1.3; Taylor et al., 2013).



584  
 585 Figure 1.4 - Generalised basin-scale model for a sediment-hosted Cu deposit, modified  
 586 after Brown (2014); Hitzman et al. (2010).

587 Cox, Lindsey, Singer, and Diggles (2003) initially proposed to sub-divide these sediment-  
 588 hosted Cu deposits based on characteristics of their host lithologies and this scheme was  
 589 adopted with slight modifications by Taylor et al. (2013) and Hayes, Cox, Bliss, Piatak, and  
 590 Seal Ii (2015). This scheme divides deposits into (i) reduced-facies Cu deposits (shale-hosted),  
 591 (ii) sandstone Cu deposits, and (iii) red bed Cu deposits. However, Kirkham (1989) and  
 592 Hitzman et al. (2005) preferred to divide these deposits based on their basinal setting into  
 593 Kupferschiefer-type and red bed-type deposits. Under this scheme, most significant deposits  
 594 are classified as Kupferschiefer-type, with red bed-type deposits being comparably minor in  
 595 tonnage (Hitzman et al., 2005). Unlike Carlow Castle, the occurrence of sediment-hosted Cu-  
 596 Co deposits in Archean terranes is extremely rare (Brown, 2014). This is generally accepted to  
 597 reflect the importance of oxidised basinal fluids to mobilise metals in sediment-hosted Cu

598 deposits, and the lack of available oxidised fluids within crustal systems during the widespread  
 599 anoxia that characterised most of the Archean (Figure 1.4; Brown, 2014; Hitzman et al., 2010).



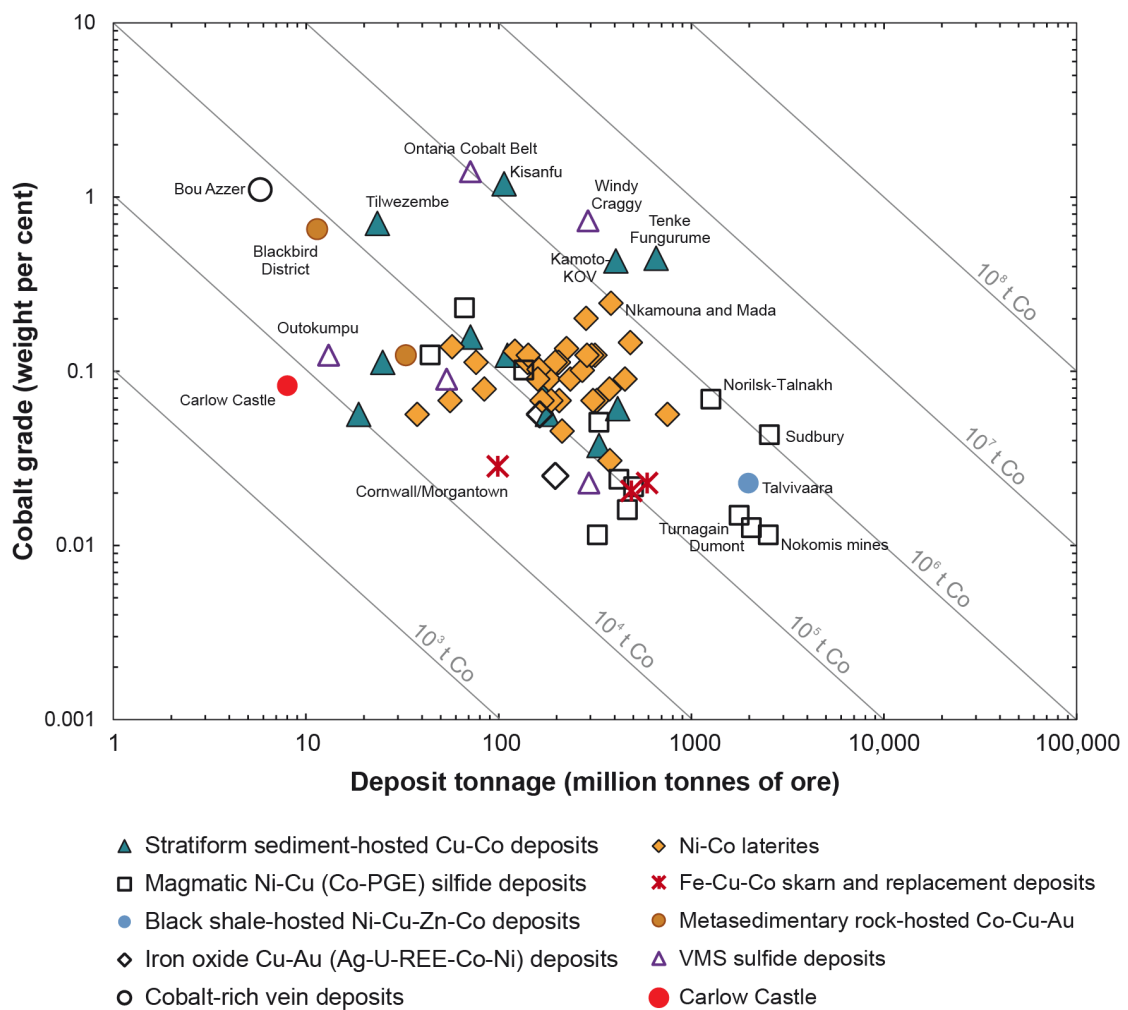
600  
 601 Figure 1.5 - Occurrence of sediment-hosted Cu deposits through geological time. Note  
 602 the lack of Archean examples, reflecting the lack of available oxidised basinal fluids to  
 603 mobilise base metals. CACB = Central African Copper Belt. Modified after Hitzman et  
 604 al. (2010).

605 Aside from sediment-hosted Cu deposits, Co-bearing examples of many hydrothermal  
 606 mineralisation styles are relatively rare. Amongst IOCG deposits, Olympic Dam (0.02% Co)  
 607 and Ernest Henry (0.05% Co) deposits in Australia contain appreciable Co concentrations.  
 608 However, Co is not currently produced from either of these deposits and Co-rich IOCG deposits  
 609 are generally uncommon (Petavratzi et al., 2019; Slack et al., 2017a). Cobalt-bearing  
 610 volcanogenic massive sulphide (VMS) deposits are similarly rare, though some pelitic-mafic-  
 611 type (Besshi-type) are relatively Co-rich; most notably Windy Craggy, Canada (0.66% Co) and  
 612 Outokumpu, Finland (0.25% Co) (Franklin, Gibson, Jonasson, & Galley, 2005; Hitzman et al.,  
 613 2017). These mineralisation styles differ significantly from the previously discussed sediment-  
 614 hosted deposits as they represent magmatic-hydrothermal systems in quite distinct geodynamic  
 615 settings, with variably significant magmatic fluid input and ore formation at significantly  
 616 higher temperatures (Franklin et al., 2005; Williams et al., 2005). Other, more obscure and  
 617 relatively minor, Co-rich styles of mineralisation include metasedimentary rock-hosted Co-Cu-

618 Au, polymetallic Co-rich vein deposits (also known as five-element vein deposits), and black  
619 shale-hosted Ni-Cu-Zn-Co deposits (Petavratzi et al., 2019; Slack et al., 2017a). Amongst  
620 metasedimentary rock-hosted Co-Cu-Au deposits, the Blackbird deposit in the Idaho Cobalt  
621 Belt represents the most significant example (Bookstrom, 2013). These deposits are poorly  
622 understood with a range of different genetic models proposed. Metasedimentary rock-hosted  
623 Co-Cu-Au deposits are typically Proterozoic, with an epigenetic origin, and a connection to  
624 proximal magmatism in many deposits has been suggested along with a possible relationship  
625 between this deposit class and Co-rich IOCG deposits (Slack et al., 2013). Polymetallic Co-  
626 rich vein deposits are a diverse group of deposits that occur as structurally-hosted Co ( $\pm$  Ag,  
627 Ni, As, Bi) bearing veins through igneous and sedimentary rocks of maximum Proterozoic age  
628 (Kissin, 1992; Lefebure, 1996). These deposits are generally mineralogically complex and are  
629 believed to reflect multiple distinct stages of mineralisation, notable examples include deposits  
630 of the Bou Azzer district of Morocco and the Cobalt district of Ontario, Canada (Ahmed, Arai,  
631 & Ikenne, 2009; Kissin, 1992; Slack et al., 2017a). Finally, black shale-hosted Ni-Cu-Zn-Co  
632 deposits are relatively rare and occur as accumulations of significant base-metal mineralisation  
633 in, typically, Proterozoic organic-rich shales (Petavratzi et al., 2019; Slack et al., 2017a).  
634 Talvivaara in Finland is the most significant Co-bearing example (Figure 1.5) and is believed  
635 to have formed due to a multistage mineralisation process during diagenesis and regional  
636 metamorphism (Loukola-Ruskeeniemi & Lahtinen, 2013; Slack et al., 2017a).

637 Carlow Castle appears to bear some broad similarities to the deposits of the polymetallic Co-  
638 rich vein and metasedimentary rock-hosted Co-Cu-Au deposits. However, the rarity of  
639 comparable Archean-hosted examples of these deposits classes is striking (Kissin, 1992; Slack  
640 et al., 2013; Slack et al., 2017a); especially given the significant endowment of Archean  
641 terranes with hydrothermal and magmatic ore deposits of other varieties (Goldfarb, Bradley, &  
642 Leach, 2010). Much like in the case of the sediment-hosted Cu-Co deposits, the temporal  
643 distribution of these Co-rich hydrothermal ore deposits is most likely a reflection of the  
644 strongly redox-controlled nature of Cu and Co solubility in hydrothermal systems at low to  
645 moderate temperatures (Rose, 1989). Similar temporal patterns are also observable for other  
646 base-metal deposits with strong redox controls (Goldfarb et al., 2010). As such, the scarcity of  
647 these deposits in the Archean most likely reflects the relative immobility of Cu and Co in  
648 reduced hydrothermal fluids and the relative rarity of oxidised near-surface environments  
649 during the Archean to supply oxidised ore fluids to hydrothermal systems. In this regard, the

650 availability of oxygen to near surface environments imparts a fundamental control on  
 651 hydrothermal Cu-Co mineralisation through geological time. Whilst acidic (pH <5 at 150°C)  
 652 dissolution of metals provides an alternative mechanism for Cu-Co dissolution within reduced  
 653 fluids, the rarity of significant Archean magmatic-hydrothermal Cu-Co deposits implies that  
 654 this pathway is of comparably limited geological significance. This pathway is also explored  
 655 in subsequent chapters within this thesis. Therefore, the occurrence of Carlow Castle in a Paleo-  
 656 Mesoarchean greenstone belt makes it a notable outlier.



657  
 658 Figure 1.6 - Grade-tonnage graph for terrestrial Co-bearing deposits. Modified after  
 659 Petavratzi et al. (2019) and Slack et al. (2017a).

## 660 1.2 Research Objectives

661 The principal objective of this project was to provide a foundational understanding of the  
 662 Carlow Castle ore system and its key genetic processes. A secondary aim of the project was to  
 663 establish a methodology for Cu isotope analysis of Cu sulphide minerals in collaboration with

664 the Metal Isotopes Group at the University of Adelaide and use the Carlow Castle deposit as a  
665 case study area for this Cu isotope analytical methodology. This serves the dual purpose of  
666 improving the currently limited understanding of Archean Cu isotope systematics whilst  
667 providing further insights into the genesis of the Carlow Castle ore deposit.

668 These objectives were to be achieved through the following research goals:

- 669 1. Characterise the nature and style of Carlow Castle Cu-Co-Au deposit's mineralisation.
- 670 2. Determine the age of Carlow Castle and its temporal relationship to major tectonic  
671 events during the formation of the Pilbara Craton.
- 672 3. Characterise the multiple sulphur isotope composition of Carlow Castle, to inform  
673 about potential ore-forming processes and the ore system's sulphur source.
- 674 4. Develop a robust method for Cu isotope analysis of Cu sulphides in collaboration with  
675 the Metal Isotopes Group, University of Adelaide.
- 676 5. Characterise the Cu isotope signature of Carlow Castle Cu-Co-Au deposit, as a test of  
677 the applicability of Cu isotope analysis to understanding Archean hydrothermal systems  
678 and to inform about potential metal source and ore-forming processes.

### 679 1.3 Thesis structure

680 This thesis presents a compilation of published papers and manuscripts prepared for publication  
681 in peer-review academic journals. As such, each of these papers represents a distinct aspect of  
682 this thesis, attempting to address the above-listed aims and objectives of the research project.  
683 However, because this thesis essentially represents a compilation of several smaller case  
684 studies there is some unavoidable repetition of content between different thesis chapters due to  
685 need for these chapters to be understood independently. Chapters Two and Three are published  
686 in peer-reviewed journals and are reproduced in this thesis with minor modifications. Chapters  
687 Four and Five have been submitted to academic journals and are currently under review.

### 688 Chapter Two – Plundering Carlow Castle: First Look at a Unique 689 Mesoarchean-Hosted Cu-Co-Au Deposit

690 This chapter provides the first analysis of Carlow Castle Cu-Co-Au deposit and a foundation  
691 for subsequent studies of the deposit. This work characterises Carlow Castle as a hydrothermal  
692 Cu-Co-Au deposit, with ore mineralisation occurring through a series of quartz-carbonate  
693 veins, hosting Cu and Co sulphides. Two distinct ore minerals assemblages are identified, with  
694 the first being composed primarily of pyrite and chalcopyrite and the second being composed

695 primarily of chalcocite, cobaltite, and electrum. The mineralisation occurs through a heavily  
696 brecciated portion of the Roebourne Group adjacent to the Regal Thrust; a regionally  
697 significant crustal-scale orogenic thrust fault. The chapter also provides a comparison of  
698 Carlow Castle to other styles of Cu-Co mineralisation, noting the relative rarity of comparable  
699 hydrothermal Cu-Co deposits in Archean terranes. Finally, it is suggested that Carlow Castle  
700 is of probable Archean age, though the need for absolute age constraints is acknowledged.

### 701 Chapter Three – Working up an Apatite: Enigmatic Mesoarchean 702 Hydrothermal Cu-Co-Au Mineralization in the Pilbara Craton

703 Chapter Three establishes the first absolute age constraints on Carlow Castle Cu-Co-Au  
704 deposit's formation using in-situ U-Pb dating of hydrothermal apatite that are interpreted to  
705 have formed coeval with mineralisation. Analyses of these apatite grains in-situ gave a  
706 discordia age of  $2957 \pm 67$  Ma, making Carlow Castle among the oldest Cu-Co deposits on  
707 Earth. This age suggests that Carlow Castle significantly post-dates its orogenic host structure,  
708 the Regal Thrust, which formed around 3070 Ma during the Prinsep Orogeny. Instead, ore  
709 formation at Carlow Castle is contemporaneous with the rifting of the De Grey Superbasin  
710 following orogenesis. A potential connection is also drawn to base-metal ore formation within  
711 the proximal Whim Creek Basin around 2955 Ma. Additionally, characterisation of the  
712 alteration mineral assemblage of Carlow Castle indicates that it is characteristic of a classic  
713 propylitic alteration assemblage; dominated by chlorite, actinolite, epidote, calcite, quartz, and  
714 K-feldspar. This is suggestive of peak temperatures of hydrothermal ore formation of  $\sim 300^\circ\text{C}$   
715 and neutral-alkaline pH ore fluids.

### 716 Chapter Four - Mesoarchean oxygenation accompanied massive copper- 717 cobalt mineralization

718 In chapter Four, analysis of the multiple sulphur isotope signature of mineralisation from  
719 Carlow Castle is provided through in-situ analysis of chalcopyrite, pyrite, and pyrrhotite via  
720 secondary-ion mass spectrometry. This sulphide mineralisation displays  $\delta^{34}\text{S}$  values between -  
721 3.4‰ and 6.1‰ and  $\Delta^{33}\text{S}$  values between -0.17‰ and 0.12‰. These  $\Delta^{33}\text{S}$  data are consistent  
722 with processes of mass-dependent sulphur fractionation, with no contamination from  
723 supracrustal 'Archean sulphur' which displays large magnitude mass-independent sulphur  
724 fractionation evident in  $\Delta^{33}\text{S}$  values greater than 0.2‰ and less than -0.2‰. However,  
725 paradoxically the  $\delta^{34}\text{S}$  signature is not consistent with a mantle sulphur source ( $\delta^{34}\text{S} = 0 \pm 2\text{‰}$ ).



726 In this regard, Carlow Castle is unique amongst Archean hydrothermal ore deposits. Carlow  
727 Castle is contemporaneous with well-established oxidative conditions at ~2950 Ma, which  
728 potentially drove Earth's first major glaciation and attenuated  $\Delta^{33}\text{S}$  values in the Archean  
729 sedimentary sulphur isotope record. As such, it is suggested that the genesis of Carlow Castle  
730 may have been associated with this brief period of increased atmospheric oxygen  
731 concentration. This is also supported by thermodynamic modelling demonstrating the  
732 importance of oxidised fluids to mobilise Cu and Co at physicochemical conditions consistent  
733 with ore formation, as established in Chapter Three.

## 734 Chapter Five - Copper isotopes in Archean hydrothermal systems: A case 735 study from the Mesoarchean Carlow Castle Cu-Co-Au deposit

736 Chapter Five provides an analysis of Carlow Castle Cu-Co-Au deposit utilising Cu isotope  
737 geochemistry of the deposit's primary Cu sulphide ore and overlying supergene Cu-bearing  
738 minerals. This study serves the dual purposes of constraining key genetic processes in the  
739 formation of Carlow Castle Cu-Co-Au deposit whilst also addressing a key gap in the current  
740 understanding of Cu isotopic analysis in ore deposits by providing the first Cu isotope analysis  
741 of an Archean Cu ore deposit. The primary sulphide ore from Carlow Castle displays lightly  
742 fractionated  $\delta^{65}\text{Cu}$  values ranging between -0.80‰ to 0.00‰ whilst supergene Cu  
743 mineralization is slightly isotopically heavier, ranging between -0.50‰ to 0.62‰. The light  
744 isotopic fractionation in primary sulphide mineralization demonstrates fractionation beyond  
745 bulk silicate Earth compositions, controlled primarily by Rayleigh-type fractionation driven by  
746 preferential  $^{63}\text{Cu}$  precipitation during ore formation. This process is modelled, and a Rayleigh  
747 fractionation model is derived from Cu grade (ppm) and isotopic data of analysed samples.  
748 From this model, a mafic igneous source is suggested as the most likely Cu source for ore  
749 metals in Carlow Castle. The limited positive isotopic fractionation of the supergene samples  
750 analysed in this study suggests limited redox cycling of Cu, this is interpreted to reflect in-situ  
751 oxidation of vein-hosted Cu sulphides to form this supergene Cu mineralisation rather than  
752 significant Cu transport and enrichment necessary to form significant supergene isotopic  
753 fractionation.

## 754 Chapter Six – Synthesis and conclusions

755 Chapter Six is the final chapter of this thesis and provides a synthesis of the findings of the  
756 preceding chapters. These findings are developed into a mineral systems model for Carlow

757 Castle Cu-Co-Au deposit. Additionally, the implications of this research are discussed, along  
758 with discussion of the limitations of the current findings and outlook for further research.

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893

## 894 Chapter 2

# 895 Plundering Carlow Castle: first look at 896 a unique Mesoarchean-hosted Cu-Co- 897 Au deposit

898

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904

### 905 2.1 Abstract

906 Economically significant and geologically complex veined Cu-Co-Au mineralization was  
907 recently discovered at Carlow Castle in the Pilbara region of north-western Western Australia.  
908 The inferred resource estimate for Carlow Castle as of May 2021 is 8 Mt at 0.7 g/t Au, 0.4%  
909 Cu, and 0.05% Co (Artemis Resources Limited, 2021), making it one of Australia's most  
910 significant known Cu-Co-Au deposits. Here we provide the first account and scientific analysis  
911 of Carlow Castle. This analysis suggests that it is a hydrothermal Cu-Co-Au deposit, with  
912 mineralization hosted in sulfide-rich quartz-carbonate veins. The ore is hosted in veins that  
913 occur within a pervasively chloritized shear zone through the regionally significant Regal  
914 Thrust. At Carlow Castle the shear zone associated with this thrust occurs within the Ruth Well  
915 and Nickol River Formations; an Archean mafic volcano-sedimentary sequence. Within the  
916 mineralized veins the dominant ore minerals are pyrite (FeS<sub>2</sub>), chalcopyrite (CuFeS<sub>2</sub>),  
917 chalcocite (Cu<sub>2</sub>S), cobaltite (CoAsS), and electrum (Au,Ag). The genesis of the Carlow Castle  
918 deposit is still under investigation; however, the origin of the Cu-Co-Au mineralization is most  
919 likely related to the migration of metalliferous fluids along the Regal Thrust. Based on Carlow

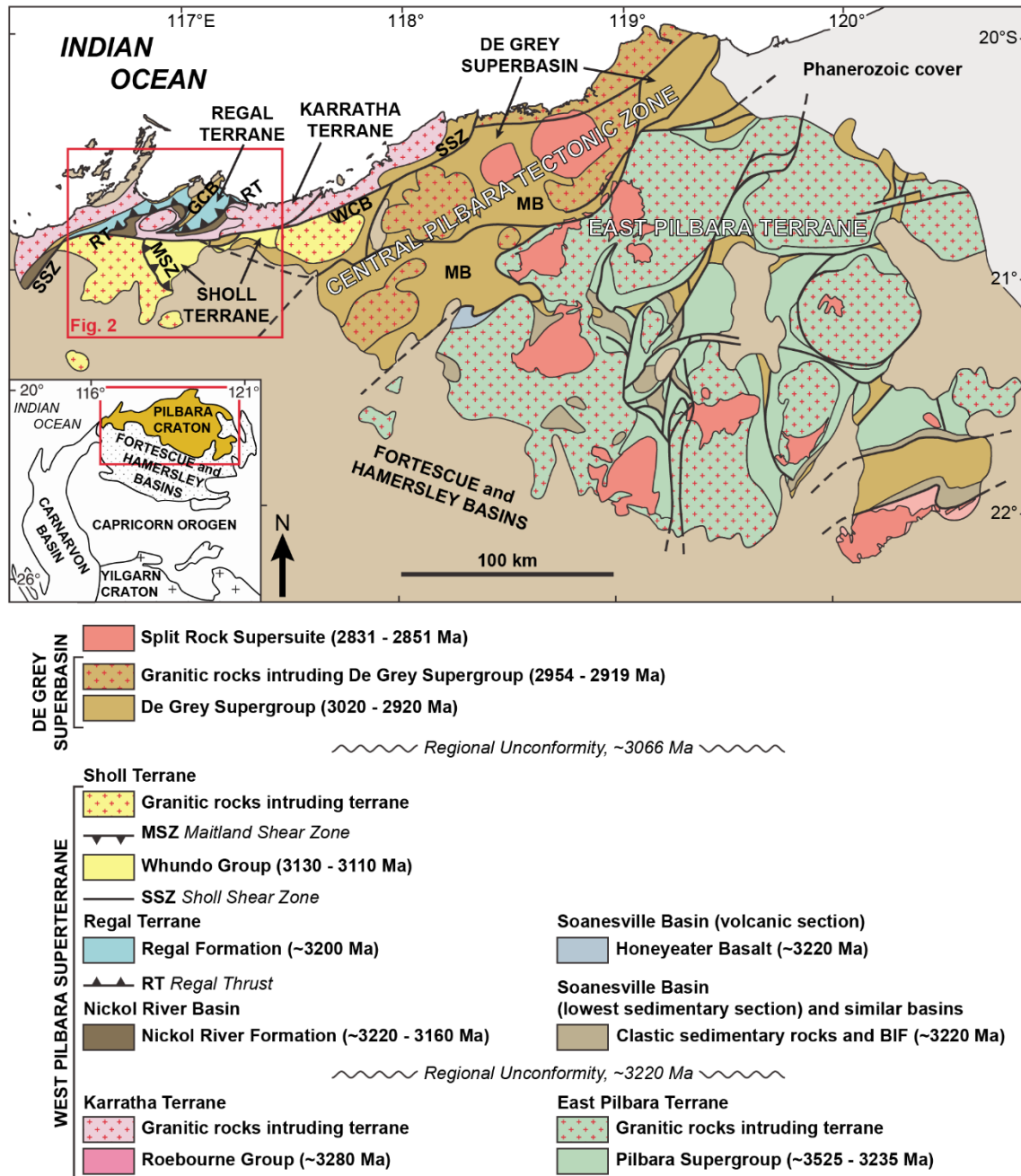


920 Castle's stratigraphic position within the Pilbara Craton and the craton's relative stability since  
921 the Archean, an Archean age of mineralization is most likely. The distinct Cu-Co-Au  
922 enrichment at Carlow Castle makes it unique among Archean ore deposits generally, as the  
923 majority of Cu-Co deposits are of maximum Proterozoic age. Therefore, understanding the  
924 genesis of the Carlow Castle deposit has important implications for understanding the unique  
925 processes through which Cu-Co-Au mineralization outside of basin-hosted ore deposits may  
926 be formed, particularly in Archean terranes.

## 927 2.2 Introduction

928 The Carlow Castle copper-cobalt-gold (Cu-Co-Au) deposit is hosted in the Roebourne  
929 greenstone belt within the West Pilbara region, Western Australia (Figure 2.1). The current  
930 inferred resource for Carlow Castle amounts to 7.7 Mt at 0.51% Cu, 0.08% Co, and 1.06 g/t  
931 Au (Artemis Resources Limited, 2019a); making it one of Australia's most well-endowed Cu-  
932 Co-Au deposits (Britt et al., 2017). These resource estimates are based on extensive exploration  
933 and drilling campaigns through 2017 and 2018, which resulted in the discovery of  
934 mineralization proximal to the previously abandoned minor Carlow Castle Cu workings. This  
935 drilling effort comprised infill drilling of historic exploration drill holes that had noted minor  
936 Cu-Au mineralization. However, despite extensive drilling targeting mineralization, the  
937 geology of the Carlow Castle deposit and the origin of its peculiar Cu-Co-Au mineralization  
938 remain poorly documented and constrained. This work is particularly topical given the  
939 remarkable endowment of Carlow Castle's mineralization in Co. Indeed, Co has been  
940 designated as a critically important metal by governmental agencies in the UK; the European  
941 Union; Australia; and the USA (British Geological Survey, 2015; European Commission,  
942 2017; Mudd et al., 2018; U.S. Department of the Interior, 2018), with increased demand from  
943 the battery sector due to the status of Co as a key metal in Li-ion batteries (US Geological  
944 Survey, 2018). Due to proliferation of electric vehicles and the challenge of stabilizing  
945 renewable energy generation, Co is a critical metal for the evolving global economy. Presently,  
946 >50% of the world's Co is supplied by the Democratic Republic of Congo, which is exposed  
947 to significant geopolitical risk (Olivetti, Ceder, Gaustad, & Fu, 2017). Beyond its economic  
948 significance, the distinctive Cu-Co enrichment of the Carlow Castle deposit makes it an oddity  
949 among Au deposits hosted in Archean greenstone belts (Groves, Goldfarb, Gebre-Mariam,  
950 Hagemann, & Robert, 1998; Kerrich, Goldfarb, Groves, Garwin, & Jia, 2000; Saunders,  
951 Hofstra, Goldfarb, & Reed, 2014). The present work provides a robust geological and

952 mineralogical background to the Cu-Co-Au mineralization at Carlow Castle to provide a basis  
 953 for different analytical approaches in subsequent chapters, including geochronology and  
 954 isotope geochemistry of the Cu-Co-Au mineralization. To this end, we provide interpretations  
 955 regarding Carlow Castle's genesis based on this geological and mineralogical characterization.



956  
 957 Figure 2.1 - Geological map of the Pilbara Craton, modified after Hickman (2016). MB  
 958 = Mallina Basin, WCB = Whim Creek Basin, GCB = Gorge Creek Basin. Note that the  
 959 extent of the area of the northwest Pilbara Craton within Figure 2.2 is indicated and the  
 960 symbology of different terranes is noted in the figure legend.

## 961 2.3 Exploration history

962 The first discovery of Cu mineralization at Carlow Castle, formerly known as Glenroebourne,  
963 dates to 1872. Minor surficial Cu and Au mining occurred intermittently until 1957 (Openpit  
964 Mining Limited, 1987; Ruddock, 1999a). It has since been the subject of several minor  
965 exploration efforts, with limited success until major mineralized intersections were  
966 encountered during drilling in 2017 and 2018 (Artemis Resources Limited, 2017, 2018a).  
967 Artemis Resources released a resource estimate of 4.5 Mt at 0.4% Cu, 0.07% Co, and 0.9g/t  
968 Au for Carlow Castle in January 2018 (Artemis Resources Limited, 2018b). The current  
969 inferred resource estimate is 7.7 Mt at 0.51% Cu, 0.08% Co, and 1.06 g/t Au, as of March 2019  
970 (Artemis Resources Limited, 2019b). A third phase of drilling was initiated in December 2018  
971 aiming for extensions of mineralization east and west along strike; results are pending (Artemis  
972 Resources Limited, 2018b).

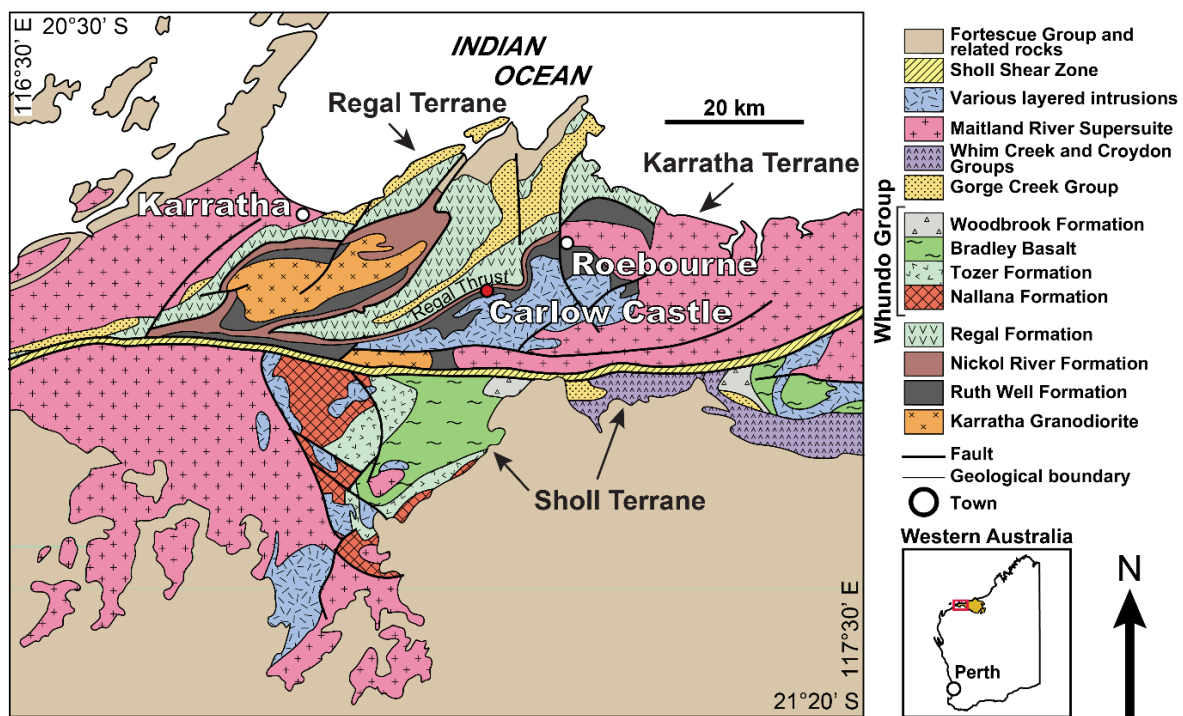
## 973 2.4 Geological Setting

974 The Carlow Castle Cu-Co-Au deposit is hosted within the Archean to Proterozoic northwest  
975 Pilbara Craton (Figure 2.1; Ruddock, 1999a; Smithies, Hickman, & Nelson, 1999). The  
976 northwest Pilbara Craton is broadly divided into two major tectonic units; the De Grey  
977 Superbasin (3066-2919 Ma) and the West Pilbara Superterrane (3280-3066 Ma) (Hickman,  
978 2016). The West Pilbara Superterrane, which hosts the Carlow Castle deposit, is further divided  
979 into the Nickol River Basin and Sholl, Regal, and Karratha Terranes (Figure 2.1; Hickman,  
980 2016; Kemp, Hickman, Kirkland, & Vervoort, 2015; Peterson et al., 2019). These Terranes  
981 form a chain of several granite-greenstone belts composed of a mix of Archean mafic volcanic  
982 rocks and intrusive granitic complexes (Hickman et al., 2006). This provides a unique  
983 geological setting for Cu-Co-Au mineralization, the majority of which is hosted in terranes of  
984 Proterozoic or younger age (US Geological Survey, 2018). The West Pilbara Superterrane is  
985 presently interpreted to have formed during the Prinsep Orogeny (3070 Ma) via the accretion  
986 of the Sholl, Regal, and Karratha Terranes. The Karratha Terrane (3280-3236 Ma) is  
987 interpreted to represent a separated portion of the early Paleoproterozoic Pilbara Craton that rifted  
988 from the East Pilbara Terrane (3530-3223 Ma), which presently crops out ~130 km east of  
989 Carlow Castle (Figure 2.1), during the initiation of horizontal plate tectonics around 3220 Ma  
990 (Hickman, 2021). Subsequently, a thin basaltic oceanic crust, that presently composes the  
991 Regal Terrane (3200-3160 Ma), developed within the Central Pilbara Basin between the rifted  
992 Karratha and East Pilbara Terranes. The Nickol River Basin developed along a passive margin

993 of the rifted Karratha Terrane as a sequence of marine sediments (Hickman, 2021; Hickman et  
994 al., 2010). Renewed convergence of the Karratha and East Pilbara Terranes from 3160 Ma  
995 initiated subduction within the Central Pilbara Basin and ultimately resulted in basin closure  
996 (Hickman, 2016). Convergence and subduction within the Central Pilbara Basin between 3130-  
997 3093 Ma initiated the formation of the Sholl Terrane as a volcanic oceanic island arc (Van  
998 Kranendonk et al., 2007). This extended period of convergence ultimately culminated in the  
999 Prinsep Orogeny at 3070 Ma and the accretion of the Nickol River Basin and Sholl, Karratha,  
1000 and Regal Terranes into the West Pilbara Superterrane (Hickman, 2016, 2021). The Prinsep  
1001 Orogeny is also generally interpreted to have produced the Regal Thrust; a regionally  
1002 significant thrust fault that separates the overlying Regal Terrane to the north from the  
1003 underlying Karratha Terrane to the south (Figure 2.2) and has been associated with Cu-Au  
1004 mineralization regionally (Hickman, 2004a, 2016; Hickman et al., 2006; 2010; Van  
1005 Kranendonk et al., 2010). The Regal Thrust likely represents the plane along which the Regal  
1006 Terrane was thrust over the Karratha Terrane and Nickol River Basin when the West Pilbara  
1007 Superterrane collided with the East Pilbara Terrane (Hickman, 2016).

1008 Following the accretion of the West Pilbara Superterrane, an extended period of post-orogenic  
1009 crustal relaxation, extension, and basin formation initiated across the Pilbara Craton (Hickman  
1010 et al., 2006). This extensional period led to the formation of the De Grey Superbasin (3066-  
1011 2919 Ma); beginning with the formation of the Gorge Creek Basin (3066-3015 Ma) as a  
1012 shallow-water basin covering much of the West Pilbara Superterrane and East Pilbara Terrane  
1013 (Hickman, 2016). Within the Gorge Creek Basin, sedimentation was dominated by a succession  
1014 of sandstones, conglomerate, and shales in addition to chert and banded iron formations  
1015 (Hickman, 2016; Van Kranendonk et al., 2007). A new phase of regional convergence resulted  
1016 in the development of a subduction zone along the northwest margin of the Pilbara Craton.  
1017 Continued development of this subduction zone ultimately resulted in the formation of the  
1018 Whim Creek Basin (3009-2991 Ma) as a volcanic arc over the Pilbara Craton to the southeast  
1019 of Carlow Castle (Figure 2.1) and significant mafic and felsic magmatism resulting in the  
1020 intrusion of the Orpheus (3023-3012 Ma) and Maitland River (3006-2982 Ma) Supersuites  
1021 (Figure 2.2) through much of the West Pilbara Superterrane (Hickman, 2021). In particular,  
1022 the layered mafic-ultramafic Andover Intrusion ( $3016 \pm 4$  Ma) is observed to crop out  
1023 immediately south of Carlow Castle and the Ruth Well Formation, where it occurs as a series  
1024 of layered ultramafic and gabbroic rocks (Hickman, 2016). During the emergence of the Whim

1025 Creek Basin volcanic arc, the Mallina Basin (3015-2931 Ma) developed to the southeast  
 1026 (Figure 2.1) as a back-arc basin (Hickman, 2016, 2021; Hickman et al., 2010). The Mallina  
 1027 Basin preserves 2-4 km of sediment fill that is composed predominantly of conglomerate,  
 1028 sandstone, and shale (Eriksson, 1982; Hickman et al., 2010). Extension within the Mallina  
 1029 Basin was eventually halted following the initiation of convergence across most of the  
 1030 northwestern Pilbara Craton during the latter stages of the North Pilbara Orogeny from 2955-  
 1031 2919 Ma (Hickman et al., 2010; Hickman & Van Kranendonk, 2012). This period was  
 1032 associated with major granitic magmatism, with the intrusion of the Sisters Supersuite (2954-  
 1033 2919 Ma) throughout the Mallina Basin (Hickman et al., 2010). The North Pilbara Orogeny  
 1034 represents the final stage of the cratonization of the Pilbara Craton and the final reconvergence  
 1035 of the West Pilbara Superterrane and the East Pilbara Terrane (Hickman, 2016).

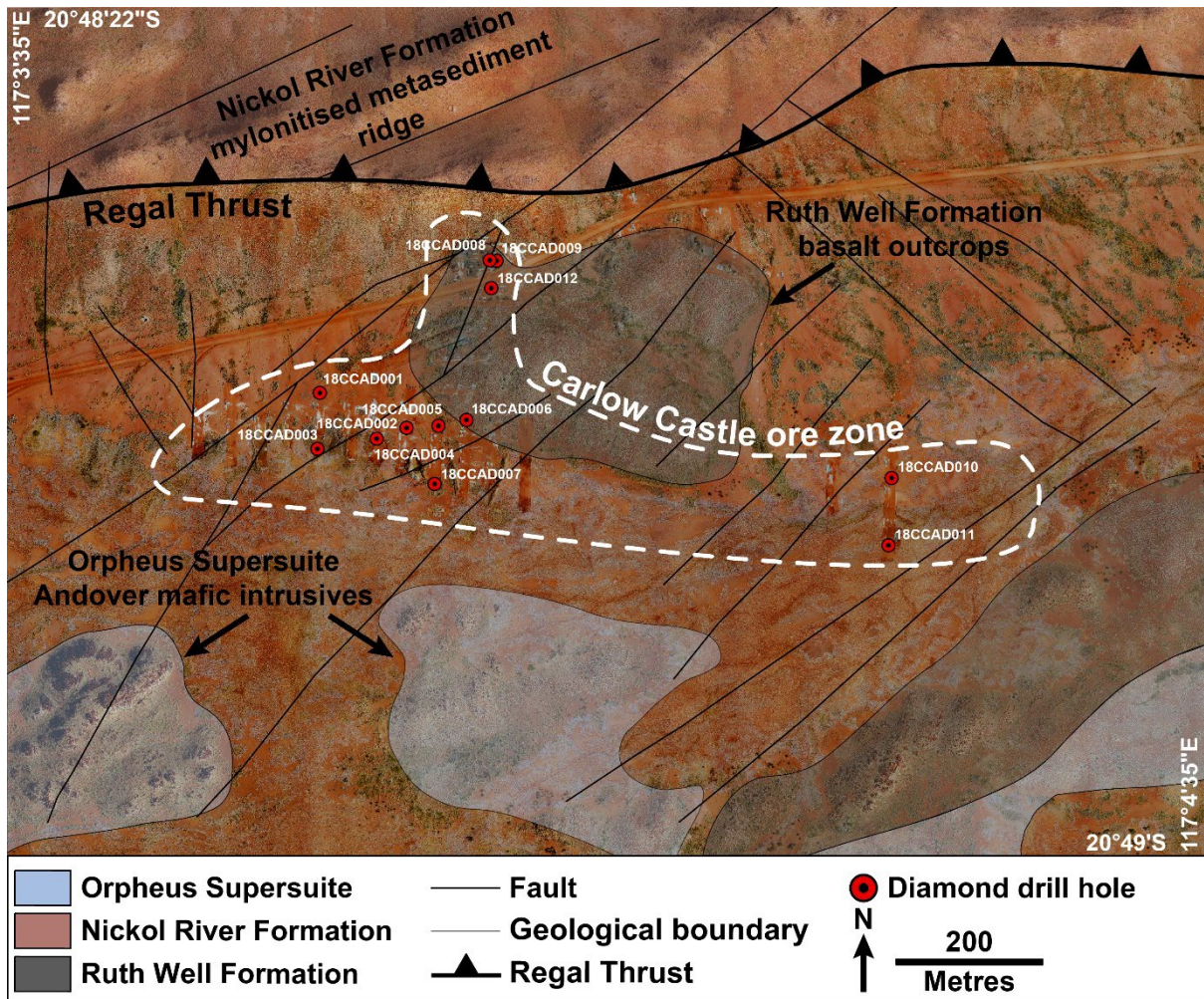


1036  
 1037 Figure 2.2 - Geological map of the northwest Pilbara Craton. The Pilbara Craton is  
 1038 highlighted in orange within the inset map. Modified after Van Kranendonk et al. (2002).

1039 Carlow Castle is hosted in a tectonized zone through the Nickol River and Ruth Well  
 1040 formations, on the southern side of the Regal Thrust; separating the Nickel River and Ruth  
 1041 Well Formations from the Regal Terrane (Figure 2.3; Hickman, 2016; Ruddock, 1999a). The  
 1042 Ruth Well Formation (~3280 Ma) is a ~1500 m thick volcano-sedimentary succession  
 1043 dominated by variably altered metamorphosed peridotite and basalt, with minor chertified shale  
 1044 (Hickman, 2016; Van Kranendonk et al., 2002). Locally pillowed basalt within the Ruth Well

1045 Formation provide evidence of submarine extrusion (Hickman, 2016). This is conformably  
1046 overlain by the Weerianna Basalt; a metamorphosed basalt unit up to 1000 m in thickness,  
1047 which lacks peridotite and chert units (Hickman, 2016). Together the Ruth Well Formation and  
1048 Weerianna Basalt constitute the Roebourne Group within the Karratha Terrane (Figure 2.2),  
1049 which was historically interpreted to have formed as an Archean mid-ocean ridge basalt (Kato  
1050 et al., 1998; Kiyokawa et al., 2002; Shibuya, Kitajima, Komiya, Terabayashi, & Maruyama,  
1051 2007). However, recently Hickman (2016) has argued that the Roebourne Group was instead  
1052 extruded onto continental crust of probable tonalite-trondhjemite-granodiorite (TTG)  
1053 composition that formed part of the original Paleoproterozoic Pilbara Craton prior to rifting. The  
1054 felsic rocks of the East Pilbara Terrane (3720-3165 Ma), which developed independently for  
1055 much of the Archean, are considered to represent a record of this original TTG-dominated  
1056 Paleoproterozoic Pilbara Craton crust (Hickman, 2016; Van Kranendonk et al., 2006). The rocks  
1057 of the Roebourne Group are overlain by the intensely sheared metasediments of the Nickol  
1058 River Formation (~3220 Ma); a succession of fine-grained marine siliciclastic rocks initially  
1059 deposited during Paleoproterozoic rifting of the East Pilbara Terrane (Hickman, 2016).

1060 On the deposit-scale, the lithologies in proximity to the Carlow Castle ore zone are covered  
1061 with a cover sequence of quaternary colluvium and alluvium (Hickman, 2002). However, rocks  
1062 of the Ruth Well Formation of probable basaltic composition are observed to crop out within  
1063 the ore zone, whilst gabbroic intrusive outcrops of the Orpheus Supersuite Andover Intrusion  
1064 are observed due south of the ore zone (Figure 2.3). The northern edge of the Carlow Castle  
1065 ore zone is defined by a ridge of mylonitized and chertified shale and minor banded iron  
1066 formations of the Nickol River Formation that strikes broadly from southwest to northeast  
1067 along the interpreted extent of the Regal Thrust across the southeastern boundary of the Regal  
1068 Terrane (Figure 2.1). The Carlow Castle ore zone occurs over a strike length of approximately  
1069 1.2 km, striking roughly east-west; parallel to the outcropping chert ridge that defines the  
1070 preserved extent of the Nickol River Formation and Regal Thrust (Figure 2.3). Within the ore  
1071 zone, mineralization is structurally-controlled and occurs through a series of Cu-, Co-, and Au-  
1072 bearing lodes within mineralized faults that strike broadly northeast-southwest. These  
1073 mineralized faults are marked on Figure 2.3 and are interpreted based on the results of a sub-  
1074 audio magnetics survey conducted by Artemis Resources Limited (Eremenco, Mortimer, &  
1075 Mead, 2019).



1076

1077 Figure 2.3 – Geological map of outcrop lithologies proximal to Carlow Castle Cu-Co-Au  
 1078 deposit, modified after (Hickman, 2002) and internally produced maps by Artemis  
 1079 Resources Limited.

1080 **2.5 Methods**

1081 For this study 44 samples were taken from core from 12 drill holes that penetrated the Carlow  
 1082 Castle ore body in 2018 (Figure 2.3; down hole geochemistry plots of drill hole data in  
 1083 Appendix 2.1; sample list in Appendix 2.2). Core from these drill holes was logged in detail  
 1084 by Geologists from Artemis Resources Limited and the authors, representative samples of  
 1085 mineralisation of varying mineralogical and geochemical characteristics were then collected.  
 1086 Eleven halved core samples from these drill holes were then polished for X-ray fluorescence  
 1087 (XRF) mapping and polished blocks were prepared for petrographic analysis. The XRF  
 1088 elemental mapping of Cu and Co was conducted using a Bruker M4 TORNADO Micro-XRF  
 1089 mapper with a beam diameter of 25 µm, point spacing of 40 µm, and a dwell time of 10 ms.

1090 Preliminary petrographic analysis was conducted through reflected light optical microscopy  
1091 using a Nikon LV100N POL petrographic microscope to roughly establish mineralogy and  
1092 mineral textures. This was followed by detailed petrographic analysis through SEM imaging  
1093 and quantitative elemental mapping, conducted using a Philips XL 40 environmental SEM  
1094 fitted with a Bruker X-Flash energy dispersive X-ray spectroscopy (EDS) detector. These  
1095 analyses were conducted using the facilities at the CSIRO Advanced Characterisation Facility  
1096 in the Australian Resources Research Centre in Perth, Western Australia.

## 1097 2.6 Results

### 1098 2.6.1 Style of mineralization

1099 Mineralization at Carlow Castle is hosted in a heavily chlorite-silica altered mafic volcano-  
1100 sedimentary sequence, within the Nickol River and Ruth Well Formations. It is often difficult  
1101 to constrain the primary host lithology of Carlow Castle mineralization due to extensive post-  
1102 depositional alteration during ore formation. However, based on unaltered equivalent sections  
1103 of the host formations observed by Hickman (2016); Ruddock (1999a); Van Kranendonk et al.  
1104 (2006), it is likely that the host lithology was dominated by a volcanic sequence of mafic to  
1105 ultramafic rocks along with carbonaceous shale and minor chert beds. It has previously been  
1106 observed outside of the Carlow Castle ore zone that these formations have been  
1107 metamorphosed to amphibolite facies, most likely during the period of tectonic convergence  
1108 preceding the Prinsep Orogeny between 3160-3070 Ma (Hickman, 2016). Within the Nickol  
1109 River and Ruth Well Formations, the hypogene Co-Cu-Au mineralization typically occurs  
1110 within 100 m of the contemporary surface (see drill hole geochemistry logs in Appendix 2.1).  
1111 This mineralization occurs as a primary sulfide ore zone with an overlying supergene oxidized  
1112 zone that varies from 25 m to 50 m thick. Within this oxidized supergene layer, there is partial  
1113 to complete replacement of the original sulfide mineralization with secondary Cu-oxides,  
1114 silicates, and carbonates (e.g., chrysocolla, malachite). The primary sulfide mineralization is  
1115 the most significant and is structurally-hosted; where mineralization occurs in quartz-carbonate  
1116 and sulfide veins through brecciated and sheared host rock.

### 1117 2.6.2 Structure

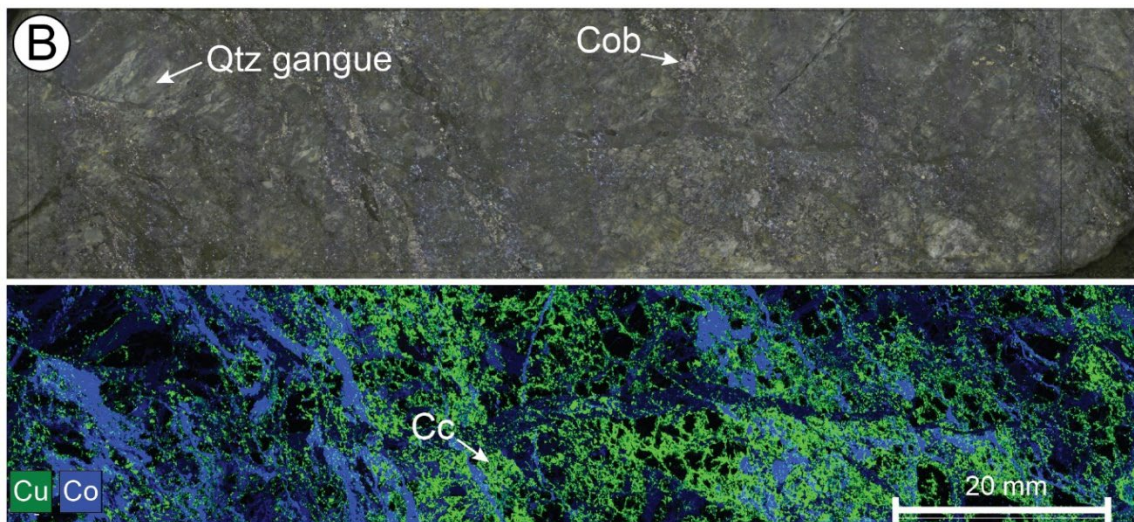
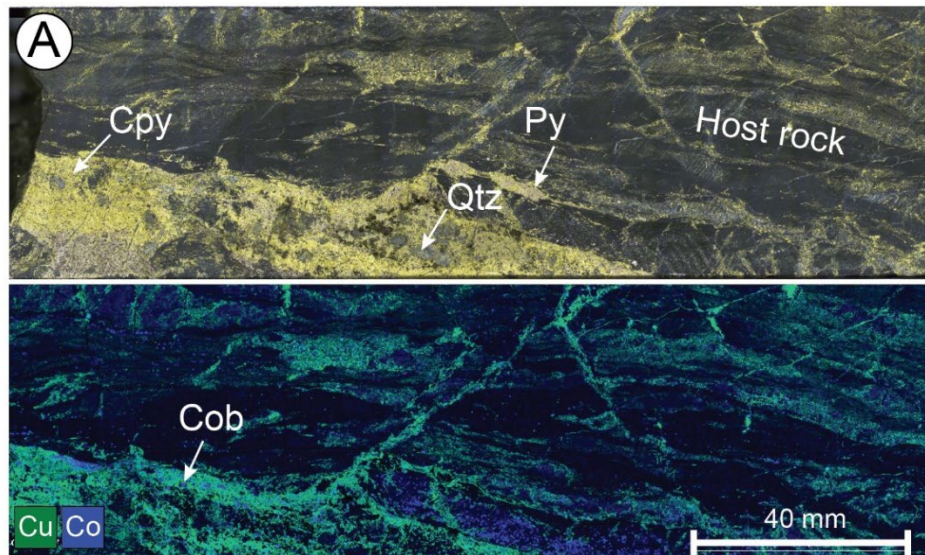
1118 The Nickol River and Ruth Well Formations at Carlow Castle bear evidence of several phases  
1119 of deformation. Brittle deformation is intense in some areas, with brecciation forming a  
1120 network of mineralized fractures within the ore zone (Figure 2.4a, b). In addition to these brittle  
1121 structures, there is also extensive evidence of ductile deformation at Carlow Castle, where



1122 quartz and sulfide veins within the Carlow Castle ore zone occur through intensely sheared and  
1123 folded host rocks. Both brittle and ductile structures at Carlow Castle host mineralization,  
1124 which occurs infilling these structures as a series of disseminated to discordant quartz-sulfide  
1125 veins. The timing of these structures relative to the sulfide mineralization at Carlow Castle is  
1126 presently unconstrained. Regardless, these structures appear to have been critical to the genesis  
1127 of the mineralization as they provided permeable structures to focus migrating mineralizing  
1128 fluids at Carlow Castle. This includes mineralizing fluids that may have flowed along the  
1129 proximal Regal Thrust.

### 1130 2.6.3 Mineralogy

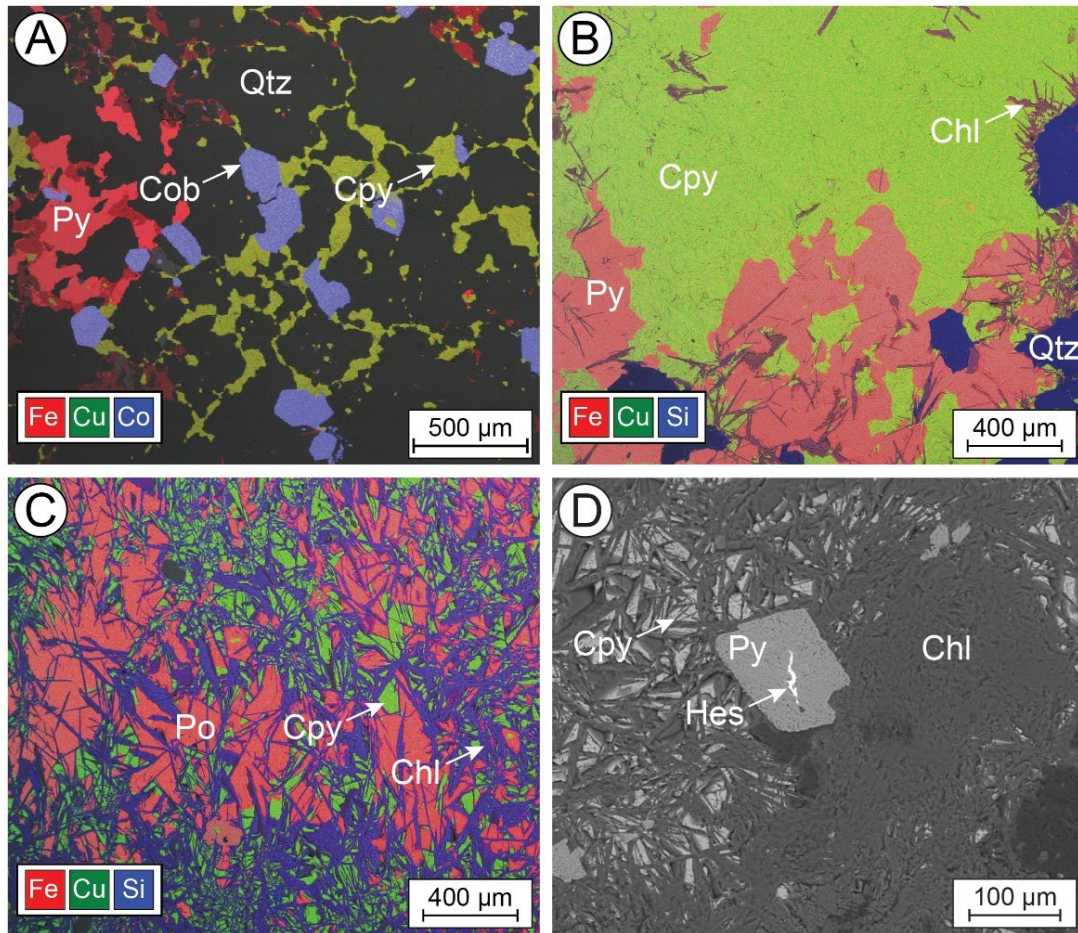
1131 The hypogene structurally-hosted mineralization at Carlow Castle comprises two sulfide-  
1132 dominated mineral assemblages (Table 2.1): (i) a pyrite-chalcopyrite assemblage (Assemblage  
1133 One hereafter; Figure 2.4a; Appendix 2.3a, b), and (ii) a cobaltite-gold-chalcocite assemblage  
1134 (Assemblage Two hereafter; Figure 2.4b, Appendix 2.3c). However, it is unclear if the presence  
1135 of these two mineral assemblages does not appear to reflect any systematic spatial zonation in  
1136 the mineralogy and geochemistry of the ore body at Carlow Castle. Assemblage One occurs  
1137 within the main shear zone at Carlow Castle and is the most volumetrically dominant  
1138 structurally-hosted sulfide mineralization at Carlow Castle, whereas Assemblage Two occurs  
1139 primarily in comparably minor vein networks.



1140  
 1141 Figure 2.4 - A. Photograph mosaic and XRF map (Sample CC010\_137) of chalcopyrite  
 1142 (green) and pyrite-rich mineralization with minor cobaltite (blue). Mineralization occurs  
 1143 as veins through pervasively chloritized and sheared host basalt (Assemblage One from  
 1144 Carlow Castle) B. Photograph mosaic and XRF map (Sample CC007\_41) of chalcocite  
 1145 (green) and cobaltite-rich (blue) ore. Mineralization occurs as veinlets through a larger  
 1146 quartz-carbonate vein (Assemblage Two from Carlow Castle). Note Cu is highlighted in  
 1147 green and Co in blue.

1148 Assemblage One is characterized by intergrown chalcopyrite, pyrite, and minor pyrrhotite at  
 1149 both microscopic and macroscopic scales. This assemblage commonly occurs as a series of  
 1150 sulfide-quartz-carbonate veins infilling a breccia network and through sheared bands of  
 1151 sulfides within the heavily chloritized host rock (Figure 2.4a). Cobaltite is observed as  
 1152 dispersed finer aggregates throughout some of the samples rather than distinct cobaltite-rich

1153 veinlets as in Assemblage Two, though cobaltite does not always occur within Assemblage  
1154 One. Where cobaltite is present in quartz-carbonate veins, pyrite and chalcopyrite tend to be  
1155 anhedral and form stringy veinlet networks infilling space between quartz (Figure 2.5a). These  
1156 veinlet networks of chalcopyrite and pyrite are occasionally associated with and crosscut  
1157 paragenetically earlier euhedral to subhedral grains of cobaltite (Figure 2.5a). Assemblage One  
1158 also occurs in texturally distinct thick (>20 cm) and particularly sulfide-rich veins where the  
1159 volume of gangue quartz and carbonate is limited (<10 vol. %) and the host rock has been  
1160 pervasively replaced by sulfides, forming a massive sulfide texture. These veins almost  
1161 exclusively comprise massive sulfide in the form of coarse-grained, intergrown anhedral pyrite,  
1162 pyrrhotite and chalcopyrite with accessory needles of chlorite and actinolite, along with epidote  
1163 and subhedral quartz (Figure 2.5b). Within these coarse-grained massive sulfide samples  
1164 (Figure 2.5b) cobaltite is less common in comparison to the more quartz-rich veinlet textured  
1165 samples (Figure 2.5a). Occasionally, pyrrhotite and chalcopyrite are observed to be extensively  
1166 and intimately intergrown with volumetrically abundant chlorite from the altered host rock  
1167 (Figure 2.5c).



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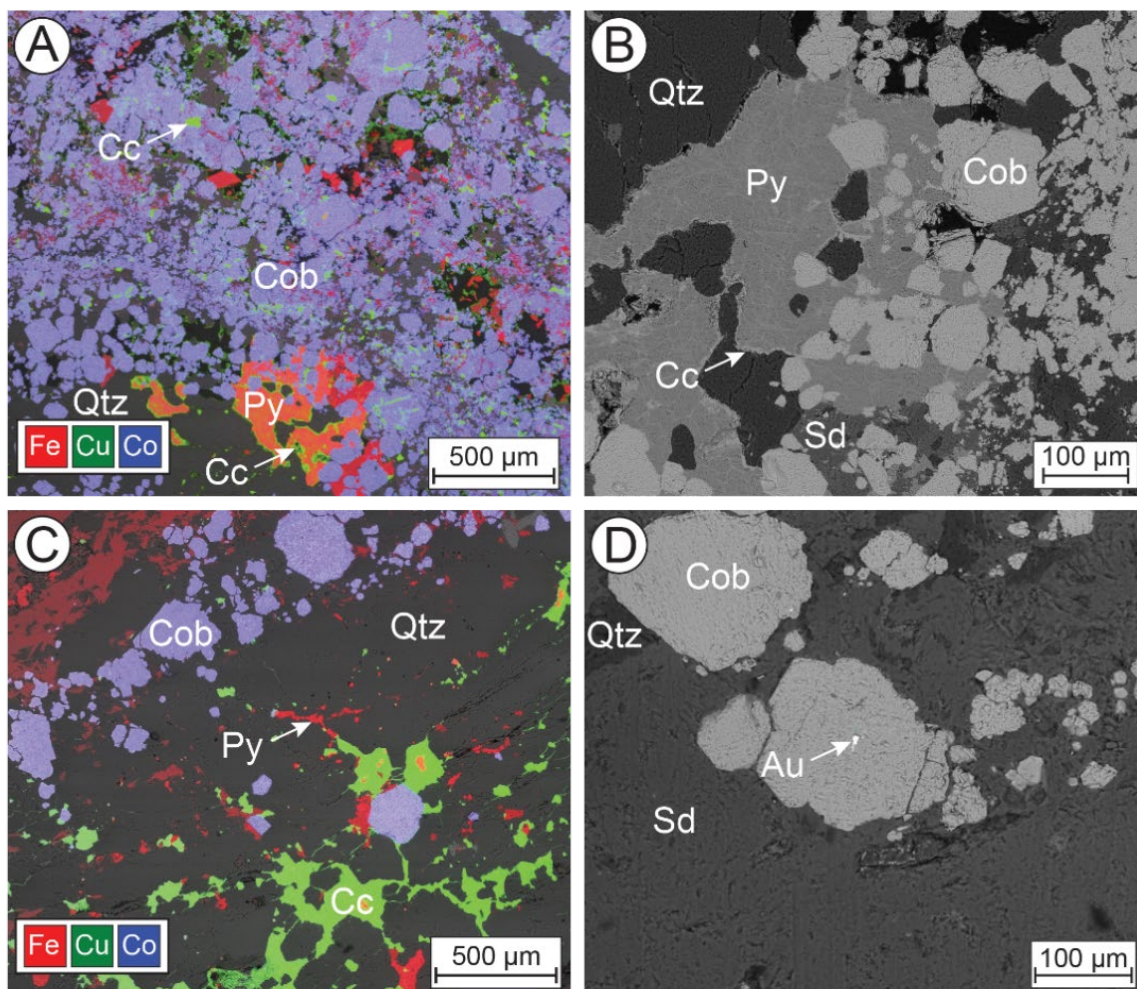
1182

1183

Figure 2.5 - SEM photomicrographs of mineralization from Assemblage One. A. Element map (Sample CC003\_61) of Fe (red), Cu (green), and Co (blue) showing the occurrence of coarse subhedral to euhedral cobaltite (cob) with veinlets of anhedral chalcopyrite (cpy) and pyrite (py) within a larger quartz (qtz) cement. B. Element map (Sample CC007\_102) of Fe (red), Cu (green), and Si (blue) showing the occurrence of coarse intergrown chalcopyrite and pyrite within massive samples, with minor quartz and coarse chlorite (chl). C. Element map (Sample CC007\_102) of Fe (red), Cu (green), and Si (blue) showing syngenetic pyrrhotite (po) and chalcopyrite intergrown with coarse chlorite. D. Secondary electron image (Sample CC007\_102) of a hessite (hes) inclusion within pyrite.

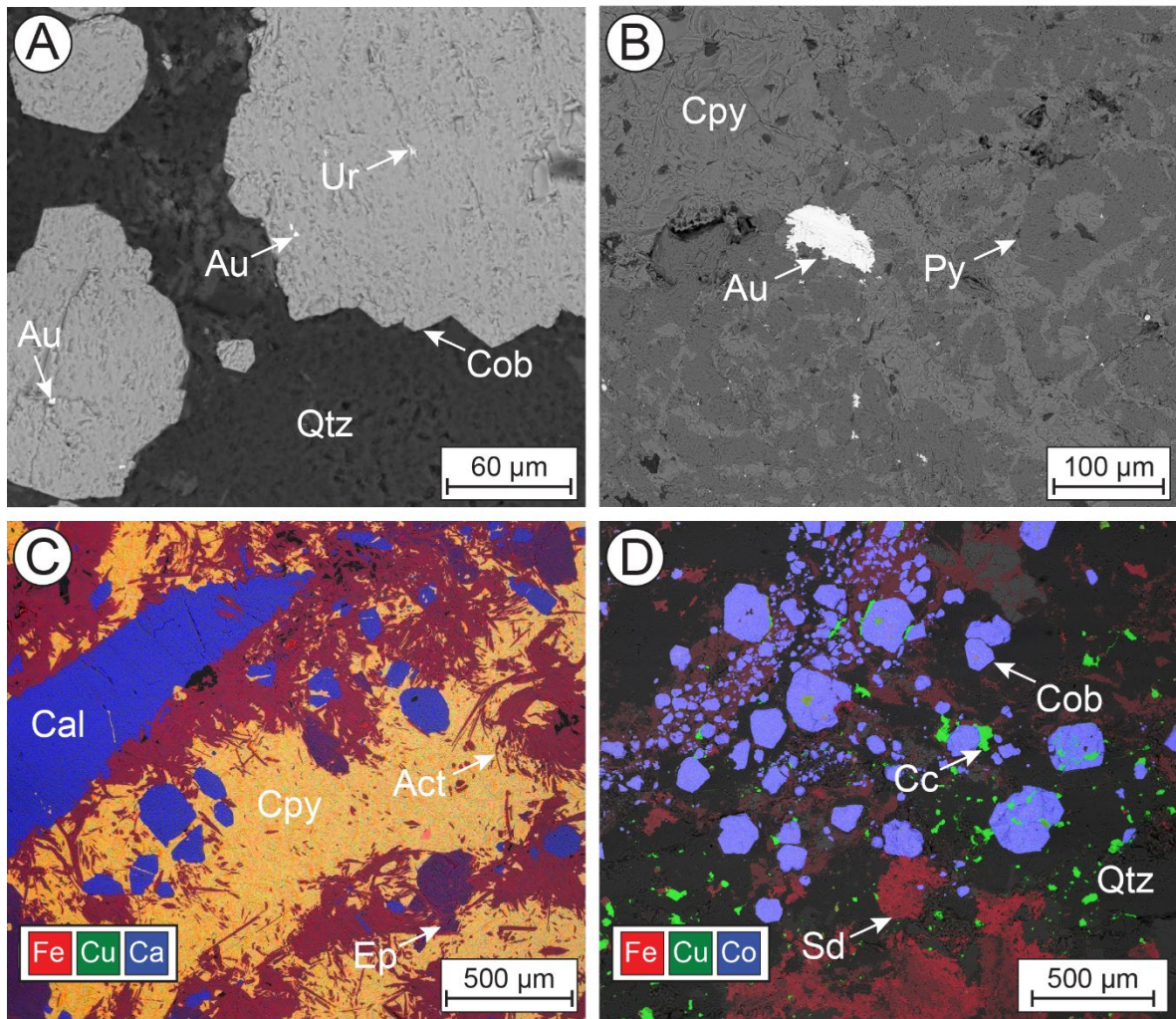
The dominant sulfide minerals within Assemblage Two are cobaltite and chalcocite. Within Assemblage Two cobaltite tends to occur as a series of discordant veinlets composed of aggregates of cobaltite grains (Figure 2.4b). However, chalcocite generally occurs as an interconnected network of thin veinlets infilling a quartz-carbonate breccia (Figure 2.4b). Within these veinlets cobaltite occurs as coarse aggregates of individual euhedral-subhedral grains cemented by quartz and siderite/calcite (Figure 2.6a) with comparably minor pyrite

1184 (<10%), which is occasionally observed enveloping grains of cobaltite (Figure 2.6a, b, c).  
 1185 Primary chalcocite often crosscuts these cobaltite-rich veinlets; here chalcocite envelopes  
 1186 cobaltite grains and forms an interconnected network through the quartz-carbonate cement.  
 1187 Additionally, this chalcocite is commonly observed to partially replace pyrite; forming around  
 1188 the edges of pyrite grains (Figure 2.6b).



1189  
 1190 Figure 2.6 - SEM photomicrographs of mineralization from Assemblage Two. A. Element  
 1191 map (Sample CC007\_41) of Fe (red), Cu (green), and Co (blue) within a quartz vein  
 1192 dominated by a cobaltite-rich aggregate with minor pyrite and paragenetically late  
 1193 chalcocite (cc). B. Secondary electron image (Sample CC007\_41) of a cobaltite-rich  
 1194 quartz-siderite (sd) vein with late pyrite being partially replaced by chalcocite. C. Element  
 1195 map (CC009\_48) of Fe (red), Cu (green), and Co (blue) from within a quartz vein hosting  
 1196 coarse blocky cobaltite and stringy veinlets of chalcocite and pyrite. D. Secondary  
 1197 electron image (Sample CC009\_48) of cobaltite within a quartz-siderite vein showing the  
 1198 common occurrence of Au as inclusions within cobaltite.

1199 Across both Assemblage One and Two, Au generally occurs as electrum (Au,Ag), typically  
1200 >75% Au. Volumetrically, gold occurs in trace abundances as micron scale inclusions within  
1201 the sulfide minerals of Assemblage One and Two (Figures 2.6d and 2.7a, b). However, in  
1202 general, Au mineralization tends to be more abundant in samples from Assemblage Two and  
1203 is comparably rare in Assemblage One. Across Carlow Castle there appears to be a clear genetic  
1204 relationship between gold and cobaltite mineralization, as within these samples electrum most  
1205 commonly occurs as micron-scale inclusions within cobaltite grains (Figures 2.6d and 2.7a).  
1206 This may account for the comparably greater abundance of electrum within Assemblage Two,  
1207 which hosts cobaltite in much greater abundances than Assemblage One. In addition to  
1208 electrum, trace inclusions of various selenide and telluride minerals (BiTe, PbSe, AgTe) and  
1209 uraninite (UO<sub>2</sub>) are observed to occur as inclusions within cobaltite in Assemblage Two (Figure  
1210 2.7a). In Assemblage One, where cobaltite is less abundant, these tellurides and selenides occur  
1211 more commonly as inclusions within pyrite (Figure 2.5d).



1212

1213 Figure 2.7 – SEM photomicrographs from Assemblages One and Two. A. Secondary  
 1214 electron image of cobaltite containing micron-scale inclusions of uraninite (Ur) and  
 1215 electrum from ore mineral Assemblage Two (Sample CC009\_48). B. Secondary electron  
 1216 image showing scattered electrum mineralisation (bright phases) within a matrix of coarse  
 1217 grained pyrite and chalcopyrite from Assemblage One (Sample CC003\_61). C. Element  
 1218 map (Sample CC012\_67) of Fe (red), Cu (green), and Ca (blue) of paragenetically early  
 1219 calcite (cal), epidote (ep), and actinolite (act) alteration phases entrained in massive  
 1220 chalcopyrite mineralisation. D. Element map (Sample CC009\_48) of Fe (red), Cu (green),  
 1221 and Co (blue) showing scattered euhedral cobaltite grains within a quartz-siderite vein  
 1222 with minor chalcocite mineralisation. Note that chalcocite appears to cross-cut or envelop  
 1223 some cobaltite grains.

1224

1225 Table 2.1 - Summary of ore mineralogy of Assemblages One and Two.

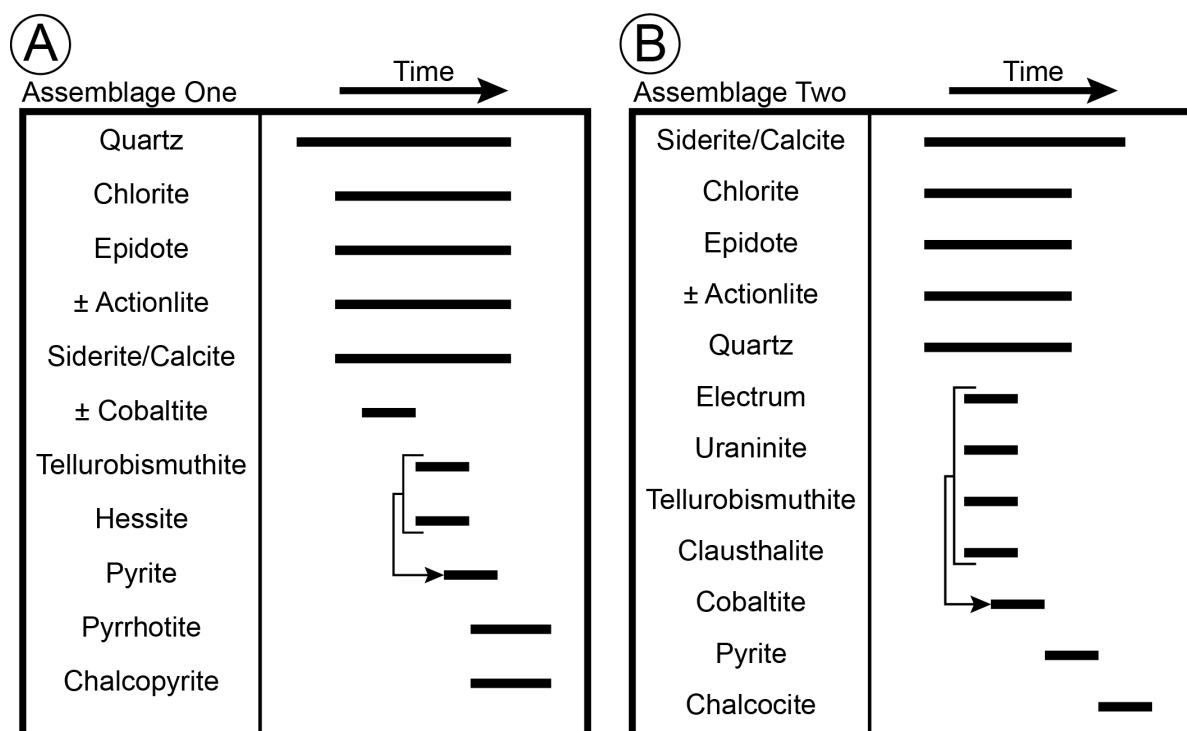
|                  | <b>Mineral Assemblage One</b>  | <b>Mineral Assemblage Two</b>  |
|------------------|--|--|
| <b>Cu phases</b> | Chalcopyrite, ± chalcocite   | Chalcocite, ± chalcopyrite   |
| <b>Co phases</b> | ± Cobaltite  | Cobaltite  |
| <b>Au phases</b> | ± Electrum   | Electrum   |
| <b>Other</b>     | Pyrrhotite, pyrite, ± tellurobismuthite(?) (BiTe), ± hessite(?) (AgTe) | ± Pyrite, ± uraninite (UO <sub>2</sub> ), ± tellurobismuthite(?), ± clausthalite(?) (PbSe) |

## 1226 2.7 Discussion

### 1227 2.7.1 Paragenetic relationships and relative geochronology

1228 The paragenetic relationships and relative geochronology of ore minerals within Assemblage  
1229 One and Assemblage Two are summarized in Figure 2.8. Quartz, epidote, chlorite, and  
1230 siderite/calcite are common paragenetically early phases across both Assemblage One and Two  
1231 (Figure 2.7c). Within these assemblages quartz and siderite/calcite represent the early  
1232 cementing minerals that fill most of the system of mineralized veins (e.g., Figures 2.5a and  
1233 2.6c, d). Within both assemblages chlorite appears to be roughly synchronous with, or slightly  
1234 predates these minerals and as such it is most likely a product of the extensive hydrothermal  
1235 alteration of ferromagnesian minerals within the mafic-ultramafic volcanics and shales of the  
1236 Nickol River and Ruth Well Formations, coinciding with the flow of the hydrothermal fluid  
1237 that produced the Cu-Co-Au mineralization. Chlorite is an extremely common alteration  
1238 mineral in hydrothermal gold systems, particularly in those hosted in mafic lithologies that are  
1239 rich in ferromagnesian minerals (De Caritat, Hutcheon, & Walshe, 1993; Saunders et al., 2014).  
1240 Extensive paragenetically early quartz-carbonate veining is similarly common in hydrothermal  
1241 gold deposits (Kerrich et al., 2000; Zhu, Fang, & Tan, 2011).





1242  
 1243 Figure 2.8 - A. Generalized paragenetic sequence of ore minerals for Assemblage One.  
 1244 B. Generalized paragenetic sequence of ore minerals for Assemblage Two. Note that  
 1245 arrows drawn between minerals denote the occurrence of one mineral exclusively as  
 1246 inclusions within another (e.g., gold occurs exclusively as inclusions within cobaltite);  
 1247 suggesting a genetic relationship.

1248 In Assemblage One, the early phase of quartz-carbonate-chlorite mineralization was followed  
 1249 by a phase of limited cobaltite mineralization, with a subsequent phase of telluride  
 1250 mineralization hosted in pyrite grains (Figure 2.5d). However, in Assemblage Two telluride  
 1251 and selenide mineralization was synchronous with electrum and uraninite mineralization, all of  
 1252 which are associated with Co mineralisation (Figure 2.7a). This is interpreted from the  
 1253 occurrence of electrum, uraninite, and these telluride/selenides almost exclusively as inclusions  
 1254 within cobaltite in this assemblage (Figures 2.6c and 2.7a). The co-occurrence of these phases  
 1255 implies they are genetically related, associated with common transport and co-precipitation. A  
 1256 similar affinity between Au and cobaltite has previously been noted in hydrothermal Au-Cu  
 1257 systems (Le Vaillant et al., 2018; Soloviev, Kryazhev, & Dvurechenskaya, 2013). This was  
 1258 followed by the growth of pyrite around pre-existing cobaltite grains (Figure 2.6b). In  
 1259 Assemblage One this phase of pyrite precipitation was followed by the co-precipitation of  
 1260 chalcopyrite and pyrrhotite within an interlocking framework of chlorite (Figure 2.5c). This  
 1261 textural relationship indicates that either the synchronous precipitation of chalcopyrite and

1262 pyrrhotite infilled this pre-existing chlorite lattice or that the chlorite, pyrrhotite, and  
1263 chalcopyrite were approximately synchronous. In Assemblage Two the pyrite precipitation  
1264 phase was instead followed by a final phase of chalcocite mineralization, which is observed to  
1265 have partially replaced the preceding pyrite and forms a thin layer around or cross-cuts some  
1266 cobaltite grains (Figures 2.6a, b, c and 2.7d).

1267 Across both Assemblage One and Assemblage Two, the paragenetic sequence of ore minerals  
1268 in Figure 2.8 appears to reflect a progressive sequence of Au, Co, and then Cu mineralization.  
1269 Within Assemblage One this is reflected in the sequential mineralization of cobaltite (Co), then  
1270 chalcopyrite (Cu). In Assemblage Two this sequence is represented by the mineralization of  
1271 electrum (Au), then cobaltite, and finally chalcocite (Cu). However, as Assemblage One and  
1272 Two are typically spatially distinct, it is difficult to ascertain the nature of their relative timing.  
1273 As such, whether the existence of these two distinct assemblages represent the progressive  
1274 geochemical evolution of a single ore fluid arising from a single mineralizing event or two  
1275 distinct phases of mineralization with two genetically distinct ore fluids is as yet unconstrained.  
1276 However, the significant difference in the prevalence of cobaltite and predominant stable  
1277 copper sulfide phase between these two assemblages suggest potential physicochemical  
1278 differences in the hydrothermal conditions under which these phases formed (Crerar & Barnes,  
1279 1976; Jansson & Liu, 2020; Liu et al., 2011; Zhao, Brugger, Ngothai, & Pring, 2014).  
1280 Additionally, the rarity of samples containing both ore mineral assemblages may suggest that  
1281 the existence of these distinct assemblages reflects a spatial geochemical and mineralogical  
1282 zonation within the Carlow Castle ore zone.

### 1283 2.7.2 Preliminary constraints on Carlow Castle ore fluid

1284 The mineralogical analyses conducted in this study provide several key insights into the genesis  
1285 of Carlow Castle's Cu-Co-Au mineralization; particularly pertaining to the nature of its ore  
1286 fluid, metal mobilization processes, and metal concentration processes. Firstly, the pervasive  
1287 carbonate veining throughout the ore zone and the surrounding tectonized zone precludes the  
1288 possibility that Carlow Castle's ore fluid was strongly acidic, due to the well-established  
1289 reactivity of carbonate minerals with acidic fluids (Golubev, Benezeth, Schott, Dandurand, &  
1290 Castillo, 2009; Mikucki & Ridley, 1993). The prevalence of carbonate veining hosting Cu-Co-  
1291 Au mineralisation at Carlow Castle suggests that dissolution of metals under acidic (pH <4)  
1292 conditions was unlikely to be a critical metal transport process within the ore fluid, unlike in  
1293 many porphyry Cu or high-sulfidation epithermal Au deposits for example (Keith, Smith,

1294 Jenkin, Holwell, & Dye, 2018; Seedorff et al., 2005). The abundance of this carbonate veining  
1295 also suggests that the ore fluid contained a significant amount of dissolved carbonate (Chi, Liu,  
1296 & Dube, 2009; Kerrich & Fyfe, 1981). Additionally, the occurrence of Cu, Co, Au  
1297 mineralization at Carlow Castle as, or in association with, base metal sulfides/sulfarsenides,  
1298 tellurides/selenides, uraninite, and siderite is clear evidence of the precipitation of these metals  
1299 under reducing conditions due to the well-documented influence of redox conditions on the  
1300 stability of these phases (Grundler et al., 2013; Keith et al., 2018; Plasil, 2014; Postma, 1983).  
1301 This may indicate that a shift in redox conditions within the ore fluid towards reducing  
1302 conditions at the site of mineralization may have been key to inducing Cu, Co, Au precipitation,  
1303 particularly if these metals were transported in  $\text{Cl}^-$  complexes within an oxidized ore fluid  
1304 (Heinrich & Candela, 2014; Pokrovski, Akinfie, Borisova, Zotov, & Kouzmanov, 2014;  
1305 Saunders et al., 2014). It is possible that the interaction of such an oxidized ore fluid with  
1306 reduced mineral phases within the black shales and mafic volcanic rocks of the Nickol River  
1307 and Ruth Well Formations at the site of mineralization may have driven the evolution of this  
1308 ore fluid toward reducing conditions. Therefore, it is feasible that the Cu, Co, Au mineralization  
1309 at Carlow Castle precipitated from a brine that was near-neutral in pH, carbonate-bearing, and  
1310 reducing at the point of mineralization.

1311 The co-occurrence of Cu, Co, and Au-bearing minerals within Carlow Castle, with no evidence  
1312 of successive overprinting phases of mineralization, indicates that the ore fluid must have been  
1313 capable of simultaneously transporting these metals. This is particularly the case for Co and  
1314 Au, which have already been demonstrated to display a clear genetic relationship in their  
1315 mineralogical occurrence. While generally  $\text{Cl}^-$ -complexation is the dominant process for metal  
1316 transport in most hydrothermal Cu-Co systems (Brown, 2014; Hitzman et al., 2010; Rose,  
1317 1989), it is generally the less common mechanism for Au transportation in hydrothermal  
1318 solutions  $<500^\circ\text{C}$  (Liu et al., 2014b; Zhu et al., 2011). Instead complexation with  $\text{HS}^-$  is  
1319 typically the preferred process for Au mobilization in most hydrothermal Au systems, though  
1320 Au may still be transported through  $\text{Cl}^-$ - or  $\text{OH}^-$ -complexation (Saunders et al., 2014; Zhu et  
1321 al., 2011). However, the presence of significant volumes of dissolved  $\text{HS}^-$  in an ore fluid would  
1322 generally preclude the transport of Cu and Co in the volumes required to form Carlow Castle,  
1323 unless under strongly acidic ( $\text{pH} < 4$ ) conditions (Gaboury, 2019; Saunders et al., 2014).  
1324 Additionally, Au dissolution via complexation with  $\text{Cl}^-$  is most effectively facilitated under  
1325 oxidizing conditions; which are similarly conducive to Cu and Co transport (Pokrovski et al.,

1326 2014; Rose, 1989). As such, within this ore fluid  $\text{Cl}^-$  or  $\text{OH}^-$ -complexation of these metals is  
1327 the most probable process for metal mobilization. While it is possible that the metal  
1328 enrichments may have been carried in multiple stages, with Au initially transported in a  $\text{HS}^-$   
1329 rich fluid and a second distinct  $\text{Cl}^-$ -rich fluid carrying Cu and Co, this interpretation is not  
1330 supported by the petrographic evidence that indicates that Co and Au appear to have co-  
1331 precipitated. This interpretation implies that the Carlow Castle ore fluid was saline and more  
1332 characteristic of an ore fluid in a basin-hosted mineral system or a magmatically-derived fluid,  
1333 rather than a low-salinity metamorphic fluid that is more typical of lode gold mineralization  
1334 (Heinrich & Candela, 2014; Saunders et al., 2014).

1335 The paragenetic sequences of ore mineralization (Figure 2.8) for Assemblage One and Two  
1336 appear to record evidence of the geochemical evolution of the Carlow Castle ore fluid. Broadly  
1337 within both sequences, but particularly Assemblage Two, there is a clear trend regarding sulfur  
1338 fugacity within the hydrothermal system. This is reflected in the sequence of initial telluride,  
1339 selenide, and alloy precipitation; followed by sulfarsenide precipitation; and finally base metal  
1340 sulfide precipitation. This sequence appears to reflect an increase in sulfur fugacity within the  
1341 system through time; this is suggested by the early occurrence of electrum in association with  
1342 tellurides and/or selenides and a simultaneous lack of sulfides. This is characteristic of many  
1343 sulfur undersaturated hydrothermal Au deposits (Hedenquist, Arribas, & Gonzalez-Urien,  
1344 2000; Saunders et al., 2014; White & Hedenquist, 1995). Following this, the gradual addition  
1345 of sulfur to the system could account for the occurrence of cobaltite; a sulfarsenide. If this  
1346 additional sulfur was supplied as  $\text{HS}^-$  or another reduced sulfur species, this could act as a  
1347 reductant to induce Co precipitation into cobaltite from the ore fluid that carried these metals  
1348 (Heinrich & Candela, 2014; Seward, Williams-Jones, & Migdisov, 2014). Following Co  
1349 precipitation, the continued addition of sulfur may have induced a final phase of Cu  
1350 mineralization hosted solely in base metal sulfides. This sequence also coincides with the  
1351 sequential Au, Co, and Cu mineralization observed at Carlow Castle and most likely reflects  
1352 the different solubility of Co, Cu, and particularly Au in the presence of  $\text{HS}^-$  (Hofmann, 1999;  
1353 Saunders et al., 2014; Saunders & Schoenly, 1995). The progressive increase in sulfur fugacity  
1354 through this sequence and the insolubility of Cu and Co in most  $\text{HS}^-$ -rich fluids suggests that  
1355 the  $\text{HS}^-$  necessary for sulfide mineralization may have been supplied to the Cu-Co-Au bearing  
1356 ore fluid at a late stage, proximal to the site of mineralization. As such, it can be tentatively  
1357 interpreted that this  $\text{HS}^-$  was supplied either by the mixing of Carlow Castle's ore fluid with a

1358 distinct HS<sup>-</sup> rich fluid that was also channeled into the Carlow Castle tectonized zone, or  
1359 potentially from the interaction of the ore fluid with a HS<sup>-</sup>-rich wall rock. The gradual addition  
1360 of HS<sup>-</sup> to the ore fluid from either of these sources could have induced the reductive  
1361 precipitation of Au, Co, and Cu and account for the progressive increase in sulfur fugacity  
1362 within the system; recorded by the increase in sulfur through the paragenetic sequence of ore  
1363 minerals (Figure 2.8). The black shales that occur through the Nickol River and Ruth Well  
1364 Formations may have provided this HS<sup>-</sup> if they formed a significant portion of the wall rock at  
1365 Carlow Castle. However, this is difficult to determine due to the extensive hydrothermal  
1366 alteration of the wall rock and sulfur isotope studies are required to fully understand the nature  
1367 of Carlow Castle's sulfur source.

### 1368 2.7.3 Probable age and fluid pathway

1369 As there are presently no geochronological constraints on Carlow Castle's formation, fitting its  
1370 genesis into the established model for the tectonic evolution of the northwest Pilbara Craton is  
1371 challenging. However, the nature of its occurrence within the stratigraphy of the Pilbara Craton  
1372 and the insights about the nature of its mineralization provided by this study may still provide  
1373 some constraints on the mineral system. Presently, the best indication of Carlow Castle's age  
1374 is provided by its occurrence within higher-order faults that are interpreted here to be a product  
1375 of deformation along the Regal Thrust. The Regal Thrust formed during the convergent  
1376 tectonic regime that preceded the Prinsep Orogeny (~3070 Ma); wherein the Regal Terrane  
1377 was obducted onto the Karratha Terrane (Hickman, 2016). This provides a maximum possible  
1378 age of 3070 Ma for Carlow Castle's mineralization. However, this provides little insight into  
1379 its minimum possible age. It has previously been recognized that the Carlow Castle area has  
1380 been affected by multiple stages of deformation, with at least one stage of major deformation  
1381 post-dating the Prinsep Orogeny and potentially coinciding with the North Pilbara Orogeny  
1382 (Hickman, 2016; Kiyokawa et al., 2002). As such, several different tectonic events during the  
1383 formation of the Pilbara Craton could have been associated with the formation of the Carlow  
1384 Castle deposit and its higher-order host structures within the Regal Thrust. Additionally, there  
1385 is also significant potential for deformation of the deposit following ore formation. With this  
1386 considered, the exceptional tectonic stability of the northwest Pilbara Craton subsequent to the  
1387 initiation of the Fortescue Rift and the formation of the Fortescue Basin ~50 km to the south  
1388 around 2775 Ma (Hickman, 2012, 2016) makes a major post-Archean ore-forming event  
1389 unlikely. Nonetheless, these constraints leave a range of ~295 Ma over which Carlow Castle

1390 may have formed during the evolution of the Pilbara Craton; with several distinct potential ore-  
1391 forming tectonic events. While the timing of the ore-forming event that produced the distinctive  
1392 Cu-Co-Au mineralization remains ambiguous, the Regal Thrust was unambiguously a key  
1393 component of the mineral system. Regionally, the Regal Thrust would have likely acted as a  
1394 major conduit to allow for the flow of Cu-Co-Au rich ore fluids. These Cu-Co-Au rich fluids  
1395 could have then been focused into and trapped within complex higher-order structures through  
1396 the brecciated and sheared zone at Carlow Castle, leading to metal precipitation.

#### 1397 2.7.4 Comparisons to other deposits

1398 Due to its distinctive Cu-Co-Au endowment, the Carlow Castle deposit bears a similar ore  
1399 element enrichment to Co-rich sediment-hosted stratiform Cu (SSC) deposits, such as those in  
1400 the Central African Copper Belt (Saintilan, Selby, Creaser, & Dewaele, 2017), which are also  
1401 typically Cu-Co sulphide-dominated (Hitzman et al., 2005). However, there are several critical  
1402 characteristics that distinguish Carlow Castle from those deposits; most importantly Carlow  
1403 Castle's geological setting, the age of its host terrane, key aspects of its genesis, and the nature  
1404 of its mineralization. First, Carlow Castle's occurrence within an accretionary greenstone belt  
1405 dominated by mafic volcanic rocks and minor shales distinguishes it from Co-rich SSC  
1406 deposits, which are most commonly hosted in rift-basins associated with thick sequences of  
1407 immature first-cycle clastic sediment (Brown, 2014). Additionally, the occurrence of Carlow  
1408 Castle within a Mesoarchean host terrane, and the probable Archean age of mineralization,  
1409 contrasts strongly with the occurrence of SSC deposits overwhelmingly in host sequences of  
1410 maximum Proterozoic age (Hitzman et al., 2010). This speaks to significant differences in the  
1411 genetic processes that produced these deposits. In SSC deposits the oxidized ore fluid is  
1412 commonly interpreted to derive from oxidized saline basinal brines, while the most common  
1413 metal source is from immature metalliferous clastic material within the thick red bed sequences  
1414 that commonly underlie these deposits (Brown, 2014; Hitzman et al., 2005; Kirkham, 1989).  
1415 These are almost certainly not the sources of ore fluids or metals within the Carlow Castle  
1416 mineral system due to the critical control that the oxidation state of the Earth's atmosphere  
1417 imparts on the formation of these, and the lack of significant oxygenation of the Earth's  
1418 atmosphere during the Archean (Lyons, Reinhard, & Planavsky, 2014). Finally, the occurrence  
1419 of this mineralization on a deposit scale further distinguishes it from an SSC deposit. At Carlow  
1420 Castle the dominantly structurally-controlled nature of the mineralization through a heavily  
1421 tectonized zone proximal to a regionally significant thrust fault, hosted in quartz-carbonate

1422 veins, with extensive chloritic hydrothermal wall-rock alteration is distinct from mineralization  
1423 in typical SSC deposit. In contrast, in SSC deposits mineralization typically occurs as a laterally  
1424 extensive stratiform ore body along the redox boundary between an underlying oxidized red  
1425 bed and an overlying organic-rich reduced bed (Lyons et al., 2014). Reflecting these  
1426 differences, Carlow Castle appears to represent an unconventional and genetically distinct class  
1427 of hydrothermal Cu-Co-Au.

1428 Although Carlow Castle does not appear to resemble a Co-rich SSC deposit beyond its  
1429 distinctive Cu-Co-Au metal endowment, it does bear some resemblance to some previously  
1430 studied unconventional Co deposits. In the Bou Azzer district, Morocco Carboniferous Co-  
1431 mineralization occurs as Co-Ni-Fe arsenide dominated orebodies hosted in quartz-carbonate  
1432 veins through a series of tectonic structures within the Neoproterozoic Bou Azzer ophiolite  
1433 complex (Le Blanc & Billaud, 1982; Oberthür et al., 2009). Additionally, Cu-Fe sulfides,  
1434 sulfosalts, and Au-Ag alloys are also observed in variable amounts in association with this Co  
1435 mineralization (Ahmed et al., 2009). This ophiolite complex is composed predominantly of a  
1436 mix of serpentinized peridotites, submarine mafic volcanics, and mafic-ultramafic cumulates  
1437 (Ahmed et al., 2009). The current genetic model for the Bou Azzer Co district posits that it was  
1438 a product of the migration of a saline brine related to felsic magmatism during the Late  
1439 Paleozoic Hercynian/Variscan orogeny (Ahmed et al., 2009; En-Naciri, Barbanson, & Touray,  
1440 1997). Within this model it is proposed that this magmatic fluid leached ore metals from the  
1441 serpentinites of the Bou Azzer ophiolite complex and precipitated these metals under  
1442 increasing pH, sulfur fugacity, and decreasing O<sub>2</sub> fugacity in response to mixing with a cool  
1443 meteoric fluid (Ahmed et al., 2009; En-Naciri et al., 1997; Essaraj, Boiron, Cathelineau, Banks,  
1444 & Benharref, 2005). In some respects, the preliminary interpretations made about the Carlow  
1445 Castle ore fluid in this study resemble those of Bou Azzer. Additionally, the paragenetic  
1446 sequence at Bou Azzer is comparable to Carlow Castle's; with a sequence reflecting increasing  
1447 sulfur fugacity and sequential Au-Co/Ni/As-Cu mineralization (Ahmed et al., 2009). Finally,  
1448 it is interesting to note that rocks of the Bou Azzer ophiolite complex are most probably of  
1449 comparable bulk composition to the dominantly mafic-ultramafic volcanics of the Ruth Well  
1450 Formation, within which Carlow Castle is hosted. If the ultramafic rocks of this ophiolite  
1451 complex were a key source of ore metals for the formation of the Bou Azzer, then this may  
1452 likewise apply to the genesis of Carlow Castle.

1453 Additional deposits that share some geological characteristics with Carlow Castle and Bou  
1454 Azzer include the five element (Ni-Co-As-Ag-Bi) deposits of Cobalt, Ontario, Canada and the  
1455 Co-Cu-Au deposits of the Idaho Co belt, USA. Most notable among these shared characteristics  
1456 are their distinctly Co-dominated ore mineralogies, vein-hosted sulfide and sulfarsenide  
1457 mineralization, and occurrence within convergent metamorphosed terranes (Bookstrom et al.,  
1458 2016; Kissin, 1992). Therefore, Carlow Castle may represent an addition to the distinct and  
1459 relatively rare class of hydrothermal and volcanogenic polymetallic deposits that Ahmed et al.  
1460 (2009) collectively grouped the above mentioned deposits into. As such, comparisons to the  
1461 established genetic models for these deposits could be useful in constraining the genesis of  
1462 Carlow Castle.

## 1463 2.8 Implications and future study

1464 The primary significance of Carlow Castle stems from its unique Cu-Co-Au metal association  
1465 and distinctive ore mineralogy given its geological setting. The fact that Carlow Castle occurs  
1466 within an Archean greenstone belt and is of probable Meso-Neoproterozoic age is particularly  
1467 notable due to the scarcity of significant hydrothermal Co mineralization outside of basin-  
1468 hosted systems and even greater scarcity of hydrothermal Co mineralization in Archean  
1469 terranes. This is because the majority of the world's Cu-Co deposits are hosted in sedimentary  
1470 basins of Proterozoic or younger age (US Geological Survey, 2018). Therefore, understanding  
1471 the unique confluence of processes necessary to form this distinctive Cu-Co-Au mineralization  
1472 is of particular importance. Understanding the genesis of Carlow Castle has important  
1473 implications for potentially uncovering other Cu-Co deposits in Archean terranes. Following  
1474 this initial study, further analysis will be undertaken to better constrain the genesis of the  
1475 Carlow Castle mineral system. This will include S-isotope analysis, to assist in constraining  
1476 the sulfur source for Carlow Castle's mineralization, and absolute geochronology to account  
1477 for Carlow Castle's formation within the established tectonic evolution of the Pilbara Craton.

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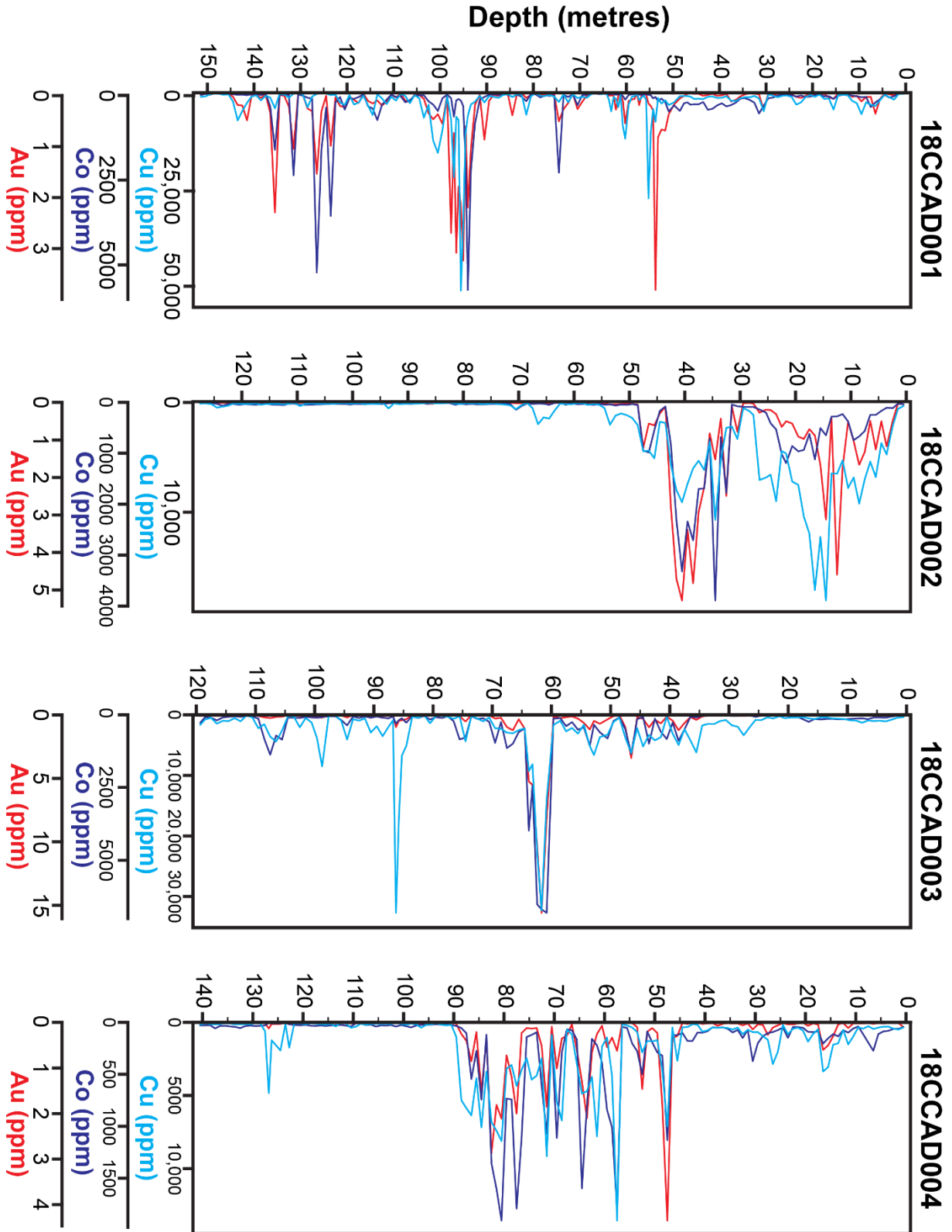
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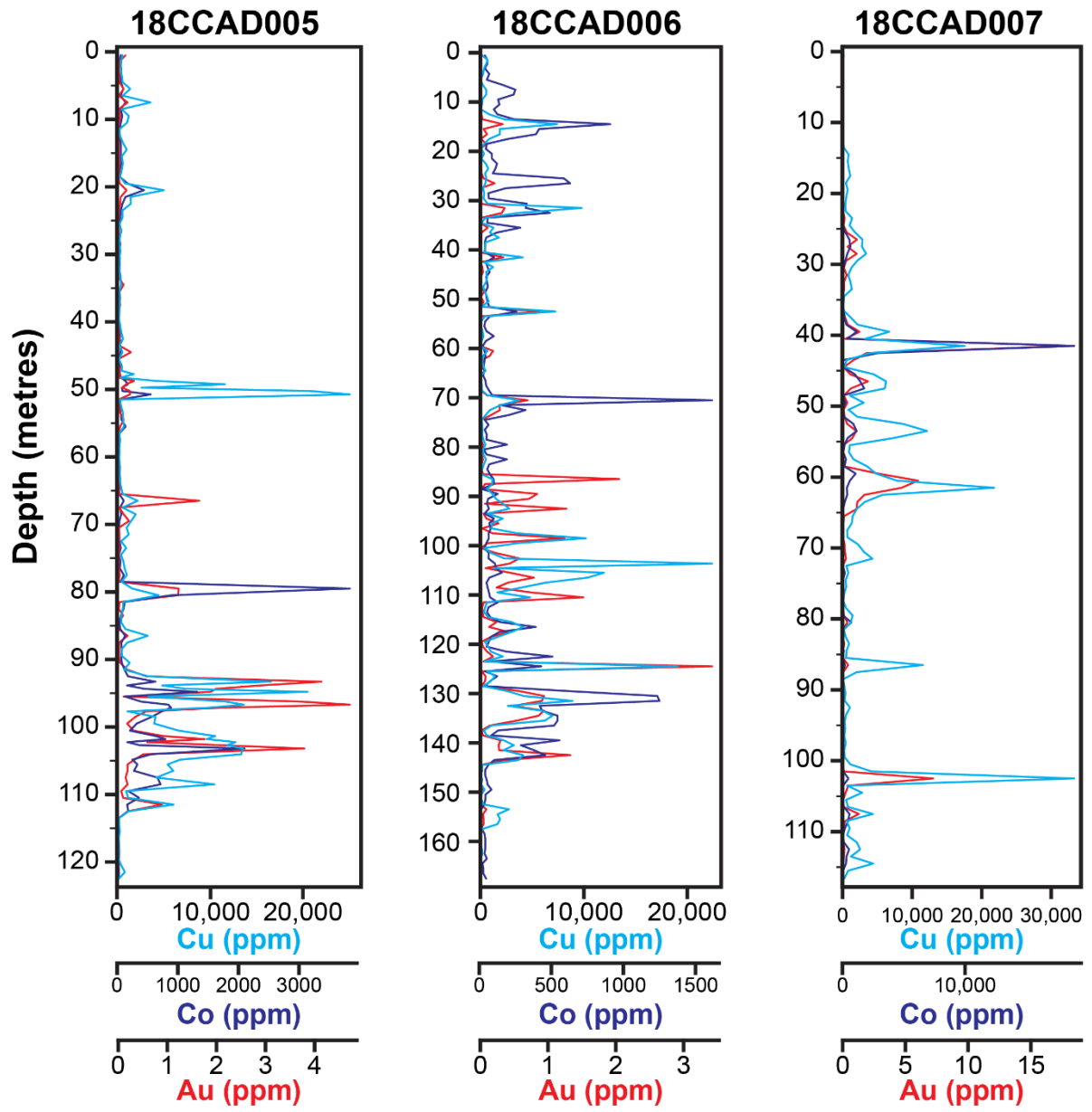
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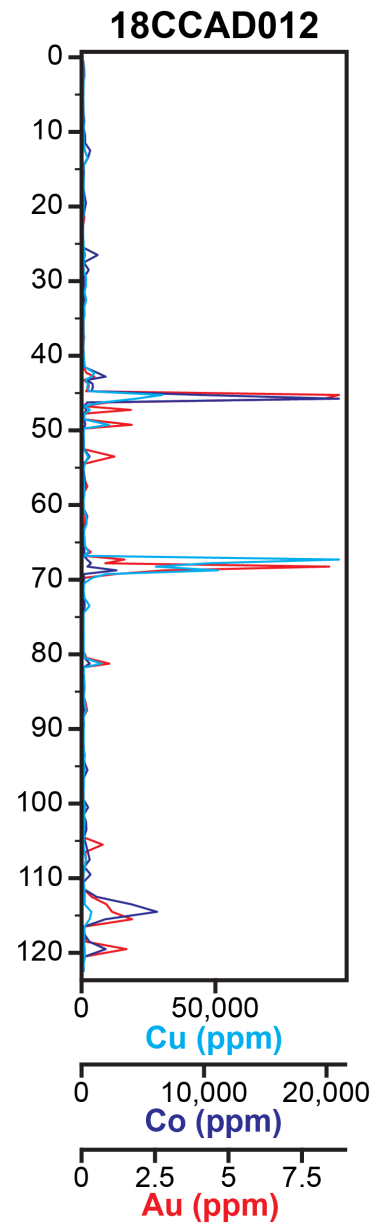
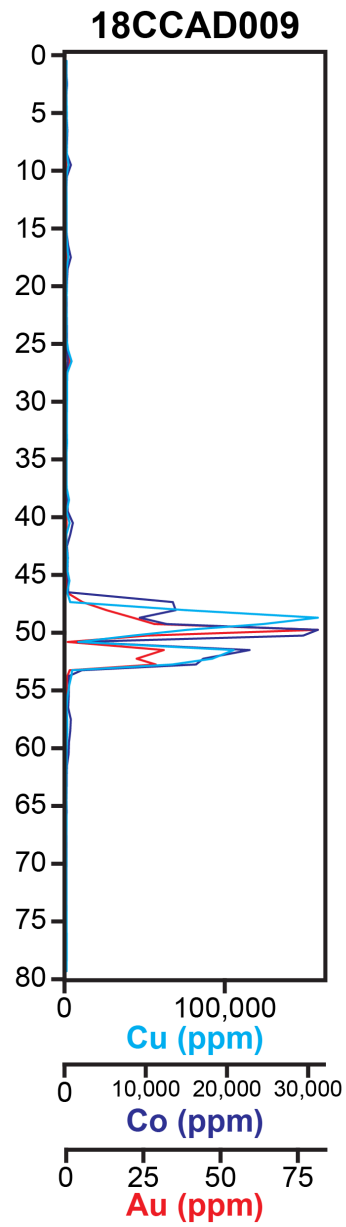
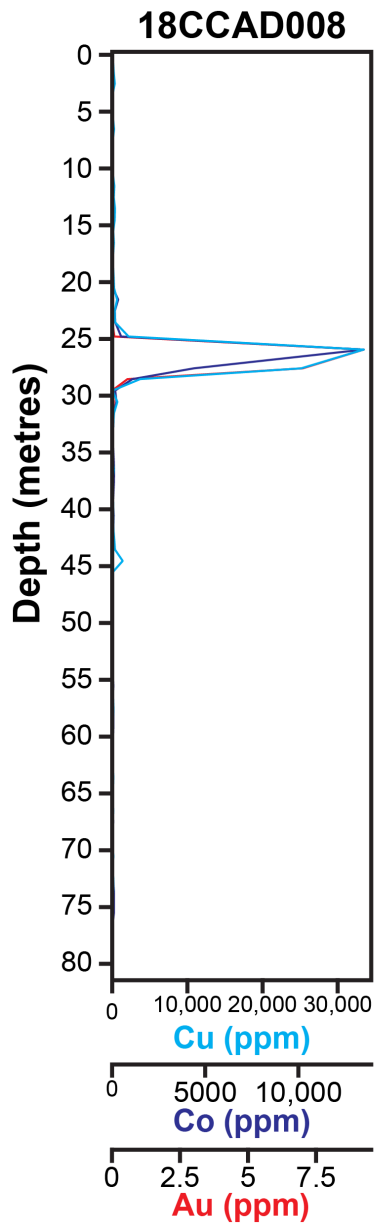
1713 Appendix 2.1 – Spatially grouped down hole geochemistry plots (Cu, Co, Au) of twelve sampled diamond drill holes from the Carlow Castle ore  
1714 body  
1715



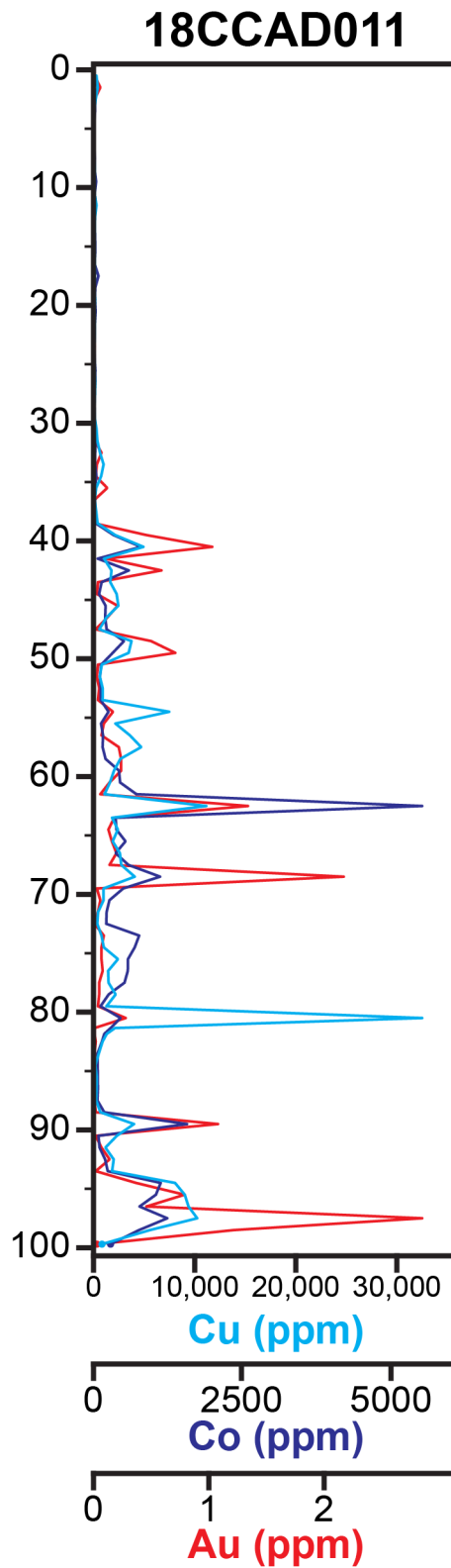
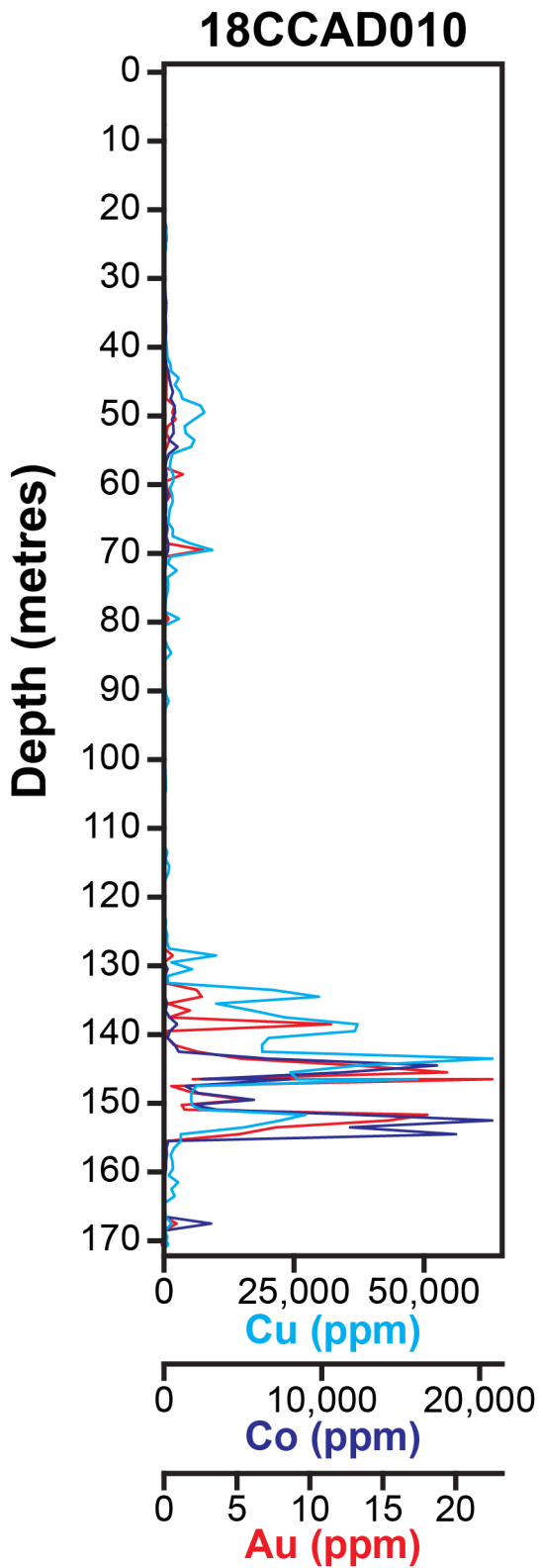
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1719

1720 Note that these plots are produced from publicly available JORC (2012) compliant drilling data

1721 released to the Australian Securities Exchange (Artemis Resources Limited, 2018c).

1722 Appendix 2.2 – List of drill core samples from Carlow Castle Cu-Co-Au  
1723 deposit

1724 Appendix 2.1 is accessible via this [link](#) and provides an .xlsx file containing a full list of drill  
1725 core samples taken from the Carlow Castle Cu-Co-Au deposit ore body

1726

1727 **Appendix 2.3 – High resolution optical microscope thin section scans of**  
1728 **select analysed sample from ore mineral Assemblage One and Two**

1729 Appendix 2.3a is accessible via this [link](#) and provides a .pdf file containing a labelled optical  
1730 thin section scan of sample CC003\_61 from ore mineral Assemblage One.

1731 Appendix 2.3b is accessible via this [link](#) and provides a .pdf file containing a labelled optical  
1732 thin section scan of sample CC012\_67 from ore mineral Assemblage One.

1733 Appendix 2.3c is accessible via this [link](#) and provides a .pdf file containing a labelled optical  
1734 thin section scan of sample CC007\_41b from ore mineral Assemblage Two.

1735

## 1736 Chapter 3

### 1737 Working up an apatite: Enigmatic

### 1738 Mesoarchean hydrothermal Cu-Co-Au

### 1739 mineralization in the Pilbara Craton

1740

1741 This chapter is published in *Economic Geology* as:

1742 *Fox, D., Spinks, S., Barham, M., Kirkland, C. L., Pearce, M. A., Aspandiar, M., Birchall, R.,*  
1743 *Mead, E. (2021). Working up an apatite: Enigmatic Mesoarchean hydrothermal Cu-Co-Au*  
1744 *mineralization in the Pilbara Craton. Economic Geology, 116(7), 1561-1573.*  
1745 *doi:<https://doi.org/10.5382/econgeo.4842>*

1746

#### 1747 3.1 Abstract

1748 Globally, significant examples of hydrothermal Cu-Co mineralization are rare within Archean  
1749 greenstone belts, especially relative to the endowment of these terranes with other world-class  
1750 hydrothermal ore deposits, particularly Au deposits. Using U-Pb geochronology of  
1751 hydrothermal apatite, this study provides the first absolute age constraints on the timing of  
1752 mineralization for the Carlow Castle Cu-Co-Au deposit. Carlow Castle is a complex shear-  
1753 zone hosted veined Cu-Co-Au mineral system situated within the Paleo-Mesoarchean  
1754 Roebourne greenstone belt of the Pilbara Craton of northwestern Western Australia. Although  
1755 U-Pb geochronology of this deposit is challenging due to low levels of radiogenic Pb in syn-  
1756 mineralization apatite, mineralization is best estimated at  $2957 \pm 67$  Ma ( $n = 61$ ). Additionally,  
1757 analysis of alteration phases associated with Carlow Castle's mineralization suggest that it is  
1758 dominated by a propylitic assemblage that is characteristic of alkaline fluid chemistry and peak  
1759 temperatures  $\sim 300^\circ\text{C}$ . Within proximal portions of the northwest Pilbara Craton, the period of  
1760 Carlow Castle's formation constrained here was associated with significant base-metal VMS  
1761 mineralization and magmatic activity related to back-arc rifting. This period of extension and

1762 rifting was the most likely source of Carlow Castle's unique Cu-Co-Au mineralization. Carlow  
1763 Castle's Mesoarchean mineralization age makes it among the oldest discovered Cu-Co-Au  
1764 deposits on Earth, and relatively unique in the broader context of hydrothermal Cu-Co-Au  
1765 deposits. Globally, hydrothermal Cu-Co mineralization occurs almost exclusively as  
1766 Proterozoic and Phanerozoic stratiform sediment-hosted Cu-Co deposits due to the necessity  
1767 of oxidized basinal ore fluids in their formation. This research therefore has implications for  
1768 exploration for atypical Cu-Co deposits and Cu-Co metallogensis through recognition of  
1769 comparably uncommon Archean hydrothermal Cu-Co-Au ore-forming processes and,  
1770 consequently, the potential for analogous Cu-Co-Au mineralization in other Archean  
1771 greenstone belts.

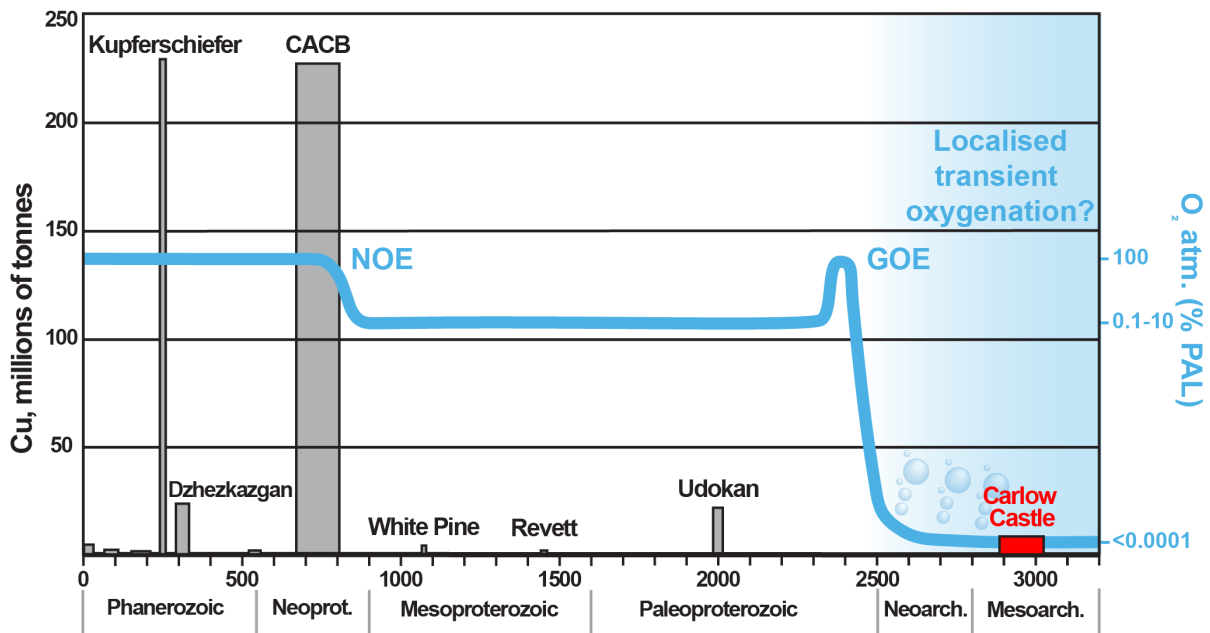
## 1772 3.2 Introduction

1773 The world's most significant sources of cobalt are sediment-hosted copper-cobalt (Cu-Co)  
1774 deposits; these deposits account for ~60% of global Co production and are the only globally  
1775 significant style of hydrothermal Cu-Co mineralization (Mudd et al., 2013; Slack et al., 2017a).  
1776 Temporally, these deposits are constrained to basins of maximum Paleoproterozoic age (Figure  
1777 3.1; Hitzman et al., 2010). This is thought to reflect the critical role of atmospheric oxygen in  
1778 the formation of these deposits, due to the redox sensitive nature of Cu and Co mobility (Brown,  
1779 2014). It is generally accepted that Earth's atmosphere was not significantly oxygenated until  
1780 the Great Oxygenation Event (GOE) at ~2400 Ma (Lyons et al., 2014).

1781 Carlow Castle is a Cu-Co-Au deposit hosted in a Paleo-Mesoarchean volcano-sedimentary  
1782 sequence in the Pilbara Craton of northwestern Western Australia and has previously been  
1783 hypothesized to be of Archean age (Fox et al., 2019; Hickman, 2016). It thus appears to sit  
1784 outside the usual models of hydrothermal Cu-Co metallogensis (Fox et al., 2019). As of  
1785 November 2019, Carlow Castle's inferred resource estimate is 8 million tonnes (Mt) at 0.51%  
1786 Cu, 0.08% Co, 1.6g/t Au (Artemis Resources Limited, 2019c), placing it among Australia's  
1787 most significant Cu-Co-Au deposits. Carlow Castle's occurrence as a structurally-controlled  
1788 hydrothermal Cu-Co-Au deposit within the Paleo-Mesoarchean Roebourne volcano-  
1789 sedimentary greenstone belt makes it unique among Cu-Co-Au deposits, with no other known  
1790 deposits of comparable style and setting (Hitzman et al., 2017). As such, Carlow Castle may  
1791 have implications for understanding the controls on Cu-Co metallogensis through Earth  
1792 history, and particularly during the Archean. To better constrain the age of Carlow Castle, we  
1793 studied the geochemistry and geochronology of what appears to be syn-mineralization



1794 hydrothermal apatite. In addition to this, the alteration assemblage of the host rocks associated  
 1795 with Carlow Castle's sulfide mineralization were also analysed and interpreted to provide  
 1796 evidence of the prevailing physicochemical conditions during hydrothermal ore formation.

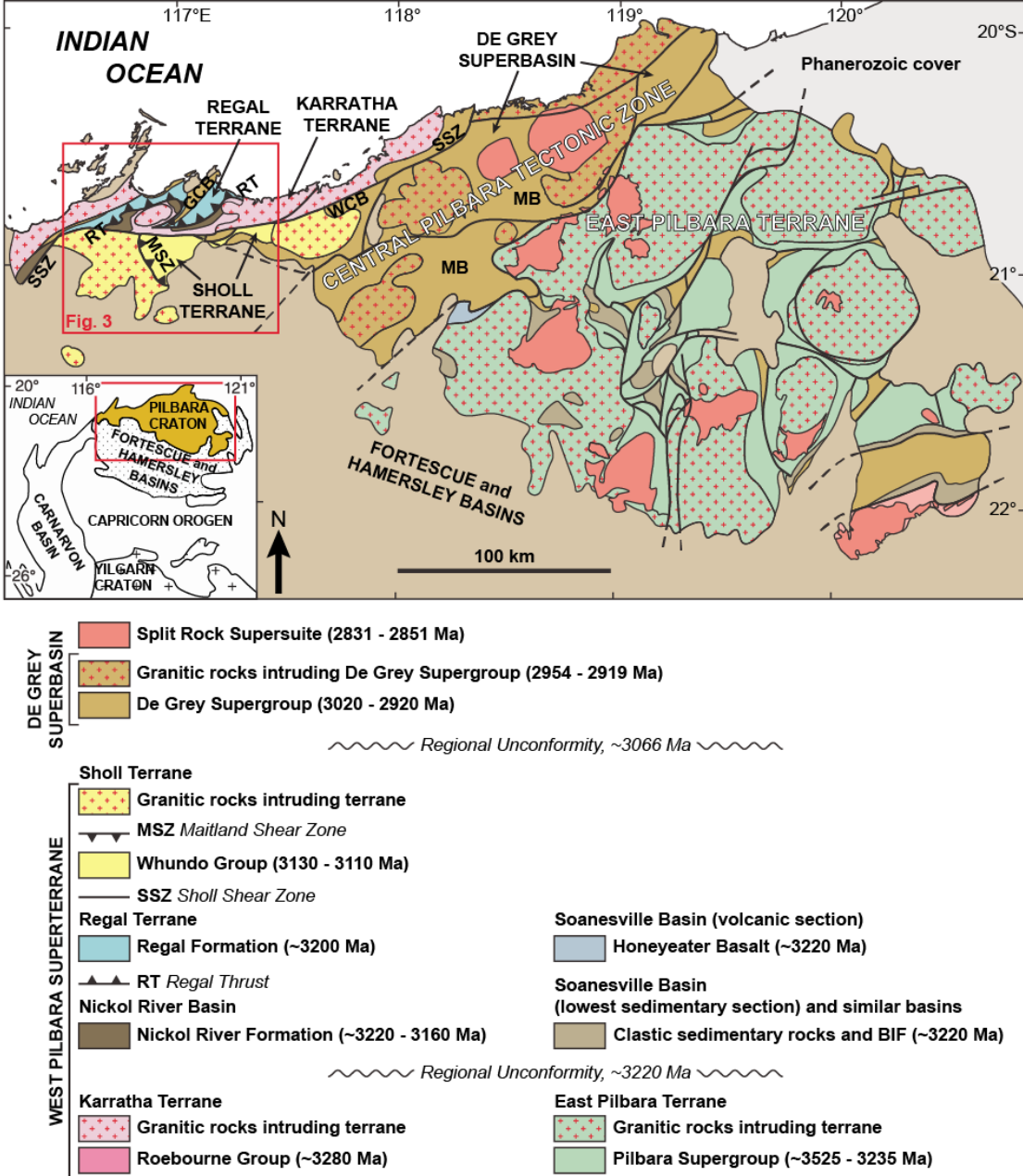


1797  
 1798 Figure 3.1 - Timing of major sediment-hosted Cu deposit formation relative to Carlow  
 1799 Castle's formation and compared with atmospheric O<sub>2</sub> evolution following the Great  
 1800 Oxygenation Event (GOE) and Neoproterozoic Oxygenation Event (NOE). Modified after  
 1801 Hitzman et al. (2010) and Lyons et al. (2014). CACB = Central African Copperbelt, PAL  
 1802 = present atmospheric level.

### 1803 3.3 Geological Setting

1804 The Carlow Castle deposit occurs in the West Pilbara Superterrane (3280-3066 Ma) in  
 1805 northwest Western Australia (Figures 3.2 and 3.3; Hickman, 2016; Van Kranendonk et al.,  
 1806 2002). The West Pilbara Superterrane comprises several Paleo-Mesoarchean volcano-  
 1807 sedimentary greenstone belts and associated intrusive granitic supersuites (Hickman et al.,  
 1808 2006). The Karratha (~3280 Ma), Regal (~3200 Ma), and Sholl (3130-3110 Ma) terranes were  
 1809 assembled into the West Pilbara Superterrane during a period of convergence from ~3160 Ma  
 1810 that culminated in the Prinsep Orogeny at ~3070 Ma (Van Kranendonk et al., 2007). This  
 1811 convergence ultimately led to the obduction of the Regal Terrane over the Karratha Terrane,  
 1812 forming the Regal Thrust (Figures 3.2 and 3.3); a regionally extensive (~2000 km<sup>2</sup>) thrust fault  
 1813 defined by a laterally continuous belt of mylonitized stratigraphy (Hickman, 2016). The

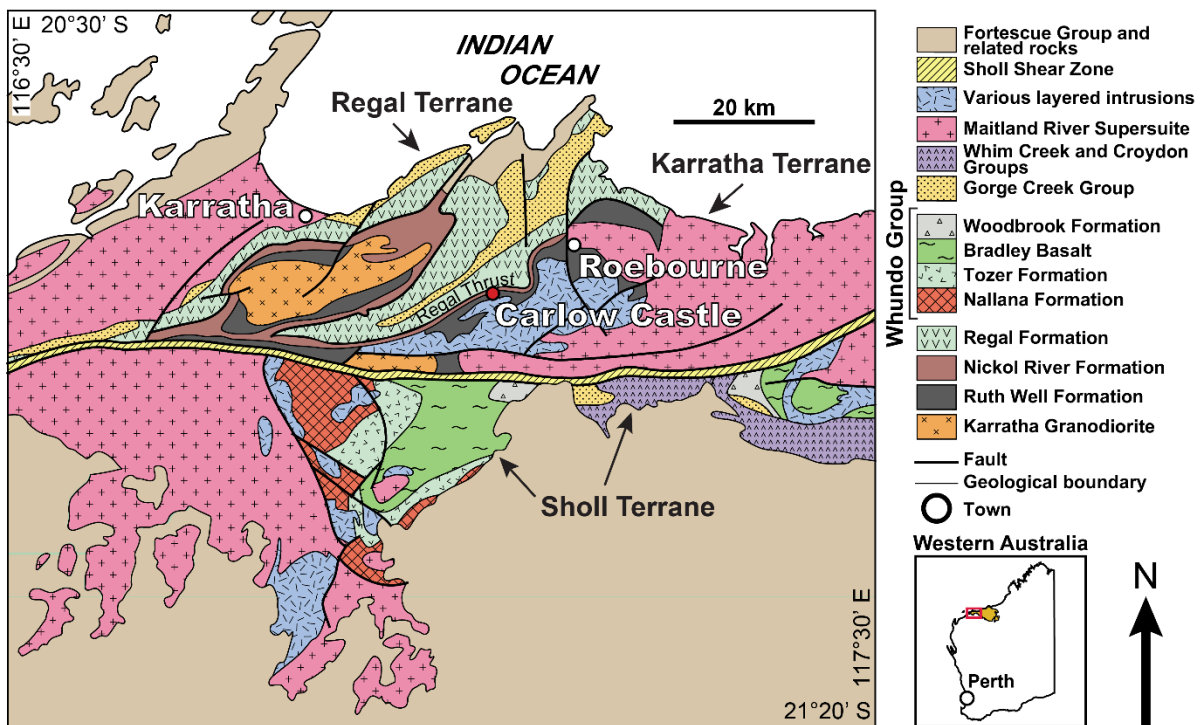
1814 Prinsep Orogeny was followed by the formation of the nearby De Grey Superbasin (3066-2919  
 1815 Ma; Figure 3.2) during a period of significant post-orogenic rifting from 3066 Ma (Hickman,  
 1816 2016). The formation of the De Grey Superbasin was initially driven by post-collisional crustal  
 1817 relaxation and extension following the Prinsep Orogeny and subsequently by the initiation of  
 1818 subduction at the northwest margin of the Pilbara Craton from 3023 Ma (Hickman, 2016;  
 1819 Hickman et al., 2010). Ultimately, the formation of the De Grey Superbasin was terminated  
 1820 during the North Pilbara Orogeny (NPO) from 2955-2919 Ma (Hickman, 2016).



1821  
 1822 Figure 3.2 - Geological map of the Pilbara Craton, modified after Hickman (2016). MB  
 1823 = Mallina Basin, WCB = Whim Creek Basin, GCB = Gorge Creek Basin. Note that the

1824 extent of the area of the northwest Pilbara Craton within Figure 3.3 is indicated and the  
 1825 extent of different terranes is noted in the figure legend.

1826 Mineralization at Carlow Castle occurs within a heavily tectonized zone on the south-eastern  
 1827 footwall of the Regal Thrust (Figure 3.2) through the Nickol River (~3220 Ma) and Ruth Well  
 1828 (~3280 Ma) Formations; the latter of which is a constituent of the Roebourne Group in the  
 1829 Karratha Terrane (Hickman et al., 2006). The Ruth Well Formation is up to 2 km thick and is  
 1830 composed predominantly of extrusive basalt and peridotite with interstratified silicified  
 1831 carbonaceous shale units. These are overlain by fine-grained marine siliciclastic sedimentary  
 1832 rocks of the Nickol River Formation (Hickman, 2016; Hickman & Van Kranendonk, 2012).  
 1833 However, within the mineralized zone of Carlow Castle, the host formation is pervasively  
 1834 chloritized, effectively obscuring the primary lithology of the host rock (Fox et al., 2019).

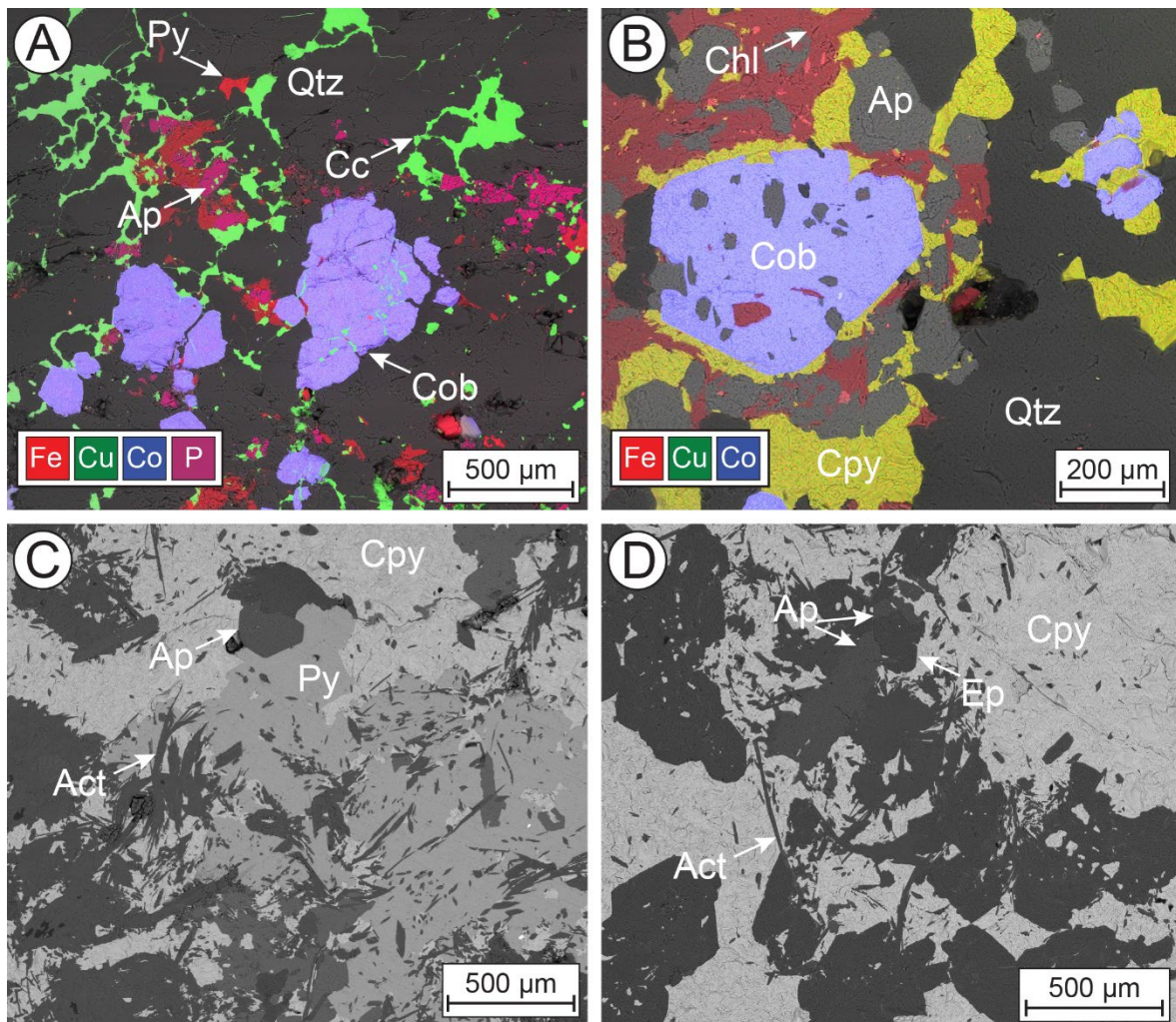


1835  
 1836 Figure 3.3 - Geologic map of the northwest Pilbara Craton. Modified after Van  
 1837 Kranendonk et al. (2002).

1838 **3.3.1 Ore deposit geology**

1839 The Cu-Co-Au ore at Carlow Castle occurs as base-metal sulfides through a network of  
 1840 pervasive quartz-carbonate veining and can be broadly separated into two distinct ore mineral  
 1841 assemblages (Fox et al., 2019). Assemblage One is more volumetrically significant and  
 1842 composed primarily of pyrite and chalcopyrite with minor cobaltite, pyrrhotite, and trace

1843 electrum. Assemblage Two is significantly more Co- and Au-rich and is dominated by cobaltite  
 1844 and chalcocite with electrum typically occurring as micron-scale inclusions within cobaltite  
 1845 grains. Throughout the mineralized zone, the sulfide ore is intergrown with chlorite and  
 1846 actinolite from the extensively hydrothermally altered wall-rock. Syn-mineralization  
 1847 hydrothermal apatite occurs plentifully throughout both ore mineral assemblages, intergrown  
 1848 with the sulfide ore and hydrothermal alteration minerals (Figures 3.4 and 3.5; Appendix 3.1).  
 1849 Detailed characterization and discussion of Carlow Castle’s style of mineralization is provided  
 1850 in Fox et al. (2019).



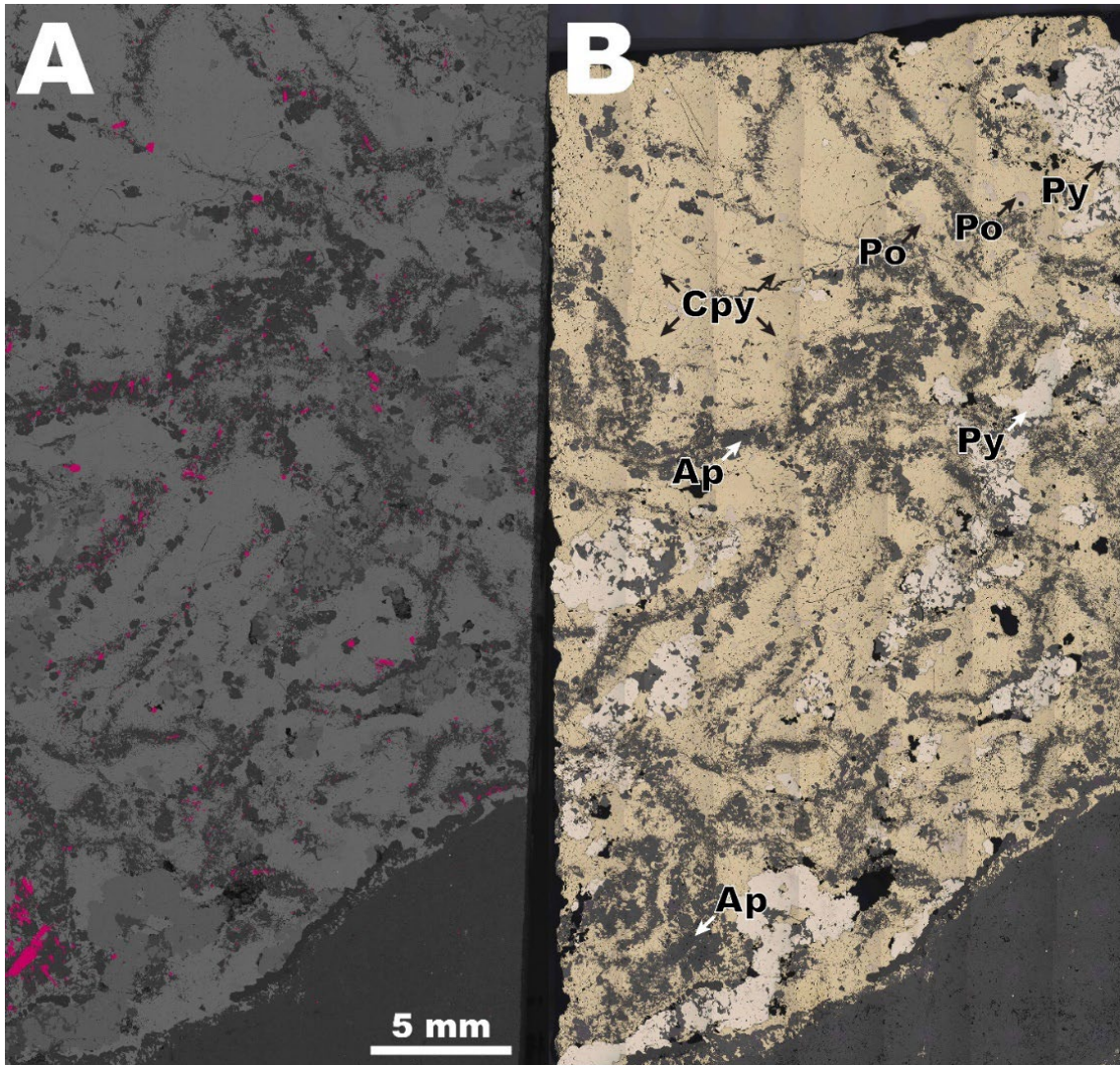
1851  
 1852 Figure 3.4 – Scanning electron microscope (SEM) photomicrographs showing syn-  
 1853 mineralization hydrothermal apatite associated with sulfide mineralization at Carlow  
 1854 Castle Cu-Co-Au deposit. A. Element map from Sample CC009\_48 of Fe (red), Cu  
 1855 (green), Co (blue), and P (pink) within a mineralized quartz (qtz) vein containing cobaltite  
 1856 (cob), chalcocite (cc), pyrite (py), and apatite (ap). Note how apatite is intergrown with

1857 chalcocite within this vein. B. Element map from Sample CC003\_61 of Fe (red), Cu  
1858 (green), Co (blue) showing a cobaltite grain rimmed by chalcopyrite and hydrothermal  
1859 apatite, suggesting apatite post-dates cobaltite within the mineral paragenetic sequence.  
1860 C. Euhedral apatite hosted within massive intergrown coarse-grained pyrite and  
1861 chalcopyrite mineralization (Sample CC012\_67). D. Apatite intergrown with epidote and  
1862 actinolite alteration minerals, hosted within massive coarse-grained chalcopyrite (Sample  
1863 CC012\_67).

## 1864 **3.4 Methods and materials**

### 1865 **3.4.1 Samples**

1866 Hydrothermal apatite within a representative mineralized polished block section of sample  
1867 CC012\_67 (Figure 3.5, Appendix 3.1) from ore mineral Assemblage One (Fox et al., 2019),  
1868 from the Carlow Castle ore body was selected for U-Pb geochronology. Sample CC012\_67  
1869 contains abundant coarse euhedral apatite grains contemporaneous with sulfide mineralization,  
1870 ideal for in-situ laser ablation analysis. This sample comprises coarse-grained intergrown pyrite  
1871 and chalcopyrite, with minor pyrrhotite and trace sphalerite and cobaltite. Gangue quartz,  
1872 calcite, siderite, chlorite, actinolite, epidote, and apatite are also common throughout the  
1873 sample. The hydrothermal apatite analyzed within this sample occur as clusters of syn-  
1874 mineralization apatite grains intergrown with other hydrothermal gangue or sulfide minerals  
1875 (Figures 3.4 and 3.5, Appendix 3.1).



1876  
 1877 Figure 3.5 - A. SEM backscattered electron mineral map of apatite grains (pink) within  
 1878 laser ablation sample CC012\_67. B. Reflected light photomosaic of sample CC012\_67  
 1879 with major mineral phases labelled. Note visible laser ablation pits within apatite grains.

1880 In addition to sample CC012\_67, polished block sections of samples CC003\_61 and CC009\_48  
 1881 were also prepared as representative samples of veined mineralization for small-scale  
 1882 mineralogical analysis. Sample CC003\_61 is composed predominantly of chalcopyrite and  
 1883 pyrite, typical of ore mineral Assemblage One, hosted within quartz-carbonate veins. Sample  
 1884 CC009\_48 is representative of ore mineral Assemblage Two and contains cobaltite and  
 1885 chalcocite hosted within a quartz-carbonate vein.

### 1886 3.4.2 Scanning electron microscopy

1887 Polished sections of samples CC012\_67, CC009\_48, and CC003\_61 were carbon coated for  
 1888 mineralogical analysis using a Tescan Integrated Mineral Analyzer (TIMA) at the

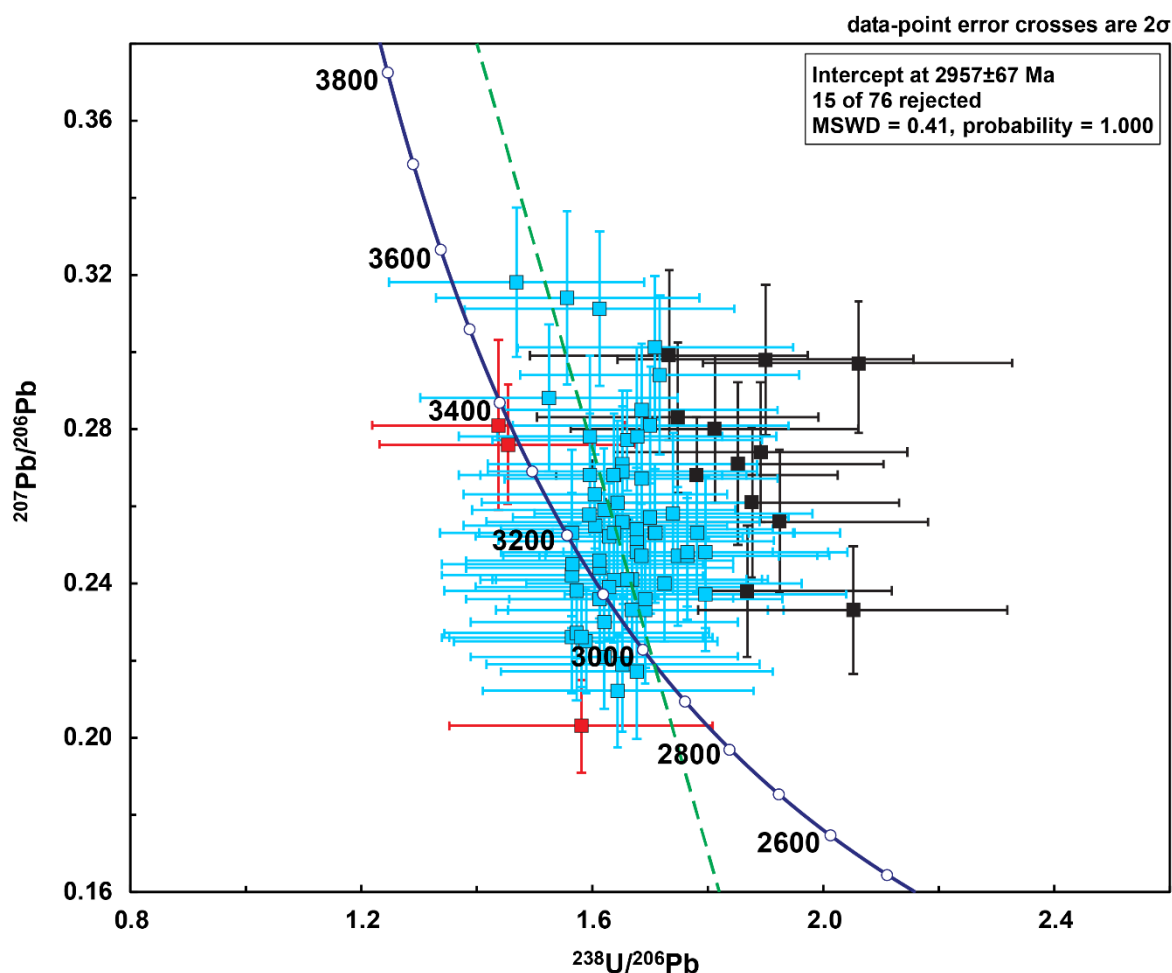
1889 Commonwealth Scientific and Industrial Research Organisation (CSIRO) Advanced Resource  
1890 Characterisation Facility at the Australian Resources Research Centre. These samples were  
1891 analyzed through collection of energy dispersive X-ray spectra (EDS) dot maps (Appendix 3.1)  
1892 of the samples with a backscattered electron (BSE) pixel size of 3  $\mu\text{m}$ , spectra were collected  
1893 every 10  $\mu\text{m}$  and/or where a change in BSE contrast was detected. These analyses used an  
1894 accelerating voltage of 25 kV, beam intensity of 18, beam current of 4.35 nA, spot size of 51  
1895 nm, and working distance of 15 mm. After the samples were imaged in BSE and EDS spectra  
1896 collected, the mineralogical composition of these samples was determined via referencing these  
1897 parameters to a customized TIMA mineral library.

### 1898 3.4.3 Apatite spot analysis

1899 Laser Ablation Inductively Coupled Plasma Mass Spectrometry (LA-ICP-MS) for U-Pb  
1900 geochronology and trace element geochemistry was performed at the GeoHistory Facility in  
1901 the John De Laeter Centre at Curtin University in Perth, Western Australia. Single spot  
1902 ablations (76) were performed across 71 in-situ apatite grains within polished section  
1903 CC012\_67, with simultaneous collection of age and geochemical data from each ablation  
1904 volume. Ablations were performed with a Resonetics RESolution M-50A-LR, utilizing a  
1905 COMpex 193 nm excimer laser in constant energy mode using a 50  $\mu\text{m}$  spot size, 5 Hz laser  
1906 repetition rate, and a laser energy of 1.80 J  $\text{cm}^{-2}$ . Isotope geochemistry was measured by an  
1907 Agilent 7700s quadrupole ICP-MS, using high purity Ar as a plasma gas. Each spot analysis  
1908 included a 30 s background acquisition and a 35 s ablation. A total of 28 isotopes were  
1909 measured  $^{29}\text{Si}$ ,  $^{31}\text{P}$ ,  $^{43}\text{Ca}$ ,  $^{44}\text{Ca}$ ,  $^{55}\text{Mn}$ ,  $^{88}\text{Sr}$ ,  $^{89}\text{Y}$ ,  $^{91}\text{Zr}$ ,  $^{139}\text{La}$ ,  $^{140}\text{Ce}$ ,  $^{141}\text{Pr}$ ,  $^{146}\text{Nd}$ ,  $^{147}\text{Sm}$ ,  $^{151}\text{Eu}$ ,  
1910  $^{157}\text{Gd}$ ,  $^{163}\text{Dy}$ ,  $^{165}\text{Ho}$ ,  $^{166}\text{Er}$ ,  $^{169}\text{Tm}$ ,  $^{172}\text{Yb}$ ,  $^{175}\text{Lu}$ ,  $^{201}\text{Hg}$ ,  $^{204}\text{Pb}$ ,  $^{206}\text{Pb}$ ,  $^{207}\text{Pb}$ ,  $^{208}\text{Pb}$ ,  $^{232}\text{Th}$ , and  $^{238}\text{U}$ .  
1911 The measurement time was 0.02 s for isotopes of Pb, Th, U, and 0.01 s for all other isotopes.  
1912 McClure Mountain syenite apatite was used as the primary reference material (Thomson,  
1913 Gehrels, Ruiz, & Buchwaldt, 2012). Duluth Complex apatite (U-Pb zircon age of  
1914  $1099.1 \pm 0.2$  Ma (Schmitz, Bowring, & Ireland, 2003)) was used as a secondary reference  
1915 material, and treated as an unknown. The age obtained, by free regression, on the secondary  
1916 reference Duluth Complex apatite ( $1123 \pm 39$ ; MSWD = 1.15; n = 13 of 14) is within analytical  
1917 uncertainty of the expected value (Appendix 3.2; Kirkland et al., 2018).

1918 Time-resolved mass spectra were reduced using Iolite (Paton, Hellstrom, Paul, Woodhead, &  
1919 Hergt, 2011) and U-Pb and Pb-Pb ages were produced and interpreted using the Isoplot Visual  
1920 Basic add-in for Microsoft Excel (Ludwig, 1991). Isoplot 4.15 was used to calculate common

1921 (non-radiogenic)  $^{207}\text{Pb}$ -corrected  $^{238}\text{U}/^{206}\text{Pb}$  ages for each apatite spot analysis based upon an  
 1922 estimate of common Pb using the Stacey and Kramers (1975) model of the Earth's Pb isotope  
 1923 evolution (Figure 3.6). An iterative approach was applied to determine the appropriate common  
 1924  $^{207}\text{Pb}/^{206}\text{Pb}$  ratio to use in  $^{207}\text{Pb}$ -correction, where an initial  $^{207}\text{Pb}/^{206}\text{Pb}$  age was used as an  
 1925 estimate for a model common Pb composition. Following this, an initial estimated  $^{207}\text{Pb}$ -  
 1926 corrected age was used to revise the common Pb model to calculate a  $^{207}\text{Pb}$ -corrected  $^{238}\text{U}/^{206}\text{Pb}$   
 1927 age. This process was repeated until no change in  $^{207}\text{Pb}$ -corrected age resulted from a change  
 1928 in common Pb value. Such  $^{207}\text{Pb}$ -corrected individual ages are used to distinguish different age  
 1929 components. A regression (green in Figure 3.6) through the population anchored from a  
 1930 common Pb value was used for the final age estimate.



1931  
 1932 Figure 3.6 - U-Pb concordia plot of 76 apatite analyses from sample CC012\_67. The  
 1933 regression line (green) through roughly concordant analyses (blue) intersects the  
 1934 concordia curve at 2957 Ma. Note rejected analyses are colored red for those left of



1935 concordia with apparent ages older than the host terrane and black for those to the far  
 1936 right of concordia that have undergone radiogenic Pb loss.

### 1937 3.5 Results

#### 1938 3.5.1 Sample mineralogy

1939 Consistent with the previous mineralogical categorization based on visual observation;  
 1940 CC012\_67 and CC003\_61 belong to the chalcopyrite and pyrite dominated ore mineral  
 1941 Assemblage One, whilst CC009\_48 belongs to the more cobaltite- and chalcocite-rich ore  
 1942 mineral Assemblage Two (Table 3.1). In addition to these ore minerals, an alteration  
 1943 assemblage composed of actinolite, chlorite (dominantly chamosite, based on EDS mapping  
 1944 results), epidote, K-feldspar, quartz, siderite, and calcite occurs within the samples; replacing  
 1945 the primary basaltic host-rock (Appendix 3.1). However, the abundance of individual alteration  
 1946 minerals within each sample varies significantly (Table 3.1). Whilst CC012\_67 is particularly  
 1947 actinolite-rich in addition to siderite; epidote; and minor K-feldspar and chlorite, the alteration  
 1948 assemblages within CC009\_48 and CC003\_61 are composed strongly of greater amounts of  
 1949 chlorite with significantly less actinolite and epidote (Appendix 3.1). The alteration assemblage  
 1950 of CC009\_48 is composed of quartz, chlorite, siderite, and minor K-feldspar and actinolite.  
 1951 Despite their different ore mineral assemblages, the alteration assemblage of CC003\_61 is  
 1952 comparable but more chloritic; being composed of quartz, chlorite, K-feldspar, calcite, and  
 1953 minor to trace siderite and actinolite. CC009\_48 and CC003\_61 are also particularly quartz-  
 1954 rich due to their quartz vein hosted occurrence, in contrast to CC012\_67, which is exceptionally  
 1955 chalcopyrite rich and more comparable to coarse-grained massive sulfide.

1956 Table 3.1 - Quantitative mineralogy of samples analysed with TIMA from Carlow Castle.

| <b>Sample/Mineral</b> | <b>CC012_67</b> | <b>CC009_48</b> | <b>CC003_61</b> |
|-----------------------|-----------------|-----------------|-----------------|
| <b>Chalcopyrite</b>   | 52.29%          | 0.04%           | 15.34%          |
| <b>Actinolite</b>     | 17.22%          | 0.21%           | 0.04%           |
| <b>[Unclassified]</b> | 10.53%          | 7.19%           | 8.06%           |
| <b>Pyrite</b>         | 8.38%           | 0.0035%         | 3.41%           |
| <b>Siderite</b>       | 3.06%           | 5.36%           | 0.14%           |
| <b>Epidote</b>        | 3.00%           | 0.0004%         | -               |
| <b>Pyrrhotite</b>     | 1.71%           | 0.0009%         | 0.1%            |
| <b>K-feldspar</b>     | 0.89%           | 0.34%           | 2.83%           |

|                                 |       |         |         |
|---------------------------------|-------|---------|---------|
| <b>Chlorite<br/>(chamosite)</b> | 0.83% | 7.36%   | 19.34%  |
| <b>Apatite</b>                  | 0.67% | 1.23%   | 1.03%   |
| <b>Quartz</b>                   | 0.60% | 69.45%  | 46.17%  |
| <b>Holes</b>                    | 0.56% | 3.07%   | 0.64%   |
| <b>Biotite (annite)</b>         | 0.10% | 0.02%   | 0.25%   |
| <b>Cobaltite</b>                | 0.01% | 4.81%   | 0.37%   |
| <b>Chalcocite</b>               | -     | 0.76%   | 0.0002% |
| <b>Calcite</b>                  | 0.01% | 0.02%   | 1.83%   |
| <b>Ilmenite/rutile</b>          | -     | 0.03%   | 0.44%   |
| <b>Other</b>                    | 0.14% | 0.1052% | 0.0098% |

### 1957 3.5.2 Geochronology

1958 U-Pb ratios in 61 measurements of the hydrothermal apatite define a broadly concordant  
1959 distribution (Figure 3.6) with some scatter towards common Pb, consistent with a common-  
1960 radiogenic mixing line. While the low U content of these grains (mean: 1.6 ppm) makes  
1961 geochronology challenging, only fifteen out of 61 data points do not conform to a single  $^{207}\text{Pb}$ -  
1962 corrected age population. Three data points (red) (*12\_67 – 9, 19, 55*) have apparent  $^{207}\text{Pb}$ -  
1963 corrected ages significantly older than the host rock's formation age (~3300 Ma) and plot left  
1964 of concordia. Additionally, these spots were sited on grain edges where they may reflect  
1965 mixtures with other phases containing radiogenic Pb. A further 12 data points (black) (*12\_67*  
1966 *– 1, 2, 3a, 25, 27, 33, 35, 45, 46, 49, 55, 57, 77*) may have undergone minor radiogenic Pb loss,  
1967 are discordant, and plot to the right of a radiogenic to common Pb mixing array. No distinct  
1968 chemical variation is noted between different components, and there is no apparent relationship  
1969 between the  $^{207}\text{Pb}$ -corrected age of a given apatite grain and its trace element geochemistry.  
1970 Nonetheless, several statistical outliers that were rejected are geochemically anomalous and  
1971 contain highly variable heavy rare earth element (REEs) concentrations, with many being  
1972 depleted relative to the concordant analyses though, some are comparably enriched. In general,  
1973 the 15 discordant outliers exhibit greater variability in HREE concentration than the concordant  
1974 population. Excluding the 15 statistical outliers, 61 measurements yield a concordia intercept  
1975 age of  $2957 \pm 67$  Ma (MSWD = 0.41,  $p = 1$ ), calculated from a common Pb composition  
1976 corresponding to a terrestrial Pb model of 2960 Ma ( $^{207}\text{Pb}/^{206}\text{Pb}$ : 0.90) as per Stacey and  
1977 Kramers (1975). The calculated apparent age is stable despite varying the common Pb value in

1978 the model up to 3300 Ma ( $^{207}\text{Pb}/^{206}\text{Pb}$ : 0.83) (corresponding to the maximum age of the host  
 1979 formation). This is due to the near concordance of the data, making the final age insensitive to  
 1980 the choice of common Pb value. All U-Pb data are provided in Appendix 3.2.

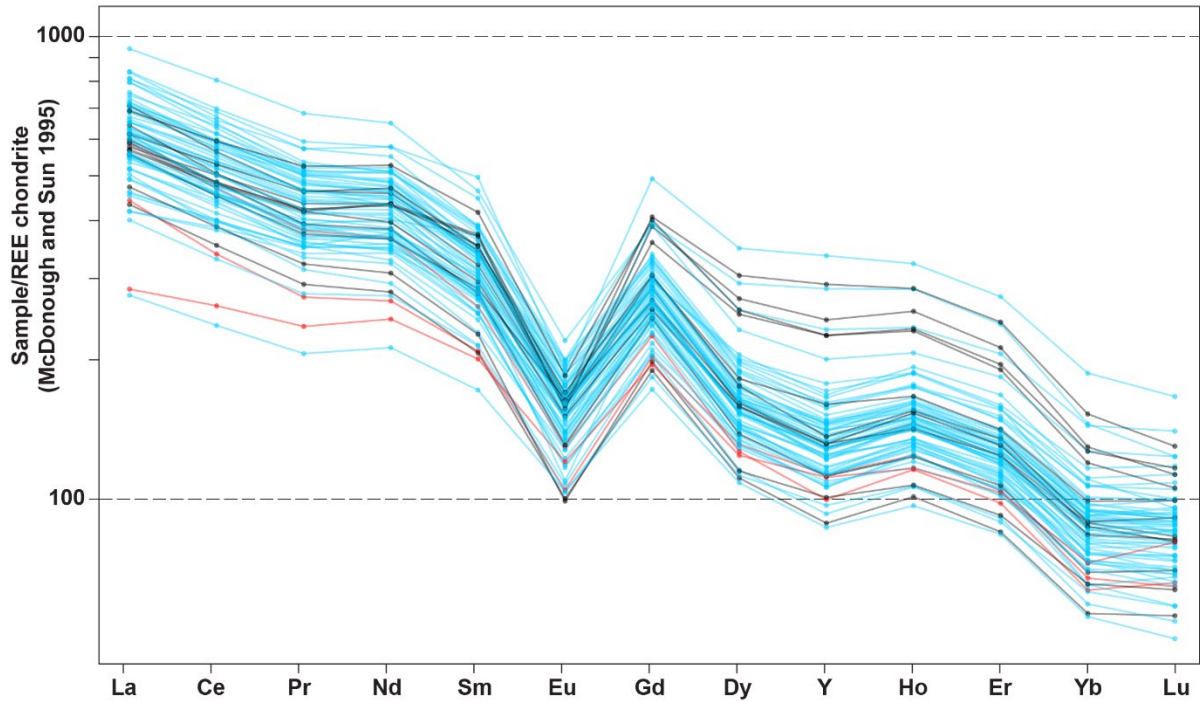
### 1981 3.5.3 Apatite geochemistry

1982 Carlow Castle apatite crystals are typical in terms of major element geochemistry relative to  
 1983 apatite in magmatic-hydrothermal ore deposits (Mao, Rukhlov, Rowins, Spence, & Coogan,  
 1984 2016), though values for P tend to vary more significantly than Ca. Si varies considerably,  
 1985 between 180-9100 ppm (mean: 1730 ppm). Mn, Sr, Zr, and Hg concentrations vary within a  
 1986 relatively narrow range (Table 3.2) and Zr and Hg values are commonly below detection limits.  
 1987 Total U, Th, and Pb concentrations within the apatite are particularly low in comparison to  
 1988 apatite associated with various ore deposits (Table 3.2) (Mao et al., 2016). There is no  
 1989 consistent relationship between total P, Ca, Mn, Sr, Si, Zr, Hg, U, Th, and Pb concentrations,  
 1990 and the ages of the apatite grains. The full geochemical data for all analyzed elements are  
 1991 provided in Appendix 3.3.

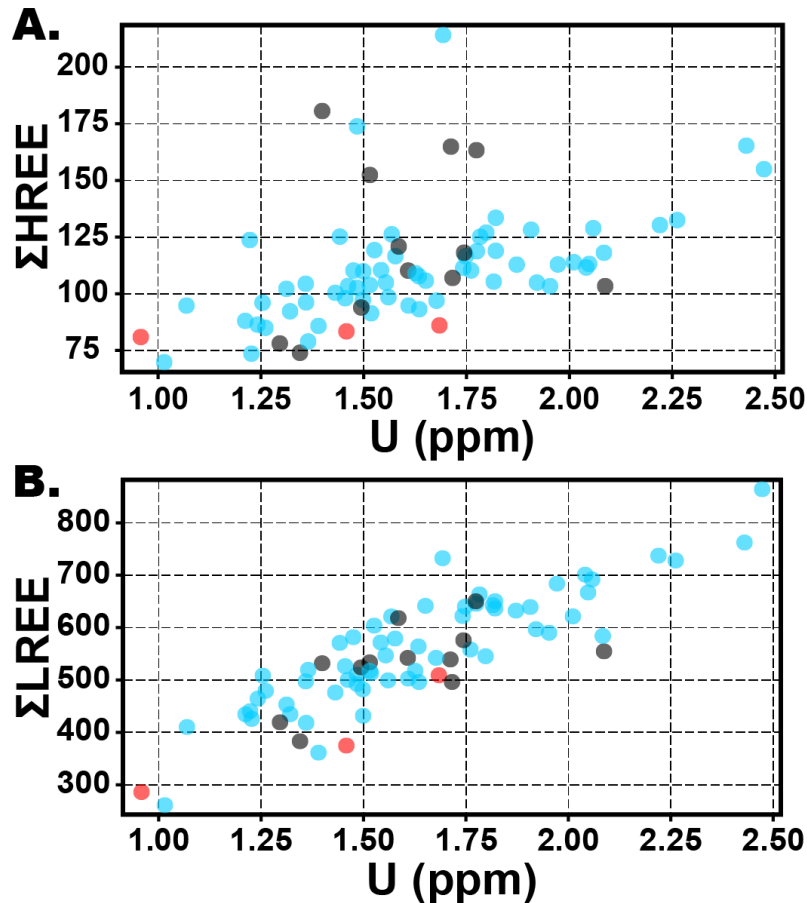
1992 Table 3.2 - Summary of apatite LA-ICP-MS geochemical data. Note values are in ppm;  
 1993 DL = detection limit, SD = standard deviation.

|                | <b>P</b> | <b>Ca</b> | <b>Mn</b> | <b>Sr</b> | <b>Si</b> | <b>Zr</b> | <b>Hg</b> | <b>U</b> | <b>Th</b> | <b>Pb</b> |
|----------------|----------|-----------|-----------|-----------|-----------|-----------|-----------|----------|-----------|-----------|
| <b>Minimum</b> | 154,600  | 384,000   | 216       | 109       | 180       | <DL       | <DL       | 1.015    | <DL       | 0.019     |
| <b>Maximum</b> | 194,200  | 395,400   | 324       | 197       | 9100      | 0.19      | 74        | 2.473    | 0.0184    | 0.305     |
| <b>Mean</b>    | 166,320  | 388,932   | 270       | 145       | 1730      | 0.17      | 18        | 1.64     | 0.0032    | 0.0736    |
| <b>Median</b>  | 165,000  | 389,000   | 268       | 147       | 890       | 0.012     | 21        | 1.608    | 0.0028    | 0.065     |
| <b>Range</b>   | 39,600   | 11,400    | 108       | 88        | 8920      | -         | -         | 1.46     | -         | 0.29      |
| <b>SD</b>      | 8980     | 2033      | 22.58     | 12.08     | 2182      | 0.036     | 27.69     | 0.302    | 0.0034    | 0.043     |

1994 Broadly the REE profiles of the apatite define a single population, characterized by a negatively  
 1995 sloped REE profile with a notable negative Eu anomaly (Figure 3.7). Similarly to U, absolute  
 1996 concentrations of REEs within these apatite are low; with  $\sum\text{REE}+\text{Y}$  ranging between 487.94-  
 1997 1545.14 ppm (mean: 929.16 ppm).  $\sum\text{LREE}$  (La + Ce + Pr + Sm + Eu) range between 261.17  
 1998 and 864.13 ppm (mean: 558.47 ppm), whilst  $\sum\text{HREE}$  (Gd + Dy + Yb) values range between  
 1999 69.79 and 214.18 ppm with some outliers skewing toward higher HREE concentrations (mean:  
 2000 112.64 ppm). Across the population there is a strong positive correlation ( $r^2 = 0.733$ ) between  
 2001 LREE and U concentrations, whilst there is a moderate positive correlation ( $r^2 = 0.222$ ) between  
 2002 HREE and U concentrations (Figure 3.8a, b).



2003  
 2004 Figure 3.7 - Rare earth element plot for 76 apatite analyses from CC012\_67, displaying a  
 2005 consistently negative slope and negative Eu anomaly characteristic of a single population.  
 2006 The coloring scheme (blue, red, black) reflects the assigned group of each analysis on the  
 2007 concordia plot in Figure 3.6. Note the disproportionately low REE values of red analyses.



2008

2009 Figure 3.8 - A.  $\Sigma$ HREE plotted against U (ppm) content of 76 apatite analyses. B.  $\Sigma$ LREE  
 2010 plotted against U (ppm) of 76 apatite analyses. Note the strong positive linear relationship  
 2011 between  $\Sigma$ LREE and U. Colors for each data point are consistent with those in Figure 3.6  
 2012 and Figure 3.7.

### 2013 3.6 Discussion

#### 2014 3.6.1 Geochronological constraints on the timing of ore formation

2015 The Mesoarchean ( $2957 \pm 67$  Ma) age of the hydrothermal apatite grains dated in this study is  
 2016 interpreted to represent the age of Carlow Castle's Cu-Co-Au mineralization, based on textural  
 2017 evidence of their syn-mineralization formation (intimate intergrowth with sulfides and ore-  
 2018 stage hydrothermal alteration phases; Figure 3.5). Although the age variation of the 61 near-  
 2019 concordant apatite analyses is higher than expected for a single population, they define a  
 2020 broadly Gaussian distribution and their trace element geochemistry defines a single population  
 2021 (Figure 3.7); supporting a single growth stage. Apatite in the population with particularly old  
 2022 ages can be accounted for through the incorporation of a radiogenic Pb component from  
 2023 mixtures with other uranium-bearing minerals on the margins of apatite inadvertently sampled

2024 by the laser during analysis. Conversely, apatite that lie to the right of a regression from the  
2025 common Pb anchor, i.e., apparently young  $^{207}\text{Pb}$ -corrected ages within the population, are  
2026 explained through Pb loss. Both of these features are frequently encountered in laser ablation  
2027 of apatite and are a routine consideration when conducting apatite geochronology (Kirkland et  
2028 al., 2018). Additionally, the anomalous REE variation within discordant outliers lying to the  
2029 right of the best fit regression may reflect the alteration of the chemistry of these apatite grains  
2030 post-crystallization via various processes that could have simultaneously altered the original  
2031 REE and U-Pb contents of some grains (Cherniak, 2005; Cherniak, Lanford, & Ryerson, 1991;  
2032 Hughes & Rakovan, 2015). Conversely, relatively REE-enriched discordant outlier grains to  
2033 the right of the best fit regression may be explained by the presence of trace U, Pb, and REE-  
2034 bearing phases as inclusions within these grains that were not noticed prior to analysis.

2035 By comparison to whole rock REE data of basaltic portions of the Ruth Well Formation  
2036 published by Smithies, Champion, Van Kranendonk, and Hickman (2007), the Carlow Castle  
2037 apatite display notable REE fractionation with a clear depletion in HREEs relative to LREEs.  
2038 Additionally, the Carlow Castle apatite also display a significantly more pronounced negative  
2039 Eu anomaly than the host Ruth Well Formation. These fractionated REE patterns along with  
2040 the notably low concentrations of U, Pb, Th, and  $\Sigma\text{REE}$  are not uncommon within apatite from  
2041 hydrothermal ore deposits, especially when compared with igneous apatite (Mao et al., 2016).  
2042 These trace element concentrations within the Carlow Castle apatite most likely reflect the  
2043 preferential partitioning of U, Pb, Th, and REEs into other ore and alteration phases. The  
2044 comparable rarity of such U-, Pb-, Th-, and REE-bearing ore and alteration phases in igneous  
2045 rocks, accounts for the commonly observed fractionation and depletion of these elements in  
2046 hydrothermal apatite relative to igneous apatite (Krneta, Ciobanu, Cook, Ehrig, & Kontonikas-  
2047 Charos, 2017; Mao et al., 2016). For example, the occurrence of uraninite in Carlow Castle's  
2048 Cu-Co-Au mineralization suggests that its hydrothermal fluid was uraniferous (Fox et al.,  
2049 2019). Initially this is difficult to reconcile with the low U concentrations in the Carlow Castle  
2050 apatite; however, the presence of uraninite is likely the reason that Carlow Castle's apatite are  
2051 so U poor. This is because any dissolved U in this fluid was most likely preferentially  
2052 incorporated into uraninite, this would leave relatively little dissolved U to be incorporated into  
2053 the hydrothermal apatite. This is supported by evidence in Figure 3.4b, that apatite appears to  
2054 post-date cobaltite paragenetically. Given that uraninite is known to be associated with  
2055 cobaltite mineralisation within Carlow Castle (Fox et al., 2019), it is likely that apatite also

2056 post-dates uraninite mineralisation. A similar process would also account for the REE  
2057 concentrations observed within these apatite, as several hydrothermal minerals observed within  
2058 Carlow Castle may incorporate REEs up to thousands of ppm; including uraninite, epidote,  
2059 chlorite, calcite, K-feldspar, and amphibole (Fryer & Taylor, 1987; Pike, Cas, & Smithies,  
2060 2002; Tillberg et al., 2019). As such, the strongly fractionated REE pattern along with the low  
2061 absolute U and REE concentrations within the Carlow Castle apatite may reflect the  
2062 synchronous or subsequent growth of these apatite crystals with other REE- and U-bearing  
2063 hydrothermal minerals and the consequent competition for these components during crystal  
2064 growth. This geochemical evidence supports previously noted textural evidence of a  
2065 hydrothermal origin for these apatite, that is contemporaneous with Carlow Castle's  
2066 hydrothermal Cu-Co-Au mineralization; suggesting that the age of the apatite measured by this  
2067 study can also be interpreted to be the age of Cu-Co-Au mineralization.

2068 Prior to this study, a maximum age constraint of 3070 Ma on Carlow Castle was provided by  
2069 the Regal Thrust host structure (Hickman, 2016). While a minimum possible age was uncertain,  
2070 it was suggested that Carlow Castle was unlikely to be significantly younger than the initiation  
2071 of rifting in the Fortescue Basin around 2775 Ma due to the paucity of obvious subsequent ore-  
2072 forming tectonothermal events in the Pilbara Craton (Fox et al., 2019). The Carlow Castle  
2073 deposit was originally hypothesized to be an orogenic Au system with atypical Cu-Co  
2074 enrichment, with the Prinsep Orogeny (~3070 Ma) being the likely event responsible for  
2075 mineralization (Fox et al., 2019; Hickman, 2016). This interpretation was based on Carlow  
2076 Castle's occurrence within the Regal Thrust, an orogenic structure, and several similarities in  
2077 style and setting to an orogenic mineral system (Fox et al., 2019). However, the age of  
2078 mineralization constrained by this study ( $2957 \pm 67$  Ma) suggests instead that Carlow Castle's  
2079 ore formation post-dates the Prinsep Orogeny. Therefore, whilst the critical structures (Regal  
2080 Thrust) that host Carlow Castle's ore are almost certainly a product of orogenesis, Carlow  
2081 Castle's mineralizing processes are most likely related to the later stages of the NW Pilbara  
2082 Craton's evolution.

### 2083 3.6.2 Tectonic process and regional metallogensis

2084 The age range over which the Carlow Castle deposit most likely formed (3024-2890 Ma)  
2085 significantly postdates the age of the host Karratha Terrane (3280-3236 Ma) and instead  
2086 corresponds to the formation of the nearby De Grey Superbasin (3066-2931 Ma) within the  
2087 Central Pilbara Tectonic Zone (CPTZ) of the NW Pilbara Craton (Hickman, 2016; Figure 3.2).

2088 Within the De Grey Superbasin, the inferred age of Carlow Castle's mineralization coincides  
2089 with the formation of three distinct sub-basins: (i) the Gorge Creek Basin (3066-3015 Ma); a  
2090 shallow water intracontinental basin of dominantly marine siliciclastic fill over the northern  
2091 Pilbara Craton, (ii) the Whim Creek Basin (3009-2991 Ma); a small fault-bounded volcanic arc  
2092 to the southeast, and (iii) the Mallina Basin (3015-2931 Ma); a back-arc basin further to the  
2093 southeast of the Whim Creek Basin volcanic arc (Hickman, 2016; Hickman et al., 2010). The  
2094 formation of the Gorge Creek and Whim Creek Basins was not associated with any  
2095 economically significant mineralization, except iron ore (Hickman, 2016; Hickman et al.,  
2096 2006).

2097 The formation of the Mallina Basin was terminated by the North Pilbara Orogeny (NPO; 2955-  
2098 2919 Ma), with peak metamorphism and deformation between 2940-2920 Ma (Hickman, 2016;  
2099 Hickman et al., 2006; Van Kranendonk et al., 2007). During the NPO, regional mineralization  
2100 is interpreted to have occurred in two discrete phases. Mineralization between 2955-2940 Ma  
2101 was dominated by volcanogenic massive sulfide (VMS) mineralization of Pb, Zn, Ag, Cu, and  
2102 V within the Whim Creek greenstone belt during continued extension and granitic magmatism  
2103 within the Mallina Basin, along with PGE and Ni-Cu mineralization associated with the  
2104 intrusion of mafic-ultramafic sills throughout the central Mallina Basin (Hickman, 2016;  
2105 Huston, Smithies, & Sun, 2000; Huston et al., 2002b; Pike et al., 2002). This period produced  
2106 the most diverse assemblage of mineralization of any terrane in the northern Pilbara Craton  
2107 (Hickman, 2016). This was followed by a subsequent period of mineralization during the  
2108 metamorphic peak of the NPO (2940-2920 Ma) dominated by shear zone hosted orogenic Au  
2109 in addition to layered intrusive Ni-Cu and PGE mineralization (Hickman, 2016; Hoatson &  
2110 Sun, 2002; Huston et al., 2002a).

2111 The high intensity of tectonism during the period of Carlow Castle's potential mineralization  
2112 makes it difficult to identify exactly which tectonic event is likely to have driven the formation  
2113 of the ore system. However, our data suggests Carlow Castle's metal enrichment likely formed  
2114 through the reactivation of pre-existing orogenic structures associated with the Regal Thrust  
2115 (~3070 Ma) during a later extensional regime that prevailed for most of the development of the  
2116 proximal CPTZ. This interpretation is supported by the observation of multiple generations of  
2117 deformation fabrics within the Carlow Shear Zone by Kiyokawa et al. (2002), with evidence  
2118 of late-stage post-orogenic deformation produced by extension and basin formation between  
2119 3020-2925 Ma. It is perhaps unsurprising that Carlow Castle formed during this period, given



2120 that the CPTZ is the portion of the Pilbara Craton most endowed in hydrothermal and magmatic  
2121 gold and base-metal ore deposits (Hickman, 2016; Hickman et al., 2006; Huston, Blewett,  
2122 Mernagh, Sun, & Kamprad, 2001). This suggests that the formation of the CPTZ was  
2123 associated with geological processes that were particularly conducive to ore formation more  
2124 broadly. In consideration of the current understanding of the northeast Pilbara Craton's major  
2125 ore-forming events, Carlow Castle was most likely produced by the major period of  
2126 hydrothermal base-metal mineralization during the latter stages of the Mallina Basin's  
2127 development between 2955-2940 Ma.

### 2128 3.6.3 Conditions of ore genesis

2129 Whilst there remains some ambiguity regarding the specific tectonic event associated with  
2130 Carlow Castle's ore formation based on age constraints alone, the hydrothermal alteration  
2131 mineral assemblages associated with Carlow Castle's Cu-Co-Au mineralization can provide  
2132 insights into key aspects of its ore genesis. The alteration assemblages present within the  
2133 samples analyzed in this study are consistent with the propylitic hydrothermal alteration of  
2134 basaltic rocks (Corbett & Leach, 1997; Reed, 1997). However, there is clear variation in the  
2135 abundance of the different characteristic propylitic alteration minerals across these samples;  
2136 most notably chlorite, epidote, and actinolite (Table 3.1). The increase in actinolite and epidote  
2137 in sample CC012\_67 relative to samples CC009\_48 and CC003\_61 at the expense of chlorite  
2138 most likely reflects a relative increase in the temperature of alteration; from the middle range  
2139 toward the upper temperature bound (250 to 350°C) of propylitic alteration according to the  
2140 scheme established by Corbett and Leach (1997) based on compiled data from various  
2141 localities. The more chloritic assemblages of CC009\_48 and CC003\_61 are consistent with  
2142 temperatures around 200-250°C, whilst the alteration assemblage of CC012\_67 is within the  
2143 realm of actinolite and epidote stability at temperatures around 280-340°C (Corbett & Leach,  
2144 1997; Henley & Ellis, 1983; Reyes, 1990). These alteration assemblages are comparable with  
2145 those of the contemporaneous base-metal deposits in the Whim Creek greenstone belt produced  
2146 through crustal extension and granitic magmatism within the Mallina Basin between 2955-  
2147 2940 Ma (Hickman, 2016; Pike et al., 2002). Additionally, this temperature estimate for Carlow  
2148 Castle is comparable with the temperature of ore formation ( $295 \pm 35^\circ\text{C}$ ) estimated by Huston  
2149 (2006) for the proximal and contemporaneous Mons Cupri VMS Cu-Zn-Pb deposit.

2150 The consistent prevalence of calc-silicate and carbonate minerals across these samples suggest  
2151 a more-or-less constant fluid chemistry; suggesting that differences in temperature were the

2152 primary factor controlling the variation in alteration mineralogy between samples (Corbett &  
2153 Leach, 1997). In this instance, this mineralogy is characteristic of alteration defined by neutral  
2154 to slightly alkaline fluids (>6 pH), with minimal H<sup>+</sup> metasomatism (Corbett & Leach, 1997;  
2155 Pirajno, 2009; Reed, 1997). The variation in alteration assemblage, and hence fluid  
2156 temperature, appears to relate to the style of sulfide mineralization; with the highest  
2157 temperature alteration assemblage being associated with the highest grade coarse-grained  
2158 ‘massive sulfide’ mineralization and the lower temperature alteration being associated with the  
2159 lower grade quartz-carbonate vein hosted mineralization, irrespective of ore mineralogy  
2160 (Appendix 3.1a, b). It may be that the higher temperature coarse-grained high-grade massive  
2161 sulfide ore formed closer to the center of the Carlow Castle ore body, with the lower  
2162 temperature vein-hosted sulfide mineralization representing more peripheral mineralization at  
2163 the outer edge of the ore body.

#### 2164 3.6.4 Copper-cobalt metallogenesis through geological time

2165 Sediment-hosted Cu-Co deposits comprise the most significant Cu-Co deposits on Earth and  
2166 are the only globally ubiquitous hydrothermal Cu-Co mineralization style (Hitzman et al.,  
2167 2017). The established genetic model for sediment-hosted Cu-Co deposits proposes that they  
2168 form via the migration of oxidizing low temperature (<200°C) meteoric fluids through  
2169 metalliferous sedimentary red beds rich in detrital mafic minerals (Brown, 2014; Hitzman et  
2170 al., 2010). These hydrothermal fluids mobilize and transport Cu and Co in chloride complexes  
2171 and Cu/Co sulfides precipitate upon interaction with a reduced sulfide-rich sedimentary unit,  
2172 such as an organic-rich shale (Brown, 2014). Due to the redox-sensitive nature of Cu and Co,  
2173 sediment-hosted Cu-Co deposits are temporally limited to terranes of maximum  
2174 Paleoproterozoic age; after which the atmosphere was sufficiently globally oxygenated to  
2175 provide a source of oxidized meteoric fluids to mobilize Cu and Co (Figure 3.1; Hitzman et al.,  
2176 2010; Lyons et al., 2014). Similarly, genetic models for other less significant Co deposit types,  
2177 such as polymetallic (Ni-Co-As-Ag-Bi) vein deposits and primary Co-vein deposits (e.g., Bou  
2178 Azzer, Morocco) (Slack et al., 2017a), also invoke oxygenated meteoric fluids as a critical  
2179 component of the mineral system (Ahmed et al., 2009; Kissin, 1992). These deposit types are  
2180 also perhaps more comparable to the Carlow Castle deposit in terms of mineralization style  
2181 and host terrane lithology than sediment-hosted Cu-Co deposits (Fox et al., 2019). However,  
2182 all of these deposits are similarly temporally limited to the Proterozoic onwards (Lefebure,  
2183 1996; Oberthür et al., 2009). Given the apparent paucity of Archean Cu-Co deposits, with the

2184 exception of a few magmatically-associated deposits discussed later in the text, Carlow Castle's  
2185 Mesoarchean age most likely makes it the oldest discovered Cu-Co deposit on Earth. Reflecting  
2186 this it is important to understand the potential metallogenic processes for Cu-Co ore formation,  
2187 and how these processes have evolved through time. Particularly as this could be critical for  
2188 understanding the genesis of, and therefore exploring for, other Archean Cu-Co deposits.

2189 The critical importance of atmospheric oxygen to mobilize metals in hydrothermal Cu-Co  
2190 deposits makes the Mesoarchean age of Carlow Castle particularly notable. One way to account  
2191 for Carlow Castle's genesis would be to suggest that it formed under the influence of  
2192 atmospherically oxygenated fluids; mirroring the younger sediment-hosted Cu-Co and vein-  
2193 hosted Co deposit types. Supporting this, there is growing evidence of temporally and spatially  
2194 isolated oxygenated surface environments prior to the GOE (Eickmann et al., 2018; Lyons et  
2195 al., 2014; Ohmoto, 2020; Planavsky et al., 2014). Additionally, several studies have postulated  
2196 the possibility of pre-GOE oxygenated conditions specifically within the Pilbara Craton during  
2197 the Archean, based on sulfur isotope and mineralogical evidence from several formations  
2198 ranging in age between 3460-2500 Ma (Hickman & Van Kranendonk, 2012; Hoashi et al.,  
2199 2009; Kato et al., 2009; Kaufman et al., 2007; Ohmoto, Watanabe, Ikemi, Poulson, & Taylor,  
2200 2006). If Carlow Castle's genesis was a product of ore-forming processes consistent with those  
2201 younger Cu-Co deposits, this could contribute further evidence of transient locally oxygenated  
2202 conditions during the Archean, which could have implications for global-scale base-metal  
2203 metallogenesis in Archean cratons.

2204 Whilst the most important pathway for Cu-Co mobilization in hydrothermal ore deposits relates  
2205 to the availability of atmospherically-derived oxygenated (sulphate-stable) ore fluids, rare  
2206 examples of hydrothermal Cu-Co mineralization in Archean terranes may provide insights into  
2207 alternative Archean Cu-Co mineralizing processes; involving potential magmatic-  
2208 hydrothermal interactions. These processes are important to consider given Carlow Castle's  
2209 characteristic differences relative to a typical sediment-hosted Cu-Co deposit. Hannan South;  
2210 a Neoproterozoic Au-Cu-(Co-Bi) oxidized endoskarn in the Eastern Goldfields Orogen of the  
2211 Yilgarn Craton in Western Australia provides one example of such processes (Mueller,  
2212 Lawrence, Muhling, & Pooley, 2012). In this instance, the critical processes for metal  
2213 mobilization relate to the transportation of metals within a magmatically-derived hydrothermal  
2214 fluid (Meinert, 1992). These fluids transport dissolved metals by virtue of their highly reactive  
2215 physicochemical properties as they are commonly hot (>400°C), moderately to highly acidic,

2216 moderately to highly saline, and/or oxidized (Meinert, Dipple, & Nicolescu, 2005; Misra,  
2217 2000). However, Carlow Castle lacks several of the defining characteristics of a skarn deposit  
2218 (Misra, 2000).

2219 Cobalt-bearing iron oxide Cu-Au (IOCG) systems of purported Archean age have also been  
2220 identified; most notably the Neoproterozoic Guelb Moghrein Cu-Au deposit in  
2221 Mauritania and the Neoproterozoic Salobo Cu-Au deposit of Carajás IOCG province in Brazil (de  
2222 Melo, Monteiro, Xavier, Moreto, & Santiago, 2019; de Melo et al., 2017; Kirschbaum &  
2223 Hitzman, 2016; Kolb & Petrov, 2016). Models for Salobo attribute its formation to high  
2224 temperature (>500°C), corrosive (HF-bearing), oxidized, and highly saline magmatically-  
2225 derived fluids produced by contemporaneous granitic magmatism (de Melo et al., 2019; de  
2226 Melo et al., 2017). Conversely, models for Guelb Moghrein generally point to a metamorphic  
2227 origin with limited contemporaneous granitic magmatism (Kirschbaum & Hitzman, 2016).  
2228 Instead mineralization was produced by fluid mixing, redox change, and cooling of a hot  
2229 (>400°C) oxidized saline metalliferous metamorphic fluid during retrograde metamorphism.  
2230 However, the mechanism by which this metalliferous metamorphic fluid attained this  
2231 chemistry has not been proposed and the possibility of derivation from an unexposed intrusive  
2232 felsic magmatic source has been acknowledged (Kirschbaum & Hitzman, 2016; Kolb et al.,  
2233 2008; Kolb et al., 2010; Kolb & Petrov, 2016; Sakellaris & Meyer, 2008). Nonetheless, Carlow  
2234 Castle's mineralization similarly lacks several of the diagnostic features of an IOCG deposit  
2235 (Williams et al., 2005).

2236 Given Carlow Castle's probable formation during a period of major rifting with associated  
2237 VMS mineralization, it is also worth considering the Cu-Co-Au mineralizing processes  
2238 associated with VMS formation. Whilst Co-rich VMS deposits are not exceptionally common,  
2239 Besshi-type (mafic-siliciclastic) VMS deposits are observed to be Co-bearing in some cases  
2240 (Hitzman et al., 2017). However, Archean examples of these Co-bearing Besshi-type deposits  
2241 are notably rare (Franklin et al., 2005; Huston, Pehrsson, Eglington, & Zaw, 2010; Schulz,  
2242 2012). The Neoproterozoic (~2.7 Ga) Noranda district, Canada is known to host Cu-Zn-Co-Au  
2243 VMS mineralization of the bimodal-mafic VMS deposit sub-type (Boucher, 2011; Morgan &  
2244 Schulz, 2012). Within the Noranda district, the Ansil deposit is known to contain Co  
2245 mineralization in the form of Co-pentlandite and cobaltite associated with pyrrhotite and  
2246 chalcopyrite, however these are relatively minor in comparison to Carlow Castle's Co  
2247 mineralization (Boucher, 2011; Fox et al., 2019). Combining elements of deposit models

2248 discussed previously, mineralization within the Noranda district has been proposed to have  
2249 formed through a combination of hot (~350-450°C) reduced sulfur-bearing acidic magmatic  
2250 fluids and cool oxidized sulfate-bearing seawater (Barrett, Maclean, & Cattalani, 1991;  
2251 Boucher, 2011; Galley, Jonasson, & Watkinson, 2000; Galley, Watkinson, Jonasson, &  
2252 Riverin, 1995; Sharman, Taylor, Minarik, Dube, & Wing, 2015). The magmatic fluid is  
2253 believed to be derived from the underlying trondhjemitic Flavrian pluton that is coeval with  
2254 mineralization (Sharman et al., 2015; Westendorp & Watkinson, 1991). Whilst a source of  
2255 oxidized Archean seawater has not been accounted for, it has been suggested that seawater  
2256 sulfate was supplied via photochemical oxidation of volcanically degassed SO<sub>2</sub> and that  
2257 seawater sulfate levels through the Neoproterozoic may have been non-negligible (Sharman et al.,  
2258 2015). Once again, magmatic fluids derived from proximal granitic melts appear to have been  
2259 critical to the formation of Noranda's VMS mineralization.

### 2260 3.6.5 Potential models for Carlow Castle Cu-Co-Au ore formation

2261 Hydrothermal Cu-Co deposits through time can be grouped broadly into two categories based  
2262 on their critical ore-forming processes; (i) deposits formed from the Proterozoic onward under  
2263 the influence of oxidized meteoric hydrothermal fluids and (ii) deposits formed from the  
2264 Archean onward as a product of magmatic-hydrothermal processes associated with significant  
2265 crustal magmatism. In light of recent studies purporting to show evidence of potentially  
2266 oxygenated environments on Earth during the Archean, it is tempting to suggest that Carlow  
2267 Castle may have formed in a manner consistent with the Earth's major post-Proterozoic Cu-Co  
2268 deposits; as a product of oxidized meteoric fluids. However, these models of the Earth's early  
2269 atmospheric evolution are far from widely accepted and the relatively high alteration  
2270 temperature provided by this study is not consistent with a relatively cool meteoric ore fluid.  
2271 As such, Carlow Castle may have formed in relation to proximal magmatism in the northwest  
2272 Pilbara; consistent with the other Archean hydrothermal Cu-Co deposits discussed. Supporting  
2273 this, Carlow Castle's formation is contemporaneous with a period of major crustal rifting with  
2274 the formation of a proximal volcanic arc and back-arc basin associated with significant crustal  
2275 melting (Hickman, 2016); a classic environment for VMS ore formation (Franklin et al., 2005).  
2276 Additionally, the formation of other significant base-metal VMS deposits within the proximal  
2277 Whim Creek greenstone belt during this period could support a magmatically-related origin for  
2278 Carlow Castle. The primary issue with this hypothesis is that there is limited evidence of  
2279 contemporaneous granitic magmatism in close spatial association with Carlow Castle. Whilst

2280 Carlow Castle is partially hosted through a tectonized portion of the dominantly mafic igneous  
2281 Ruth Well Formation, these host rocks predate the deposit by ~330 Ma (Hickman, 2016).  
2282 However, the geology underlying Carlow Castle at depth is not well known and the deep crustal  
2283 scale of the pre-existing Regal Thrust host structure makes it a potentially effective conduit for  
2284 regionally and deeply sourced fluids during rifting and magmatism on a broad scale; such as  
2285 from the marine succession of the overlying Gorge Creek Basin and proximal Whim Creek  
2286 greenstone belt and Mallina Basin during melting, extension, and associated VMS formation  
2287 around 2950 Ma. As such, a magmatic association for Carlow Castle-may be the most likely  
2288 mechanism for its formation; in the context of its age, regional setting, and regional  
2289 metallogenesis during its formation. This magmatic-hydrothermal connection to mineralization  
2290 is consistent with the suggestion that Carlow Castle's formation was most likely associated  
2291 with the period (2955-2940 Ma) of major base-metal VMS formation in the proximal Whim  
2292 Creek greenstone belt during the late stages of the development of the De Grey Superbasin.

### 2293 3.7 Conclusions

2294 Newly presented apatite U-Pb geochronology indicate that the Carlow Castle deposit formed  
2295 at  $2957 \pm 67$  Ma, hundreds of millions of years before any comparable Cu-Co-Au deposit; most  
2296 of which are of maximum Proterozoic age. Due to the critical role that atmospherically oxidized  
2297 ore fluids play in mobilizing these ore metals to form these deposits, the formation of the  
2298 Carlow Castle deposit in the Mesoarchean before widespread oxygenation of the atmosphere,  
2299 provides evidence for alternative and comparably uncommon processes for hydrothermal Cu-  
2300 Co metallogenesis. Additionally, new constraints on the physiochemical properties of Carlow  
2301 Castle's hydrothermal ore fluid are provided; suggesting that it formed at peak temperatures  
2302 ~300°C under near neutral fluid conditions. In this instance a magmatic origin is tentatively  
2303 suggested given Carlow Castle's regional setting, Mesoarchean age, and broad similarities to  
2304 proximal and contemporaneous magmatic-hydrothermal ore deposits. However, it is  
2305 acknowledged that direct evidence of a magmatic source is not apparent. Carlow Castle's  
2306 mineralization was most likely related to the significant contemporaneous magmatic activity  
2307 and associated base-metal VMS mineralization within the proximal Whim Creek greenstone  
2308 belt between 2955-2940 Ma. A magmatic-hydrothermal origin is consistent with other rare  
2309 Archean hydrothermal Cu-Co deposits. Carlow Castle's occurrence suggests that similar  
2310 hydrothermal Cu-Co-Au deposits in Archean greenstone belts could represent novel targets for  
2311 future Cu-Co exploration, which has traditionally focused on Proterozoic and Phanerozoic

2312 basins. Nonetheless, the uncertainties surrounding Carlow Castle's ore-forming processes  
2313 suggest that further research is required to understand its genesis more fully.

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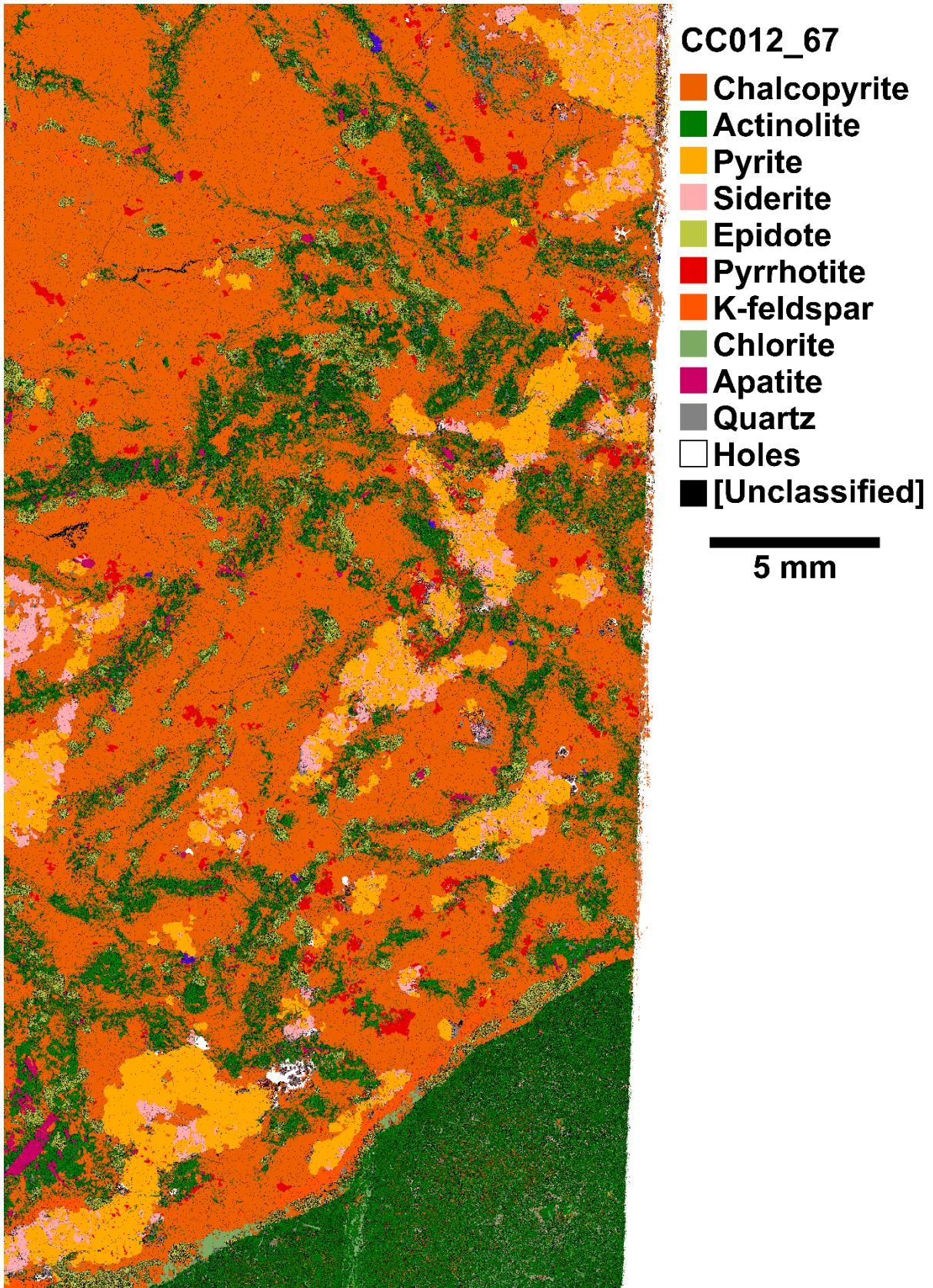
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2550 *Volume*: Society of Economic Geologists.  
2551

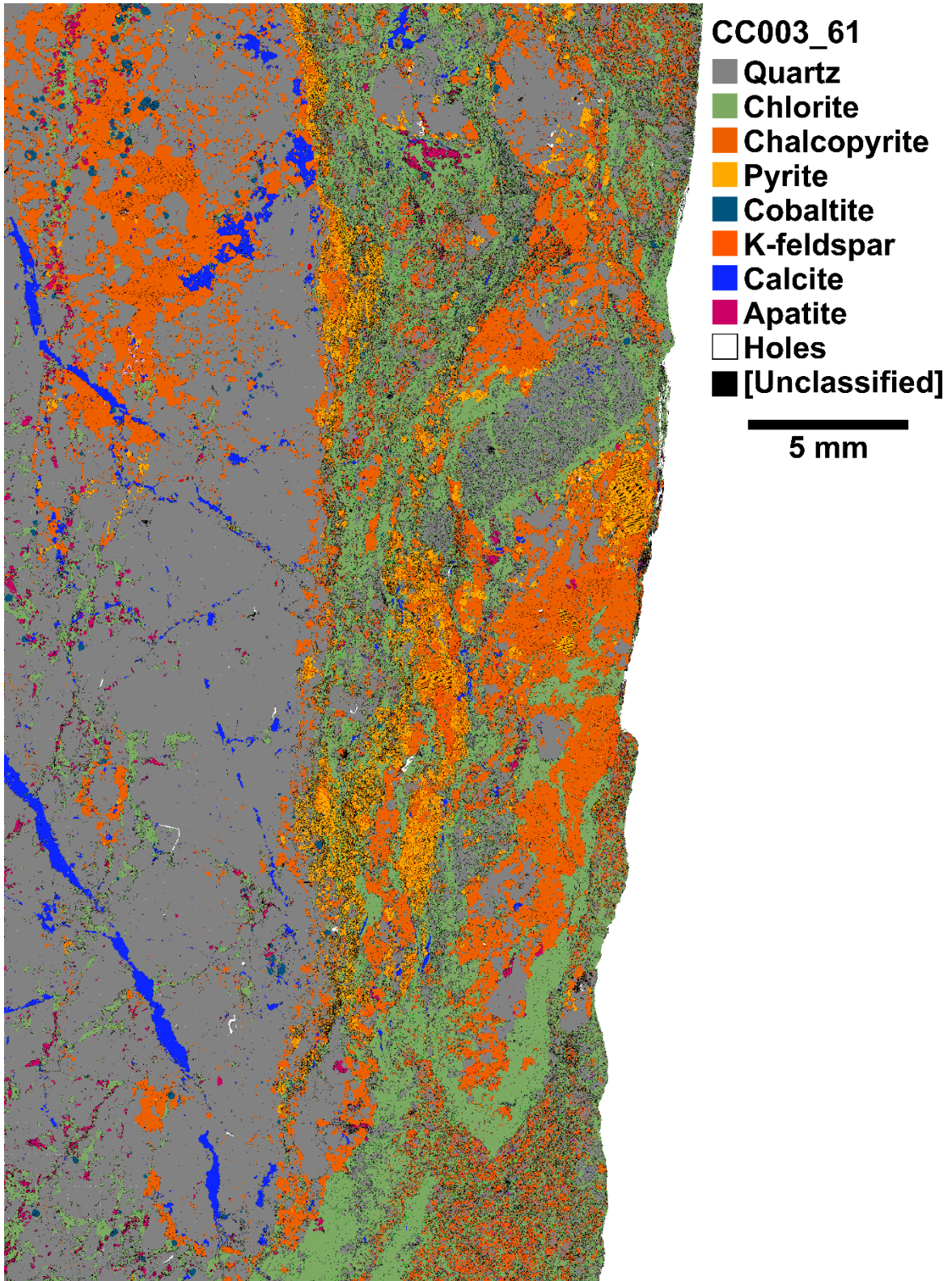
2552 Appendix 3.1 – Quantitative mineralogy of representative samples from  
2553 Carlow Castle Cu-Co-Au deposit



2554



2555 Appendix 3.1a – Automated mineralogy map of sample CC012\_67 from ore mineral  
2556 Assemblage One of Carlow Castle. Chalcopyrite, pyrite, and pyrrhotite are the  
2557 predominant ore minerals within this sample. Within this sample, these sulphide minerals  
2558 are massive and relatively coarse-grained. Sulphide minerals are intergrown with  
2559 actinolite, epidote, and minor chlorite throughout the sample, which occur as  
2560 hydrothermal alteration minerals. Additionally, these alteration minerals are commonly  
2561 intergrown with apatite, which tends to form either as isolated euhedral grains throughout  
2562 the massive sulphide or as clusters of elongate grains that are commonly intergrown with  
2563 actinolite and epidote.  
2564



2565

2566 Appendix 3.1b – Automated mineralogy map of sample CC003\_61 from ore mineral

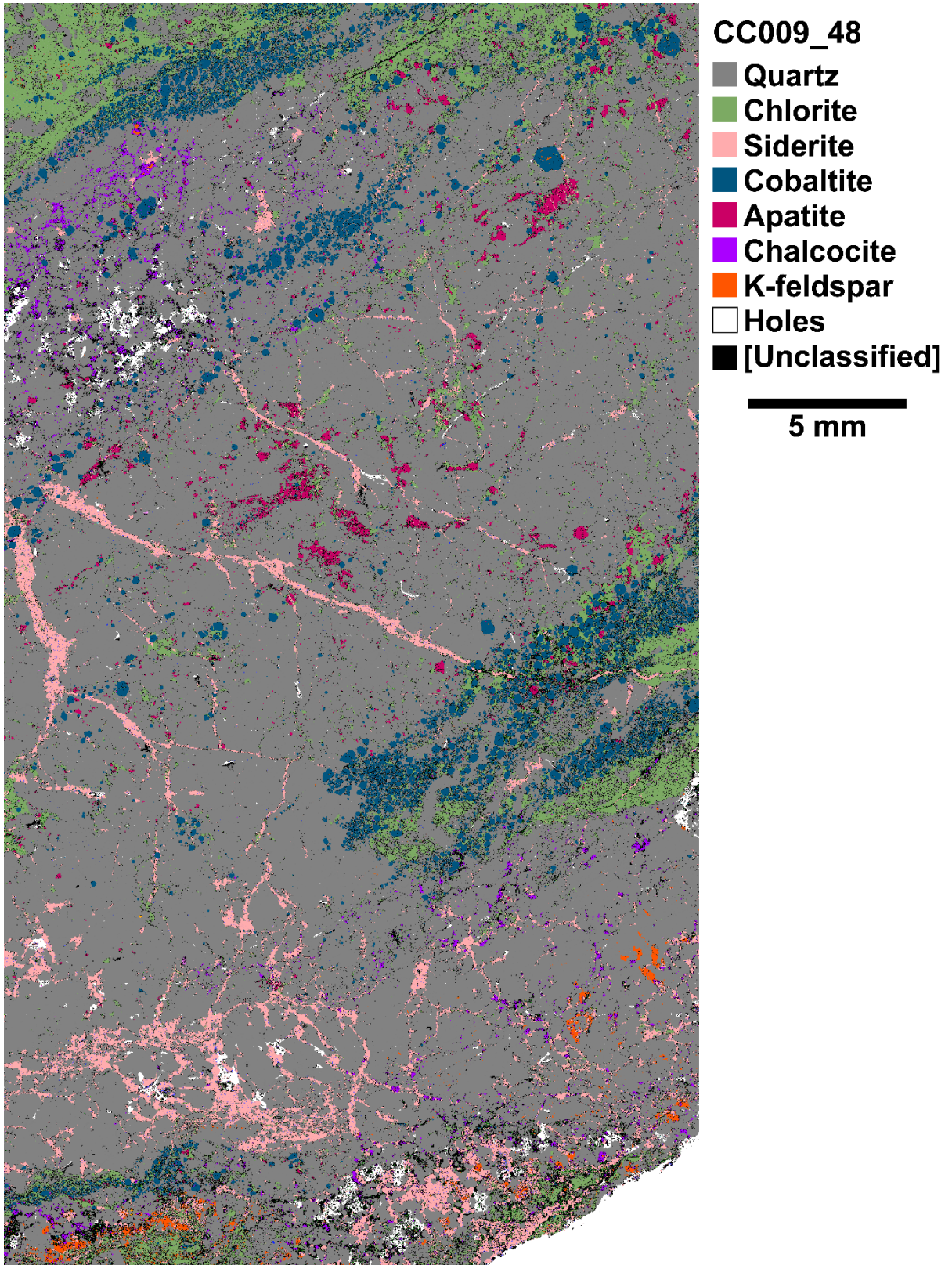
2567 Assemblage One of Carlow Castle. Chalcopyrite and pyrite constitute the primary ore

2568 minerals within this sample, with cobaltite occurring as a minor constituent. Note the

2569 finer-grained nature of these sulphides and lower grade by comparison to CC012\_67.

2570 Additionally, CC003\_61 is comparably richer in quartz, calcite, and chlorite and contains  
2571 comparably little actinolite or epidote. Apatite appears to commonly occur in associated  
2572 with chlorite.

2573



2574

2575 Appendix 3.1c – Automated mineralogy map of sample CC009\_48 from ore mineral  
 2576 Assemblage Two of Carlow Castle. By contrast to samples CC012\_67 and CC003\_61,  
 2577 the sulphide mineralogy of sample CC009\_48 is composed predominantly of cobaltite

2578 and chalcocite with minimal pyrite or chalcopyrite. However, similarly to sample  
2579 CC003\_61, chlorite is the predominant alteration mineral along with minor K-feldspar  
2580 and apatite. These ore and alteration minerals are hosted within a quartz-siderite vein.

2581

2582 **Appendix 3.2 – Apatite U-Pb LA-ICP-MS data**

2583 Appendix 3.2 is accessible via this [link](#) and provides an .xlsx file containing U-Pb  
2584 geochronology data for hydrothermal apatite analysed in this study.

2585

2586 **Appendix 3.3 – Apatite major and trace element LA-ICP-MS data**

2587 Appendix 3.3 is accessible via this [link](#) and provides an .xlsx file containing major and trace  
2588 element geochemical data for hydrothermal apatite analysed in this study.

2589

## 2590 Chapter 4

# 2591 Mesoarchean oxygenation 2592 accompanied massive copper-cobalt 2593 mineralization

2594

2595 This chapter is currently under review for publication in *Science Advances*.

2596

### 2597 4.1 Abstract

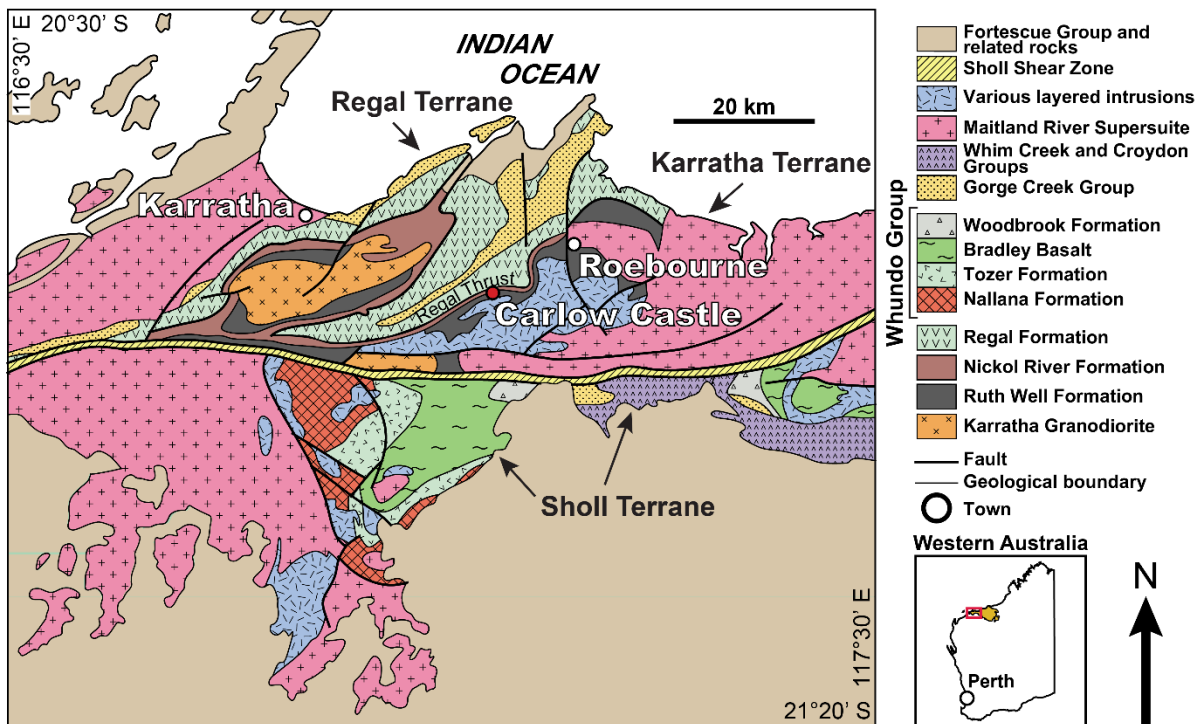
2598 Major hydrothermal Cu-Co mineralization in sedimentary rocks commenced following the  
2599 Great Oxidation Event, 2.4 billion years ago. The emergence of Cu-Co deposits reflects the  
2600 primary processes by which these metals are mobilized in Earth's crust, principally by oxidized  
2601 fluids. However, a recently discovered significant Mesoarchean (~2.95 Ga) Cu-Co-Au deposit  
2602 in the Pilbara Craton of Western Australia is difficult to reconcile within the traditional view  
2603 of a pervasively and persistently anoxic Archean world. Here we show sulfur isotope signatures  
2604 in the deposit lack mass-independent fractionation of anoxic Archean atmospheric sulfur. In-  
2605 situ triple sulfur isotope ( $^{32}\text{S}$ ,  $^{33}\text{S}$ ,  $^{34}\text{S}$ ) analysis of sulfide mineralization displays  $\delta^{34}\text{S}$  values  
2606 between -3.4 and 6.1‰, and critically,  $\Delta^{33}\text{S}$  values between -0.17‰ and 0.12‰; within the  
2607 defined bounds of mass-dependent sulfur fractionation. These sulfur isotopic data are atypical  
2608 for Archean sulfides and, combined with thermodynamic modelling, suggest that Cu-Co  
2609 mobilization occurred under oxidized fluid conditions. Cu-Co ore formation coincided with  
2610 well-established ephemeral oxygenic conditions around 2.95 Ga, which directly preceded and  
2611 potentially drove atmospheric changes that resulted in Earth's earliest major glaciation. Our  
2612 findings are thus consistent with an emerging model that suggests the Mesoarchean atmosphere  
2613 was subject to localized, transiently oxidized conditions following the emergence of oxygenic  
2614 photosynthesis, sufficient to enable large-scale base-metal ore formation.



## 2615 4.2 Introduction

2616 There are very few examples of volumetrically large hydrothermal Cu-Co mineralization in the  
2617 Archean, with the most significant examples of this mineralization style typically limited to  
2618 basins no older than the Proterozoic (Hitzman et al., 2017; Hitzman et al., 2010). A reducing  
2619 atmosphere meant that basinal fluids during the Archean were generally unable to mobilize  
2620 strongly redox-sensitive metals like Cu and Co (Brown, 2014). However, the formation of the  
2621 recently discovered Carlow Castle Cu-Co-Au deposit within the Pilbara Craton during the  
2622 Mesoarchean necessitates large-scale Cu-Co mobility well into the Archean (Fox et al., 2021;  
2623 Fox et al., 2019).

2624 Carlow Castle Cu-Co-Au deposit is hosted in a Mesoarchean (3.07 Ga) orogenic thrust fault  
2625 (Regal Thrust) through 3.28-3.16 Ga volcano-sedimentary rocks (peridotite, basalt, chert,  
2626 carbonaceous shale) of the Roebourne greenstone belt of the West Pilbara Superterrane (3.28-  
2627 3.066 Ga); an accretionary superterrane in northwest Western Australia (Figure 4.1; Hickman,  
2628 2016). Recent U-Pb apatite dating demonstrated that the deposit formed around ~2.955 Ga  
2629 during major post-orogenic rifting and the formation of the proximal De Grey Superbasin  
2630 (3.066-2.919 Ga) (Fox et al., 2021). The Carlow Castle ore deposit's exceptionally old age of  
2631 mineralization and distinctive ore deposit character make it unique in the global context of Cu-  
2632 Co mineralization (Fox et al., 2021). Detailed characterization of the Carlow Castle ore  
2633 mineralization is given in Fox et al. (2019).

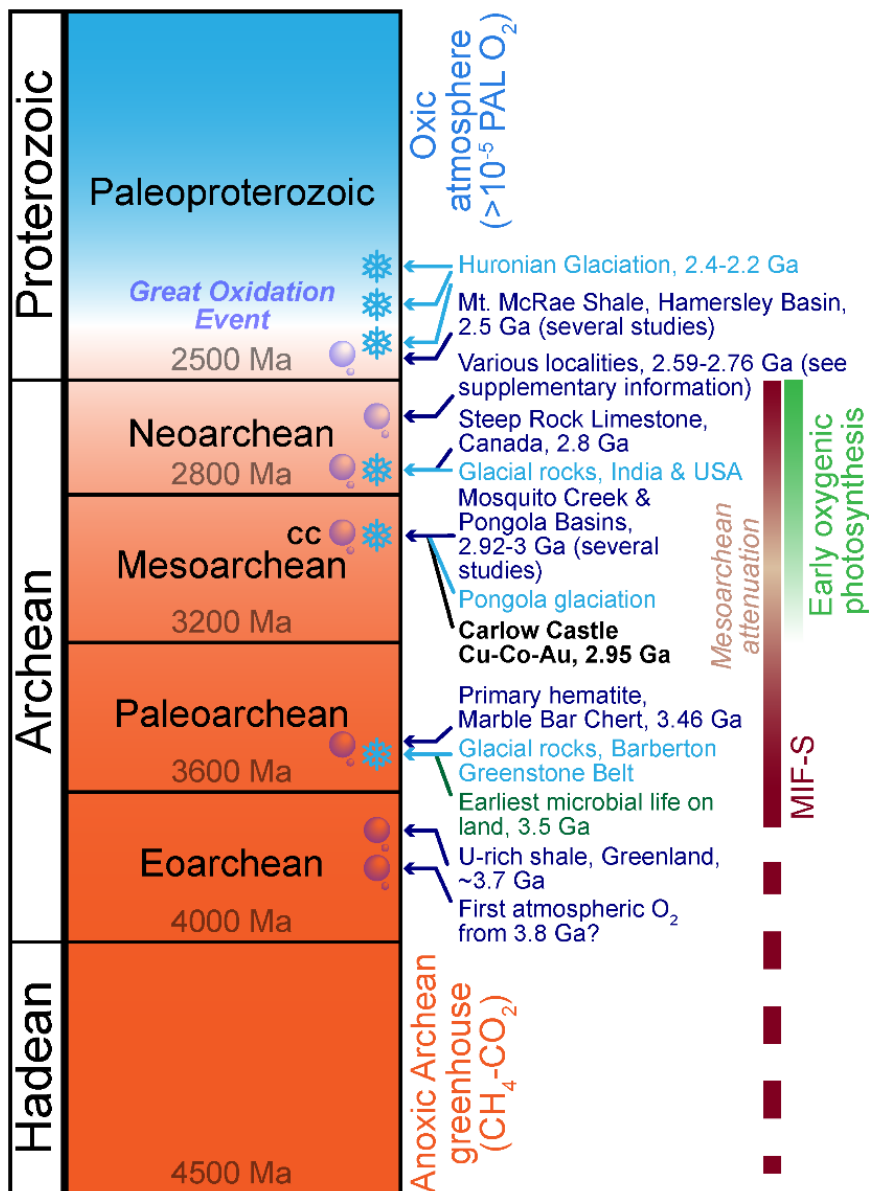


2634  
 2635 Figure 4.1 - Geologic map of the northwest Pilbara Craton. The Carlow Castle ore deposit  
 2636 site is indicated by the red circle. The Pilbara Craton is highlighted in orange within the  
 2637 inset map. Modified after Van Kranendonk et al. (2002).

2638 Given difficulties explaining Cu-Co ore formation within the context of an anoxic Archean  
 2639 atmosphere, the concentration of Cu and Co at ~2.95 Ga (Fox et al., 2021) within the Carlow  
 2640 Castle deposit may support a model of spatially and temporally limited periods of oxygenation  
 2641 during the Mesoarchean. Localized transient oxygenated conditions during the Archean have  
 2642 been demonstrated recently in several studies from the Pilbara Craton and other Archean  
 2643 terranes (Figure 4.2; Appendix 4.1; Anbar et al., 2007; Catling & Zahnle, 2020; Eickmann et  
 2644 al., 2018; Frei, Gaucher, Poulton, & Canfield, 2009; Hickman & Van Kranendonk, 2012;  
 2645 Hoashi et al., 2009; Kato et al., 2009; Kaufman et al., 2007; Lyons & Gill, 2010; Lyons et al.,  
 2646 2014; Ono, Beukes, Rumble, & Fogel, 2006; Ossa Ossa et al., 2016; Ostrander, Johnson, &  
 2647 Anbar, 2021; Planavsky et al., 2014; Reinhard, Raiswell, Scott, Anbar, & Lyons, 2009; Wille  
 2648 et al., 2007), with some studies even suggesting the possibility of extended periods of pervasive  
 2649 Archean oxygenation (Ohmoto, 2020; Ohmoto et al., 2006). Whilst it is possible for Cu-Co  
 2650 transport to occur under reduced hydrothermal conditions, this would typically require high  
 2651 fluid temperatures or acidic fluid pH as an alternative mechanism for metal solubility (Brown,  
 2652 2014; Rose, 1989), however such conditions are not supported by evidence from alteration  
 2653 patterns at Carlow Castle (Fox et al., 2021).

2654 Multiple sulfur isotope analysis has been a key tool in many studies of the evolution of Earth's  
2655 atmosphere and sulfur cycle, especially during the transition between the Archean and  
2656 Proterozoic (Farquhar, Bao, & Thiemens, 2000). This is due to mass-independent fractionation  
2657 (MIF-S) signatures imparted on atmospheric sulfur before the development of an atmospheric  
2658 ozone shield, evidenced by the large  $\Delta^{33}\text{S}$  fractionation within Archean sulfides (Farquhar &  
2659 Wing, 2003). However, the attenuation of this  $\Delta^{33}\text{S}$  MIF-S signal during the Mesoarchean,  
2660 coinciding with the early emergence of oxygenic photosynthesis and the Pongola Glaciation  
2661 (Catling & Zahnle, 2020; Kasting & Ono, 2006) has been interpreted to reflect a volatile,  
2662 evolving early atmosphere that may have accumulated sufficient oxygen, whether ephemerally  
2663 or persistently, to dampen MIF-S signals and displace the methane-rich Archean atmospheric  
2664 greenhouse (Ono et al., 2006).

2665 Given the importance of atmospheric oxygen for tracking the evolution of life, the biosphere,  
2666 climate, element cycling, and ore deposit genesis, we utilized multiple sulfur isotopic analysis  
2667 to resolve atmospheric conditions during the formation of Cu-Co-Au mineralization at Carlow  
2668 Castle. This isotopic investigation was supported by thermodynamic modelling to constrain the  
2669 conditions of the Carlow Castle hydrothermal ore-forming system.



2670

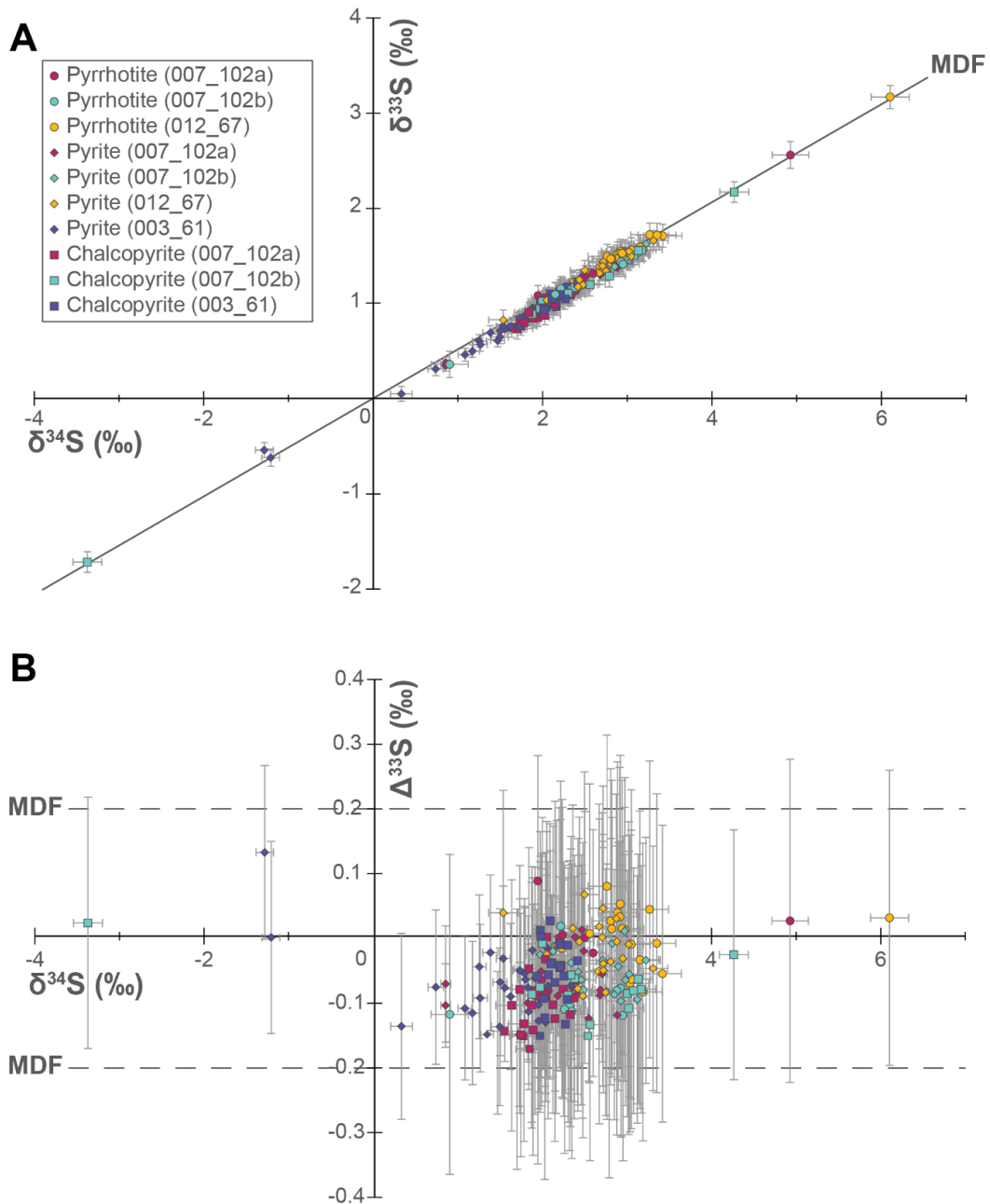
2671 Figure 4.2 - Compilation of studies purporting to show evidence of 'whiffs' of oxygen  
 2672 before the Great Oxidation Event and Carlow Castle's relative temporal position. See full  
 2673 list of studies compiled in Appendix 4.1. PAL = present atmospheric level, MIF-S = mass-  
 2674 independent sulfur fractionation.

2675 **4.3 Results**

2676 **4.3.1 Sulfur Isotope characteristics of Carlow Castle Cu-Co-Au deposit**

2677 The  $\delta^{34}\text{S}$  values of all measured mineral phases across four samples define a relatively wide  
 2678 range, between -3.4‰ and 6.1‰ (Figure 4.3a). The mean and median values are 2.3‰, and  
 2679 standard deviation is 0.86‰. Fractionation between samples is limited, with CC012\_67  
 2680 showing the highest mean  $\delta^{34}\text{S}$  of 2.9‰ and CC003\_61 showing the lowest mean  $\delta^{34}\text{S}$  of

2681 2.06‰. Fractionation between phases within a given sample is also limited; the largest  
2682 variation in mean  $\delta^{34}\text{S}$  values between phases in a given sample is 0.8‰ between chalcopyrite  
2683 (2.2‰) and pyrite (1.4‰) in CC003\_61. The  $\Delta^{33}\text{S}$  values show minimal variance; between -  
2684 0.17‰ and 0.12‰, with a mean of -0.05‰, median of -0.06, and standard deviation of 0.05‰.  
2685 All 183 measurements fall within the traditional bounds of mass-dependent sulfur fractionation  
2686 (MDF-S) of  $0 \pm 0.2\text{‰}$  and plot along the MDF-S line (Figure 4.3b). Full results are provided  
2687 in Appendix 4.2.



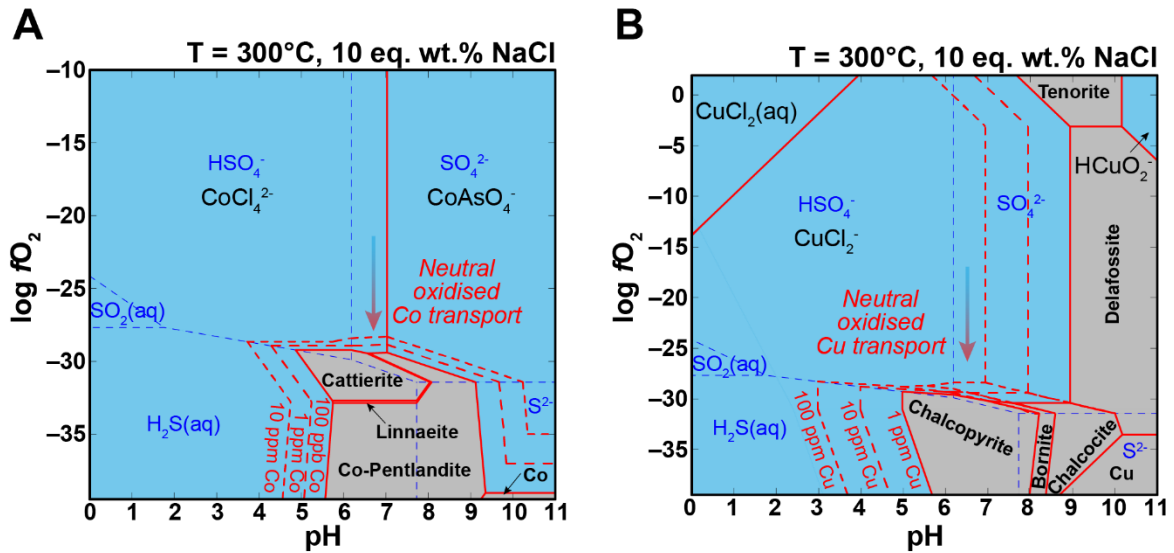
2688

2689 Figure 4.3 - (a, b) Multiple sulfur isotopes composition of Carlow Castle ore minerals  
 2690 (pyrrhotite, pyrite, and chalcopyrite). Note the clear mass-dependent fractionation pattern  
 2691 of  $\Delta^{33}\text{S}$ . Error bars are  $2\sigma$ , as calculated according to method documented in LaFlamme  
 2692 et al. (2016).

### 2693 4.3.2 Hydrothermal Cu-Co mobility

2694 At a peak fluid temperature of  $300^\circ\text{C}$  and neutral pH (6.0 at  $300^\circ\text{C}$ ), consistent with estimated  
 2695 conditions of Carlow Castle deposit's ore formation based on its propylitic hydrothermal  
 2696 alteration assemblage (Fox et al., 2021), the solubility of Co and Cu are controlled by  $\log f(\text{O}_2)$

2697 (Figure 4.4). Under oxidising conditions, at equilibrium with  $\text{HSO}_4^-/\text{SO}_4^{2-}$ , Co is mobile as  
 2698  $\text{CoCl}_4^{2-}$  and Cu is mobile as  $\text{CuCl}_2^-$  complexes. Under increasingly reducing conditions Cu and  
 2699 Co become progressively less soluble and immobile at equilibrium with  $\text{H}_2\text{S}$  or  $\text{S}^{2-}$ , where they  
 2700 are stable as sulfides/sulfosalts. As such, rapid precipitation of Cu and Co from a 300°C,  
 2701 neutral, oxidized fluid could be induced through a sudden change in redox conditions and  
 2702 reduction of  $\text{SO}_4^{2-}$  to  $\text{H}_2\text{S}$ .



2703  
 2704 Figure 4.4 - Stability fields for selected minerals and dissolved species along with  
 2705 solubility contours for Co (a) and Cu (b) as a function of pH and oxygen fugacity at 300°C  
 2706 and water-saturated pressure. Both Co and Cu are mobile (as  $\text{Cl}^-$  complexes) under  
 2707 oxidized conditions with metal concentration increasing with acidity. Note that the  
 2708 dominance of propylitic hydrothermal alteration acts as a limit on minimum ore fluid pH.  
 2709 Neutral pH at 300°C is 6.0. Metals precipitate as sulfides/sulfosalts upon reduction. Fluid  
 2710 compositions:  $\text{S}_{\text{total}} = 0.01$  molal,  $\text{Cl}_{\text{total}} = 2.0$  molal,  $\text{As}_{\text{total}} = 0.001$  molal,  $\text{Fe}_{\text{total}} = 0.01$   
 2711 molal.

## 2712 4.4 Discussion

### 2713 4.4.1 Sulfur source and absence of 'Archean sulfur'

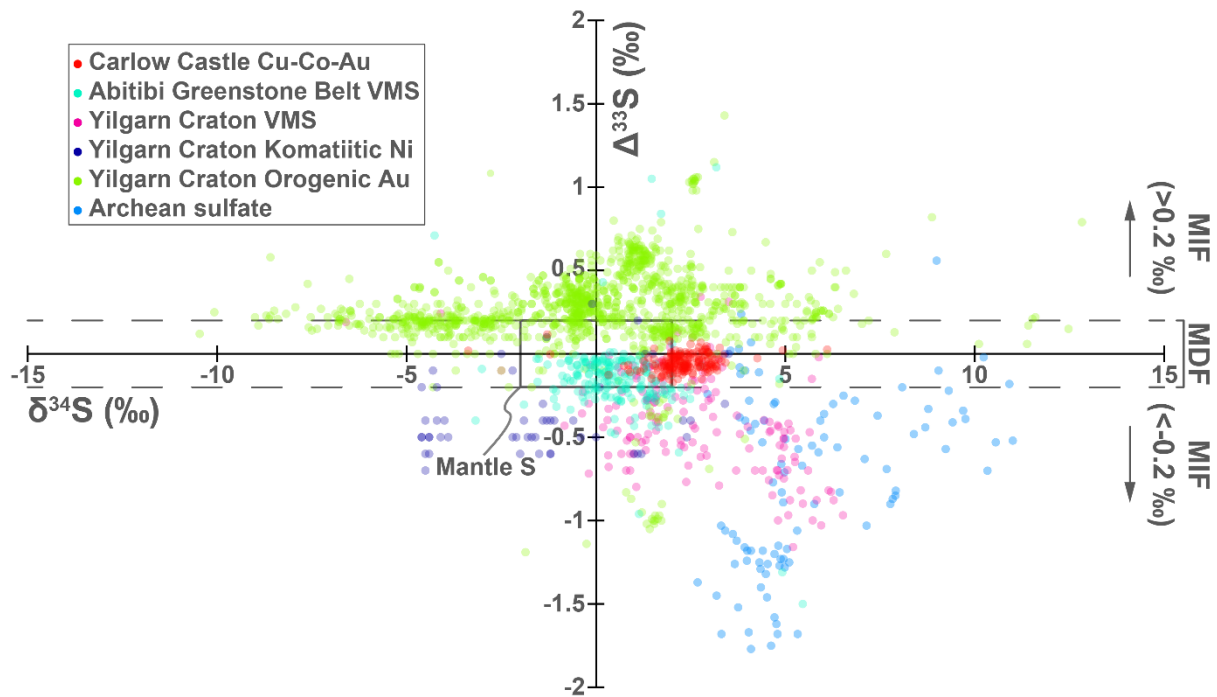
2714 The sulfur source most commonly implicated in the formation of typical post-Archean  
 2715 sediment-hosted Cu-Co deposits is a combination of evaporite-derived seawater sulfate and  
 2716 sedimentary sulfide produced by microbial sulfate reduction. Both sources tend to produce  
 2717 highly fractionated  $\delta^{34}\text{S}$  signatures in derived sulfide ores (Hitzman et al., 2005). Conversely,  
 2718 the  $\delta^{34}\text{S}$  signature of Archean sulfides generally exhibit extremely limited fractionation as

2719 Mesoarchean seawater sulfate and sedimentary sulfides had yet to diverge significantly from  
2720 bulk mantle compositions (Canfield, 2004). Reflecting this, the spread of most  $\delta^{34}\text{S}$  signatures  
2721 around 2‰ in the Carlow Castle deposit's sulfide mineralization is typical of Archean sulfides  
2722 from various environments, including in sulfide ore deposits, sedimentary sulfides, and  
2723 magmatic sulfides (Farquhar, Wu, Canfield, & Oduro, 2010; Johnston, 2011). This could be  
2724 interpreted to reflect a homogenous mantle origin for sulfur in the ore deposit ( $\delta^{34}\text{S} = 0 \pm 2\text{‰}$   
2725 (Seal, 2006)). Alternatively, the spread of the majority of data points beyond the typical bounds  
2726 of mantle sulfur (130 of 183 data points), combined with the otherwise attenuated  $\delta^{34}\text{S}$   
2727 signatures of Archean sulfur reservoirs, could be indicative of sulfur input from a slightly heavy  
2728 Archean seawater sulfate source ( $\delta^{34}\text{S} = 3$  to  $8\text{‰}$ ) (Muller, Philippot, Bard-Rollion, & Cartigny,  
2729 2016) into the Carlow Castle ore body. This latter interpretation of the  $\delta^{34}\text{S}$  signature would  
2730 support a model more akin to conventional post-Archean Cu-Co mineralization; necessitating  
2731 an oxidized seawater sulfate-bearing ore fluid to enable Cu-Co mobility.

2732 The  $\Delta^{33}\text{S}$  signatures in our data are characteristic of mass-dependent fractionation processes  
2733 and there is no evidence of any MIF-S signature indicative of typical Archean sulfur. By  
2734 comparison, it is relatively rare to observe Archean hydrothermal ore deposits with no  
2735 incorporation of MIF-S (Figure 4.5; Appendix 4.3). It is unlikely that the Carlow Castle  
2736 deposit's lack of MIF-S can be explained purely through a magmatic sulfur source. Particularly  
2737 as there are no mantle-derived rocks that are contemporaneous with, and proximal to the  
2738 orebody mineralization, to allow for direct input of mantle sulfur without contamination from  
2739 other Archean sulfur sources during ore fluid migration (Fox et al., 2021). However, sulfur  
2740 derived from Archean sedimentary sulfide or seawater sulfate sources would be expected to  
2741 exhibit significant MIF-S signatures, unless such oxidized ore fluids were also associated with  
2742 a weakly oxygenated atmosphere that temporarily inhibited production or preservation of large  
2743 magnitude MIF-S. This adapts the model presented by Kasting and Ono (2006); Ono et al.  
2744 (2006) to account for attenuated MIF-S signatures evident within the Mesoarchean sedimentary  
2745 record. As noted by Wang et al. (2018), the otherwise very small atmospheric  $\text{O}_2$  reservoir of  
2746 the Archean would have likely been highly dynamic and subject to significant variability  
2747 following the evolution of oxygenic photosynthesis. As such, even subtle oxygen fluxes may  
2748 have been sufficient to significantly alter Mesoarchean atmospheric chemistry and temporarily  
2749 attenuate MIF-S signatures. With this considered, the small but consistently negative  $\Delta^{33}\text{S}$   
2750 values (156 of 183 data points) displayed by the Carlow Castle deposit's ore sulfides are



2751 consistent with the slightly heavy  $\delta^{34}\text{S}$  and negative  $\Delta^{33}\text{S}$  signals that are characteristic of  
2752 Archean sulfate (Figure 4.5), but with heavily attenuated  $\Delta^{33}\text{S}$  values (Johnston, 2011).



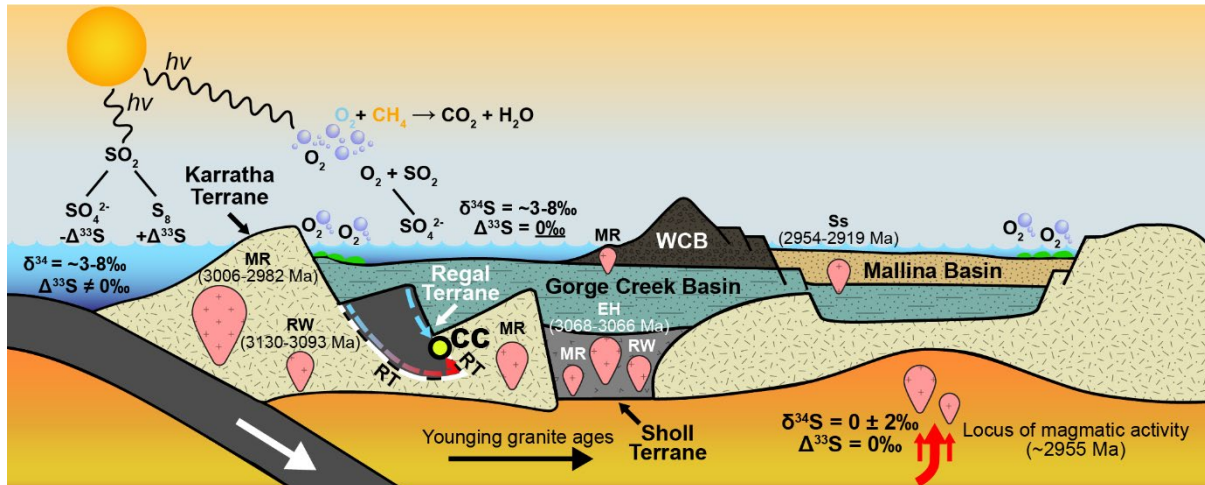
2753

2754 Figure 4.5 - Comparison of Archean greenstone-hosted hydrothermal and magmatic ore  
2755 deposits with Carlow Castle deposit (red). Note that Carlow Castle is the only deposit  
2756 with no clear MIF-S. Mantle sulfur range is marked by the grey box. Range of Archean  
2757  $\text{SO}_4^{2-}$  sulfur isotope compositions from Muller et al. (2016). Ore deposit sulfur isotope  
2758 data compiled from various sources in Appendix 4.3.

#### 2759 4.4.2 Cu-Co mobility during Mesoarchean rifting

2760 Given the established physicochemical constraints on Carlow Castle's hydrothermal system,  
2761 our modelling shows that oxidized sulfate-bearing ore fluids were the fundamental chemical  
2762 pathway to enable Cu-Co mobility prior to reductive precipitation of metal sulfide ore minerals  
2763 upon interaction with a suitable redox buffer, i.e., mafic volcanics or carbonaceous shales  
2764 (Figure 4.4). Therefore, if the Mesoarchean Pilbara Craton's atmosphere was sufficiently  
2765 oxidized to inhibit the generation of MIF-S, we argue an oxidized seawater origin for the  
2766 Carlow Castle deposit is possible. The shallow water margins of the overlying marine Gorge  
2767 Creek Basin and Whim Creek volcanic arc within the De Grey Superbasin would be a potential  
2768 source of seawater sulfate-bearing oxygenated fluids during a period of major crustal extension  
2769 and increased crustal heat flow (Fox et al., 2021; Hickman, 2016). In support of this  
2770 paleogeographic setting, there is a preponderance of evidence that supports the existence of

2771 isolated photosynthetically oxygenated shallow marine environments from at least ~2.95 Ga  
 2772 (Eickmann et al., 2018; Planavsky et al., 2014; Wang et al., 2020). Following this, an  
 2773 atmosphere briefly oxygenated sufficiently to inhibit or attenuate MIF-S signatures could  
 2774 account for the atypical lack of MIF-S in the Carlow Castle deposit's ore sulfides in a manner  
 2775 that is also consistent with conventional post-Archean models of Cu-Co metallogenesis and the  
 2776 results of thermodynamic modelling in this study; relying on redox processes as a mechanism  
 2777 to mobilize Cu and Co. This model is summarised schematically in Figure 4.6.

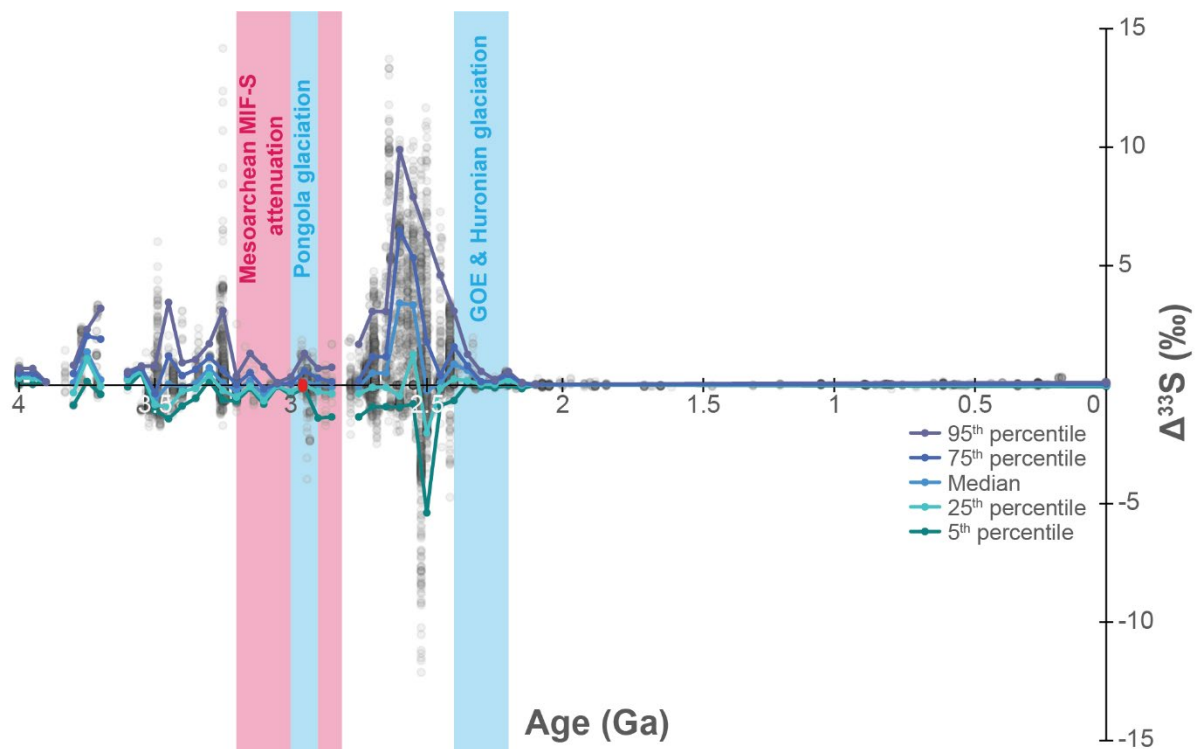


2778  
 2779 Figure 4.6 - Conceptual schematic model of the northwest Pilbara Craton during the  
 2780 formation of Carlow Castle (CC) and the De Grey Superbasin around 2.955 Ga. Note:  
 2781 MR = Maitland River Supersuite, RW = Railway Supersuite, EH = Elizabeth Hill  
 2782 Supersuite, Ss = Sisters Supersuite, WCB = Whim Creek Basin.

#### 2783 4.4.3 The Mesoarchean MIF-S record

2784 The broad attenuation of MIF-S signatures throughout the Mesoarchean (Figure 4.7) is a widely  
 2785 discussed aspect of the Archean sedimentary sulfur isotopic record (Johnston, 2011). This  
 2786 phenomenon has been attributed to a variety of processes including variability in volcanic SO<sub>2</sub>  
 2787 output (Ohmoto et al., 2006), the oxidation state of volcanic sulfur volatiles (Halevy, Johnston,  
 2788 & Schrag, 2010), changes in anoxic atmospheric chemistry (Domagal-Goldman, Kasting,  
 2789 Johnston, & Farquhar, 2008; Farquhar et al., 2007; Thomazo, Ader, Farquhar, & Philippot,  
 2790 2009), dilution by crustal MDF-S sources (Guy et al., 2012), and the potential for transient or  
 2791 prolonged oxygenation of the Mesoarchean atmosphere associated with the early emergence  
 2792 of oxygenic photosynthesis (Ohmoto et al., 2006; Ono et al., 2006; Wang et al., 2018). The  
 2793 latter argument has been disputed as some studies have demonstrated the preservation of MIF-  
 2794 S signatures through this period (Farquhar et al., 2007; Guy et al., 2012). Nonetheless, a notable

2795 reduction in the magnitude of MIF-S is evident through the Mesoarchean, with significant MIF-  
2796 S values generally being outliers during this period (Figure 4.7). This also coincided with a  
2797 shift in uranium isotope systematics observed by Wang et al. (2018), which is similarly  
2798 attributed to the evolution of oxygenic photosynthesis. It should be noted that once a large  
2799 reservoir of MIF-S had been established in the Archean crust it is likely to have been recycled  
2800 and therefore evident within the sulfur isotope record long after the generation of new large  
2801 MIF-S anomalies had been attenuated due to a rise in atmospheric O<sub>2</sub> levels toward the critical  
2802 10<sup>-5</sup> PAL (Philippot et al., 2018; Reinhard, Planavsky, & Lyons, 2013). This ‘crustal memory  
2803 effect’ of MIF-S is also evident immediately following the GOE and implies that any  
2804 attenuation of MIF-S evident in the sulfur isotope record, such as the one displayed through  
2805 the Mesoarchean and especially around 2.95 Ga, is likely to understate the relative decrease in  
2806 the magnitude of new attenuated MIF-S signatures generated due to contamination with the  
2807 existing crustal reservoir of large MIF-S signatures. As such, the occurrence of some larger  
2808 magnitude outlier MIF-S signatures within the Mesoarchean rock record doesn’t necessarily  
2809 discount the possibility of changes in atmospheric chemistry that are suggested by the general  
2810 attenuation of these MIF-S values. Further, in those studies purporting to show evidence of the  
2811 persistence of MIF-S through the Mesoarchean (Farquhar et al., 2007; Guy et al., 2012), it is  
2812 notable that data from rocks coeval with glacial tillite and diamictites from ~2.95 Ga also  
2813 generally show significantly attenuated MIF-S values broadly consistent with those presented  
2814 by Ono et al. (2006). This would effectively support a connection between attenuated MIF-S  
2815 signatures, changing atmospheric chemistry, and the Pongola Glaciation around 2.95 Ga.



2816

2817 Figure 4.7 - Compilation of sedimentary sulfur isotope data through geological time (gray  
 2818 dots) compared to Carlow Castle sulfur isotope data (red dots). The colored lines here  
 2819 indicate the percentile values of these data, binned in periods of 50 My from 4 Ga through  
 2820 to 2.1 Ga. A clear minimum in the magnitude of MIF-S values is observable through the  
 2821 Mesoarchean, most likely coinciding with the first evolution of oxygenic photosynthesis.  
 2822 Note gaps here are a result of no data being available for a given 50 My period.  
 2823 Additionally, some time periods are biased by relatively small data sets. Modified from  
 2824 data compiled by Killingsworth et al. (2019) and Selvaraja, Caruso, Fiorentini, and  
 2825 LaFlamme (2017).

#### 2826 4.4.4 A whiff of oxygen at ~2.95 Ga

2827 Here we suggest that the Carlow Castle deposit's mineralization and MDF-S signature is  
 2828 consistent with a 'whiff' of oxygen sufficient to destabilise Mesoarchean atmospheric anoxia.  
 2829 This is asserted based on (i) the absence of proximal and contemporaneous mantle-derived  
 2830 rocks to provide a mantle MDF-S source; (ii) the spread of  $\delta^{34}\text{S}$  values beyond the generally  
 2831 accepted bounds of mantle sulfur ( $0 \pm 2\text{‰}$ ), and predominantly small negative  $\Delta^{33}\text{S}$  values  
 2832 consistent with an attenuated Mesoarchean seawater  $\text{SO}_4^{2-}$  source (Figure 4.3); (iii) the rarity  
 2833 of pure MDF-S signatures in other Archean hydrothermal ore deposits (Figure 4.5), especially  
 2834 those without a proximal magmatic source; (iv) the thermodynamic necessity for oxidized

2835 hydrothermal fluids, at equilibrium with  $\text{HSO}_4^-/\text{SO}_4^{2-}$ , to mobilize Cu/Co under neutral fluid  
2836 conditions at 300°C (Figure 4.4); and (v) the close temporal correlation of Carlow Castle  
2837 mineralization with glacial rocks of the Pongola and Witwatersrand basins that display  
2838 similarly muted MIF-S signatures interpreted to relate to an ephemeral oxidation event  
2839 preceding the Pongola Glaciation.

2840 Ephemeral oxygenation within the northwest Pilbara Craton during the formation of the Carlow  
2841 Castle deposit at ~2.955 Ga would have been contemporaneous with widely postulated  
2842 photosynthetically oxygenated conditions during the deposition of the Pongola Supergroup in  
2843 South Africa at ~2.95 Ga (Figure 4.2; Crowe et al., 2013; Eickmann et al., 2018; Kasting &  
2844 Ono, 2006; Luskin, Wilson, Gold, & Hofmann, 2019; Ono et al., 2006; Ossa Ossa et al., 2019;  
2845 Ossa Ossa et al., 2016; Planavsky et al., 2014). In addition to evidence from a variety of isotopic  
2846 systems (Appendix 4.1), sedimentary pyrite in the Mozaan Group of the Pongola Supergroup  
2847 also display muted MIF-S signatures. This has been interpreted as evidence of deposition under  
2848 a weakly oxidized atmosphere (Ono et al., 2006). It has been suggested that during this time  
2849 the Pilbara and Kaapvaal Cratons may have been connected as part of the supercontinent  
2850 Vaalbara; however, this hypothesis remains contentious (de Kock, Evans, & Beukes, 2009;  
2851 Evans & Muxworthy, 2018). Regardless of geographic connections between the northwest  
2852 Pilbara Craton and Pongola Basin, the two regions demonstrate widespread development of  
2853 oxidative conditions at that time based on these data. The independent coevolution of two  
2854 oxygenated environments around 2.95 Ga could suggest that oxygenic metabolic processes  
2855 were more globally widespread and productive than previously understood (Catling & Zahnle,  
2856 2020). This interpretation is consistent with previous hypotheses of an abrupt but ephemeral  
2857 increase in atmospheric  $\text{O}_2$  that destabilized a reduced methane-rich Mesoarchean atmosphere  
2858 leading to the Pongola Glaciation around 2.95 Ga, evidenced by glacial diamictites of the  
2859 Mozaan Group (Pongola Supergroup) in South Africa (Catling & Zahnle, 2020; Kasting &  
2860 Ono, 2006; Lyons et al., 2014; Ono et al., 2006). As such, we argue that the significant oxygen  
2861 required to form the Carlow Castle deposit's Cu-Co mineralization, may be a symptom of a  
2862 global phenomenon associated with a broad increase in  $\text{O}_2$  production around 2.95 Ga  
2863 coinciding with, or immediately following, the early evolution of oxygenic photosynthesis  
2864 (Crowe et al., 2013; Ono et al., 2006; Planavsky et al., 2014; Wang et al., 2018). This was  
2865 sufficient to destabilize a methane-rich greenhouse Mesoarchean atmosphere, significantly  
2866 reduce atmospheric temperatures, suppress MIF-S signatures, and modify large-scale metal

2867 mobilization processes in the crust, leading to the formation of large base-metal deposits. These  
2868 findings may also support previous studies, which suggest that fluctuations in the magnitude  
2869 of  $\Delta^{33}\text{S}$  signatures in the Archean rock record reflect the evolving chemistry of the atmosphere  
2870 throughout the Archean (Ohmoto et al., 2006; Ono et al., 2006; Wang et al., 2018). These  
2871 findings contradict the traditional model of pervasive and persistent Archean atmospheric  
2872 anoxia and are consistent with more nuanced emerging models of ephemerally oxidized  
2873 conditions in Archean surface environments following the evolution of oxygenic  
2874 photosynthesis and the potential for associated fluctuations in atmospheric chemistry following  
2875 this critical evolutionary milestone.

## 2876 4.5 Materials and methods

### 2877 4.5.1 Samples

2878 The sulfur isotope compositions of four mineralized samples (CC012\_67, CC007\_102a,  
2879 CC007\_102b, CC003\_61) from the Carlow Castle Cu-Co-Au deposit were characterized in  
2880 this study (Appendix 4.4). These samples are composed predominantly of chalcopyrite, pyrite,  
2881 pyrrhotite and minor cobaltite; hosted in quartz-calcite veins and intergrown with hydrothermal  
2882 chlorite, actinolite, and epidote alteration minerals. Discs (~8 mm in diameter) of these samples  
2883 were mounted into epoxy pucks and trimmed into a semi-circular shape to be accompanied by  
2884 another semi-circular puck of standard reference materials in the sample holder. These samples  
2885 were gold coated to a thickness of ~30 nm to prevent charging.

### 2886 4.5.2 Secondary-ion mass spectrometry

2887 In-situ triple sulfur isotope ( $^{32}\text{S}$ ,  $^{33}\text{S}$ ,  $^{34}\text{S}$ ) analyses were conducted using a CAMECA 1280  
2888 large geometry secondary-ion mass spectrometer (SIMS) at the Centre for Microscopy and  
2889 Microanalysis, University of Western Australia. After thorough cleaning in ethanol and  
2890 distilled water, samples were coated with 20 nm of Au. A total of 183 analyses of pyrite,  
2891 pyrrhotite, and chalcopyrite were conducted across the four mineralized samples.

2892 The sample surfaces were sputtered over a 10x10  $\mu\text{m}$  area using a 10 kV, Gaussian  $\text{Cs}^+$  beam  
2893 with an intensity of 3.5 nA and impact energy of 20 keV. Secondary ions were admitted to the  
2894 mass spectrometer via a 90  $\mu\text{m}$  entrance slit within a 3000  $\mu\text{m}$  field aperture at  $\times 133$  field  
2895 magnification. The NMR magnetic field controller locked the axial mass at the beginning of  
2896 each session, and the mass spectrometer operated at a mass resolution ( $M/\Delta M$ ) of ~2500. The  
2897 isobaric interference between  $^{33}\text{S}$  and  $^{32}\text{S}$  1H peaks is not fully resolved in these conditions.  
2898 The magnetic field was thus offset slightly on the low mass side to avoid the interference on

2899 the  $^{33}\text{S}$  peak. Three sulfur isotopes were detected simultaneously in multi-collection mode by  
2900 three Faraday cups with  $10^{11} \Omega$  (L2,  $^{32}\text{S}$ ) and  $10^{10} \Omega$  (L1 and H1,  $^{33}\text{S}$  and  $^{34}\text{S}$ ) resistors. Each  
2901 analysis included 30 s of pre-sputtering followed by 20 cycles of 4 s counting time. Analytical  
2902 drift and calibration were monitored through bracketing of analyses with matrix matched  
2903 standards for pyrite (Sierra), chalcopyrite (Nifty-b), and pyrrhotite (Alexo). Reference values  
2904 for these standards and data reduction procedures utilized in this study are documented in  
2905 LaFlamme et al. (2016).

#### 2906 4.5.3 Thermodynamic modelling

2907 Thermodynamic modelling of the hydrothermal solubility of Cu and Co was conducted using  
2908 The Geochemist's Workbench (Bethke, 1996) and methodologies utilized by Jansson and Liu  
2909 (2020). Metal solubility, speciation, and phase stability were modelled and  $\log f_{\text{O}_2}$ -pH  
2910 diagrams were constructed for a peak ore fluid temperature of  $300^\circ\text{C}$ , consistent with  
2911 temperature estimates of the Carlow Castle deposit's ore formation (Fox et al., 2021), and  
2912 salinity of 10% NaCl. Thermodynamic properties for Cu- and Co-bearing minerals and aqueous  
2913 species were taken from various published sources (Brugger et al., 2007; Helgeson, Delany,  
2914 Nesbitt, & Bird, 1978; Liu et al., 2011; Liu & McPhail, 2005; Midgisov, Zezin, & Williams-  
2915 Jones, 2011; Mills, 1974; Shock, Sassani, Willis, & Sverjensky, 1997; Vaughan, 1978; Wolery,  
2916 1992). However, the lack of thermodynamic data for cobaltite necessitates estimating stability  
2917 based on chemically comparable solid Co-bearing phases.

2918

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2923

2924 **References**

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3197 Appendix 4.1 – Compilation of studies purporting to show evidence of  
 3198 Archean oxygen oases

| <b>Age (Ma)</b> | <b>Lithostratigraphic unit</b>                | <b>Terrane</b>                         | <b>Method/evidence</b>  | <b>Reference</b>                                   |
|-----------------|---|--|---|--|
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | C and S isotopes, wt. % C and S   | (Kaufman et al., 2007)                             |
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | C and N isotopes, wt. % N, TOC, Fe and Mo concentrations                  | (Garvin, Buick, Anbar, Arnold, & Kaufman, 2009)    |
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | Mo isotopes   | (Duan et al., 2010)                                |
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | Re-Os isotopes  | (Kendall, Creaser, Reinhard, Lyons, & Anbar, 2015) |
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | U isotopes  | (Kendall, Brennecke, Weyer, & Anbar, 2013)         |
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | Se isotopes   | (Stüeken, Buick, & Anbar, 2015)                    |
| 2500            | Mount McRae Shale, Hamersley Group            | Hamersley Basin, Pilbara Craton        | Re-Os isotopes, TOC, wt. % S, major and trace metal concentrations        | (Anbar et al., 2007)                               |
| 2590            | Ghaap Group, Transvaal Supergroup             | Griqualand West Basin, Kaapvaal Craton | Fe speciation, trace metal concentrations, wt. % C and S, TOC, S isotopes | (Kendall et al., 2010)                             |
| 2640            | Ghaap & Pretoria Groups, Transvaal Supergroup | Griqualand West Basin, Kaapvaal Craton | Mo concentrations and isotopes  | (Wille et al., 2007)                               |
| 2640            | Ghaap Group, Transvaal Supergroup             | Griqualand West Basin, Kaapvaal Craton | Mo isotopes   | (Voegelin, Nägler, Beukes, & Lacassie, 2010)       |
| 2670            | Ghaap Group, Transvaal Supergroup             | Griqualand West Basin, Kaapvaal Craton | Fossil biomarkers   | (Waldbauer, Sherman, Sumner, & Summons, 2009)      |
| 2680            | Ghaap Group, Transvaal Supergroup             | Griqualand West Basin, Kaapvaal Craton | Fe and Mo isotopes, Fe wt. %, Mo concentrations                           | (Czaja et al., 2012)                               |

|           |   |   |  |   |
|-----------|---|---|--|---|
| 2700      | Marra Mamba Formation, Hamersley Group and Maddina and Jeerinah Formations, Fortescue Group | Hamersley and Fortescue Basins, Pilbara Craton      | Fossil biomarkers  | (Brocks, Logan, Buick, & Summons, 1999)   |
| ~2700     | Carajas Formation, Grão Pará Group  | Carajas Province                                    | Mo isotopes  | (Cabral et al., 2013)                     |
| 2700      | Platberg Group, Ventersdorp Supergroup  | Hartbeesfontein Basin, Kaapvaal Craton              | Bacterial microfossils                                     | (Wilmeth et al., 2019)                    |
| 2700      | Unnamed exhalite proximal to Hart Ni-Cu-PGE deposit   | Shaw Dome, Abitibi greenstone belt                  | S and Fe isotopes, whole rock geochemistry (Mn enrichment) | (Hiebert, Bekker, Houlié, & Rouxel, 2018) |
| <2720     | Hamersley and Fortescue Groups  | Hamersley Basin, Pilbara Craton                     | Fossil biomarkers  | (Eigenbrode, Freeman, & Summons, 2008)    |
| 2763      | Apex basalt, Warrawoona Group   | Marble Bar greenstone belt, Pilbara Craton          | Primary hematite   | (Kato et al., 2009)                       |
| <2800     | Various banded iron formations  | Various global localities                           | Cr isotopes  | (Frei et al., 2009)                       |
| 2800      | Mosher Carbonate Formation, Steep Rock Group  | Wabigoon Subprovince, Canadian Shield               | REE analyses   | (Riding, Fralick, & Liang, 2014)          |
| 2760-2920 | Hardey and Mosquito Creek Formations, Fortescue, and Nullagine Groups                       | Fortescue and Mosquito Creek Basins, Pilbara Craton | S isotopes   | (Ohmoto et al., 2006)                     |
| 2950      | Sinqeni Formation, Mozaan Group   | Pongola Basin, Kaapvaal Craton                      | Mo and Fe isotopes   | (Planavsky et al., 2014)                  |
| 2950      | Several formations, Mozaan Group  | Pongola Basin, Kaapvaal Craton                      | C and S isotopes   | (Ono et al., 2006)                        |
| 2950      | Sinqeni and Ntombe Formations, Mozaan Group   | Pongola Basin, Kaapvaal Craton                      | Primary Mn   | (Ossa Ossa et al., 2016)                  |
| 2950      | Sinqeni Formation, Mozaan Group   | Pongola Basin, Kaapvaal Craton                      | Mo, Fe, and U isotopes                                     | (Albut et al., 2019)                      |
| 2950      | Several formations, Mozaan Group  | Pongola Basin, Kaapvaal Craton                      | U isotopes   | (Wang et al., 2020)                       |
| 2950      | Several formations, Mozaan Group  | Pongola Basin, Kaapvaal Craton                      | C, O, Fe, Mo isotopes                                      | (Ossa Ossa et al., 2018)                  |
| 2960      | Gabarintha Formation, Luke Creek Group  | Youanmi Terrane, Yilgarn Craton                     | Ce anomalies   | (Kerrick & Said, 2011)                    |
| 2970      | Chobeni Formation, Nzuse Group  | Pongola Basin, Kaapvaal Craton                      | Fe & S isotopes  | (Eickmann et al., 2018)                   |
| 2980      | Sinqeni and Agatha Formations, Mozaan and Nzuse Groups                                      | Pongola Basin, Kaapvaal Craton                      | Cr isotopes  | (Crowe et al., 2013)                      |

|       |  |  |  |                         |
|-------|--|--|--|-------------------------|
| 3200  | Various marine halites and black shales  | Various global localities                  | Oxygen in halite fluid inclusions and Se/Co ratios in sedimentary pyrite | (Steadman et al., 2020) |
| 3460  | Marble Bar Chert, Warrawoona Group       | Marble Bar greenstone belt, Pilbara Craton | Primary hematite   | (Hoashi et al., 2009)   |
| >3700 | Turbiditic and pelagic sedimentary rocks | Isua supracrustal belt, West Greenland     | Pb isotopes  | (Rosing & Frei, 2004)   |
| 3800  | Various submarine hydrothermal deposits  | Various global localities                  | S isotope mass balance calculations and modelling                        | (Ohmoto, 2020)          |

3199

3200

3201 **Appendix 4.2 – Carlow Castle multiple sulfur isotope data**

3202 Appendix 4.2 is accessible via this [link](#) and provides an .xlsx file containing SIMS multiple  
3203 sulfur isotope data of four analysed samples from Carlow Castle.

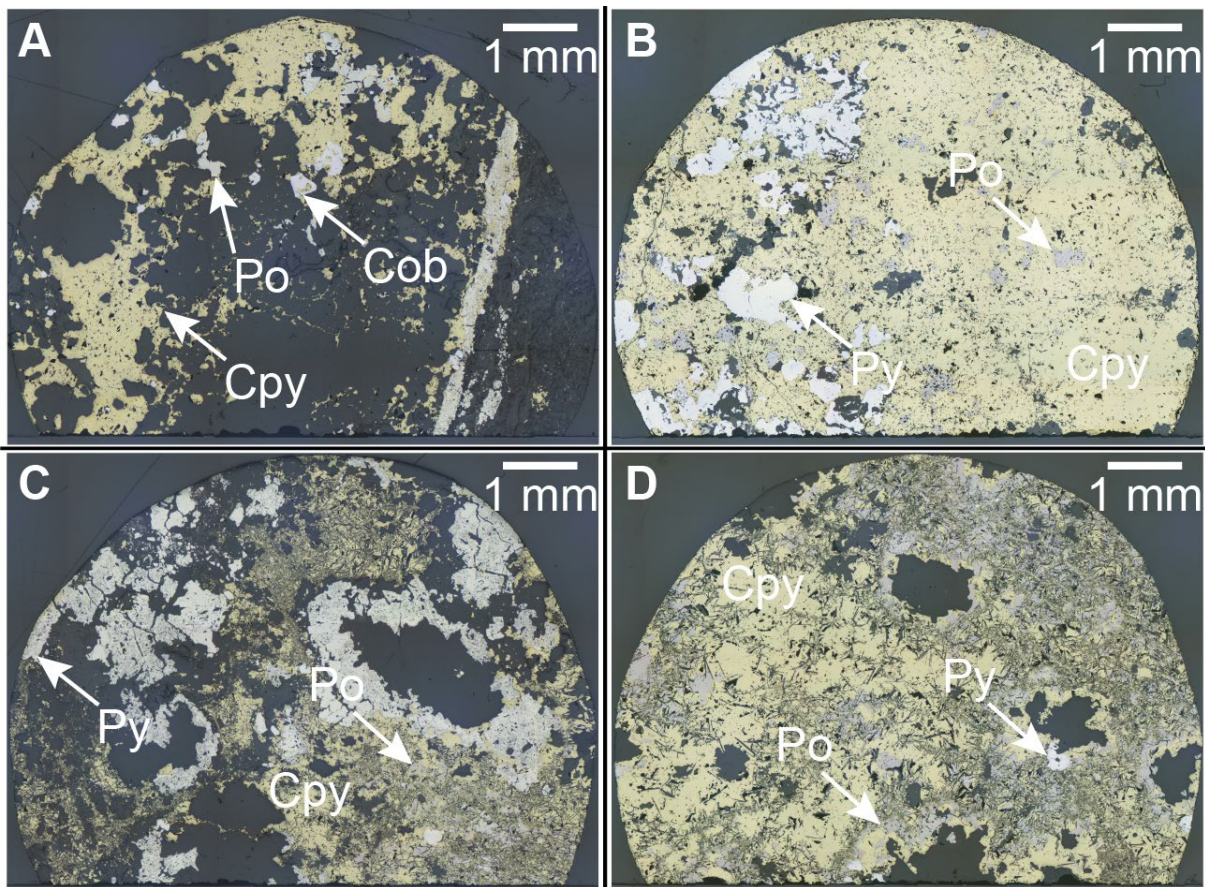
3204

3205 **Appendix 4.3 – Compilation of multiple isotope data from various ore**  
3206 **deposits in Archean greenstone belts**

3207 Appendix 4.3 is accessible via this [link](#) and provides an .xlsx file containing a compiled data  
3208 set of multiple sulfur isotope data from various Archean hydrothermal and magmatic ore  
3209 deposits, used to construct Figure 4.5.

3210

3211 Appendix 4.4 – Analyzed samples from Carlow Castle Cu-Co-Au deposit



3212

3213 Appendix 4.4 - Carlow Castle deposit sulphide ore sample mounts prepared for  
3214 secondary-ion mass spectrometry (SIMS) sulfur isotope analysis. Sample numbers are as  
3215 follows: a. CC003\_61, b. CC012\_67, c. CC007\_102a, d. CC007\_102b.

3216

3217 Chapter 5  
3218 Copper isotopes in Archean  
3219 hydrothermal systems: A case study  
3220 from the Mesoarchean Carlow Castle  
3221 Cu-Co-Au deposit

3222

3223 This chapter is currently under review for publication in *Geochimica et Cosmochimica Acta*.

3224

3225 **5.1 Abstract**

3226 Copper isotope analysis of ore minerals has emerged as a promising tool for understanding  
3227 genetic processes in Cu ore deposits. However, applications of this analytical technique to  
3228 Archean Cu deposits have been very limited, even though Archean terranes are among some  
3229 of the most economically endowed on Earth. As such, this study presents the first Cu isotope  
3230 analysis of an Archean Cu deposit; the Mesoarchean Carlow Castle hydrothermal Cu-Co-Au  
3231 deposit. In this study, 19 Archean primary Cu sulphide ore samples and 11 secondary  
3232 supergene Cu ore samples were analysed. Primary ore samples are isotopically light, with  
3233  $\delta^{65}\text{Cu}$  values ranging between -0.80 and 0.00‰, whilst supergene samples are isotopically  
3234 heavier and range between -0.50 and 0.62‰. In primary ore samples, a relationship is observed  
3235 between Cu isotope signature, ore grade, and alteration assemblage that records the isotopic  
3236 and physicochemical evolution of the Carlow Castle deposit's hydrothermal ore-forming  
3237 system. This isotopic evolution is suggested to reflect a Rayleigh fractionation process due to  
3238 the preferential precipitation of  $^{63}\text{Cu}$  during ore formation, which is modelled using isotopic  
3239 and Cu ore grade data. Based on this modelling, a mafic igneous source is suggested as a likely  
3240 metal source in the Carlow Castle Cu-Co-Au deposit. The relatively limited heavy isotopic  
3241 fractionation of supergene Cu ore samples in this study is interpreted to reflect limited redox

3242 cycling of Cu due to in-situ oxidative weathering of vein-hosted Cu sulphides in the overlying  
3243 supergene system. This differs from previously studied deposits where significant Cu transport  
3244 and multiple stages of isotopic enrichment that are often evident in supergene Cu enrichment  
3245 layers. The results of this study suggest that Cu isotope analysis could be valuable in  
3246 understanding genetic processes in hydrothermal Cu deposits, including in Archean ore  
3247 deposits and terranes.

## 3248 5.2 Introduction

3249 Copper isotope analysis has emerged as a novel geochemical tool in understanding Cu mobility  
3250 and cycling in geological systems (Mathur et al., 2009; Zhu et al., 2002; Zhu, O'Nions, Guo,  
3251 Belshaw, & Rickard, 2000). Because Cu is a redox-sensitive base-metal, Cu isotope analysis  
3252 is particularly promising as a potential palaeoredox proxy and for understanding ore-forming  
3253 processes (Asael, Matthews, Bar-Matthews, & Halicz, 2007; Kendall, 2021). However,  
3254 significant gaps remain in the current understanding of Cu isotope systematics and applications  
3255 of Cu isotope analysis in different geological environments. This includes the understanding  
3256 of processes that may induce Cu isotope fractionation in Archean geological environments,  
3257 where the study of Cu isotope systematics has been very limited. This is most likely due to the  
3258 fact that the most significant Cu isotope fractionation in geological environments is typically a  
3259 product of redox processes, during changes in oxidation state between  $\text{Cu}^+$  and  $\text{Cu}^{2+}$  (Markl,  
3260 Lahaye, & Schwinn, 2006). Since the availability of oxidising agents was limited throughout  
3261 most of the Archean, it is reasonable to expect that Cu isotope fractionation should be minimal  
3262 in Archean Cu deposits and that any Cu ore minerals in these deposits should record an  
3263 unfractionated Cu isotope signature characteristic of the bulk silicate Earth, around  $\delta^{65}\text{Cu} =$   
3264 0‰ (Liu et al., 2015). However, because Cu isotope studies in Archean ore-forming systems  
3265 have been very limited in comparison to younger deposits, this hypothesis has not been tested  
3266 to any significant extent.

3267 In this study we provide the first Cu isotope characterisation of an Archean Cu deposit by  
3268 analysing hydrothermal chalcopyrite and chalcocite in the Mesoarchean Carlow Castle Cu-Co-  
3269 Au deposit in the Pilbara Craton of northwestern Australia (Fox et al., 2021; 2019). In addition  
3270 to this, we also analyse the younger supergene Cu-bearing layer that overlies Carlow Castle's  
3271 sulphide mineralisation. This was done to determine whether the previously established  
3272 exploration applications of Cu isotope analysis could still be applied to Cu mineralisation in  
3273 Archean terranes (Braxton & Mathur, 2011; Mathur et al., 2009; Mathur & Wang, 2019). Given



3274 that Archean terranes are among the most economically endowed on Earth (Goldfarb et al.,  
3275 2010), understanding how Cu isotope exploration methods can be applied in these terranes  
3276 could enhance exploration efforts for Archean Cu-rich ore deposits. In addition, given the  
3277 paucity of Cu isotope data from Archean Cu ore deposits, this study was also conducted to  
3278 determine whether Cu isotope analysis may have potential for understanding geological  
3279 processes in Archean hydrothermal systems given the comparably limited potential for redox  
3280 fractionation.

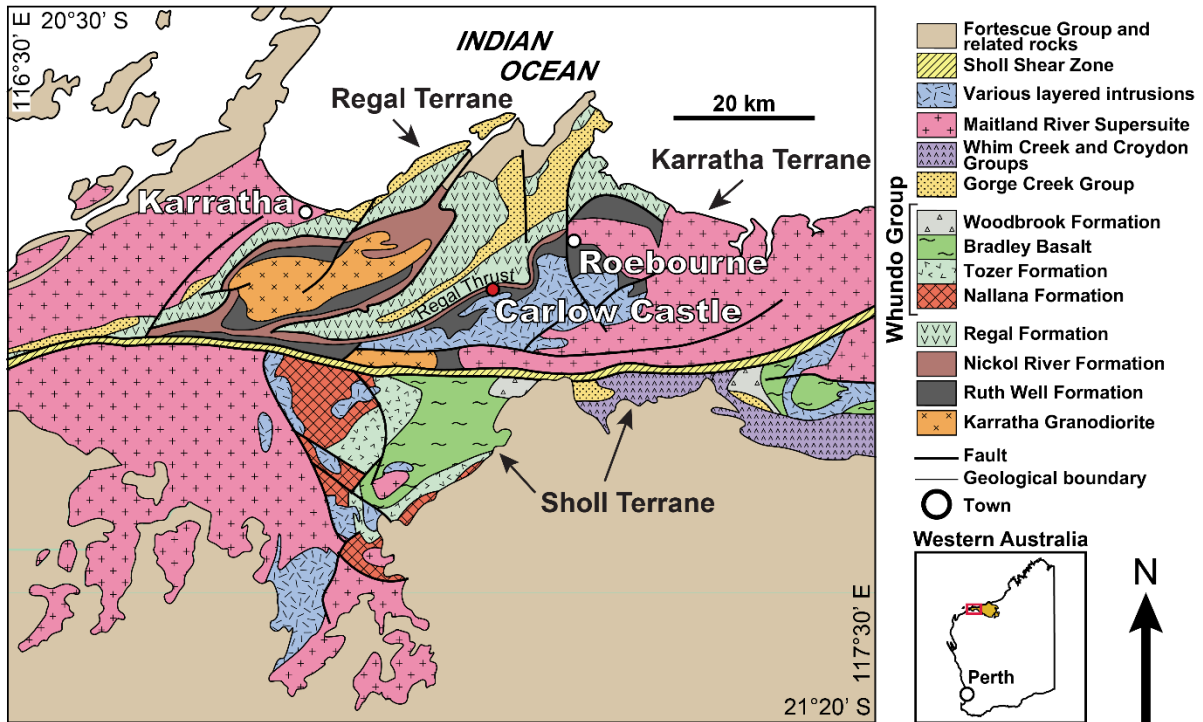
### 3281 5.3 Geological Setting

3282 Carlow Castle is a structurally-hosted hydrothermal Cu-Co-Au deposit situated in the Archean  
3283 Pilbara Craton of northwest Western Australia (Hickman, 2016). Specifically, Carlow Castle  
3284 is hosted in the West Pilbara Superterrane (3280–3066 Ma) of the northwest Pilbara Craton; a  
3285 superterrane produced through the accretion of the Karratha (c. 3280 Ma), Regal (c. 3200 Ma),  
3286 and Sholl (3130–3110 Ma) terranes during the Prinsep Orogeny (c. 3070 Ma) following ~90  
3287 Myr of crustal convergence (Hickman, 2016; Van Kranendonk et al., 2007). These terranes are  
3288 composed dominantly of a series of greenstone belts and several large intrusive granitic  
3289 complexes (Hickman et al., 2010). During the accretion of the West Pilbara Superterrane, the  
3290 Regal Terrane was thrust over the Karratha Terrane along the Regal Thrust. As such, the Regal  
3291 Thrust defines the boundary between these terranes across the northwest Pilbara Craton  
3292 (Hickman, 2016). A heavily tectonised portion of the Regal Thrust, within the Roebourne  
3293 greenstone belt, acts as the key host structure for Carlow Castle’s Cu-Co-Au mineralization  
3294 (Fox et al., 2019; Hickman, 2016). Due to its occurrence within the Regal Thrust, an orogenic  
3295 crustal structure, Carlow Castle was initially interpreted to be a Cu-Co enriched orogenic Au  
3296 deposit produced by the Prinsep Orogeny (Fox et al., 2019; Hickman, 2016). However, recent  
3297 U-Pb dating of syn-ore apatite from Carlow Castle has revealed that it is significantly younger  
3298 than initially thought. Carlow Castle most likely formed via the reactivation of the Regal Thrust  
3299 around c. 2955 Ma, coinciding with a period of major precious- and base-metal volcanogenic  
3300 massive sulphide (VMS) mineralisation during rifting and formation of the proximal De Grey  
3301 Superbasin (3066–2919 Ma; Figure 5.1; Fox et al., 2021). The De Grey Superbasin is  
3302 composed of three distinct sub-basins; (i) the Gorge Creek Basin (3066-3015 Ma), (ii) Whim  
3303 Creek Basin (3009-2991 Ma), and (iii) Mallina Basin (3015-2919) (Hickman, 2016). Carlow  
3304 Castle’s formation is interpreted to have been contemporaneous with the evolution of the

3305 Mallina Basin as a back-arc basin to the southeast of the West Pilbara Superterrane (Fox et al.,  
3306 2021).

3307 On a deposit-scale, Carlow Castle's mineralisation is hosted by a series of faults through a  
3308 hydrothermally altered (chlorite, actinolite, and epidote dominated) section of the Ruth Well  
3309 (c. 3280 Ma) and Nickol River (3220–3160 Ma) formations on the south-eastern footwall of  
3310 the Regal Thrust (Figure 5.1; Fox et al., 2021; Hickman, 2016). The Ruth Well Formation is a  
3311 sub-unit of the Roebourne Group in the Karratha Terrane and is a typical greenstone  
3312 succession; composed primarily of extrusive mafic rocks, carbonaceous shale, and chert  
3313 (Hickman & Van Kranendonk, 2012). The Roebourne Group is unconformably overlain by the  
3314 predominantly fine-grained clastic marine sedimentary rocks of the Nickel River Formation  
3315 (Hickman, 2016). The Roebourne Group and Nickel River Formation significantly pre-date the  
3316 De Grey Superbasin, and are interpreted to have formed during volcanism on and rifting of a  
3317 nascent Paleoarchean Pilbara Craton (Hickman, 2016). Carlow Castle's mineralisation occurs  
3318 as a stockwork of quartz-carbonate veins hosting Cu-Co-Au bearing base-metal sulphides and  
3319 sulphosalts (Fox et al., 2019). Fox et al. (2019) identified that Carlow Castle's mineralisation  
3320 occurs in two distinct ore mineral assemblages. The first and more abundant assemblage is  
3321 composed predominantly of chalcopyrite and pyrite with minor pyrrhotite and cobaltite, and  
3322 trace electrum. Assemblage Two contains a significantly higher abundance of cobaltite which  
3323 co-occurs with chalcocite, rather than chalcopyrite, and a slightly higher abundance of  
3324 electrum. Based on analyses of its characteristically propylitic hydrothermal alteration  
3325 assemblage, it has been established that Carlow Castle's hydrothermal ore fluid was most likely  
3326 neutral to slightly alkaline in pH, reaching peak temperatures of ~300°C (Fox et al., 2021).  
3327 However, a transition in the temperature of this propylitic alteration assemblage is evident and  
3328 defined by a decrease in actinolite and epidote relative to chlorite. Fox et al. (2021) interpreted  
3329 that this may reflect zoning of the alteration assemblage within the Carlow Castle deposit  
3330 associated with fluid cooling at the periphery of the deposit during the evolution of the  
3331 hydrothermal system. Fox et al. (2019) suggested that within this fluid Cu, Co, and Au were  
3332 most likely transported in chloride complexes, before they co-precipitated at the site of Carlow  
3333 Castle's ore mineralisation due to the reaction of this hydrothermal fluid with wall-rocks within  
3334 the Regal Thrust. Additionally, supergene Cu mineralisation occurs primarily within 50 metres  
3335 of the present-day surface at Carlow Castle and generally overlies the primary sulphide ore.  
3336 However, due to the structurally complex nature of Carlow Castle, this supergene

3337 mineralisation does not necessarily occur at uniform depths and some oxidised altered sulphide  
 3338 veins are observed as deep as ~75 metres in drill core. This supergene Cu mineralisation is  
 3339 mineralogically diverse; containing native Cu, secondary chalcocite, plancheite  
 3340 ( $\text{Cu}_8\text{Si}_8\text{O}_{22}(\text{OH})_4 \cdot (\text{H}_2\text{O})$ ), agardite  $((\text{REE}, \text{Ca})\text{Cu}_6(\text{AsO}_4)_3(\text{OH})_6 \cdot 3\text{H}_2\text{O})$ , and Cu-bearing  
 3341 pyrolusite ( $\text{MnO}_2$ ).



3342  
 3343 Figure 5.1 - Geological map of northwest Pilbara Craton near Carlow Castle Cu-Co-Au  
 3344 deposit, modified after Van Kranendonk et al. (2002). Inset map of Western Australia  
 3345 depicts portion of Pilbara Craton shown in main image.

## 3346 5.4 Methods and materials

### 3347 5.4.1 Analysed samples

3348 A total of 30 samples were analysed in this study, with 19 samples from Carlow Castle's deep  
 3349 primary sulphide mineralisation, and 11 samples from supergene Cu mineralisation (sample  
 3350 list in results). The majority of these samples were retrieved from drill core, with the exception  
 3351 of four which were retrieved from mine tailings at the present-day surface of Carlow Castle.  
 3352 Some mineralogically or texturally heterogeneous drill core samples were sub-sampled twice  
 3353 to account for potential isotopic variation within a given sample, e.g., samples CC012\_67(1)  
 3354 and CC012\_67(2).

3355 Sulphide samples were taken from core by drilling and then weighing out ~10 mg of material  
3356 using a small handheld rotary multi-tool equipped with a diamond drill bit and a precision  
3357 balance. These 10 mg samples were dissolved in 4 mL aqua regia solution (3.2 mL 12M HCl  
3358 + 0.8 mL 15M HNO<sub>3</sub>) at room temperature, or in concentrated HF + 7M HNO<sub>3</sub> at 140°C for  
3359 samples containing SiO<sub>2</sub> (USGS rock standards and plancheite). Samples were evaporated on  
3360 a hotplate once the material was completely dissolved. These samples were then redissolved in  
3361 6M HCl to create a sample solution and diluted to a concentration of approximately 10 µg Cu  
3362 in 1.5 ml of 6M HCl + 0.001% H<sub>2</sub>O<sub>2</sub> to load onto the columns. Dilution factors varied between  
3363 phases, depending on the stoichiometric Cu concentration in each mineral.

#### 3364 5.4.2 Copper isotope analysis

3365 In order to develop a robust method for Cu isotope analysis of Cu sulphides, experiments were  
3366 conducted by varying the concentration of Cu in the loading solution, volume of loading  
3367 solution, volume of resin, column size, and acid volume in each step of the matrix and Cu  
3368 elution. To observe how effective the elution of Cu from other elements under these different  
3369 parameters was, the eluate was collected every 1-2 mL for each of these experiments. The  
3370 successful calibrations are shown in Appendix 5.1.

3371 Upon evaluation of the results for the different experiments, the final parameters and  
3372 methodology chosen for Cu separation from Cu sulphides and Cu-bearing silicates are listed in  
3373 Appendix 5.1. As only Cu isotopes were being analysed in this study, the final methodology  
3374 did not involve any further elution stages for collection of Fe or Zn, these additional stages  
3375 were only undertaken in the previous experiments to test how effective the separation of Cu  
3376 from these elements was. As such, they were deemed unnecessary for the final method.

#### 3377 5.4.3 Multicollector-inductively coupled plasma mass spectrometry for isotopic 3378 analysis

3379 Samples were diluted to 200 ppb Cu in 3 mL 2% HNO<sub>3</sub> for analysis on a Thermo Scientific  
3380 Neptune multi collector inductively coupled plasma mass spectrometer (MC-ICP-MS) with an  
3381 autosampler at CSIRO Land and Water on the University of Adelaide Waite Campus. Standard  
3382 sample bracketing with NIST976 was used for all analyses, and this was diluted to 200 ppb  
3383 with the same batch of 2% HNO<sub>3</sub> as the samples. Delta values in per mil for Cu isotopes are  
3384 calculated as follows:

3385 
$$\delta^{65}\text{Cu} \text{ ‰} = \left[ \frac{\left(\frac{65\text{Cu}}{63\text{Cu}}\right)_{\text{sample}}}{\left(\frac{65\text{Cu}}{63\text{Cu}}\right)_{\text{NIST976}}} - 1 \right] \cdot 1000$$

3386 Part way through this study, Ga spiking of samples (after Hou et al. (2016)) was introduced  
 3387 into the measurement routine to improve precision (see Appendix 5.1). The sample  
 3388 introduction system used was a dual Scott/cyclonic spray chamber with a 100 µl/min nebulizer,  
 3389 Ni sampler cone, and Ni X skimmer cone. Analyses used low resolution mode, with a typical  
 3390 signal strength of 12V on samples and NIST976. A single block of 60 four second integrations  
 3391 was run for each analysis. Raw <sup>65</sup>Cu/<sup>63</sup>Cu and Ga mass bias corrected <sup>65</sup>Cu/<sup>63</sup>Cu ratios were  
 3392 returned with mass bias correction undertaken for each integration within the block. The mean  
 3393 of the analysis was used for sample-standard bracketing. At least two analyses of each sample  
 3394 were obtained, and the precision calculated as the standard deviation between these (2SD). A  
 3395 procedural blank was prepared with each sample batch and contained less than 3 mV <sup>65</sup>Cu.

3396 Internal reference materials were developed from a supergene native Cu sample (BHNatCop)  
 3397 from Broken Hill Pb-Zn-Ag deposit (long term average δ<sup>65</sup>Cu = 1.20‰ ± 0.16) and a coarse-  
 3398 grained massive chalcopyrite sample from Carlow Castle (DFCu) (long term average δ<sup>65</sup>Cu =  
 3399 -0.79‰ ± 0.09). The long-term repeatability of these reference materials is demonstrated in  
 3400 data from multiple analytical sessions over the course of roughly two years. These data are  
 3401 provided in Appendix 5.2. However, as these materials do not yet have consensus values, two  
 3402 digestions each of USGS BIR-1 basalt and AGV-2 andesite reference materials were also used  
 3403 to monitor Cu isotope analyses after development by Liu et al. (2014a). BIR-1 was measured  
 3404 as -0.01 ± 0.02 and AGV-2 as -0.01 ± 0.05, which are within uncertainty of the values for a  
 3405 recent compilation of reference materials by Wang et al. (2019) (BIR-1: 0.02 ± 0.02, AGV-2:  
 3406 0.05 ± 0.02).

## 3407 5.5 Results

### 3408 5.5.1 Copper isotope signature of Carlow Castle

3409 The results of Cu isotope analyses from Carlow Castle's primary and supergene Cu  
 3410 mineralisation are summarised in Table 5.1 and Figure 5.2. Primary mineralisation is lightly  
 3411 fractionated, whilst the supergene mineralisation is predominantly heavily fractionated, with  
 3412 some lightly fractionated values. The 19 primary samples exhibit light δ<sup>65</sup>Cu values: ranging  
 3413 between -0.80 and 0.00‰, with a mean of -0.36‰. The 11 supergene samples exhibit  
 3414 predominantly heavy δ<sup>65</sup>Cu values but range between -0.50 and 0.62‰, with a mean of 0.12‰.

3415 The values of both primary and supergene Cu mineralisation appear fractionated relative to the  
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 3417

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3484            doi:[https://doi.org/10.1016/S0009-2541\(99\)00076-5](https://doi.org/10.1016/S0009-2541(99)00076-5)  
3485



3486 Table 5.1 - Sample details and Cu isotope composition of mineralised samples from Carlow Castle. Note that Cu, Co, and Au grades (ppm)  
3487 here are for the metre intersection from which a given sample was taken and are sourced from publicly available JORC (2012) compliant  
3488 exploration drill core data released to the Australian Securities Exchange (Artemis Resources Limited, 2018c). Note: Samples analysed  
3489 with Ga mass bias correction are marked with an asterisk (\*).

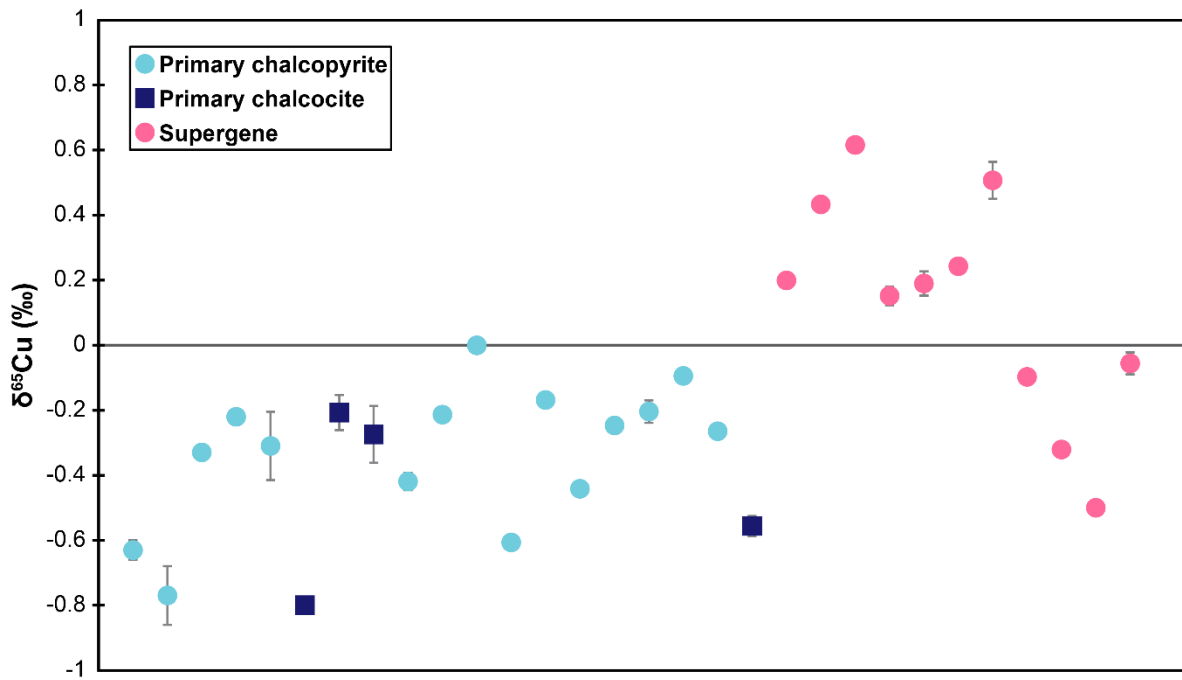
| Sample                        | Drill hole | Drill hole easting (MGA 50) | Drill hole northing (MGA 50) | Measured sample depth (m) | Drill hole plunge (°) | True vertical depth (TVD) (m) | Cu (ppm) | Co (ppm) | Au (ppm) | Analysed phase | $\delta^{65}\text{Cu}$ (‰) | 2 $\sigma$ error |
|-------------------------------|------------|-----------------------------|------------------------------|---------------------------|-----------------------|-------------------------------|----------|----------|----------|----------------|----------------------------|------------------|
| <b>Primary mineralisation</b> |            |                             |                              |                           |                       |                               |          |          |          |                |                            |                  |
| <b>DF2</b><br>(CC012_67(2))   | 18CCAD012  | 506935                      | 7698900                      | 67                        | 60                    | 58                            | 95,600   | 338      | 1.41     | Chalcopyrite   | -0.63                      | 0.03             |
| <b>DFCu</b><br>(CC012_67(1))* | 18CCAD012  | 506935                      | 7698900                      | 67                        | 60                    | 58                            | 95,600   | 338      | 1.41     | Chalcopyrite   | -0.79                      | 0.09             |
| <b>DF3</b><br>(CC007_102)     | 18CCAD007  | 506857.87                   | 7698633.3                    | 102                       | 60                    | 89                            | 33,300   | 496      | 7.21     | Chalcopyrite   | -0.33                      | 0.001            |
| <b>DF4</b><br>(CC003_61)(1)   | 18CCAD003  | 506698.19                   | 7698681                      | 61                        | 75                    | 59                            | 13,900   | 6800     | 8.13     | Chalcopyrite   | -0.22                      | 0.001            |
| <b>DF5</b><br>(CC009_52)      | 18CCAD009  | 506942.27                   | 7698937.2                    | 52                        | 60                    | 45                            | 66,300   | 16,000   | 29       | Chalcopyrite   | -0.31                      | 0.11             |
| <b>DF6</b><br>(CC007_41)      | 18CCAD007  | 506857.87                   | 7698633.3                    | 41                        | 60                    | 36                            | 17,550   | 18,900   | 18.45    | Chalcocite     | -0.80                      | 0.02             |
| <b>DF7</b><br>(CC000_2)       | -          | -                           | -                            | -<br>(Surface             | -                     | -                             | -        | -        | -        | Chalcocite     | -0.21                      | 0.05             |

|                                |           |           |           |                          |    |     |        |        |       |              |            |       |      |
|--------------------------------|-----------|-----------|-----------|--------------------------|----|-----|--------|--------|-------|--------------|------------|-------|------|
|                                |           |           |           | sample)                  |    |     |        |        |       |              |            |       |      |
| <b>DF8<br/>(CC000_3)</b>       | -         | -         | -         | -<br>(Surface<br>sample) | -  | -   | -      | -      | -     | -            | Chalcocite | -0.27 | 0.09 |
| <b>DF9<br/>(CC007_62)*</b>     | 18CCAD007 | 506857.87 | 7698633.3 | 62                       | 60 | 53  | 21,800 | 365    | 4.7   | Chalcopyrite | -0.42      | 0.027 |      |
| <b>DF10<br/>(CC003_85)*</b>    | 18CCAD003 | 506698.19 | 7698681   | 85                       | 75 | 82  | 17,000 | 190    | 0.45  | Chalcopyrite | -0.21      | 0.008 |      |
| <b>DF11<br/>(CC010_154)*</b>   | 18CCAD010 | 507480.5  | 7698641.4 | 154                      | 60 | 133 | 3240   | 18,500 | 5.15  | Chalcopyrite | 0.00       | 0.007 |      |
| <b>DF12<br/>(CC001_55)*</b>    | 18CCAD001 | 506701.45 | 7698757.3 | 55                       | 60 | 47  | 27,700 | 69     | 0.21  | Chalcopyrite | -0.61      | 0.017 |      |
| <b>DF13<br/>(CC005_102)*</b>   | 18CCAD005 | 506863.16 | 7698712.4 | 102                      | 60 | 89  | 12,700 | 167    | 0.59  | Chalcopyrite | -0.17      | 0.010 |      |
| <b>DF14<br/>(CC010_137)*</b>   | 18CCAD010 | 507480.5  | 7698641.4 | 137                      | 60 | 118 | 23,200 | 367    | 0.35  | Chalcopyrite | -0.44      | 0.003 |      |
| <b>DF15<br/>(CC009_50)*</b>    | 18CCAD009 | 506942.27 | 7698937.2 | 50                       | 60 | 43  | 41,500 | 29,200 | 26.2  | Chalcopyrite | -0.25      | 0.003 |      |
| <b>DF16<br/>(CC003_61)(2)*</b> | 18CCAD003 | 506698.19 | 7698681   | 61                       | 75 | 59  | 13,900 | 6800   | 8.13  | Chalcopyrite | -0.20      | 0.034 |      |
| <b>DF17<br/>(CC001_95)*</b>    | 18CCAD001 | 506701.45 | 7698757.3 | 95                       | 60 | 82  | 7160   | 192    | 1.84  | Chalcopyrite | -0.09      | 0.002 |      |
| <b>DF18<br/>(CC009_48)*</b>    | 18CCAD009 | 506942.27 | 7698937.2 | 48                       | 60 | 41  | 74,900 | 13,500 | 13.05 | Chalcopyrite | -0.27      | 0.011 |      |
| <b>DF19</b>                    | 18CCAD008 | 506932.99 | 7698937.9 | 25                       | 60 | 22  | 33,500 | 13,500 | 9.25  | Chalcocite   | -0.56      | 0.031 |      |

|                                 |           |           |           |                          |    |    |        |      |      |                                   |           |       |       |
|---------------------------------|-----------|-----------|-----------|--------------------------|----|----|--------|------|------|-----------------------------------|-----------|-------|-------|
| (CC008_25)                      |           |           |           |                          |    |    |        |      |      |                                   |           |       |       |
| <b>Supergene mineralisation</b> |           |           |           |                          |    |    |        |      |      |                                   |           |       |       |
| <b>DF20<br/>(CC000_1)</b>       | -         | -         | -         | -<br>(Surface<br>sample) | -  | -  | -      | -    | -    | -                                 | Native Cu | 0.20  | 0.01  |
| <b>DF21<br/>(CC006_14)(1)*</b>  | 18CCAD006 | 506901.24 | 7698720.4 | 14                       | 60 | 12 | 7410   | 909  | 0.33 | Pyrolusite<br>(15% Cu)            | 0.43      | 0.005 |       |
| <b>DF22<br/>(CC006_14)(2)*</b>  | 18CCAD006 | 506901.24 | 7698720.4 | 14                       | 60 | 12 | 7410   | 909  | 0.33 | Plancheite                        | 0.62      | 0.005 |       |
| <b>DF23<br/>(CC000_4)*</b>      | -         | -         | -         | -<br>(Surface<br>sample) | -  | -  | -      | -    | -    | -                                 | Native Cu | 0.15  | 0.029 |
| <b>DF24<br/>(CC011_97)*</b>     | 18CCAD011 | 507476.27 | 7698549.7 | 97                       | 50 | 74 | 10,250 | 1260 | 2.85 | Plancheite +<br>Agardite-<br>(Ce) | 0.19      | 0.037 |       |
| <b>DF25<br/>(CC011_56)*</b>     | 18CCAD011 | 507476.27 | 7698549.7 | 56                       | 50 | 42 | 3590   | 186  | 0.07 | Plancheite                        | 0.24      | 0.007 |       |
| <b>DF26<br/>(CC010_44)*</b>     | 18CCAD010 | 507480.5  | 7698641.4 | 44                       | 60 | 38 | 2830   | 351  | 0.2  | Chalcocite                        | 0.51      | 0.057 |       |
| <b>DF27<br/>(CC001_50)*</b>     | 18CCAD001 | 506701.45 | 7698757.3 | 50                       | 60 | 43 | 3200   | 560  | 0.42 | Plancheite +<br>Agardite-<br>(Ce) | -0.10     | 0.015 |       |
| <b>DF28<br/>(CC002_16)*</b>     | 18CCAD002 | 506778.93 | 7698694.9 | 16                       | 60 | 13 | 17,100 | 1120 | 0.88 | Plancheite                        | -0.32     | 0.008 |       |
| <b>DF29</b>                     | 18CCAD006 | 506901.24 | 7698720.4 | 52                       | 60 | 45 | 7260   | 259  | 0.87 | Plancheite                        | -0.50     | 0.014 |       |

3490

|                                      |           |           |           |    |    |    |      |     |      |                   |       |       |
|--------------------------------------|-----------|-----------|-----------|----|----|----|------|-----|------|-------------------|-------|-------|
| <b>(CC006_52)(1)*</b>                |           |           |           |    |    |    |      |     |      |                   |       |       |
| <b>DF30</b><br><b>(CC006_52)(2)*</b> | 18CCAD006 | 506901.24 | 7698720.4 | 52 | 60 | 45 | 7260 | 259 | 0.87 | Agardite-<br>(Ce) | -0.06 | 0.034 |



3491

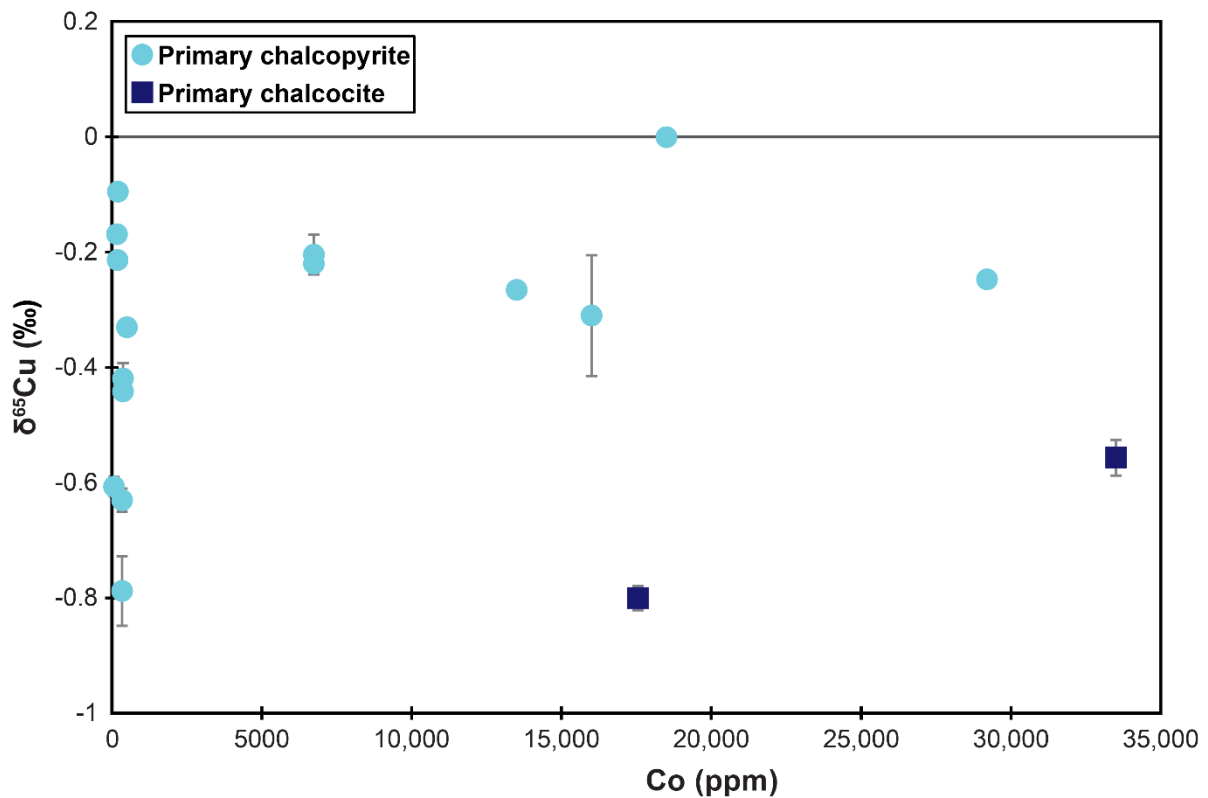
3492 Figure 5.2 - Copper isotope composition of primary Cu sulphides (blue) and supergene  
 3493 Cu minerals (pink) from Carlow Castle. Note errors bars are  $2\sigma$ .

## 3494 5.6 Discussion

### 3495 5.6.1 Carlow Castle Cu-Co-Au deposit copper isotope composition

3496 Carlow Castle's primary Cu sulphide mineralisation is lightly fractionated; with  $\delta^{65}\text{Cu}$  values  
 3497 ranging between -0.80‰ and 0.00‰ across 19 samples. Across this sample suite there is  
 3498 minimal difference in isotopic signature between different phases, chalcocite ranges between -  
 3499 0.80‰ and -0.21‰, whilst chalcopyrite ranges between -0.79‰ and 0.00‰ (Figure 5.2).  
 3500 Similarly, there is no systematic variation in isotopic signature between those samples from  
 3501 Co-poor and Co-rich intersections of the mineralization (Figure 5.3), designated *Assemblage*  
 3502 *One* and *Assemblage Two* by Fox et al. (2019). Note that Cu, Co, and Au grade values are  
 3503 taken from the one metre intersection of a given sample as reported by Artemis Resources  
 3504 Limited (2018c). Given the lack of observed spatial association between these mineral  
 3505 assemblages it has previously been difficult to determine whether they are cogenetic or if they  
 3506 represent two distinct stages of mineralisation in a multiphase mineral system. The similar Cu  
 3507 isotope signatures of the distinct Co-poor and Co-rich mineral assemblages suggests that they  
 3508 represent mineralogically distinct portions of a single evolving ore-forming system, or that at  
 3509 very least they represent ore formation from a similar Cu source and under a comparable  
 3510 precipitation mechanism. Comparable Cu isotope signatures within the analysed chalcocite and

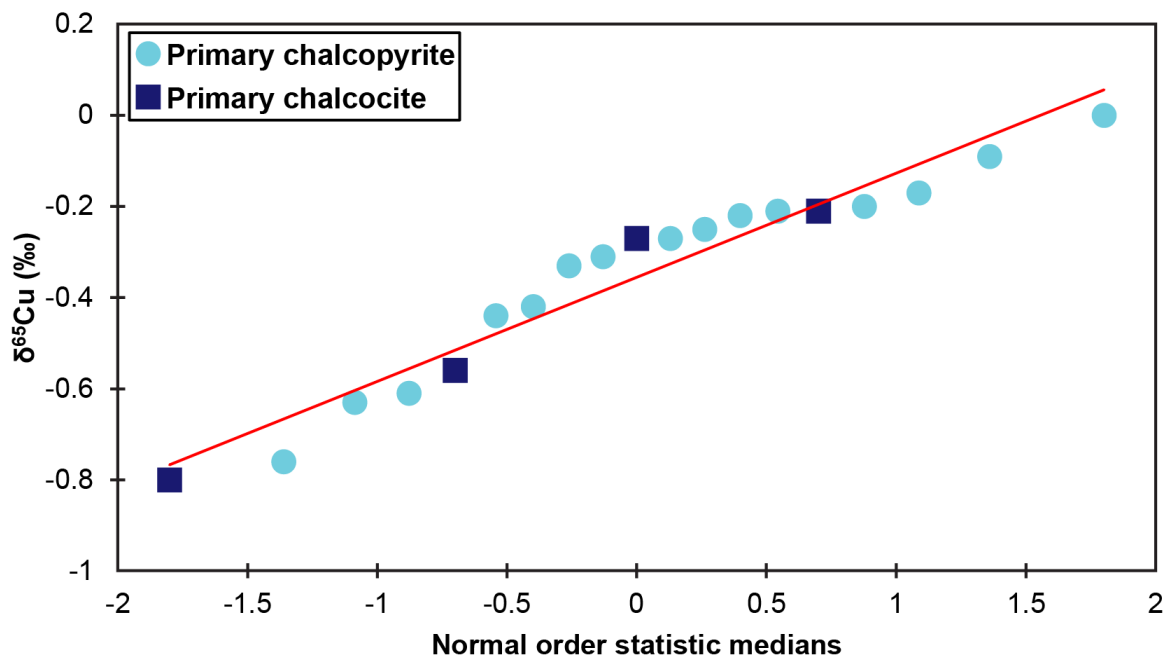
3511 chalcopyrite imply a common genesis with a similar metal source rather than derivation from  
 3512 two distinct stages of mineralisation in a complex multiphase ore-forming system. There is no  
 3513 clear systematic spatial zonation in Cu isotope composition evident in this data set, as has  
 3514 previously been reported in some porphyry Cu deposits (Gregory & Mathur, 2017; Li, Jackson,  
 3515 Pearson, & Graham, 2010). However, the structurally-controlled vein-hosted nature of Carlow  
 3516 Castle's Cu mineralisation, and potential for subsequent deformation of this mineralisation  
 3517 along these structures would make discerning any spatial zonation in Cu isotope signatures  
 3518 highly complex.



3519  
 3520 Figure 5.3 - Cross-plot of Cu isotope signature and Co grade in analysed primary Cu  
 3521 samples. Note errors bars are 2σ.

3522 The isotopically light Cu isotope signature of Carlow Castle's Cu sulphide mineralisation is  
 3523 broadly consistent with known pathways of Cu isotope fractionation based on Cu isotope data  
 3524 from other hydrothermal Cu sulphide ore deposits and experimental studies. It is generally  
 3525 established that hydrothermal ore-forming processes tend to induce preferential incorporation  
 3526 of the lighter (<sup>63</sup>Cu) isotope into solid Cu sulphide phases during reductive ore precipitation  
 3527 (Asael et al., 2007; Fujii, Moynier, Abe, Nemoto, & Albarède, 2013; Mathur et al., 2018;  
 3528 Pękala, Asael, Butler, Matthews, & Rickard, 2011; Zhu et al., 2002). The broadly normally

3529 distributed (Figure 5.4) Cu isotope signature of these Cu sulphide phases is consistent with  
 3530 formation in an ore system with a single Cu source progressively evolving toward  $\delta^{65}\text{Cu} = 0\text{‰}$ ,  
 3531 rather than from the mixing of several Cu-bearing fluids of different isotopic composition or  
 3532 episodic input of isotopically distinct Cu-bearing fluids. Additionally, the lack of any heavy  
 3533 Cu fractionation in Carlow Castle's mineralisation is somewhat striking, as it is relatively  
 3534 common for Cu isotope signatures in hydrothermal Cu deposits of many varieties to display a  
 3535 range of isotopic compositions; including some  $\delta^{65}\text{Cu}$  values  $>0\text{‰}$  (see Kim, Lee, Oyungerel,  
 3536 Jargal, and Tsedenbal (2019) for a recent review and comprehensive compilation of Cu isotope  
 3537 data from various ore deposits). Nonetheless, Carlow Castle's Cu isotope signature is within  
 3538 the established constraints of most hypogene ( $>150^\circ\text{C}$ ) ore deposits which typically cluster  
 3539 around  $0\text{‰}$  (Mathur et al., 2018). This is consistent with Carlow Castle's ore formation at peak  
 3540 temperatures of  $\sim 300^\circ\text{C}$  (Fox et al., 2021).



3541  
 3542 Figure 5.4 - Normal probability plot of Cu isotopes composition of primary Cu sulphide  
 3543 mineralisation from Carlow Castle, showing a broadly normal distribution indicative of a  
 3544 single homogenous Cu source.

### 3545 5.6.2 Copper isotope fractionation in Archean ore-forming systems

3546 Although the application of Cu isotopic analysis in ore-forming systems has expanded  
 3547 significantly in recent years with improvements in analytical capabilities (Zhu et al., 2000),  
 3548 studies of Archean ore deposits have been extremely sparse. As such, very little is known about  
 3549 if and how Archean ore-forming processes induce Cu isotope fractionation. This is also true of

3550 Archean Earth systems more broadly due to the limited number of studies applying Cu isotopic  
3551 analysis in Archean terranes generally, with some exceptions where Cu isotopic analysis has  
3552 been applied as a palaeoredox proxy due to the manner in which redox processes induce  
3553 significant Cu isotope fractionation (Chi Fru et al., 2016; Zavina-James et al., 2021). Given the  
3554 globally anoxic conditions that prevailed for much of the Archean and the consequently limited  
3555 role of redox processes in most Archean ore-forming systems, it may previously have been  
3556 assumed that Cu isotopic analysis would be of limited utility when applied to these systems.  
3557 As such, the fractionated and consistently observable pattern of light  $\delta^{65}\text{Cu}$  in Carlow Castle's  
3558 Cu sulphide mineralisation relative to an unfractionated mantle reservoir ( $0.06 \pm 0.20\text{‰}$  based  
3559 on estimates by Liu et al. (2015)) is notable. This suggests that Cu isotopic analysis could have  
3560 novel applications in understanding Archean ore-forming systems and that previously proposed  
3561 exploration techniques utilising this analytical technique may be similarly applicable in  
3562 Archean terranes (Mathur & Wang, 2019). However, Carlow Castle's Cu isotope fractionation  
3563 is still relatively limited by comparison to other hydrothermal Cu deposits, which have been  
3564 observed to exhibit  $\delta^{65}\text{Cu}$  values  $<-3\text{‰}$  and  $>4\text{‰}$  (Kim et al., 2019; Li et al., 2010; and  
3565 references therein). (Fox et al., 2021; Fox et al., 2019)

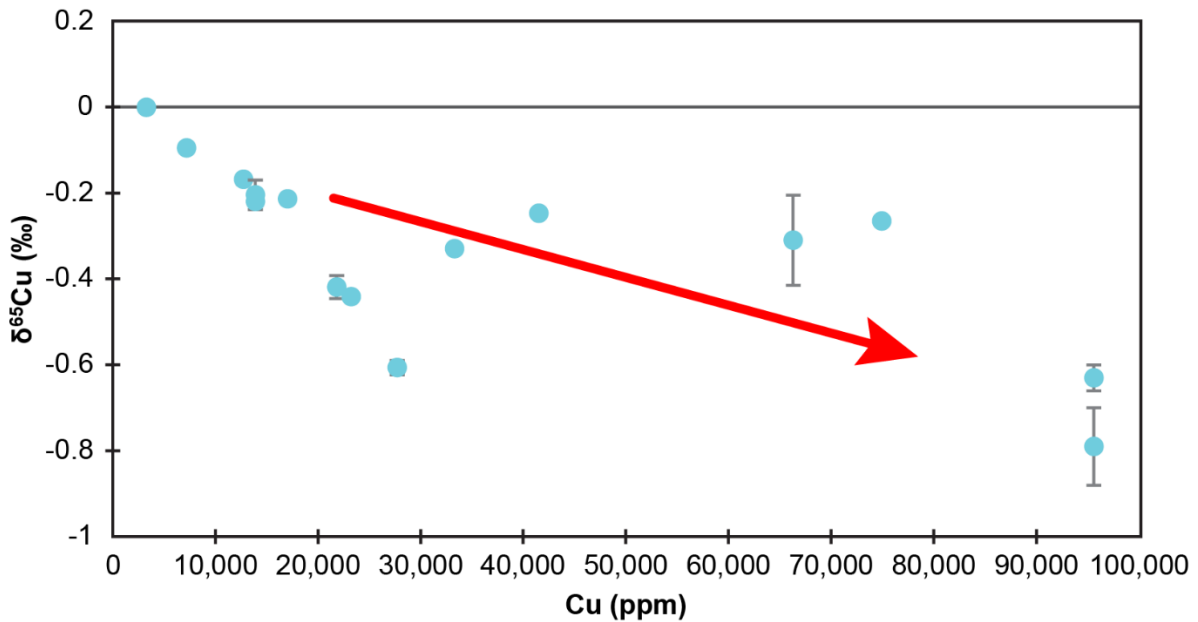
3566 The absence of any  $\delta^{65}\text{Cu}$  values  $>0\text{‰}$  is a relatively unique aspect of the Cu isotope signature  
3567 of the primary Cu sulphide mineralisation at Carlow Castle. The current lack of knowledge of  
3568 Archean Cu isotope systematics makes it challenging to explain this feature of Carlow Castle's  
3569 Cu mineralisation. Nonetheless, it is well established that oxidative weathering of Cu-bearing  
3570 minerals tends to preferentially mobilise  $^{65}\text{Cu}$ ; leading to enriched  $\delta^{65}\text{Cu}$  values in weathering  
3571 product solutions (Mathur et al., 2012; Moynier, Vance, Fujii, & Savage, 2017). Given this,  
3572 Chi Fru et al. (2016) observed an evolution toward heavier  $\delta^{65}\text{Cu}$  signatures in black shales  
3573 across 2.66-2.08 Ga, from slightly isotopically light around 2.66 Ga ( $\delta^{65}\text{Cu} = -0.55\text{‰}$ ) to  
3574 slightly isotopically heavy ( $\delta^{65}\text{Cu} = 0.74\text{‰}$ ) by 2.3 Ga. This isotopic shift is primarily  
3575 explained by increased oxidative weathering and delivery of  $^{65}\text{Cu}$  to the marine environment  
3576 because of the significant increase in atmospheric oxygen following the Great Oxygenation  
3577 Event (GOE) around 2.45 Ga. Following this logic, the negative  $\delta^{65}\text{Cu}$  signature of Carlow  
3578 Castle's Cu sulphide mineralisation may be explained by the lack of any significant  $^{65}\text{Cu}$ -  
3579 enriched crustal Cu isotope reservoirs to provide a source of isotopically heavy Cu to Carlow  
3580 Castle's mineral system. This is due to the comparably limited potential for significant  
3581 oxidative weathering of Cu-bearing minerals in surface environments during the Mesoproterozoic



3582 to develop any fractionated and isotopically heavy surface reservoirs. Contamination by  
3583 fractionated crustal material is a commonly cited source of heavy Cu isotope enrichment and  
3584 heterogeneity in hypogene systems (Kim et al., 2019; Ripley, Dong, Li, & Wasylenki, 2015).  
3585 By contrast,  $\delta^{65}\text{Cu}$  values in modern seawater were measured to range from 0.9 to 1.5‰ by  
3586 Vance et al. (2008) whilst modern oceanic sediments from the Pacific, Atlantic and, Indian  
3587 oceans were reported by Dekov, Rouxel, Asael, Hålenius, and Munnik (2013) to range between  
3588 0.41 and 0.95‰. Although some evidence of oxidative weathering has been observed in  
3589 terrestrial environments following the emergence of oxidative photosynthesis, most likely  
3590 during the Mesoarchean (Catling & Zahnle, 2020; Wang et al., 2018), this weathering was  
3591 likely produced by isolated mildly oxygenated environments during periods of transient  
3592 oxygenation prior to the GOE (Johnson et al., 2021; Lyons et al., 2014). Because of the limited  
3593 extent and magnitude of Archean oxidative weathering, it would have been unlikely to induce  
3594 heavy Cu isotope fractionation of crustal reservoirs on a significant scale. Additionally, the  
3595 intensity of such oxidative weathering processes would have been relatively limited by  
3596 comparison to modern (likely Cenozoic) supergene processes and the transient nature of such  
3597 conditions would preclude the multiple cycles of oxidation and reduction necessary to generate  
3598 significant heavy Cu isotope fractionation in crustal reservoirs (Johnson et al., 2021; Mathur &  
3599 Fantle, 2015). Similarly, given the physicochemical conditions of Carlow Castle's ore  
3600 formation (Fox et al., 2021), it is unlikely that redox changes (oxidation of  $\text{Cu}^{1+}$  to  $\text{Cu}^{2+}$ ) driven  
3601 by hydrothermal ore-forming processes would have induced significant heavy Cu isotope  
3602 fractionation as  $\text{CuCl}_2^-$  is the predominant stable Cu-Cl species in solution around 300°C, even  
3603 under oxidising conditions (Brugger et al., 2016; Liu, Brugger, McPhail, & Spiccia, 2002;  
3604 Maher, Jackson, & Mountain, 2011; Syverson, Borrok, Niebuhr, & Seyfried, 2021; Zhu et al.,  
3605 2002). Whilst aspects of the Archean crustal environment may account for the notable lack of  
3606 any  $>0\%$  fractionation of Carlow Castle's Cu isotope signature, the pattern of Cu isotope  
3607 fractionation in Carlow Castle's mineralisation most clearly reflects the evolution of an ore  
3608 system defined by Rayleigh-type fractionation due to preferential incorporation of  $^{63}\text{Cu}$  into  
3609 solid Cu phases during ore precipitation. This was associated with a consequent increase in  
3610  $\delta^{65}\text{Cu}$  in the fluid phase from an initially isotopically light composition ( $\delta^{65}\text{Cu} = <0\%$ );  
3611 evolving toward a progressively heavier composition around  $\delta^{65}\text{Cu} = 0\%$ . Similar behaviour  
3612 has previously been invoked to account for Cu isotope fractionation in various hypogene Cu  
3613 ore systems (Gregory & Mathur, 2017; Li et al., 2010; Wu, Zheng, Wang, Chang, & Tan,  
3614 2017).

### 3615 5.6.3 Copper isotope record of ore system evolution and Rayleigh fractionation

3616 The  $\delta^{65}\text{Cu}$  values of the analysed Cu sulphide samples in this study appear to record the isotopic  
3617 and physicochemical evolution of Carlow Castle's hydrothermal ore-forming system. In these  
3618 data, there is a negative relationship between  $\delta^{65}\text{Cu}$  of a given sample and the Cu grade  
3619 (measured as Cu ppm for a given intersection of drill core as reported within (Artemis  
3620 Resources Limited, 2018c)) within chalcopyrite-bearing samples (Figure 5.5). However, no  
3621 such relationship is evident for  $\delta^{65}\text{Cu}$  and Co or Au grade in the same samples. Additionally,  
3622 the relationship between  $\delta^{65}\text{Cu}$  and Cu grade doesn't seem to extend to primary chalcocite  
3623 samples that were analysed, though the number of chalcocite analyses is comparably limited.  
3624 This may be due to the influence of different stoichiometric Cu contents of chalcocite and  
3625 chalcopyrite and differences in Cu isotope fractionation during Cu precipitation into  
3626 chalcopyrite and chalcocite (Liu et al., 2021). Phase-based differences in Cu isotope  
3627 fractionation are certainly evident in the reverse case; where oxidation of chalcocite produces  
3628 heavier Cu isotope signatures in solution than chalcopyrite (Mathur et al., 2005; Mathur &  
3629 Wang, 2019). Similar behaviour can also be observed during comparisons of bornite and  
3630 chalcocite weathering (Wall et al., 2011a; Wall, Mathur, Post, & Heaney, 2011b). In addition,  
3631 the different chemical and thermodynamic stability conditions between chalcocite and  
3632 chalcopyrite almost certainly influence Cu isotope fractionation behaviour between the fluid  
3633 and solid mineral phases (Maher et al., 2011). As such, it is likely that chalcocite and  
3634 chalcopyrite within the deposit do not necessarily record the isotopic evolution of this ore  
3635 system in a comparable manner and, in general, different phases should be considered  
3636 separately.



3637

3638 Figure 5.5 - Cross-plot of Cu grade and Cu isotope signature of primary chalcopyrite  
 3639 samples from Carlow Castle ( $r = -0.7$ ,  $r^2 = 0.49$ ,  $p = 0.004$ ). Note errors bars are  $2\sigma$ .

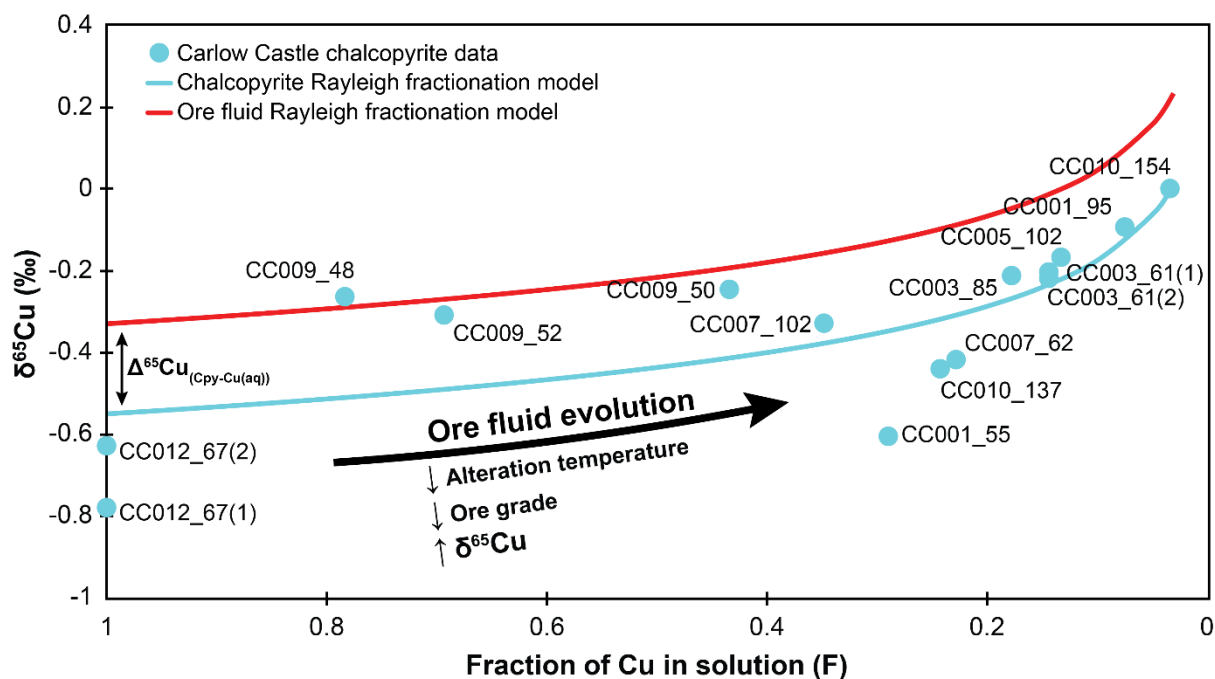
3640 A linear least squares regression (Figure 5.5) of Cu (ppm) and  $\delta^{65}\text{Cu}$  for the analysed  
 3641 chalcopyrite samples produces a moderate negative correlation ( $r = -0.7$ ,  $r^2 = 0.49$ ,  $p = 0.004$ ).  
 3642 This relationship is interpreted to represent a record of the evolution of Carlow Castle's  
 3643 hypogene ore system following a Rayleigh-type fractionation process. It is interpreted that  
 3644 sample CC012\_67(1) (Cu (ppm) = 95,600;  $\delta^{65}\text{Cu} = 0.79\text{‰}$ ), as the most Cu-rich and  
 3645 isotopically lightest sample, represents an early stage in the evolution of this ore system and  
 3646 the initiation of Cu-sulphide precipitation. As it is well established that  $^{63}\text{Cu}$  preferentially  
 3647 precipitates into the solid phase during Cu sulphide ore formation within hydrothermal systems  
 3648 (Asael et al., 2007; Bornhorst & Mathur, 2017; Ehrlich et al., 2004; Zhu et al., 2002), it is  
 3649 reasonable to interpret that the initial stages of mineralisation would be both the lightest  
 3650 isotopically and most Cu-rich. As the system evolved according to a Rayleigh fractionation  
 3651 process, progressive removal of total dissolved Cu and specifically  $^{63}\text{Cu}$  from the fluid phase  
 3652 and into the ore would be reflected in a concomitant decrease in ore grade and increase in  $\delta^{65}\text{Cu}$   
 3653 over time. It should be noted however that this hypothetical model with a finite Cu-bearing  
 3654 fluid reservoir that is almost certainly a broad simplification of the evolution of what was no  
 3655 doubt a complex hydrothermal system.

3656 Fox et al. (2021) suggested that changing abundances of actinolite, chlorite and epidote across  
 3657 the deposit's propylitic alteration assemblage indicated variations in alteration temperatures.

3658 Higher abundances of actinolite were taken to indicate higher peak temperatures of alteration  
3659 (Corbett & Leach, 1997). Notably, these variations in alteration assemblage also appear to  
3660 broadly correlate with the relationship observed in this study between Cu grade and  $\delta^{65}\text{Cu}$ . Fox  
3661 et al. (2021) observed that sample CC012\_67 displays the highest temperature propylitic (most  
3662 actinolite-rich) alteration assemblage of those samples analysed and therefore posited that it  
3663 may represent a high-temperature portion of a thermally-zoned deposit compared to lower-  
3664 temperature (actinolite and epidote-poor, chlorite-rich) mineralisation, e.g., samples  
3665 CC009\_48 and CC003\_61. This observation supports the interpretation in the current study;  
3666 that the trend in Cu grade and  $\delta^{65}\text{Cu}$  reflects the progressive evolution of Carlow Castle's  
3667 hydrothermal system during ore precipitation. There are three lines of converging evidence  
3668 from alteration temperatures, Cu grade of the samples, and Cu isotope compositions that  
3669 support this interpretation. In this case, the early stages of mineralisation characterised by an  
3670 isotopically light and Cu-rich fluid would also be expected to have been the hottest. This is  
3671 supported by the comparably high temperature alteration assemblage of CC012\_67. Here the  
3672 isotopic evolution of this ore fluid may reflect its coeval physicochemical evolution,  
3673 precipitating Cu as it equilibrated with the host wall-rock of the Ruth Well and Nickol River  
3674 Formations; becoming progressively cooler and more reduced. This is suggested by the  
3675 decreased abundance of actinolite and epidote relative to chlorite in isotopically heavier and  
3676 lower Cu grade samples, e.g., CC009\_48 and CC003\_61. The alteration assemblage here  
3677 appears to be evolving from a relatively high-temperature to comparably low-temperature  
3678 propylitic assemblage dominated by calc-silicate minerals formed at neutral to slightly alkaline  
3679 pH (6.0 at 300°C) (Reed, 1997). Therefore, the primary physicochemical control on Cu isotope  
3680 fractionation appears to be temperature (Corbett & Leach, 1997; Pirajno, 2009). As such,  
3681 changes in pH can be ruled out as a significant influence on Cu isotope fractionation in this  
3682 system. This is important to note given the influence that pH can have on Cu isotope  
3683 fractionation, due to the impact of pH on Cu solubility and fluid-mineral fractionation factors  
3684 (Gregory & Mathur, 2017; Maher et al., 2011; Mathur et al., 2013).

3685 The parameters of Carlow Castle's Rayleigh fractionation process can be estimated using the  
3686 chalcopyrite Cu grade and  $\delta^{65}\text{Cu}$  data in this study as a record of this process. It is assumed that  
3687 the chalcopyrite samples analysed in this study represent the full spectrum of Carlow Castle's  
3688 ore system evolution with isotopically lightest/heaviest and most Cu-rich/poor samples  
3689 representing the initial and final stages of this ore fluid evolution process respectively.

3690 Therefore,  $\alpha$  (fractionation factor) and  $\delta_o$  (initial isotopic composition) can be estimated  
 3691 relatively simply with a logarithmic regression through these data using a similar method to  
 3692 Scott, Lu, Cavanaugh, and Liu (2004). Within this Rayleigh fractionation model the Cu grade  
 3693 of the samples is used to represent  $F$  (fraction of remaining dissolved Cu). As sample  
 3694 CC012\_67 is interpreted to represent the earliest phase of mineralisation in this system (Figure  
 3695 5.6), when the ore fluid was most Cu-rich, Cu grades of all samples are normalised relative to  
 3696 sample CC012\_67's Cu grade (95,600 ppm) by the formula;  $F = \frac{\text{Sample grade}}{\text{Sample CC012_67 grade}}$ . A  
 3697 logarithmic expression that defines the Rayleigh fractionation model of Carlow Castle's  
 3698 chalcopyrite can then be simply derived through a least squares linear regression of the  $\delta^{65}\text{Cu}$   
 3699 data and corresponding values for  $\ln(F)$ . This gives a best fit line defined by the expression  
 3700  $y = -0.164x - 0.55$  or  $\delta^{65}\text{Cu} = -0.164(F) - 0.55$ . From this  $\alpha - 1$  can be derived as  
 3701  $\frac{-0.164}{1000} = -0.000164$  and  $\delta_o = -0.55\text{‰}$ . This defines a best fit ( $r = -0.75$ ,  $r^2 = 0.56$ ,  $p =$   
 3702  $0.0014$ ) Rayleigh fractionation model for Carlow Castle's Cu mineralisation of  $\delta^{65}\text{Cu} =$   
 3703  $(1000 - 0.55) \cdot F^{-0.000164} - 1000$ . Whilst this method allows for estimation of the Rayleigh  
 3704 fractionation model for Carlow Castle's chalcopyrite mineralisation, it is necessary to know  
 3705 the degree of isotopic fractionation between dissolved Cu in the fluid phase and precipitated  
 3706 Cu in chalcopyrite, given as  $\Delta^{65}\text{Cu}_{(\text{chalcopyrite-Cu(aq)})}$ . It has previously been assumed in some  
 3707 studies that isotopic compositions of Cu-sulphides closely reflect the isotopic composition of  
 3708 their parent ore fluid, i.e.,  $\Delta^{65}\text{Cu}_{(\text{chalcopyrite-Cu(aq)})} \approx 0$  (Asael, Matthews, Oszczepalski, Bar-  
 3709 Matthews, & Halicz, 2009). However, Syverson et al. (2021) recently used experimental data  
 3710 and natural analogues from submarine hydrothermal vent systems to demonstrate that  
 3711  $\Delta^{65}\text{Cu}_{(\text{chalcopyrite-Cu(aq)})} = -0.22 \pm 0.16\text{‰}$  under comparable hydrothermal conditions to Carlow  
 3712 Castle's formation. If it is assumed that a similar value for  $\Delta^{65}\text{Cu}_{(\text{chalcopyrite-Cu(aq)})}$  applies for  
 3713 Carlow Castle, this would give a starting fluid composition as  $\delta_o = -0.55 + 0.22 = -0.33\text{‰}$   
 3714 (Figure 5.6).



3715

3716 Figure 5.6 - Rayleigh fractionation model for Carlow Castle chalcopyrite and ore fluid  
 3717 based on estimated best-fit logarithmic regression of chalcopyrite data.

3718 Though generally the goodness-of-fit for the logarithmic regression used to construct the above  
 3719 Rayleigh fraction model for Carlow Castle's chalcopyrite Cu isotope data is acceptable ( $r^2 =$   
 3720  $0.56$ ), some outlier data points exist, e.g., CC009\_48 or CC001\_55. Whilst it may be tempting  
 3721 to remove these outliers and consider them as part of a separate system, there is little  
 3722 compelling geological evidence to suggest that this is valid. Instead, these deviations are more  
 3723 likely to reflect a variety of possible limitations of the method utilised here and associated  
 3724 assumptions. For example, the vein-hosted nature of Carlow Castle's mineralisation means that  
 3725 Cu grade measurements over a given metre sample of mineralised core are likely to vary  
 3726 significantly due to a 'nugget effect' that arises from the heterogeneous distribution of this  
 3727 mineralisation. This effect would be less pronounced if this mineralisation was instead evenly  
 3728 disseminated over a given mineralised intersection. This is illustrated by the difference in Cu  
 3729 grade between samples CC009\_48 (74,900 ppm Cu) and CC009\_50 (41,500 ppm Cu) which  
 3730 occur 2 metres in depth apart in the same drill hole but vary in Cu grade by 33,400 ppm  
 3731 ( $\sim 80.5\%$ ) whilst only varying  $0.02\%$  in  $\delta^{65}\text{Cu}$ . Additionally, the complex mineralogy of some  
 3732 samples, which may contain several different Cu-bearing phases with different stoichiometric  
 3733 Cu concentrations (e.g., chalcocite and chalcopyrite), may complicate these grade  
 3734 measurements over a given metre intersection. Finally, it may simply be that the Rayleigh  
 3735 fractionation model developed here to represent Carlow Castle's ore-forming system is a

3736 necessarily idealised model of a complex natural system and as such fails to account for  
3737 isotopic fractionation in all samples perfectly. Nonetheless, this still provides an approximate  
3738 model of Carlow Castle's ore system evolution through the lens of Cu isotopes.

3739 The initial Cu isotopic composition of Carlow Castle's ore fluid ( $\delta^{65}\text{Cu} = -0.304\text{‰}$ ) is broadly  
3740 consistent with derivation of Cu from a minimally fractionated igneous source, which would  
3741 be expected to display an isotopic composition in the range of  $\delta^{65}\text{Cu} = 0.06 \pm 0.20\text{‰}$  (Liu et  
3742 al., 2015). In this instance, derivation from an unfractionated and homogeneous igneous source  
3743 may account for the relatively limited fractionation and broadly normally distributed Cu  
3744 isotope signature evident in Carlow Castle's Cu sulphide mineralisation. However, there is no  
3745 evidence of large-scale magmatic activity in the West Pilbara Superterrane that was both coeval  
3746 with Carlow Castle's ore formation (c. 2.95 Ga) and proximal to the deposit to provide an  
3747 igneous driver of ore formation and source of metal (Fox et al., 2021). Nonetheless, an igneous  
3748 metal source would be consistent with previous hypotheses, as it has been suggested that the  
3749 greenstone belts that form much of the surrounding terranes could provide a potential source  
3750 of metalliferous mafic rocks from which ore metals could be liberated (Fox et al., 2021; Fox et  
3751 al., 2019). This is consistent with the established association between mafic igneous rocks and  
3752 significant hydrothermal Co deposits, including in non-magmatic-hydrothermal deposits  
3753 (Hitzman et al., 2017). In this regard, the Cu isotopic signature of Carlow Castle's ore is  
3754 potential evidence of a mafic igneous metal source, most likely from the surrounding  
3755 greenstone belts which are rich in mafic and ultramafic igneous rocks.

#### 3756 5.6.4 Copper isotope fractionation in supergene Cu samples

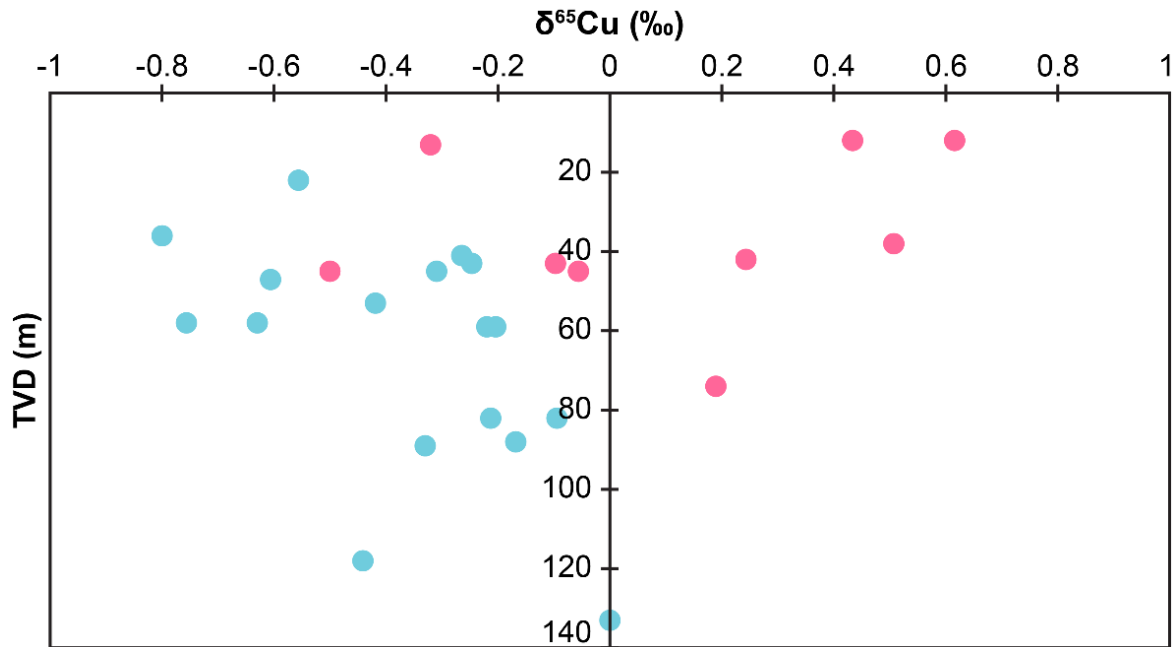
3757 Generally, the most significant fractionation of Cu isotopes in ore-forming systems is observed  
3758 in secondary minerals produced through oxidative weathering of underlying primary ore  
3759 sulphides (Larson et al., 2003; Sherman, 2013), where redox reactions generate significant  
3760 enrichment of  $^{65}\text{Cu}$  ( $\Delta^{65}\text{Cu}_{(\text{supergene Cu-primary Cu})}$ ) in resulting supergene Cu minerals up to  $>10\text{‰}$   
3761 (Kim et al., 2019). This is due to preferential mobilisation of  $^{65}\text{Cu}$  into solution during oxidative  
3762 weathering of primary Cu sulphide minerals (Markl et al., 2006; Mathur et al., 2005; Mathur  
3763 & Wang, 2019). However, in the current study the enrichment of  $^{65}\text{Cu}$  in supergene phases is  
3764 somewhat limited, only ranging up to a maximum of  $0.62\text{‰}$  in sample CC006\_14(2). Between  
3765 the isotopically lightest primary Cu sample (CC007\_41,  $\delta^{65}\text{Cu} = -0.80\text{‰}$ ) and the isotopically  
3766 heaviest supergene Cu sample this gives a relatively limited  $\Delta^{65}\text{Cu}_{(\text{supergene Cu-primary Cu})} = 0.62\text{‰}$   
3767  $- -0.80\text{‰} = 1.42\text{‰}$ . Generally, large magnitude isotopic fractionation of supergene Cu minerals

3768 is interpreted to result from multiple cycles of isotopic enrichment from successive oxidative  
3769 weathering and reprecipitation of Cu, due to the comparably limited isotopic fractionation  
3770 produced by single-cycle weathering experiments (Braxton & Mathur, 2011; Mathur, Dendas,  
3771 Titley, & Phillips, 2010; Mathur et al., 2005; Moynier et al., 2017). Reflecting this, the limited  
3772 Cu isotope fractionation within Carlow Castle's supergene Cu mineralisation suggests that this  
3773 supergene Cu has undergone limited redox cycling (Mathur et al., 2010). Although the  $^{65}\text{Cu}$   
3774 enrichment in these supergene phases is relatively moderate by comparison to other studies,  
3775 these supergene phases are isotopically distinct from Carlow Castle's primary Cu  
3776 mineralisation as the majority of the supergene Cu isotope data display  $\delta^{65}\text{Cu}$  values  $>0\text{‰}$ . In  
3777 contrast, Carlow Castle's primary mineralisation exclusively displays  $\delta^{65}\text{Cu}$  values  $\leq 0\text{‰}$ .  
3778 Additionally, it is likely that even the isotopically lightest supergene sample in this study  
3779 (CC006\_52(1);  $\delta^{65}\text{Cu} = -0.50\text{‰}$ ) is still heavily fractionated relative to the primary Cu sulphide  
3780 ore from which it was derived. This is reflected by the fact that the lightest supergene sample  
3781 is still isotopically heavier than the lightest primary Cu sulphide sample (CC007\_41;  $\delta^{65}\text{Cu} =$   
3782  $-0.80\text{‰}$ ), though the magnitude of heavy fractionation is clearly relatively limited.

3783 A significant portion of the existing literature on Cu isotope analysis of supergene materials in  
3784 ore deposit settings has focussed on porphyry Cu deposits. Here spatially distinct Fe-oxide  
3785 leach caps, underlying supergene Cu enrichment zones, and deep-lying hypogene Cu sulphide  
3786 mineralisation form well-developed, spatially extensive, discrete Cu isotope reservoirs (Asadi,  
3787 Mathur, Moore, & Zarasvandi, 2015; Braxton & Mathur, 2011; Mathur et al., 2010; Mathur et  
3788 al., 2009; Palacios, Rouxel, Reich, Cameron, & Leybourne, 2011). Very little of this literature  
3789 is focused on the effect of supergene processes on Cu isotope fractionation in structurally  
3790 complex vein-hosted Cu ore systems and even less literature is focused on Cu fractionation  
3791 between Archean Cu sulphide ores and contemporary supergene Cu mineralisation in Archean  
3792 terranes. The complex structurally-hosted nature of Carlow Castle's mineralisation means that  
3793 the development of discrete supergene and hypogene sub-zones is less spatially uniform. The  
3794 development of these zones is likely to be strongly influenced by the effect of individual  
3795 structures on regional permeability and groundwater flow, which would have preferentially  
3796 focussed fluids and therefore chemical weathering along certain permeable structures (Babiker  
3797 & Gudmundsson, 2004; Ferrill et al., 1999). As a result, the depth of supergene Cu samples  
3798 varies significantly across different drill holes. To illustrate this, samples CC011\_97 and  
3799 CC002\_16 are both supergene Cu samples that contain plancheite; however, they were sampled



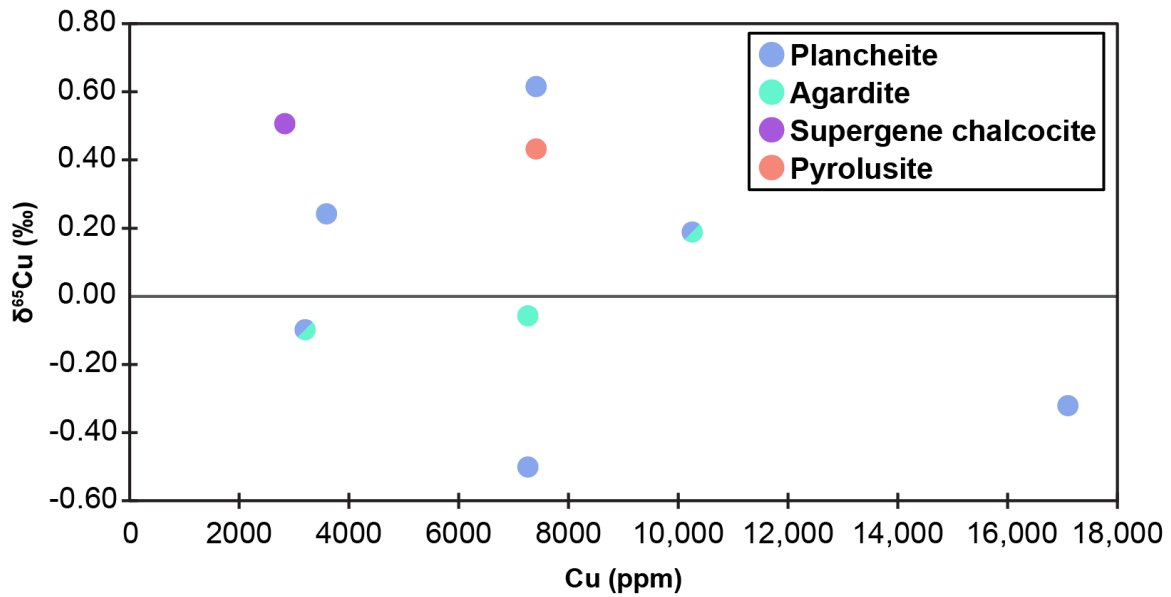
3800 from different drill holes and occur 61 metres apart in true vertical depth (TVD). Further,  
 3801 sample CC008\_25 is a primary chalcocite sample and occurs at only 22 metres TVD; 52 metres  
 3802 shallower than the deepest supergene sample. Reflecting this, there is no discernible trend  
 3803 evident between the  $\delta^{65}\text{Cu}$  values and depths of different samples in this study (Figure 5.7).



3804  
 3805 Figure 5.7 - Cross-plot of Cu isotope signature and true vertical depth of analysed  
 3806 samples. There appears to be no clear correlation between sample depth and isotopic  
 3807 signature for supergene samples. Also, some supergene Cu samples were retrieved from  
 3808 greater depths than some of the primary Cu ore samples, reflecting Carlow Castle's  
 3809 structurally complex vein-hosted mineralisation. Note that samples that were not  
 3810 retrieved from drill core have an unknown original depth and therefore are not presented  
 3811 here.

3812 Unlike the clear relationship between Cu grade and  $\delta^{65}\text{Cu}$  that is evident across the primary  
 3813 chalcopyrite data from Carlow Castle, no comparable relationship is evident across the  
 3814 supergene Cu isotope data in this study (Figure 5.8). This precludes the construction of a  
 3815 Rayleigh fractionation model for the supergene data set. These supergene data are complicated  
 3816 by mineralogical complexity comprising several different Cu-bearing phases, which almost  
 3817 certainly undergo distinct phase-specific isotopic fractionation processes during Cu  
 3818 precipitation and contain different stoichiometric proportions of Cu (Liu et al., 2021). This is  
 3819 exemplified by the relatively significant difference in isotopic composition between plancheite  
 3820 in sample CC006\_52(1) ( $\delta^{65}\text{Cu} = -0.50\text{‰}$ ) and agardite in sample CC006\_52(2) ( $\delta^{65}\text{Cu} = -$

3821 0.06‰), even though these samples were extracted from the same core sample. In addition, it  
3822 is likely that the range of supergene samples essentially represent several independent, small-  
3823 scale supergene systems along a given fluid pathway via in-situ alteration of primary sulphide  
3824 ore in a given vein with spatially limited Cu transport. This is implied by (i) the significant  
3825 difference in depth between supergene samples analysed in this study due to in-situ alteration  
3826 of primary sulphide veins with limited lateral Cu transport, (ii) spatially distinct and  
3827 disconnected supergene Cu isotope reservoirs, (iii) probable channelling of regional oxidised  
3828 groundwater flow along permeable mineralised structures with minimal fluid interaction  
3829 between hydrogeologically distinct structures. If these supergene samples represent multiple  
3830 small hydrogeologically heterogeneous oxidised vein systems, it is not geologically valid to  
3831 attempt to construct a Rayleigh fractionation model based on these data. The spatial restriction  
3832 of these supergene Cu minerals to small vein-hosted systems rather than an extensive supergene  
3833 enrichment blanket, as is commonly observed in analyses of supergene Cu minerals from  
3834 porphyry deposits (Asadi et al., 2015; Braxton & Mathur, 2011), may have limited repeated Cu  
3835 transport and redox cycling in these supergene minerals. This most likely accounts for the  
3836 limited isotopic enrichment in Carlow Castle's supergene Cu mineralisation. Given the wide  
3837 interest in the potential use of Cu isotope analysis of ore minerals for potential exploration  
3838 applications (Braxton & Mathur, 2011; Mathur et al., 2013; Mathur et al., 2009; Mathur &  
3839 Wang, 2019), this may be an important consideration when exploring in structurally complex  
3840 terranes or for structurally-hosted ore deposits.



3841

3842 Figure 5.8 - Cross-plot of Cu grade (ppm) and isotopic signature of supergene Cu samples.

3843 Note that error bars are not visible as they are smaller than the sample symbols.

### 3844 5.7 Conclusions and implications

3845 In this study we provide the first Cu isotope analysis of an Archean Cu ore deposit; the  
 3846 Mesoarchean Carlow Castle Cu-Co-Au deposit in the Pilbara Craton, Western Australia.

3847 Copper isotope fractionation in Carlow Castle's primary Cu sulphide ore is relatively limited,  
 3848 consistent with its Archean age and hypogene origin; ranging between  $\delta^{65}\text{Cu} = -0.8$  and  $0\text{‰}$ .

3849 Nonetheless, these values are fractionated beyond the bounds of a bulk silicate Earth  
 3850 composition. This isotopic fractionation is attributed to Rayleigh fractionation due to

3851 preferential incorporation of  $^{63}\text{Cu}$  into the solid phase during Cu sulphide precipitation. We

3852 observe that, based on an apparent relationship with Cu grade in analysed ore samples, the Cu

3853 isotopic signature of chalcopyrite samples appear to record the physicochemical evolution of

3854 Carlow Castle's ore system during this process. This is further supported by an apparent

3855 relationship between alteration mineral assemblage and Cu isotope signature in the analysed

3856 samples. Modelling of this Rayleigh fractionation process based on the isotopic composition

3857 of analysed chalcopyrite gives an estimated initial fluid Cu isotopic composition around  $\delta^{65}\text{Cu}$

3858  $= -0.3\text{‰}$ . This is interpreted as evidence of a probable mafic igneous metal source with minimal

3859 input from a fractionated crustal source and supports previous hypotheses for deposit

3860 formation. These findings suggest that Cu isotope analysis also has potential utility in

3861 understanding genetic processes in Archean ore-forming systems and Cu transport in Archean

3862 Earth systems more broadly. However, the current understanding of Archean Cu isotope

3863 systematics suffers from a relatively small body of existing literature. In addition, Carlow  
3864 Castle represents a relatively unique variety of Archean structurally-hosted hydrothermal Cu-  
3865 Co-Au deposit (Fox et al., 2021; Fox et al., 2019). As such, this may limit the wider  
3866 applicability of these findings. Nonetheless, this underlines the importance of expanding the  
3867 existing literature applying copper isotopes to understand Archean geological processes.  
3868 Analysis of supergene Cu minerals from Carlow Castle display relatively limited heavy  
3869 fractionation from their primary Cu sulphide source; ranging between  $\delta^{65}\text{Cu} = -0.50$  to  $0.62\%$ .  
3870 This is suggestive of limited redox cycling of Cu during the formation of these supergene Cu  
3871 minerals and is interpreted to reflect the lack of a large, shallow, hydrogeologically uniform  
3872 supergene enrichment blanket produced by multiple cycles of Cu enrichment due to oxidative  
3873 groundwater flux. Instead, the occurrence of these supergene Cu minerals within primary ore  
3874 veins due to partial or complete in-situ replacement of primary Cu sulphide minerals suggests  
3875 that they represent several hydrogeologically disconnected supergene systems with minimal  
3876 intermixing. This may be an important consideration when utilising Cu isotopes for exploration  
3877 for structurally complex Cu ore deposits and potentially warrants further study. Nonetheless  
3878 the fractionation of these supergene phases relative to their primary Cu ore may still suggest  
3879 that Cu isotope analysis could be a promising tool for exploration in Archean terranes.

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3889

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- 4108

4109 **Appendix 5.1 – Chromatography procedure**

4110 Appendix 5.1 is accessible via this [link](#) and provides an .xlsx file containing the final procedure,  
4111 with results, used for Cu separation for Cu isotope analysis.

4112

4113 **Appendix 5.2 – Long term analyses of reference materials**

4114 Appendix 5.2 is accessible via this [link](#) and provides an .xlsx file containing results of analyses  
4115 of reference materials DFCu and BHNatCop from several analytical sessions between February  
4116 2020 and October 2022, demonstrating the long-term reproducibility of reference values.

4117

# 4118 Chapter 6

## 4119 Synthesis and conclusions

4120 The primary objective of this research project was to provide insights into the nature and  
4121 genesis of the recently discovered Carlow Castle Cu-Co-Au ore system. Whilst a secondary  
4122 objective was to establish a working methodology for Cu isotope analysis in collaboration with  
4123 the Metal Isotopes Group, University of Adelaide and use Carlow Castle as a case study to test  
4124 this methodology. As this thesis is presented as a series of individual papers, each chapter has  
4125 dealt with different aspects of these objectives, with the conclusions and implications of each  
4126 chapter stated at the end of that chapter. As such, this concluding chapter focusses on  
4127 synthesising these findings to construct a simplified mineral system model for Carlow Castle,  
4128 along with a discussion of the broader implications of this research for the understanding of  
4129 Cu-Co-Au metallogenesis through geological time, the Mesoarchean sulphur isotope record,  
4130 and the application of Cu isotope analysis to understanding the genesis of and exploring for  
4131 Archean hydrothermal Cu deposits. This is accompanied by discussion of limitations of the  
4132 research along with avenues for future research.

### 4133 6.1 New constraints on the genesis of Carlow Castle Cu-Co-Au deposit

4134 The mineral systems approach to holistically defining the confluence of geological processes  
4135 at various scales that produce significant ore deposits was first formally proposed by Wyborn,  
4136 Heinrich, and Jaques (1994), though elements of this approach predate formalisation in the  
4137 cited study (Hayes & Einaudi, 1986). This concept has since become the standard framework  
4138 to define ore-forming systems in studies of ore deposit petrology (Hagemann, Lisitsin, &  
4139 Huston, 2016; Huston et al., 2016; McCuaig, Beresford, & Hronsky, 2010; McCuaig et al.,  
4140 2014). Therefore, the findings of the different analytical approaches utilised to investigate

4141 Carlow Castle Cu-Co-Au deposit in this thesis are most logically synthesised through the  
4142 development of a Carlow Castle Cu-Co-Au mineral system model. Although the components  
4143 of the mineral systems framework vary slightly depending on mineralisation style, there are  
4144 generally five key elements; (i) an energy source to drive the formation of the mineral system,  
4145 (ii) ligands to complex with dissolved metals and allow for hydrothermal metal mobility, (iii)  
4146 a fertile metalliferous source to provide ore metals for concentration into the ore deposit, (iv)  
4147 a pathway through which ore fluids can be focussed and metals transported to their site of  
4148 deposition, and (v) a physical or chemical trap to induce metal deposition and ore formation  
4149 (Hagemann, Cassidy, Hagemann, & Brown, 2000; Hagemann et al., 2016; Huston, Stevens,  
4150 Southgate, Muhling, & Wyborn, 2006). The components of this mineral systems model follow  
4151 in the proceeding text and are summarised in Table 6.1 at the end of this section.

4152 Apatite U-Pb geochronology (Chapter Three) suggests that Carlow Castle formed around 2955  
4153 Ma during the late stages of formation of the De Grey Superbasin (3066-2919 Ma).  
4154 Specifically, this age of mineralisation coincided with the final stages of extension during the  
4155 formation of the Mallina Basin (3015-2931), following subduction at the northwest margin of  
4156 the Pilbara Craton. Whilst a genetic connection to the Prinsep Orogeny (~3070 Ma) had  
4157 previously been suggested (Hickman, 2016), the age constraints provided in this thesis indicate  
4158 that Carlow Castle significantly post-dates this orogenic event. As such, it is likely that the  
4159 underlying energy source that drove the formation of Carlow Castle was related to increased  
4160 heat flow during proximal rifting of the De Grey Superbasin and extensional reactivation of  
4161 the Regal Thrust. Ultimately a magmatic source for the Carlow Castle mineral system is not  
4162 evident, unlike coeval base-metal VMS mineralisation in the Whim Creek Basin, which is  
4163 contemporaneous with proximal granitic magmatism of the Sisters Supersuite (2954-2919 Ma).  
4164 Consequently, large-scale convective flow of hydrothermal fluids driven by increased

4165 geothermal gradients due to regional extension and rifting are likely to have been the primary  
4166 driver of the Carlow Castle mineral system.

4167 The source and nature of metal complexing ligands in the Carlow Castle ore system is difficult  
4168 to constrain without detailed analyses of ore-stage fluid inclusions. However, the most likely  
4169 ligand for complexation with Cu and Co, was  $\text{Cl}^-$ , in addition to  $\text{OH}^-$  in the case of Au. This is  
4170 suggested given the predominance of  $\text{Cl}^-$  complexation to transport Cu and Co in hydrothermal  
4171 ore systems (Heinrich & Candela, 2014; Seward et al., 2014), and absence of evidence of  
4172 alternative halide ligands such as  $\text{F}^-$  (Chapter Two). Although reduced sulphur species, such  
4173 as  $\text{HS}^-$ , commonly transport Au in hydrothermal solutions, the coeval paragenesis of Cu and  
4174 Co with Au mineralisation within Carlow Castle imply synchronous transport of these ore  
4175 metals. As it is unlikely that Cu and Co would be simultaneously soluble with Au in a  $\text{HS}^-$ -rich  
4176 fluid (Heinrich & Candela, 2014), also evidenced by the thermodynamic modelling conducted  
4177 in Chapter Four,  $\text{Cl}^-$  and/or  $\text{OH}^-$  is suggested as the most likely ligand to complex with Au.  
4178 Additionally, considering that Carlow Castle would most likely have been situated at the base  
4179 of the Gorge Creek Basin at the time of ore formation, saline basinal fluids from the overlying  
4180 sedimentary sequence may have contributed significantly to the Carlow Castle ore system. Any  
4181 seawater contribution to such basinal fluids would have been relatively saline (6-24 wt. % NaCl  
4182 equiv.), given the comparably greater salinity of Archean and Proterozoic seawaters relative to  
4183 those of the Phanerozoic (Huston et al., 2010). This interpretation is supported by the slightly  
4184 heavy  $\delta^{34}\text{S}$  values of sulphide mineralisation at Carlow Castle (Chapter Four), which may be  
4185 indicative of an isotopically heavy Archean seawater sulphate ( $\delta^{34}\text{S} = 3-8\%$ ) contribution from  
4186 a basinal brine to the hydrothermal system (Crowe et al., 2013). The propylitic alteration  
4187 assemblage that characterises the Carlow Castle deposit is interpreted to reflect a neutral pH  
4188 ( $\sim 6.0$ ) ore fluid and peak fluid temperatures of  $\sim 300^\circ\text{C}$  (Chapter Three). In consideration of



4189 this evidence and thermodynamic modelling of the hydrothermal mobility of Cu and Co, it is  
4190 argued that these conditions would have been consistent with metal mobilisation under  
4191 oxidised conditions (Chapter Four). These relatively moderate fluid conditions ( $T \approx 300^{\circ}\text{C}$ ,  $\text{pH}$   
4192  $\approx 6.0$ ) further support the above interpretation of derivation of the Carlow Castle ore fluid from  
4193 a basinal brine.

4194 Unequivocally constraining metal sources in ore-forming systems is often challenging  
4195 (Goldfarb & Groves, 2015; Groves et al., 2022). However, it is suggested herein that the mafic  
4196 rocks that constitute a significant portion of the Roebourne greenstone belt surrounding Carlow  
4197 Castle are the most probable source of ore metals for the Carlow Castle mineral system. This  
4198 was initially suggested based on the strong metallogenic relationship between mafic igneous  
4199 rocks and Co-rich ore deposits (Chapter Two), due to the significantly greater concentration of  
4200 Co in mafic and ultramafic igneous rocks relative to other lithologies (Slack, Kimball, & Shedd,  
4201 2017b). This interpretation is also supported by Cu isotope analysis and Rayleigh fractionation  
4202 modelling (Chapter Five). This Rayleigh fractionation modelling gave an estimated initial ore  
4203 fluid Cu isotope composition of  $\delta^{65}\text{Cu} = 0.33\%$ , which is consistent with an igneous rock  
4204 reservoir (Liu et al., 2015). The lack of significant variation in Cu isotope composition between  
4205 Co-rich and Co-poor ore mineral assemblages within the Carlow Castle ore body was  
4206 interpreted to reflect a consistent metal source between these assemblages.

4207 Given the occurrence of Carlow Castle through higher-order structures of the crustal-scale  
4208 Regal Thrust, it is suggested to represent the critical fluid pathway that facilitated metal  
4209 transport and concentration. Given the significant spatial extent of the Regal Thrust over the  
4210 present-day northwest Pilbara Craton (Hickman, 2016), it is likely that it acted as a major  
4211 regional conduit for fluid flow following extensional reactivation during formation of the De  
4212 Grey Superbasin. Channelling of regional fluid flow, potentially from the overlying Gorge

4213 Creek Basin, along the Regal Thrust and through metal-rich mafic rocks of the Roebourne  
4214 greenstone belt would have allowed these ore fluids to be focussed into the mylonitised Carlow  
4215 Castle shear zone. Following migration into the heavily tectonised Carlow Castle shear zone,  
4216 metal precipitation most likely resulted from gradual equilibration of the Carlow Castle ore  
4217 fluid with the cold, reduced wall-rocks of the Ruth Well and Nickol River Formations  
4218 (Chapters Three and Five). This process is recorded through the lens of Cu isotopes analysis  
4219 (Chapter Five), from an initially high alteration temperature ( $\sim 300^{\circ}\text{C}$ ), isotopically light, and  
4220 Cu-rich fluid to a comparably low temperature ( $<300^{\circ}\text{C}$ ), isotopically heavy, and Cu-poor  
4221 fluid. This is evidenced by a transition from actinolite- and epidote-dominated to chlorite-  
4222 dominated alteration assemblages, with a corresponding increase in  $\delta^{65}\text{Cu}$  and decrease in Cu  
4223 grade (ppm) in analysed samples due to Rayleigh isotope fractionation. Additionally, the  
4224 progressive increase in sulphur fugacity within the ore mineral paragenetic sequence noted in  
4225 Chapter Two may also be a symptom of this fluid evolution process. A progressive decline in  
4226 temperature and oxygen fugacity and consequent decline in sulphur solubility within the  
4227 system (Tornos & Heinrich, 2008), especially during the transition from sulphate- to sulphide-  
4228 stable conditions, would be reflected in an apparent increase in sulphur precipitation into solid  
4229 mineral phases. Whilst it was initially suggested that this observed increase in sulphur fugacity  
4230 could reflect mixing with a sulphide-bearing sulphur source at the site of mineralisation, this is  
4231 not evident in the sulphur isotope analysis conducted in Chapter Four. If this increase in sulphur  
4232 fugacity was a product of the progressive addition of externally derived reduced sulphur to the  
4233 system, this would most likely be reflected in distinct sulphur isotope compositions for low  
4234 sulphur fugacity and high sulphur fugacity ore minerals. However, the indistinguishable  
4235 sulphur isotope compositions of pyrrhotite (lower  $f\text{S}_2$ ) and pyrite (higher  $f\text{S}_2$ ) analysed in  
4236 Chapter Four suggest that the progressive increase in sulphur fugacity within solid mineral

4237 phases reflects declining sulphur solubility due to declining temperature and oxygen fugacity,  
4238 rather than the progressive addition of external sulphur to the hydrothermal system.

4239 Table 6.1 - Summary of mineral system model components for Carlow Castle Cu-Co-Au deposit.

| <b>Mineral system component</b> | <b>Carlow Castle Cu-Co-Au deposit</b>  | <b>Rationale</b>   |
|---------------------------------|--|--|
| Driver/energy source            | <ul style="list-style-type: none"> <li>• Major crustal extension and rifting of the proximal/overlying De Grey Superbasin.</li> <li>• Reactivation of crustal-scale Regal Thrust host structure around 2950 Ma.</li> <li>• Enhanced geothermal gradient related to crustal rifting, driving fluid flow.</li> <li>• Lack of proximal magmatic driver, unlike contemporaneous base-metal deposits in the Mallina and Whim Creek basins.</li> </ul>               | <ul style="list-style-type: none"> <li>• Syn-ore apatite U-Pb dating, suggests a probable age around 2950 Ma, significantly post-dating Prinsep Orogeny (~3070 Ma).</li> <li>• Coincident crustal extension and formation of De Grey Superbasin.</li> <li>• Moderate peak alteration temperature (~300°C), not consistent with high temperature magmatic system.</li> <li>• Lack of proximal contemporaneous magmatic rocks associated with ore-formation in other deposits, i.e., Sisters Supersuite (2954-2919 Ma).</li> </ul> |
| Metal complexing ligands        | <ul style="list-style-type: none"> <li>• Cu, Co, and Au complexation with Cl<sup>-</sup> the most likely pathway for metal mobilization.</li> <li>• OH<sup>-</sup> also a viable ligand for Au.</li> <li>• Complexation of metals with reduced sulphur species (HS<sup>-</sup>) or alternative halide (F<sup>-</sup>) ligands unlikely.</li> <li>• Cl<sup>-</sup> potentially sourced from basinal brines of proximal/overlying De Grey Superbasin.</li> </ul> | <ul style="list-style-type: none"> <li>• Predominance of Cl<sup>-</sup> complexation as the primary pathway for metal mobilization in hydrothermal Cu-Co deposits.</li> <li>• Evidence from thermodynamic and experimental data on hydrothermal Cu, Co, and Au mobility.</li> <li>• Coeval mineralization of Cu, Co, and Au precludes transportation of Au within reduced sulphur complexes.</li> <li>• Ore formation coincident with formation of De Grey Superbasin; potential source of saline basinal brines.</li> </ul>     |

|               |   |   |
|---------------|---|---|
|               |   | <ul style="list-style-type: none"> <li>Moderate temperature, neutral pH ore fluid consistent with basinal brine.</li> </ul>   |
| Metal source  | <ul style="list-style-type: none"> <li>Mafic igneous rocks in surrounding Roebourne greenstone belt.</li> </ul>   | <ul style="list-style-type: none"> <li>Established association between mafic/ultramafic rocks and Co deposits due to Co-rich composition.</li> <li>Significant volume of mafic and ultramafic igneous rocks is surrounding host terrane and host formation.</li> <li>Cu isotope data suggest an initial fluid isotopic signature consistent with a mafic igneous source.</li> </ul> |
| Fluid pathway | <ul style="list-style-type: none"> <li>Regal Thrust and connecting higher-order structures within Carlow Castle shear zone.</li> </ul>  | <ul style="list-style-type: none"> <li>Occurrence of Carlow Castle Cu-Co-Au deposit through shear zone within Regal Thrust.</li> <li>Strongly structurally-controlled nature of Cu-Co-Au mineralisation.</li> <li>Significant regional extent and crustal-scale of Regal Thrust, through mafic rocks of Roebourne greenstone belt.</li> </ul>                                       |
| Metal trap    | <ul style="list-style-type: none"> <li>Channeling of ore fluids into complex trap structures within Carlow Castle shear zone.</li> <li>Interaction between Carlow Castle ore fluid and wall-rocks.</li> <li>Equilibration of hot (~300°C) oxidised metal-bearing ore fluid with cold and reduced host rocks within Carlow Castle shear zone.</li> </ul> | <ul style="list-style-type: none"> <li>Transition in propylitic alteration assemblage from actinolite and epidote dominated to chlorite dominated records change in temperature but not fluid pH or composition.</li> <li>Progressive ore fluid evolution recorded by correlation between Cu isotope signature, Cu grade, and alteration assemblage.</li> </ul>                     |

|  |  |  |
|--|--|--|
|  | <ul style="list-style-type: none"><li>• Transition from oxidised <math>\text{SO}_4^{2-}</math>-stable to reduced <math>\text{HS}^-</math>-stable fluid conditions.</li></ul> | <ul style="list-style-type: none"><li>• Progressive increase in sulphur fugacity reflecting decreasing sulphur solubility during cooling and declining oxygen fugacity from <math>\text{SO}_4^{2-}</math>-stable to <math>\text{HS}^-</math>-stable conditions.</li><li>• Thermodynamic modelling supports a transition from oxidised to reduced fluid conditions to induce metal precipitation.</li></ul> |
|--|--|--|

4240

## 4241 6.2 Research implications, limitations, and outlook

4242 Whilst Carlow Castle bears some similarities to previously observed Cu-Co(-Au) deposits,  
4243 especially those of the relatively obscure polymetallic/Co-rich vein and metasedimentary rock-  
4244 hosted Co-Cu-Au deposit types, e.g., Bou Azzer, Morocco; Cobalt, Ontario; Blackbird district,  
4245 Idaho (Slack et al., 2017b). The Archean age of Carlow Castle clearly distinguishes it from  
4246 these deposit types, of which no Archean examples are known. In fact, Carlow Castle is the  
4247 oldest dated Cu-Co(-Au) deposit on Earth. As such, it is suggested that Carlow Castle may  
4248 represent a distinct Archean analogue of the polymetallic/Co-rich vein deposit class due to  
4249 similarities in geological characteristics between Carlow Castle and deposits of this class.

4250 The rarity of Archean Cu-Co deposits in general is primarily due to the previously noted redox-  
4251 sensitive nature of Cu and Co stability in hydrothermal environments and the limited supply of  
4252 necessary oxidative fluids to crustal environments to mobilise these metals during the Archean.  
4253 Therefore, the Mesoarchean age of Carlow Castle is significant as it suggests that the onset of  
4254 hydrothermal Cu-Co ore-forming processes under relatively moderate hydrothermal conditions  
4255 stretches back into the Archean, rather than from the Proterozoic following the Great  
4256 Oxygenation Event. Importantly, this implies that analogous Cu-Co-Au deposits could exist  
4257 throughout Archean terranes. However, the understanding of ore-forming processes in many  
4258 of the more obscure hydrothermal Co-rich deposit types, i.e., non-sediment-hosted Cu-Co  
4259 deposits, is currently relatively lacking. Whilst this thesis attempts to provide some insights  
4260 into the genesis of a single Archean hydrothermal Cu-Co-Au deposit, further study of the  
4261 relatively rare processes that produce significant Co enrichment in some hydrothermal ore  
4262 deposits is still required, especially amongst Archean examples. Further study of the genesis  
4263 of obscure Co-rich deposit classes (e.g., polymetallic/Co-rich vein, black shale-hosted Ni-Cu-  
4264 Zn-Co, and metasedimentary rock-hosted Co-Cu-Au deposits) may provide insights into the

4265 nature of these ore-forming processes. This may also include further study of the Carlow Castle  
4266 mineral system, as there are still aspects of the mineral system to which further study would be  
4267 beneficial. To this end, a fluid inclusion study of Carlow Castle would form a logical  
4268 progression of the analysis conducted in this thesis and enhance the understanding of the  
4269 Carlow Castle mineral system, particularly regarding the nature of the ligands and ore fluid  
4270 that enabled metal mobilisation. However, comparative analysis of critical genetic differences  
4271 between Co-poor and Co-rich deposits amongst relatively well understood mineralisation  
4272 styles (e.g., sediment-hosted deposits and volcanogenic massive sulphide deposits) may also  
4273 be effective in identifying the distinct ore-forming processes that impart significant Co  
4274 endowments in some hydrothermal ore deposits of a given class, but not others.

4275 On a more regional-scale, the probable genetic relationship of Carlow Castle Cu-Co-Au deposit  
4276 to the latter stages of development of the Mallina Basin within the De Grey Superbasin around  
4277 2950 Ma reaffirm previous observations that this period produced the greatest variety of  
4278 mineralization of any period in the development of the north west Pilbara Craton (Hickman,  
4279 2016). This observation is also strengthened by recent discoveries of significant Au  
4280 mineralisation within the Mallina Basin; most notably the 6.8 million ounce Hemi Au deposit  
4281 discovered in 2019 (De Grey Mining Limited, 2021). This may be of significance for the  
4282 purpose of understanding regional metallogenesis during the formation of the Pilbara Craton  
4283 and for broad-scale exploration targeting. However, this thesis has focussed primarily on the  
4284 mineralogy, geochronology, and geochemistry of the Carlow Castle ore deposit. A study  
4285 centred around both the broad structure and microstructural evidence for potential late  
4286 deformation within the Carlow Castle, that could be correlated with regional studies on the  
4287 structural evolution of the Pilbara Craton (Kiyokawa, Aihara, Takehara, & Horie, 2019;



4288 Kiyokawa et al., 2002), would potentially provide useful insights into the formation of Carlow  
4289 Castle and metallogenesis within the Pilbara Craton more broadly.

4290 Results from multiple sulphur isotope analysis and thermodynamic modelling conducted in  
4291 Chapter Four inform the understanding of the evolution of the Archean atmosphere and its  
4292 potential relationship with Archean metallogeny. In particular, the notable absence of evidence  
4293 of mass-independent sulphur isotope fractionation (MIF-S) supports the potential existence of  
4294 a MIF-S minimum within during the Mesoarchean (Kasting & Ono, 2006). This is believed to  
4295 have coincided with the early evolution of oxygenic photosynthesis and the development of  
4296 one of Earth's first weakly ephemerally oxygenated surface environments, within the Pongola  
4297 Basin around 2950 Ma. Evidence of oxygenation in the Pongola Basin has been implicated in  
4298 triggering the Pongola Glaciation through oxidation of the methane-rich Archean atmosphere  
4299 (Crowe et al., 2013; Eickmann et al., 2018; Ono et al., 2006; Planavsky et al., 2014). An  
4300 oxidative origin for Carlow Castle is further supported through thermodynamic modelling of  
4301 Cu-Co solubility under the established conditions of ore formation and consideration of the  
4302 geological context of the deposit. As such, Carlow Castle could contribute distal evidence of a  
4303 partially oxygenated atmosphere. The coincident timing of the formation of Carlow Castle and  
4304 the Pongola Glaciation could demonstrate that oxygenic photosynthesis was already relatively  
4305 widespread by the Mesoarchean. Though a potential mantle model is discussed and contrary  
4306 reasoning presented in Chapter Four, some readers may find the oxidative model for the origin  
4307 of Carlow Castle to be contentious as the potential for ephemeral oxygenation of surface  
4308 environments prior to the Great Oxygenation Event is a relatively modern concept (Lyons et  
4309 al., 2014). This interpretation reflects an evolution of ideas from Chapter Three, where a  
4310 magmatic origin for Carlow Castle was tentatively suggested. Further research could focus on  
4311 characterising the multiple sulphur isotope composition of broadly coeval hydrothermal ore

4312 deposits and sedimentary rocks across the northwest Pilbara Craton and particularly within the  
4313 Mallina Basin. Whilst several studies of Archean hydrothermal ore deposits within the adjacent  
4314 Yilgarn Craton have focussed on utilising multiple sulphur isotopes to understand ore-forming  
4315 processes, research applying this technique to understanding ore genesis within the Pilbara  
4316 Craton is non-existent beyond the study presented here. As such, the application of multiple  
4317 sulphur isotope analysis to understanding ore formation within the northwest Pilbara Craton  
4318 would be novel and would provide useful insights into the crustal development of the northwest  
4319 Pilbara Craton and regional metallogeny through the Archean.

4320 Whilst the potential of Cu isotope analysis to understand Cu ore-forming processes and for Cu  
4321 ore deposit exploration has been demonstrated in previous studies (Asael et al., 2007; Gregory  
4322 & Mathur, 2017; Mathur & Wang, 2019), Chapter Five presents the first application of Cu  
4323 isotope analysis to an Archean Cu ore deposit. Critically, the limited potential for redox cycling  
4324 to induce isotopic fractionation in Archean ore-forming systems may imply that the isotopic  
4325 signature of Archean Cu ores should display an unfractionated signature consistent with a bulk  
4326 silicate Earth composition (Liu et al., 2015; Zhu et al., 2002). However, Cu mineralisation from  
4327 Carlow Castle display a fractionated isotopically light composition relative to the bulk silicate  
4328 Earth reservoir. Additionally, a correlation between Cu isotope signature, Cu ore grade, and  
4329 hydrothermal alteration assemblage appears to record the physicochemical evolution of the  
4330 Carlow Castle ore system. This suggests that Cu isotope analysis may be a useful tool in  
4331 understanding Archean Cu ore systems and analysis of Cu ore minerals could be a novel tool  
4332 to record the evolution of hydrothermal systems, especially when coupled with existing  
4333 mineralogical and geochemical tools. However, the existing literature applying Cu isotope  
4334 analysis to Archean geological systems is scant and the current understanding of the Cu isotope  
4335 systematics of the Archean Earth is very limited. Therefore, whilst this study provides a

4336 compelling proof of concept for the use of Cu isotope analysis to understand Archean Cu ore-  
4337 forming systems, there is a need for further research in this space. Additionally, the relatively  
4338 limited heavy isotopic fractionation of supergene Cu mineralisation from Carlow Castle bears  
4339 implications for the use of Cu isotope analysis as a geochemical exploration tool. Previous  
4340 studies primarily relied upon the significant heavy isotopic fractionation of Cu during  
4341 weathering of Cu sulphide ores as a potential means to vector towards primary Cu ore deposits  
4342 (Mathur & Wang, 2019). However, the majority of these studies focussed on porphyry Cu  
4343 deposits with well-developed supergene leach caps and enrichment zones overlying primary  
4344 Cu mineralisation. As such, the limited supergene isotopic fractionation reported here is an  
4345 important consideration for future studies implementing Cu isotope analysis of supergene  
4346 materials in structurally complex terranes, where supergene zonation around Cu sulphide  
4347 deposits may be less defined. However, the study presented here is only focussed on a single  
4348 case study area. The wider applicability of these interpretations would be strengthened by  
4349 further research on supergene Cu isotope systematics in a wider variety of ore deposit settings.

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